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Overview:

The eSTARlight Monte Carlo models photon-Pomeron interactions in electron-ion collisions. The physics approach for the photon-Pomeron interactions is described in Lomnitz and Klein, arXiv:1803.06420 (2018).

eSTARlight has several input files, all of which are expected to be in the same directory as the estarlight code. User-specified input parameters are read from a file named "slight.in"; these parameters are described below in Input.

The simulated events are written to an ASCII file named "slight.out", which is described below in Output.

What's new/changed:

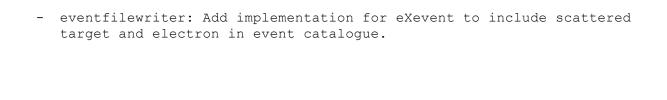
Although eSTARlight inherits many methods from STARlight, it required sufficient changes to motivate a separate distribution. As such, the source code includes several components from STARlight as well as many changes and new classes. A brief summary of the changes is as follows:

New classes:

- gammaeluminosity: Generates look-up tables for electron-ion collisions.
- e_wideResonanceCrossSection: Coherent vector meson production using wide resonance in eX collisions.
- e_narrowResonanceCrossSection: Coherent vector meson production using narrow resonance in eX collisions.
- e_starlightStandalone: Similar to STARlight case, calls methods to initialize and produce events and decay them.
- e starlight: Reads in inputParameters and checks their validity.
- e main: Driver program, initiates and calls e starlightStandalone.
- eXevent: Contained for e+X event information (analogous to upcEvent in STARlight).

Modified classes

- beambeamsystem: Add support for electron beam.
- readinluminosity: Methods to read the look up tables for _ep_ and eA have been added.
- photonNucleusCrossSection: Methods to calculate photon flux and $\gamma X \rightarrow VX$ cross-section.
- nucleus: Add support for electron (Z=1, A=0).
- inputParameters: Add inputParamters for eSTARlight, including: $\min(\max)$ γ energy(k) and virtuality(Q^2), number of k and Q^2 bins, electron energy, ...
- _
- gammaavm: Add methods to generate kinematic variables and momenta to finite virtuality. Also generate outgoing electron and target.
- filewriter: Increased precision for event catalogue. Necessary for scattered target and electron.



Installation:

To obtain the latest version:
-git clone git@github.com:mlomnitz/eSTARlight.git

Alternatively:

- -Visit https://estarlight.hepforge.org/trac/browser
- -Download the trunk [click on the download symbol in the Size column]
- -Unpackage the zip file. The trunk/ represents <PathToSource>

To build eSTARlight:

- First create your build directory <BUILDDIR> (e.g. mkdir bin)
- \$ cd <BUILDDIR>
- \$ cmake <PathToSource>
- \$ make

This creates an executable file, e starlight, in the build directory.

To clean the build:

- \$ make clean

To run eSTARlight, a configuration file, slight.in, is needed. Examples of

slight.in may be found in the config/ directory.

To run:

\$./e starlight

Enabling Pythia:

To simulate the $\eta,~\eta',~$ and η_c channels, you need Pythia v8.2 or higher to handle their decays. To enable Pythia support you need to run cmake with the option -DENABLE_PYTHIA=ON and have \$PYTHIADIR pointing to the top directory of Pythia8. [Note: when building Pythia, be sure to enable shared libraries(.so). ./configure --enable-shared before compiling Pythia.]

- \$ setenv PYTHIADIR /my/local/pythia8
- \$ cmake <PathToSource> -DENABLE PYTHIA=ON

Note: v8.2+ is necessary since the Pythia directory structure changed[trunk/cmake_modules/FindPythia8.cmake depends on the structure layout], libhapdfdummy was removed, and Standalone:allowResDec was removed.

To enable DPMJET, please see the passage on DPMJET

Input:

The input parameters are listed below with typical values for e-Au collisions at the proposed eRHIC collider. Optional parameters are denoted with *, and <u>Legacy parameter</u> is used to indicate options in the set up file that have been inherited from STARLight but might not be accurately implemented in eSTARlight.

