

Numerical Methods for Hyperbolic Equations

Sébastien Leclaire, ing., Ph.D.

Based on the following references:

- Gary A. Sod (1985) Numerical Methods in Fluid Dynamics. [<https://doi.org/10.1017/CBO9780511753138>]
- Eleuterio F. Toro (1999) Riemann solvers and numerical methods for fluid dynamics. [<https://doi.org/10.1007/b79761>]
- Randall J. LeVeque (2002) Finite Volume Methods for Hyperbolic Problems. [<https://doi.org/10.1017/CBO9780511791253>]
- Course notes given by late Prof. Paul Arminjon for MAT6165 (2006) UdeM.



Hyperbolic equations

- Classification
- Scalar Advection Equation
 - Characteristics
 - Finite Difference Scheme
 - Local Truncation Error
 - Stability Analysis
 - Courant-Friedrichs-Lewy condition (CFL condition)
 - Upwind Scheme
 - Donor-Cell Scheme
 - Lax-Friedrichs Scheme
 - Lax-Wendroff Scheme
- Two-Dimensional Scalar Advection Equation
 - Characteristics
 - Donor-Cell Upwind Scheme (2D)
 - Corner-Transport Upwind Scheme (2D)
 - Lax-Wendroff Scheme with dimensional splitting
 - Lax-Wendroff Scheme (2D)
 - Examples:
 - Bump function
 - Square wave
- Linear Hyperbolic Systems
 - Domain of dependence
 - Characteristics
 - Discontinuous Initial Condition
 - Scalar Advection Equation
 - Linear hyperbolic systems
- Conservation Law
 - Domain of Determinacy and Range of Influence
 - Rankine-Hugoniot Conditions
 - One-Dimensional Integral Form
 - Euler equations
 - One-Dimensional Euler equations
- **The Riemann Problem for the Euler equations**
 - **Problem Description and Form of the Solution**
 - **Solution for the pressure star and velocity star**
 - **Solution for all Possible States**
 - **Complete Solution**
 - **Godunov Scheme**
 - **Examples**



Hyperbolic equations

- Classification
- Scalar Advection Equation
- Two-Dimensional Scalar Advection Equation
- Linear Hyperbolic Systems
- Conservation Law
- The Riemann Problem for the Euler equations
 - Problem Description and Form of the Solution
 - Solution for the pressure star and velocity star
 - Solution for all Possible States
 - Complete Solution
 - Godunov Scheme
 - Examples

Riemann Problem Description

- The Riemann problem for the Euler equations is an initial value problem:

$$\mathbf{U}_t + \mathbf{F}(\mathbf{U})_x \equiv \frac{\partial}{\partial t} \begin{pmatrix} \rho \\ \rho u \\ E \end{pmatrix} + \frac{\partial}{\partial x} \begin{pmatrix} \rho u \\ \rho u^2 + p \\ u(E + p) \end{pmatrix} = \begin{pmatrix} 0 \\ 0 \\ 0 \end{pmatrix}, \quad -\infty < x < \infty, t > 0$$

