Lecture 20

April 8, 2024

1 Special Relativity

1.1 Lorentz Transformations

Let's start by considering 2 intertial reference frames, S and S. The S prime moves in the positive x direction with a constant velocity u relative to S. An object has coordinates:

$$S': (x', y', z')$$
 at t'

$$S:(x,y,z)$$
 at t

Finally, we'll choose the origins of our frames to coincide at t' = t = 0.

The y and z coordinates are not affected by the motion, since we've chosen our frame to only have a velocity in the x direction. As such,

$$y = y'; z = z'$$

The transformations between these frames are thus given by the Lorentz Transformations:

$$x' = \gamma(x - ut) \tag{1}$$

$$y' = y \tag{2}$$

$$z' = z \tag{3}$$

$$t' = \gamma(t - \frac{ux}{c^2})\tag{4}$$

where

$$\gamma = \frac{1}{\sqrt{1 - \frac{u^2}{c^2}}}$$

We can get the inverse transformations very simply by switching the primed variables and unprimed variables, and replacing u with -u.

$$x = \gamma(x' + ut') \tag{5}$$

$$y = y' \tag{6}$$

$$z = z' \tag{7}$$

$$t = \gamma(t' + \frac{ux'}{c^2})\tag{8}$$

(9)

1.2 Simultaneity

Now consider an observer in frame **S** who observers two light bulbs to flash simultaneously, but which are at different x coordinates, x_1 and x_2 - that is, they are separated by some distance. An observer in frame **S'** will measure a time interval of $t'_2 - t'_1$

$$t'_1 - t'_2 = \gamma \left((t_1 - t_2) - \frac{u(x_1 - x_2)}{c^2} \right) = \gamma \left(\frac{u\Delta x}{c^2} \right)$$

So this means, for the observer in S', the flash bulbs do not go off at the same time! This means events observed to be simultaneous in one interital frame will not occur simultaneously in another inertial frame!

1.3 Time Dialation

So how we use these relations? Consider first a time interval $\Delta t'$ measured by a clock which is at rest in the **S**' frame. The reason that it's at rest is important because it let's us set $\Delta x' = 0$ (that is, the clock does not move in the primed frame over this interval. If we want to figure out how much time has passed in **S**, then we can use:

$$t_2 - t_1 = \gamma \left((t_2' - t_1') + \frac{u(x_2' - x_1')}{c^2} \right)$$
 (10)

$$\Delta t = \gamma \left((\Delta t') + \frac{u\Delta x'}{c^2} \right) \tag{11}$$

$$\Delta t = \gamma \Delta t' \tag{12}$$

(13)

This leads us to the equation for "time dialation"

$$\Delta t = \frac{\Delta t'}{\sqrt{1 - \frac{u^2}{c^2}}}$$

We should also note that the clock has moved in **S**:

$$x_2 - x_1 = \gamma((x_2' - x_1') + u(t_2' - t_1'))$$
(14)

$$\Delta x = \gamma u \Delta t' \tag{15}$$

It's important to remember that this is only true for a clock which is at rest in the **S'** frame! If it weren't at rest, then the $\Delta x'$ frame being 0 would not be true. Now, a quick discussion of what this equation actually means. Since $\sqrt{1-u^2/c^2} \le 1$, then $\gamma \ge 1$. As such, $\Delta t \ge \Delta t'$. Because the clock is at rest in **S'**, this means the time measured by the clock in **S**, where the clock is moving, is longer.

Proper Time

The shortest time interval is measured by a clock which is at rest relative to two events ("proper time").

1.4 Length Contraction

Now, consider an object which is travelling at velocity u, while we are stationary. In this scenario, the object is moving with the **S**' frame, and we are in the **S** frame. We want to measure the length of the object. To do so, we must measure the positions of two ends, x_2 and x_1 at the same time in the **S** frame (so that $t_2 = t_1$, $\Delta t = 0$). We can use

$$x_2' - x_1' = \gamma((x_2 - x_1) - u(t_2 - t_1)) \tag{16}$$

$$\Delta x' = \gamma \Delta x \tag{17}$$

$$\Delta x = \frac{\Delta x'}{\gamma} \tag{18}$$

(19)

Which leads to "length contraction":

$$\Delta x = \sqrt{(1 - \frac{u^2}{c^2})} \Delta x'$$

Since $\gamma \geq 1$, then $\Delta x \leq \Delta x'$. Since we are in the unprimed frame, then we measure the object to be shorter than it actually is.

Proper Length

The longest length of an object is measured in the frame in which it is at rest ("proper length")

1.5 Example: Cosmic rays

Cosmic rays are very energetic particles which enter Earth's atmosphere and subsequently react with atoms to produce muons. These muons decay after an average life time of $\tau = 2.2 \times 10^{-6}$ s as measured in the lab, where the muons are at rest. This means that the number of muons (N) in a sample at time t should go as:

$$N(t) = N_0 \exp\left(-t/\tau\right)$$

where N_0 is the initial number of muons. Now, a detector on top of a mountain has measured the rate of muons at a height of 1907 m above sea level to be 563 muons/hr, travelling downward with a velocity of u = 0.9952c. What is the muon rate at sea level? First, the time for the muons to reach sea level is:

$$t = \frac{1907\text{m}}{0.9925(3 \times 10^8 \text{m/s})} = 6.39 \times 10^{-6} \text{s}$$

If this sample of muons were at rest, then after this amount of time, we would expect to have

$$N(t) = 563 \exp\left(-\frac{6.39 \times 10^{-6}}{2.2 \times 10^{-6}}\right) \sim 31$$

left. However, the rate at sea level has been found to be 407 muons/hr - so, we need to use the above relations to figure out what's occurring here. Since the muons are moving, their lifetime in our frame is different to what it is in the rest frame. Now, assuming we are in the unprimed frame, and the muons are travelling in the primed frame with a velocity u=0.9952:

$$\tau = \gamma \tau' = 10.22(2.2 \times 10^{-6}) = 2.24 \times 10^{-5} s$$

So now using the above expression, but with this new life time:

$$N(t) = 563 \exp\left(-\frac{6.39 \times 10^{-6}}{2.24 \times 10^{-5}}\right) \sim 423$$

which is much closer to the observed value.

We can look at this from a different angle as well. Imagine we are riding alongside the muons, which we are now going to call the unprimed frame - what is the distance between the top of the mountain and sealevel? In the rest frame of mountain (the primed frame), this is 1907 m. However, to an observer in the muon frame, the mountain is moving with a velocity of 0.9952. And so the distance is contracted to:

$$\Delta x = \frac{\Delta x'}{\gamma} = \frac{1907}{10.22} = 189m$$

The time to reach sea level in the muon rest frame is then:

$$t = \frac{189\text{m}}{0.9925(3 \times 10^8 \text{m/s})} = 6.3 \times 10^{-7} \text{s}$$

And so, the number of muons left by then is:

$$N(t) = 563 \exp\left(-\frac{6.3 \times 10^{-7}}{2.2 \times 10^{-6}}\right) \sim 423$$

So this demonstrates quite nicely that the time dialation in one frame is equivalent to length contraction in the other frame.