baseFileName	# The name of the output files. eSTARlight will copy the input slight.in to baseFileName.in, and produce output files baseFileName.txt and baseFileName.out. (slight)
BEAM_1_Z = 1	# Charge of beam one projectile. For electron-ion collisions, beam 1 must be the electron
BEAM 1 A = 0	# Atomic number of beam one projectile (0 for e).
$BEAM_2 = 79$	# Charge of beam two projectile (Au = 79).
$BEAM_2 = 197$	# Atomic number of beam two projectile (Au = 197).
$BEAM_1_GAMMA = 35295$	# Lorentz boost for beam one projectile(pz>0).
	These are 18 GeV electrons.
$BEAM_2_GAMMA = 106.6$	# Lorentz boost for beam two projectile(pz<0). These are 100 GeV/n Au ions.
W MAX = -1	# Maximum value for the gamma-pomeron center of
_	mass energy, in GeV. Setting $W_MAX = -1$ tells
	eSTARlight to use the default value specified in
	<pre>inputParameters.cpp (recommended for single meson</pre>
	production). For single mesons, the default $\mbox{W_MAX}$
	is the particle mass plus five times the width.
	For <u>lep</u> ton pairs, the default W_MAX is given by
	$2\hbar c \sqrt{\frac{\gamma_1 \gamma_2}{R_1 R_2}}$. These are defined in
	<pre>src/inputParameters.cpp</pre>
$W_MIN = -1$	#Min value of w. Minimum value for the gamma-
	pomeron center of mass energy, W, in GeV. Setting
	$W_{MIN} = -1$ tells eSTARlight to use the default
	value specified in inputParameters.cpp
	(recommended for single meson production). The
	default W_MIN is the larger of the kinematic limit
	($e.g.~2\text{m}_{\pi}$ for ρ decays) or the particle mass
	minus five times the width.
$W_N_BINS = 40$	#Bins W maximum and minimum values for W and the
	number of w bins in the lookup tables.
$RAP_MAX = 8.$	# Legacy parameter: Maximum rapidity of produced
	particle. Mostly left over from STARlight
	implementation.
$RAP_N_BINS = 80$	# Legacy parameter: Number of rapidity bins used
	in the cross section calculation. Mostly left over
	from STARlight implementation.

