

## 5.5 Multidimensional splines

---

### Multidimensional splines

Each of the approaches discussed in this weeks course material has multidimensional analogues.

Let now assume that  $X \in \mathbf{R}^2$  and we have a basis functions  $h_{1k}(X_1)$ ,  $k = 1, \dots, M_1$  for representing functions of coordinate  $X_1$ , and likewise a set of  $M_2$  functions  $h_{2k}(X_2)$  for  $X_2$ .

Then the  $M_1 \times M_2$  dimensional tensor product basis is defined by

$$g_{jk}(X) = h_{1j}(X_1)h_{2k}(X_2), \quad j = 1, \dots, M_1, k = 1, \dots, M_2$$

and it can be rewritten as a two-dimensional function:

$$g(X) = \sum_{j=1}^{M_1} \sum_{k=1}^{M_2} \theta_{jk} g_{jk}(X).$$

One dimensional smoothing splines generalize to higher dimensions, too. For  $x_i, y_i$  with  $x_i \in \mathbf{R}^d$  we seek a  $d$ -dimensional regression function  $f(x)$ . The idea is to set up the problem

$$\min_f \sum_{i=1}^N \{y_i - f(x_i)\}^2 + \lambda J[f],$$

where  $J$  is an appropriate penalty functional for stabilizing a function  $f$  in  $\mathbf{R}^d$ .

For example, in  $\mathbf{R}^2$  we have

$$J[f] = \int \int_{\mathbf{R}^2} \left[ \left( \frac{\partial^2 f(x)}{\partial x_1^2} \right)^2 + 2 \left( \frac{\partial^2 f(x)}{\partial x_1 \partial x_2} \right)^2 + \left( \frac{\partial^2 f(x)}{\partial x_2^2} \right)^2 \right] dx_1 dx_2.$$

Optimizing the above criterion leads to a smooth two-dimensional surface, known as **thin-plate spline**. The solution here has the form

$$f(x) = \beta_0 + \beta^\top x + \sum_{j=1}^N \alpha_j h_j(x),$$

where  $h_j(x) = \|x - x_j\|^2 \log \|x - x_j\|$ . Here  $h_j$  is an example of **radial basis functions**. The coefficients  $\alpha_j$  are found by plugging this solution back into the optimization criterion.