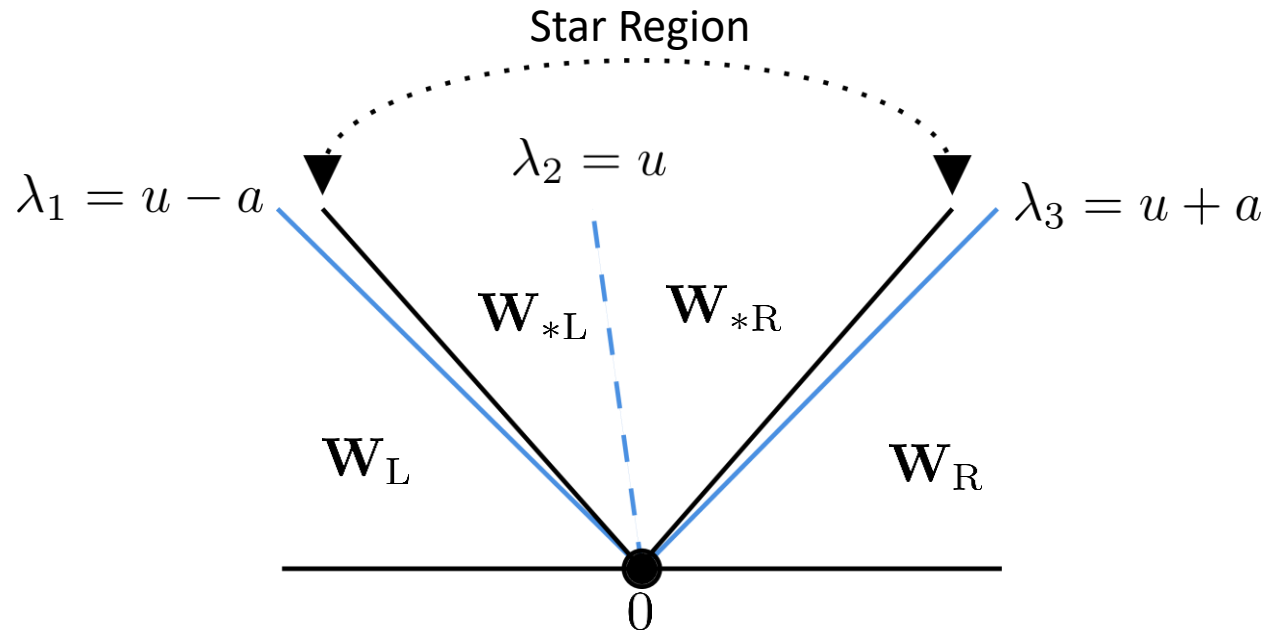
- The initial value is discontinuous:

$$\mathbf{U}(x, t = 0) = \begin{cases} \mathbf{U}_l & x < 0 \\ \mathbf{U}_r & x > 0 \end{cases}$$

- Let's define the primitive variable:

$$\mathbf{W} = (\rho, u, p)^\top$$

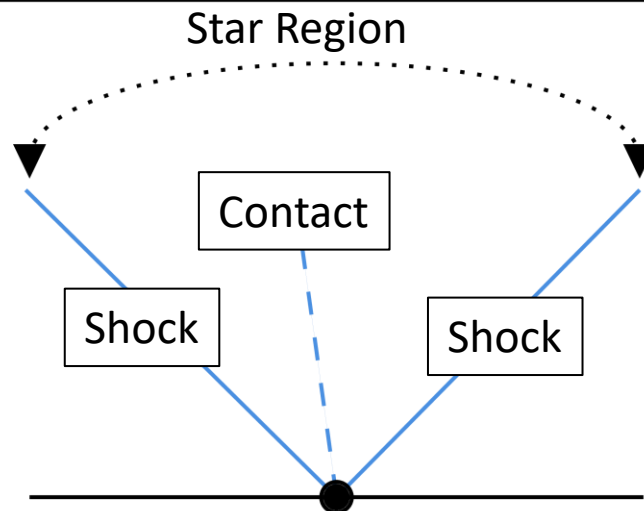
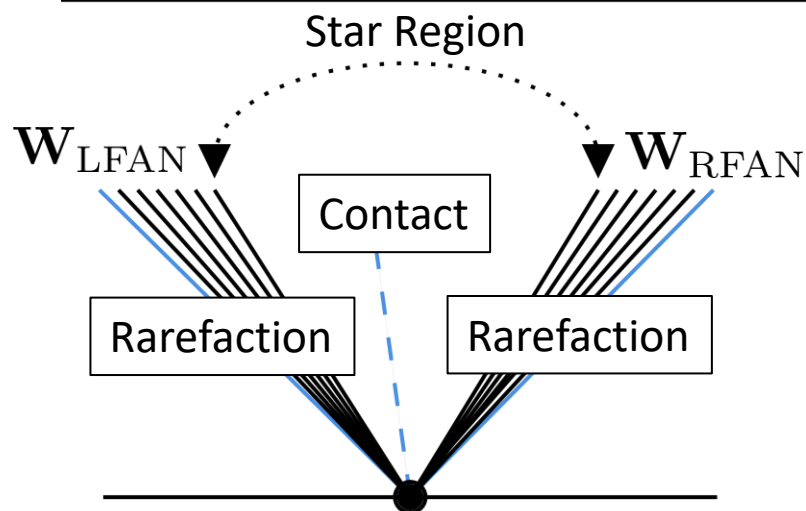
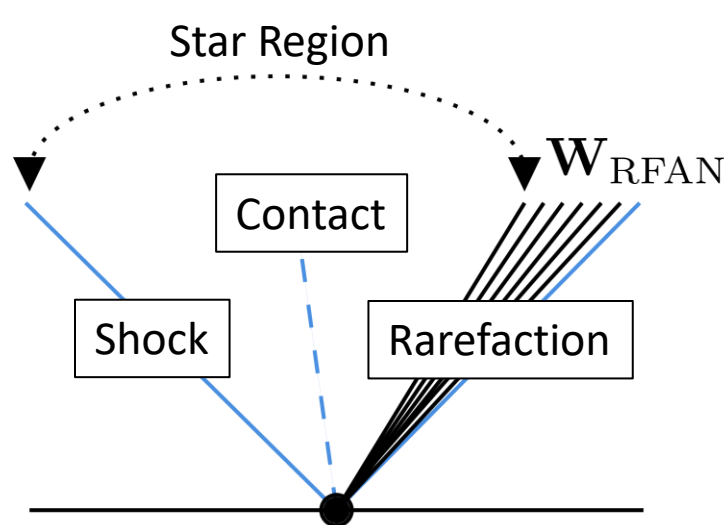
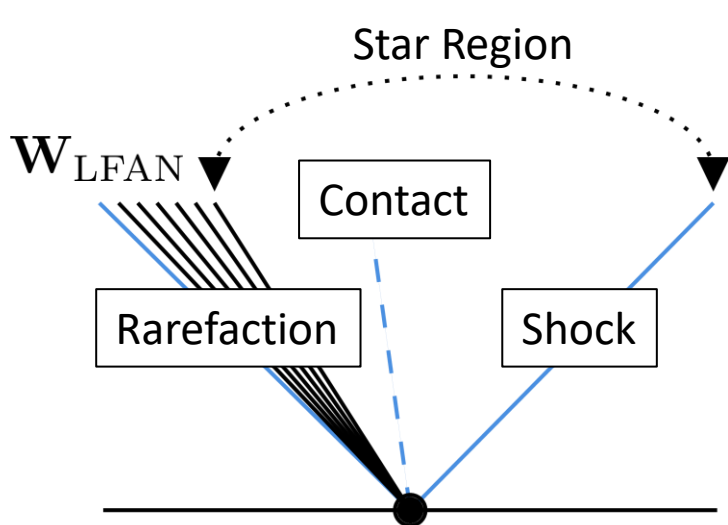
Form of the Solution



- Depending on the initial condition, the wave λ_1 and λ_3 could either be a shock wave or an expansion (rarefaction) wave.
- Between wave #1 and #3 there is the “star” region.
- The wave #2 always exist and is called a contact discontinuity.
- There is a jump in the solution at the contact discontinuity.



Form of the Solution – Principal Wave Patterns



Form of the Solution

- An analysis of the eigenstructure of the Euler equations could prove that the pressure as well as the velocity are continuous across the contact discontinuity:



$$p_{*L} = p_{*R} = p_*$$
$$u_{*L} = u_{*R} = u_*$$

- We thus need to find the values for:
 - p_* ;
 - u_* ;
 - ρ_{*L} ;
 - ρ_{*R} ; and if the case present
 - \mathbf{W}_{LFAN} and/or \mathbf{W}_{RFAN}

Solution for the pressure star and velocity star

- It is difficult and long to study in detail the proof of the following statement, so we will take it as granted.
- The solution for p_* with an ideal gas equation of state is given by the root of the following pressure equation [1]:

$$- f(p, \mathbf{W}_L, \mathbf{W}_R) \equiv f_L(p, \mathbf{W}_L) + f_R(p, \mathbf{W}_R) + u_R - u_L = 0 \text{ with:}$$

$$\bullet f_i(p, \mathbf{W}_i) = \begin{cases} (p - p_i) \left[\frac{A_i}{p + B_i} \right]^{\frac{1}{2}} & p > p_i \text{ (shock)} \\ \frac{2a_i}{\gamma - 1} \left[\left(\frac{p}{p_i} \right)^{\frac{\gamma - 1}{2\gamma}} - 1 \right] & p \leq p_i \text{ (rarefaction)} \end{cases}, i=L,R$$