CUT PT* = 0	# Specifies whether the user chooses to place
_	restrictions on the transverse momentum of the
	decay products. 0= no, 1 = yes. (0)
PT MIN* = 1.0	# If a transverse momentum cut is applied, this
11_11111 1.0	specifies the minimum value produced, in GeV/c.
	(1.0)
$PT_MAX* = 3.0$	# If a transverse momentum cut is applied, this
	specifies the maximum value produced, in GeV/c.
	(3.0)
CUT ETA* = 0	# Specifies whether the user chooses to place
	restrictions on the pseudorapidity of the decay
	products. 0= no, 1 = yes. (0)
$ETA_MIN* = -10$	# If a pseudorapidity cut is applied, this
	specifies the minimum value produced. (-10)
$ETA_MAX* = 10$	# If a pseudorapidity cut is applied, this
_	specifies the maximum value produced. (10)
PROD MODE = 12	# Specifies the production mechanism to be used in
FROD_MODE = 12	
	event generation. At present eSTARlight includes
	two options:
	<pre>PROD MODE=12: Coherent photonuclear vector meson</pre>
	production assuming narrow resonances.
	PROD MODE=13: Coherent photonuclear vector meson
	production assuming wide resonances. This option
	should be used for exclusive \$\rho^0\$ production.
1.	
N_EVENTS = 10	#Number of events produced.
$PROD_PID = 443013$	# This selects the channel to be produced, in PDG
	notation. Currently supported options are list
	below.
RND SEED = 34533	# Seed for random number generator.
MIN_GAMMA_Q2*	# Specifies whether the user desires to set a
MIN_GAMMA_Q2	
	minimum value for the photon mass. By default,
	eSTARlight will set this to physical limit $Q_{min}^2 =$
	$m_e^2 k^2 / E_e (E_e - k)$.
MAX GAMMA Q2*	# Specifies whether the user desires to set the
11111_01111111_022	maximum value for the photon mass. By default the
	value is set to $Q^2_{max} = 4E_e(E_e - k)$ with the added
	requirement that the individual photons satisfy
	longitudinal coherence $l_c = 2k/(Q^2 + M_V^2)$
INT GAMMA Q2 BINS*	# Specifies whether the user desires to change the
1111_01111111_\$2_21110	" Specifies wheeler one abor actives to change the
	\sim
	number of Q^2 bins used when preparing the look-up
	tables, the default value is 400.
QUANTUM_GLAUBER = 1	
QUANTUM_GLAUBER = 1	tables, the default value is 400. # Species whether a quantum or classical Glauber
QUANTUM_GLAUBER = 1	<pre>tables, the default value is 400. # Species whether a quantum or classical Glauber extarpolation is to be used for nuclear targets. 1</pre>
_	<pre>tables, the default value is 400. # Species whether a quantum or classical Glauber extarpolation is to be used for nuclear targets. 1 = Quantum Glauber, 0 = Classical Glauber.</pre>
QUANTUM_GLAUBER = 1 SELECT_IMPULSE_VM = 0	<pre>tables, the default value is 400. # Species whether a quantum or classical Glauber extarpolation is to be used for nuclear targets. 1 = Quantum Glauber, 0 = Classical Glauber. # Species whether the impulse approximation is to</pre>
SELECT_IMPULSE_VM = 0	<pre>tables, the default value is 400. # Species whether a quantum or classical Glauber extarpolation is to be used for nuclear targets. 1 = Quantum Glauber, 0 = Classical Glauber. # Species whether the impulse approximation is to be used. 1 = Use impulse approximation, 0 = don't.</pre>
_	<pre>tables, the default value is 400. # Species whether a quantum or classical Glauber extarpolation is to be used for nuclear targets. 1 = Quantum Glauber, 0 = Classical Glauber. # Species whether the impulse approximation is to be used. 1 = Use impulse approximation, 0 = don't. # Used for proton and nucleon (i. e. incoherent</pre>
SELECT_IMPULSE_VM = 0	<pre>tables, the default value is 400. # Species whether a quantum or classical Glauber extarpolation is to be used for nuclear targets. 1 = Quantum Glauber, 0 = Classical Glauber. # Species whether the impulse approximation is to be used. 1 = Use impulse approximation, 0 = don't.</pre>
SELECT_IMPULSE_VM = 0	<pre>tables, the default value is 400. # Species whether a quantum or classical Glauber extarpolation is to be used for nuclear targets. 1 = Quantum Glauber, 0 = Classical Glauber. # Species whether the impulse approximation is to be used. 1 = Use impulse approximation, 0 = don't. # Used for proton and nucleon (i. e. incoherent nuclear) collisions to set the t-spectrum,</pre>
SELECT_IMPULSE_VM = 0	<pre>tables, the default value is 400. # Species whether a quantum or classical Glauber extarpolation is to be used for nuclear targets. 1 = Quantum Glauber, 0 = Classical Glauber. # Species whether the impulse approximation is to be used. 1 = Use impulse approximation, 0 = don't. # Used for proton and nucleon (i. e. incoherent nuclear) collisions to set the t-spectrum, dN/dt=exp(-bt). When BSLOPE_DEFINITION=1, then the</pre>
SELECT_IMPULSE_VM = 0	<pre>tables, the default value is 400. # Species whether a quantum or classical Glauber extarpolation is to be used for nuclear targets. 1 = Quantum Glauber, 0 = Classical Glauber. # Species whether the impulse approximation is to be used. 1 = Use impulse approximation, 0 = don't. # Used for proton and nucleon (i. e. incoherent nuclear) collisions to set the t-spectrum, dN/dt=exp(-bt). When BSLOPE_DEFINITION=1, then the slope is determined by BSLOPE_VALUE (below). When</pre>
SELECT_IMPULSE_VM = 0	<pre>tables, the default value is 400. # Species whether a quantum or classical Glauber extarpolation is to be used for nuclear targets. 1 = Quantum Glauber, 0 = Classical Glauber. # Species whether the impulse approximation is to be used. 1 = Use impulse approximation, 0 = don't. # Used for proton and nucleon (i. e. incoherent nuclear) collisions to set the t-spectrum, dN/dt=exp(-bt). When BSLOPE_DEFINITION=1, then the slope is determined by BSLOPE_VALUE (below). When BSLOPE_DEFINITION=2, the slope is calculated as a</pre>
SELECT_IMPULSE_VM = 0	<pre>tables, the default value is 400. # Species whether a quantum or classical Glauber extarpolation is to be used for nuclear targets. 1 = Quantum Glauber, 0 = Classical Glauber. # Species whether the impulse approximation is to be used. 1 = Use impulse approximation, 0 = don't. # Used for proton and nucleon (i. e. incoherent nuclear) collisions to set the t-spectrum, dN/dt=exp(-bt). When BSLOPE_DEFINITION=1, then the slope is determined by BSLOPE_VALUE (below). When</pre>

 $b=4.63/\text{GeV}^2 + 4\alpha \ln (W_{yp}/90 \text{ GeV})$

The default value, BSLOPE_DEFINITION=0 has no

effect.

Note that this affects the t-slope only; it does

not affect the total cross-section

WHEN BSLOPE_DEFINITION=1, this determines the
exponential slope for dN/dt=exp(-BSLOPE VALUE*t)

The following parameters are used only when interfacing with the PYTHIA and/or DPMJET interfaces:

MIN_GAMMA_ENERGY = 6 #Allows the user to set the minimum photon energy (in GeV) in the rest frame of the target nucleus. The default is 6.0 GeV and it should never be set below this value since DPMJET was not designed to handle low energy interactions.

