

# Photometric Properties of the RR Lyrae Star SS Piscium

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## Abstract

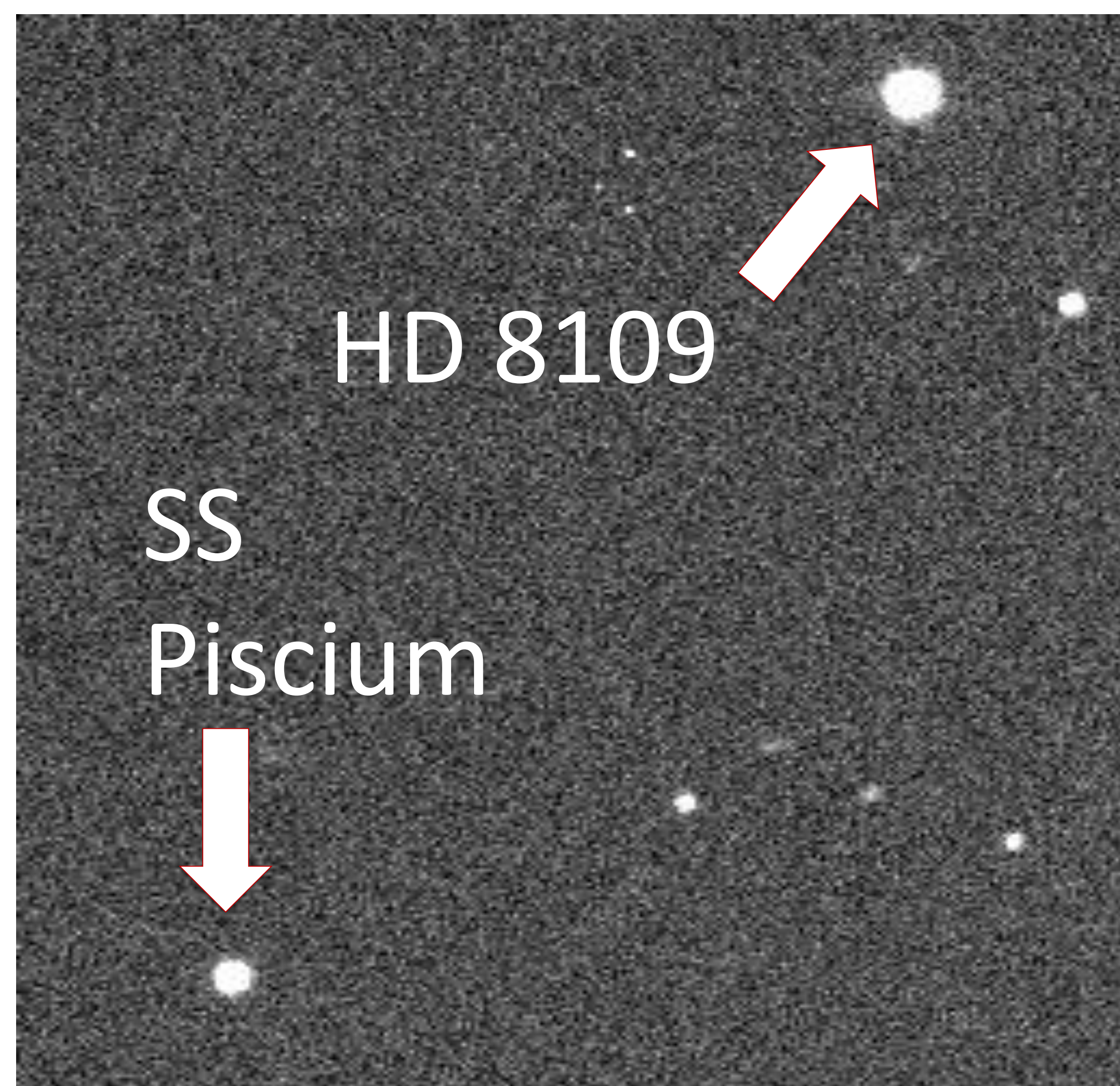
Standard candles are useful tools in astronomy. These are objects whose intrinsic luminosity can be determined from observable properties. RR Lyrae stars are standard candles due to a relation between luminosity and brightness variation. We present V-band photometric observations of the RR Lyrae star SS Piscium. We have constructed a magnitude light curve from three observing nights. From the period of variation and a Fourier series fit to this light curve, we can determine the star's iron abundance and absolute magnitude. From the latter, we determine its distance.

## Observations

Observed on three nights: 10/28, 11/4, 11/6

### Equipment

- 14in Meade LXF2000-ACF telescope
- SBIG STL-1001E CCD camera
- V-band CCD filter



