N-body Project

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Introduction

Goal: Calculation of gravitational force on a system of collisionless particles

Possible methods:

- Direct calculation
- · Iterative methods on a mesh
- · Fourier methods on a mesh
- Tree methods (hierarchical multipole methods)

In this project, I implemented a direct calculation and a tree method.

Dataset

The dataset for which to compute the forces is a set of $\sim 50'000$ particles aranged according to the spherically symmetric "Hernquist model":

$$\begin{split} \rho(r) &= \frac{M}{2\pi} \frac{a}{r} \frac{1}{(r+a)^3} \\ M(r) &= M \frac{r^2}{(r+a)^2} \qquad \Rightarrow M(a) = \frac{M}{4} \\ \phi(r) &= -\frac{GM}{r+a} \end{split}$$

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Choice of Units

Use dimensionless units, with $G \equiv 1$

- ⇒ Reduce number of multiplications necessary
- \Rightarrow Reduce effect of finite floating point precision by moving problem to a better suited order of magnitude

Define a scale for every physical quantity:

$$a_{phys} \equiv A_0 a_{code}$$

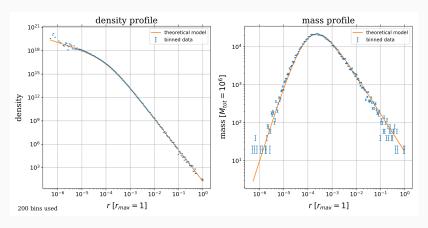
Setting G=1 restricts the scale for either time, mass or distance. I chose

$$M_0=M_{tot} \qquad \qquad \text{such that} \qquad M_{tot,code}=1$$

$$R_0=R_{max} \qquad \qquad \text{such that} \qquad R_{max,code}=1$$

$$\Rightarrow \quad T_0=\sqrt{\frac{R_0^3}{GM_0}}$$

Dataset



Only in this plot: $M_{tot} \equiv 10^6$ so that the errorbars don't dominate the plot.

Direct forces calculation

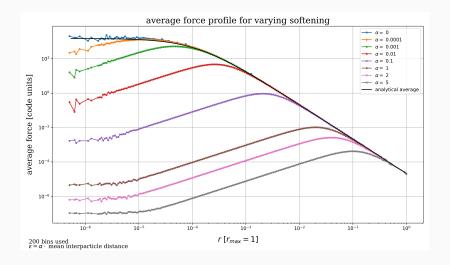
Acceleration of particle i:

$$\ddot{\mathbf{r}}_i = -G\sum_{j=1}^N \frac{m_j}{[(\mathbf{r}_i - \mathbf{r}_j)^2 + \epsilon^2]^{\frac{3}{2}}} (\mathbf{r}_i - \mathbf{r}_j)$$

 ϵ is the softening length. It's purposes are:

- · computational efficiency
- · avoid large angle scatterings
- avoid expense to calculate orbits in a singular potential
- prevent the possibility of formation of bound particle pairs

Direct force calculation results for varying softening



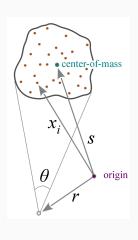
Hierarchical Multipole Method

Central idea: use the multipole expansion of a distant group of particles to describe its gravity, instead of summing up the forces from all individual particles. (For close groups, I use direct force calculation.)

Multipole expansion of the potential gives:

$$\begin{split} \Phi(\mathbf{r}) &= -G\left(\frac{M}{|\mathbf{y}|} + \frac{1}{2}\frac{\mathbf{y}^T \mathbf{Q} \mathbf{y}}{|\mathbf{y}|^5}\right) \\ \mathbf{y} &= \mathbf{r} - \mathbf{s} \\ Q_{ij} &= \sum_k m_k [3(\mathbf{s} - \mathbf{x}_k)_i (\mathbf{s} - \mathbf{x}_k)_j - \delta_{ij} (\mathbf{s} - \mathbf{x}_k)^2] \end{split}$$

The expansion is valid for $heta pprox rac{l}{y} \ll 1$



The particles are grouped hierarchically and the multipole moments are pre-computed for later use.

I used the Barnes-Hut oct-tree: Assume particles are in a cube. The cube is then recursively subdivided into 8 sub-cubes of half the size in each spatial dimension, until each sub-cube contains only a single particle (or some other user-set limit).

Force calculation:

For all leaf cells:

For all root cells:

walk tree (this leaf cell, root cell)

Walking the tree (target cell, source cell):

If source cell is leaf cell: Use direct force calculation, end walk for this source.

If target (leaf) cell is inside this source cell:

for all children of the source cell:

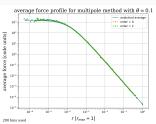
walk the tree (target, child of source)

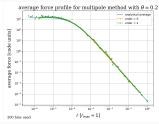
If target (leaf) cell is *not* inside this source cell:

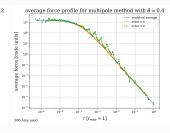
if $\theta < \theta_{max}$: Calculate multipole force, stop walk for this source

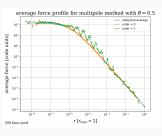
else: for all children of the source cell:

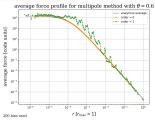
walk the tree (target, child of source)

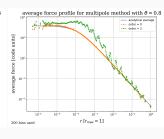










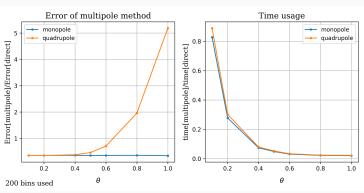


Comparison Direct vs multipole calculation

Estimating accuracy with the L1 norm:

 $L1=rac{1}{N}\sum_i|ar{F}_i-ar{F}_{analytical}(r_i)|$, where $ar{F}_{analytical}(r_i)$ is the analytical average force at r_i .

For both the direct forces and multipole method, $\epsilon=0.01\cdot mid$. Shown are the relative values for both monopole and quadrupole oders, relative to the values for the direct forces calculation.



Program, plotting scripts and this presentation available on https://bitbucket.org/mivkov/computational_astrophysics