

### 1. Plasma parameters

Find the Debye lengths,  $d_e$ ,  $d_i$ , the plasma parameters,  $N_{de}$ ,  $N_{di}$ , the plasma frequencies,  $\omega_{pe}$ ,  $\omega_{pi}$ , and the thermal velocities,  $v_{Te}$ ,  $v_{Ti}$ , for electrons and ions (protons) under the following conditions:

- (a) A fusion device:  $n_e=n_i=10^{16} \text{ cm}^{-3}$ ,  $T_e=T_i=10^7 \text{ K}$ ;
- (b) The Earth's magnetosphere:  $n_e=n_i=10^4 \text{ cm}^{-3}$ ,  $T_e=T_i=10^3 \text{ K}$ ;
- (c) The center of the Sun:  $n_e=n_i=10^{26} \text{ cm}^{-3}$ ,  $T_e=T_i=10^{7.2} \text{ K}$ ;
- (d) The solar corona:  $n_e=n_i=10^8 \text{ cm}^{-3}$ ,  $T_e=T_i=10^6 \text{ K}$ ;
- (e) The solar wind:  $n_e=n_i=10 \text{ cm}^{-3}$ ,  $T_e=T_i=10^5 \text{ K}$ ;
- (f) The atmosphere of a neutron star:  $n_e=n_i=10^{12} \text{ cm}^{-3}$ ,  $T_e=T_i=10^7 \text{ K}$ .

Can all of these cases qualify to be described as "plasmas"?

### 2. Particle interactions

Show that, in a nonrelativistic plasma ( $kT \ll m_e c^2$ ), the mutual Coulomb (electrostatic) force between two typical particles is much more important than the mutual Lorentz (magnetic) force.

### 3. Debye shielding

A spherical conductor of radius  $a$  is immersed in a plasma and charged to a potential  $\phi_0$ . The electrons remain Maxwellian and move to form a Debye shield, but the ions are stationary during the time frame of the experiment. Assuming  $\phi_0 \ll kT_e/e$ , derive an expression for the potential as a function of  $r$  in terms of  $a$ ,  $\phi_0$ , and  $d_e$ . (Hint: assume a solution of the form  $e^{-kr}/r$ .)

### 4. Gravitational drift

What is the gravitational drift speed of an electron in a tokamak, with  $|B_0|=10 \text{ kG}$ ? How about a proton? Does either of these drifts make it hard to confine a plasma in a volume of order  $1 \text{ m}^3$  for a time of order  $1 \text{ s}$ ?