## Analysis Steps

- The star HD 8109 was used as a magnitude calibrator
- The magnitude of SS Piscium is calculated as
$$m_{SS\,Psc} = m_{HD\,8109} - 2.5 \log_{10} \left( \frac{f_{SS\,Psc}}{f_{HD\,8109}} \right)$$
- Each night's data is combined into a continuous magnitude light curve
- We fit a Fourier series <sup>[1]</sup> to the combined curve

$$m_V(t) = a_0 + \sum_{n \geq 1} a_n \cos(2\pi n x + \varphi_n)$$

where

$$x \equiv \frac{1}{T} (t - t_{\text{brightness maximum}})$$

- The variation period (T) and the Fourier parameters ( $a_n, \varphi_n$ ) give the star's iron abundance <sup>[2]</sup>

$$[\text{Fe}/\text{H}] = 52.466 T^2 - 30.075 T + 0.131 \varphi_{3-1}^2 + 0.982 \varphi_{3-1} - 4.198 \varphi_{3-1} T + 2.424$$

and its absolute visual magnitude <sup>[3]</sup>

$$M_V = -0.961 T - 0.044 \varphi_{2-1} + 4.447 a_4 + 1.061$$

- The heliocentric distance to the star is

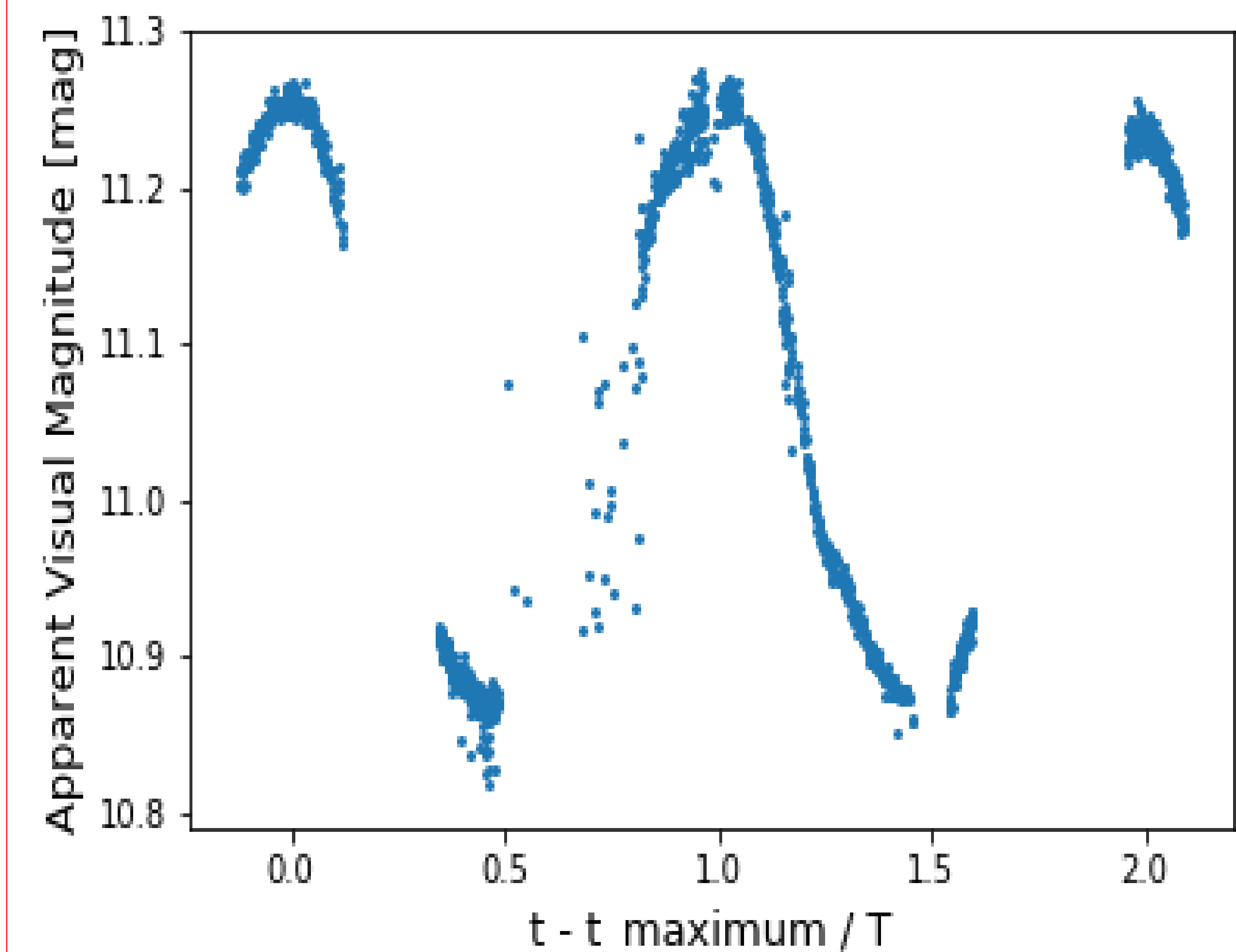
$$d = 10^{\frac{1}{5}(\langle m_V \rangle - M_V + 5)} [\text{pc}]$$

## References

- [1] arXiv:1210.7886 [astro-ph.GA]
- [2] Morgan, S., et al. 2007 MNRAS, 374, 1421
- [3] Kovács, G. 1998, MmSAI 69,49

## Results

The resulting light curve of SS Piscium



### Measurements

- Variation period  
 $T = 0.292 \pm 0.002$  day
- Apparent V-magnitude  
 $m_V = 11.09 \pm 0.15$  mag
- Absolute magnitude  
 $M_V = 1.13 \pm 0.01$  mag
- Iron Abundance  
 $[\text{Fe}/\text{H}] = -0.896 \pm 0.181 [\text{Fe}/\text{H}]_{\odot}$
- Heliocentric Distance  
 $d = 984.7 \pm 70.3$  pc