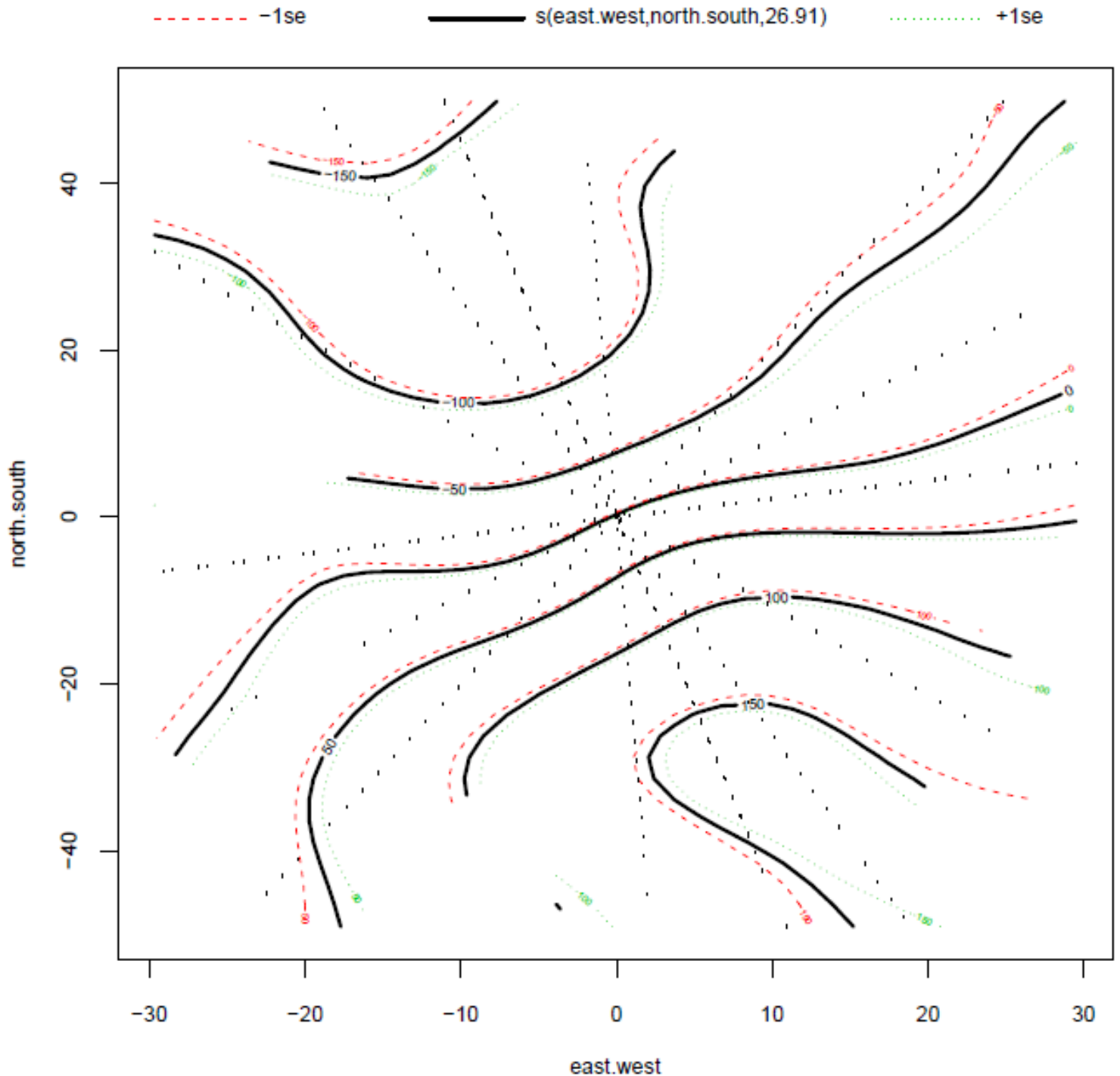
**Example: Fitting a thin plate spline in R**

A thin-plate spline fit is implemented in the package **mgcv**:

```
require(mgcv)

galaxy <- read.delim("/course/data/galaxy.data", sep=",")

fit = gam(formula = velocity ~ s(east.west, north.south), data = galaxy)
plot(fit)
```



Fits in higher dimensions may be calculated as well, but are not visualised as easily. Alternatively, a loess smooth can be fitted:

```
galaxy <- read.delim("/course/data/galaxy.data", sep=",")
require(mgcv)
```