MAX GAMMA ENERGY = 600000

#Allows the user to set the maximum photon energy (in GeV) in the rest frame of the target nucleus. The default is 60000.0 GeV.

PYTHIA PARAMS = ""

BSLOPE VALUE*

#Used to supply input parameters to the PYTHIA interface. This takes a string to pass on semicolon separated parameters to PYTHIA 6. eg: "mstj(1)=0;paru(13)=0.1" (the default is a blank string " ")

 $PYTHIA_FULL EVENT RECORD = 1$

#Determines whether the full event record from PYTHIA is written to slight.out. true = yes, false = no (false). The additional information added is as follows: daughter production vertex (x [mm], y [mm], z [mm], t [mm/c]), mother1, mother2, daughter1, daughter2, PYTHIA particle status code. PYTHA 8 Particle Properties page describes in more detail the properties of mother, daughter, and status code designations.

Channels of Interest:

Pomeron-Photon Channels

At present only the photon-pomeron channels have been included in eSTARlight (production modes 12 and 13). The channels included are:

jetset id	particle
110	
113	rho0
223	omega
333	phi
443011	J/psi> e+e-
443013	J/Psi> mu+mu-
444011	Psi(2S)> e+e-
444013	Psi(2S)> mu+mu-
553011	Upsilon(1S)> e+e-

553013	Upsilon(1S)> mu+mu-
554011	Upsilon(2S)> e+e-
554013	Upsilon(2S)> mu+mu-
555011	<pre>Upsilon(3S)> e+e-</pre>
555013	<pre>Upsilon(3S)> mu+mu-</pre>
913	rho0 + direct pi+pi- (with interference). The direct
	pi+pi- fraction is from the ZEUS results, EPJ C2 p247
	(1998)
999	<pre>four-prong final states (rho'-like to pi+pi-pi+pi-)</pre>

DPMJET:

Simulation of photonuclear interactions with eSTARlight is possible through an interface with DPMJet. These interfaces can be enabled through options passed to cmake during the configuration process. [Depreciated: Using Pythia 6 as a substitute for DPMJet]

The gfortran compiler is required to use the photonuclear interfaces.

========= 1. Photonuclear interactions with DPMJet =========

----- 1.1. Obtaining and installing DPMJet -----

The DPMJet package can be obtained by contacting the authors as explained here: http://sroesler.web.cern.ch/sroesler/dpmjet3.html

Once you have the code proceed with these steps:

Change the line containing the OPT variable in the DPMJet Makefile:

OPT = -c -C -std=legacy -O -O3 -g -fexpensive-optimizations -funroll-loops -fno-automatic -fbounds-check -v -fPIC

----- 64-bit -----

 $\,$ Make sure that all -m32 options are removed from the Makefile.

Unfortunately, the DPMJet package depends on a floating point exception trap implementation, and only a 32-bit version of that is included in the package, which needs to be replaced. An example implementation can be found here:

http://www.arsc.edu/arsc/support/news/hpcnews/hpcnews376/

Under "Fortran Floating Point Traps for Linux" there is a code example. A file based on this, fpe.c, can be found in the external/ directory in eSTARlight. Move that to your DPMJet directory to replace the original file and run:

\$ gcc -o fpe.o fpe.c

Note: if the above command returns the following error: /usr/lib/../lib64/crt1.o: In function `_start': (.text+0x20): undefined reference to `main' /tmp/ccs2CQsd.o: In function `enable_exceptions_': fpe.c:(.text+0xe): undefined reference to `feenableexcept'

```
collect2: error: ld returned 1 exit status
     Try: gcc fpe.c -Wall -g -c
           feenableexcept is a gcc extension and gcc may need all of the
     headers present.
                 ----- End 64-bit -----
     Then in the DPMJet directory run:
           $ make
     Note: When compiling at RCAS(BNL), needed to change g77 \rightarrow
qfortran, needed to install fluka and setenv FLUPRO /path/to/fluka, and
modify phojet before compiling. The changes for phojet is at line 29875,
from:
        PRINT LO, 'PHO DIFSLP: ERROR: this option is not installed !'
     to:
        WRITE (LO, '(/1X, A, I2)')
       & 'PHO DIFSLP:ERROR: this option is not installed
       & !', ISWMDL(13)
----- 1.2. Compiling eStarlight with DPMJet interface ------
           To enable the compilation of the DPMJet interface please
     follow these steps:
            CMake uses an environment variable $DPMJETDIR to locate the
            DPMJet object files, so define it.
           $ export DPMJETDIR=<path to dpmjet>
           Then create a build directory for eSTARlight
           $ mkdir <build-dir>
           and change into it
           $ cd <build-dir>
           Run CMake with the option to enable DPMJet
           $ cmake <path-to-estarlight-source> -DENABLE DPMJET=ON
```

Note: When compiling at RCAS(BNL), needed to add the gfortran library to the CMakeLists.txt and left it there.