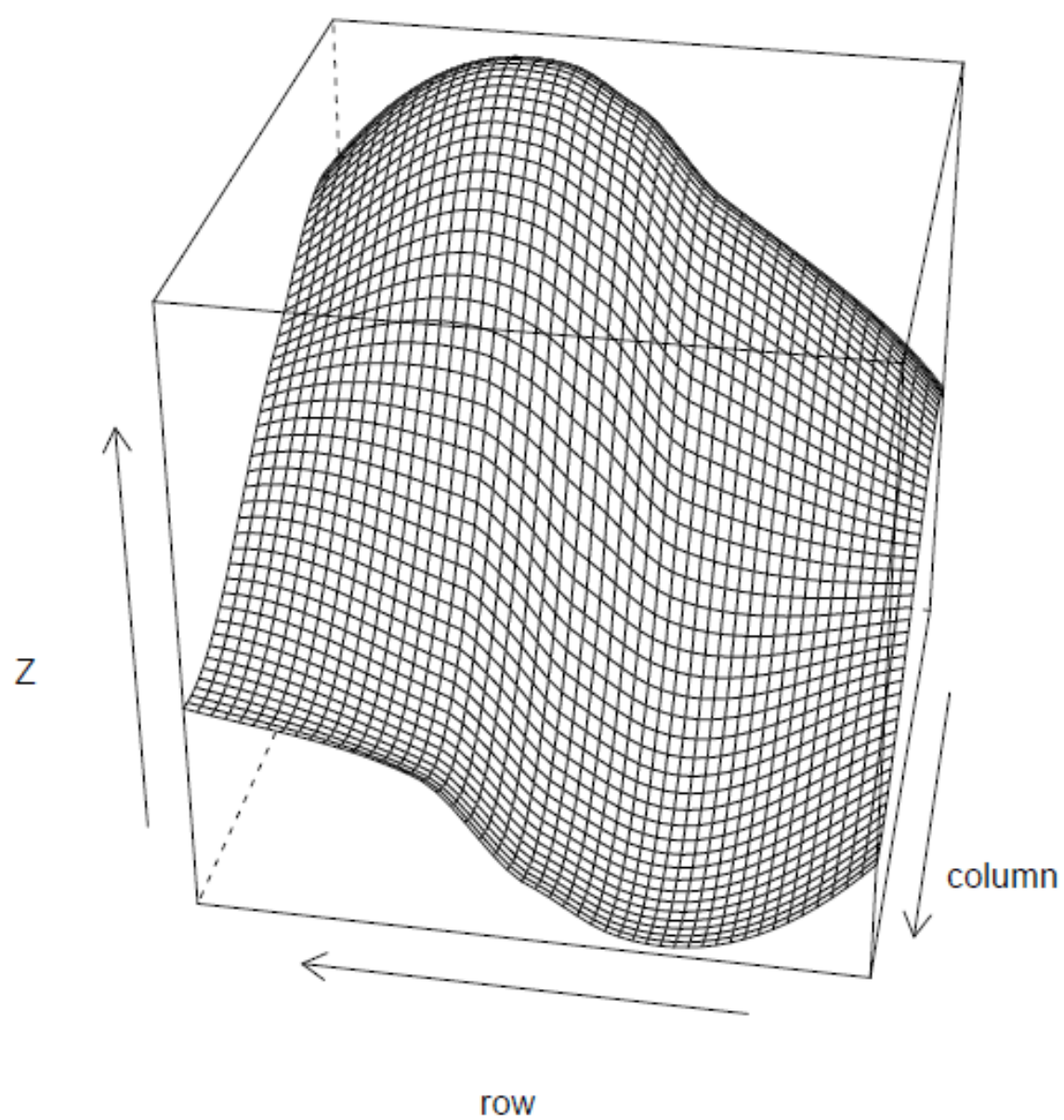
```
loc.lin.reg = loess(formula = velocity ~ north.south +
                    east.west, data = galaxy, span = 0.7, degree = 2,
                    normalize = T, family = "gaussian")
attach(galaxy)

xgrid = seq(min(north.south), max(north.south), length.out = 50)
ygrid = seq(min(east.west), max(east.west), length.out = 50)

new.coords = matrix(data = NA, nrow = 2500, ncol = 2)
new.coords[, 1] = as.vector(outer(rep(1, 50), xgrid))
new.coords[, 2] = as.vector(outer(ygrid, rep(1, 50)))

new.coords = as.data.frame(new.coords)
names(new.coords) <- c("north.south", "east.west")

Z = predict(loc.lin.reg, newdata = new.coords)
Z = matrix(data = Z, ncol = 50)
library(lattice)
wireframe(Z, screen = list(z = 170, x = -60))
```



---

## Activity in R: 3D visualisation

Consider the **galaxy** dataset:

```
galaxy <- read.csv("../data/galaxy.data")  
head(galaxy,n = 2)
```

	row.names	east.west	north.south	angle	radial.position	velocity
1	3	8.462789	-38.17317	102.5	39.1	1769
2	4	7.964978	-35.92769	102.5	36.8	1749

Generate a plot of the three-dimensional point cloud **east.west,north.south,velocity** using the *cloud()* function in the **lattice** library in R.