Then build it

\$ make

----- 1.3. Running eSTARlight with DPMJet interface ------

To run eSTARlight with the DPMJet interface a couple of files are needed in the directory where you want to run eSTARlight.

The files needed are:

slight.in (eSTARlight config file. An example suitable for
DPMJet can be found in config/slight.in.dpmjet)
my.input (DPMJet config file. An example can be found in
config/my.input)
dpmjet.dat (Can be found in the DPMJet source directory)

In the slight.in file the relevant production modes (PROD_MODE) for DPMJET is:

5: A+A single excitation6: A+A double excitation7: p+A single excitation

In addition the minimum and maximum gamma energies must be set. These must be within the interval set in the my.input file.

To run:

\$./e_starlight < my.input
[DPMJET reads from direct input/interactive]</pre>

Output

eSTARlight outputs an ASCII file named slight.out. The first 4 lines are used to store some of the important configuration options used to produce the event. The information contained in these lines is as follows:

CONFIG_OPT: prod_mod particle_id nevents q_glauber impulse seed where prod_mod indicates if a wide or narrow resonance has been used, particle_id specifies the vector meson species (and decay channel) being produced, nevents indicates the total number of events in the simulation, q_glauber indicates if a quantum (=1) or classical (=0) Glauber has been selected, impulse indicates if the nuclear effects are being modelled (=0) or a simple impulse approx. is employed, and finally seed records the random number seed used when initializing the Monte Carlo. The config opt line is followed by two lines with brief descriptions of beams in the collision, with the format:

BEAM 1(2): beam1(2) Z beam1(2) A beam1(2) LorentzGamma

where beam1(2)Z is the is the charge of the particles in beam 1(2), beam1(2)A indicates the atomic number of beam 1(2) and beam1(2)LorentzGamma is the Lorentz gamma factor associated to beam 1(2)

These lines are then followed by a brief description of the user settings for the exchanged photons in the collisions, as follows:

PHOTON: nkbins fixed02 n02bins min02 max02

where nkbins is the number of steps (with exponential steps) used for the photon energy look-up tables, fixedQ2 indicates whether the user selected a fixed(=1) range for the photon virtuality or used the physical limits (=0). nQ2bins states the number of bins in Q^2 used for the look-up tables, minQ2 and maxQ2 indicate the minimum and maximum value of Q^2 used to generate events. If the Q^2 range was not selected by the user (i.e. fixedQ2 = 0) minQ2 and maxQ2 are displayed as 0.

These lines are followed by the event catalogue. For each event, a summary line is printed, with the format

EVENT: n ntracks nvertices,

where n is the event number (starting with 1), ntracks is the number of tracks in the event, and nvertices is the number of vertices in the event (eSTARlight does not produce events with more than one vertex).

EVENT line is followed by a description of the vertex, with the format

VERTEX: x y z t nv nproc nparent ndaughters ,

where x, y, z and t are the 4-vector components of the vertex location, nv is the vertex number, nproc is a number intended to represent physical process (always set to 0), nparent is the track number of parent track (0 for primary vertex) and ndaughters is the number of daughter tracks from this vertex.

This is followed by a line describing the kinematics of the photon in the reference frame where the target (p or A) is at rest.

GAMMA: k Q2,

where k is the energy of the photon and Q2 is the invariant mass of the virtual photon.

This is followed by information related to the scattered target (X = p or A) which emerges from the collision.

t: event t ,

where, as expected, event_t is the four momentum transfer squared at the target vertex.

TARGET: px py pz E,

where px, py and pz are the three vector components of the scattered target three vector and E is it's energy.

The information related to the scattered target is followed by the scattered electron or source.

SOURCE: px py pz E,

where, again, px, py and pz are the components of the outgoing electron three vector and E is it's energy.

This is followed by a series of lines describing each of the daughter tracks emanating from this vertex. Each track line has the format

TRACK: GPID px py pz nev ntr stopv PDGPID,

where GPID is the Geant particle id code, px, py and pz are the three vector components of the track's momentum, nev is the event number, ntr is the number of this track within the vertex (starting with 0), stopy is the vertex number where track ends (0 if track does not terminate within the event), and PDGPID is the Monte Carlo particle ID code endorsed by the Particle Data Group.