- The left and right ($i=L,R$) speed of sound: $a_i = \sqrt{\frac{\gamma p_i}{\rho_i}}$
- The left and right ($i=L,R$) coefficients: $A_i = \frac{2}{(\gamma + 1)\rho_i}$ and $B_i = \frac{(\gamma - 1)}{(\gamma + 1)}p_i$
- The value for $u_* = \frac{1}{2} (u_L + u_R) + \frac{1}{2} [f_R(p_*, \mathbf{W}_R) - f_L(p_*, \mathbf{W}_L)]$

Numerical scheme to find the pressure star

- To solve the pressure equation for p_* a Newton-Raphson iterative procedure can be used:

$$p_{(k+1)} = p_{(k)} - \frac{f(p_k, \mathbf{W}_L, \mathbf{W}_R)}{f'(p_k, \mathbf{W}_L, \mathbf{W}_R)}$$

- A possible initial guess could be: $p_{(0)} = (p_L + p_R)/2$
- Numerical difficulty may arise and pressure could become negative after the first iteration if the initial pressure guess is too large.
- If $p_{(k+1)} < 0$, one possibility is to overwrite $p_{(k+1)}$ with another smaller initial guess:

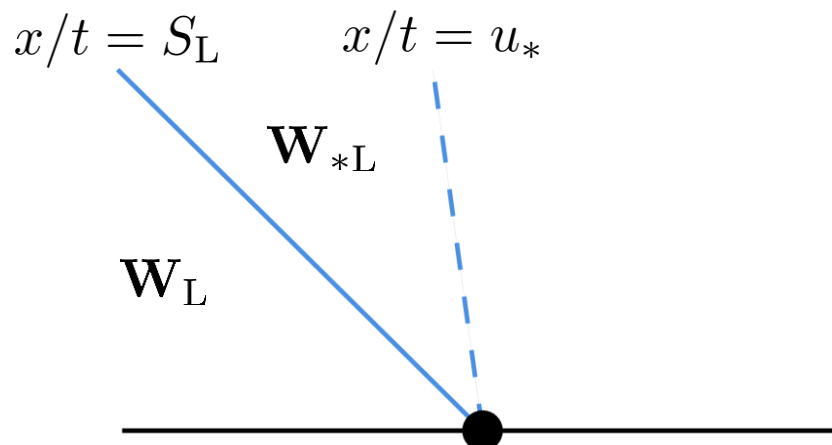
$$p_{(k+1)} \leftarrow \frac{1}{2^k} \left[\frac{a_L + a_R - (\gamma - 1)(u_R - u_L)/2}{\frac{a_L}{(p_L)^{\frac{\gamma-1}{2\gamma}}} + \frac{a_R}{(p_R)^{\frac{\gamma-1}{2\gamma}}}} \right]^{\frac{2\gamma}{\gamma-1}} \equiv \frac{1}{2^k} p_{\text{FAN}}$$

- The pressure p_{FAN} is the exact solution for p_* when wave #1 and wave #3 are rarefaction. A small enough initial guess is sufficient to always converges to the real solution.

Solution for all Possible States – Left Shock Wave

Left shock wave

$$p_* > p_L$$



- The left shock speed:
$$S_L = u_L - a_L \left[\frac{(\gamma + 1)}{2\gamma} \frac{p_*}{p_L} + \frac{(\gamma - 1)}{2\gamma} \right]^{\frac{1}{2}}$$
- The left density star for a shock:
$$\rho_{*L}^{\text{sho}} = \rho_L \left[\frac{\frac{p_*}{p_L} + \frac{(\gamma - 1)}{(\gamma + 1)}}{\frac{(\gamma - 1)}{(\gamma + 1)} \frac{p_*}{p_L} + 1} \right]$$

Solution for all Possible States – Left Rarefaction Wave

- The left density star for a rarefaction:

$$\rho_{*L}^{\text{rar}} = \rho_L (p_*/p_L)^{\frac{1}{\gamma}}$$

- The left speed of sound star:

$$a_{*L} = a_L (p_*/p_L)^{\frac{\gamma-1}{2\gamma}}$$

- The left tail speed:

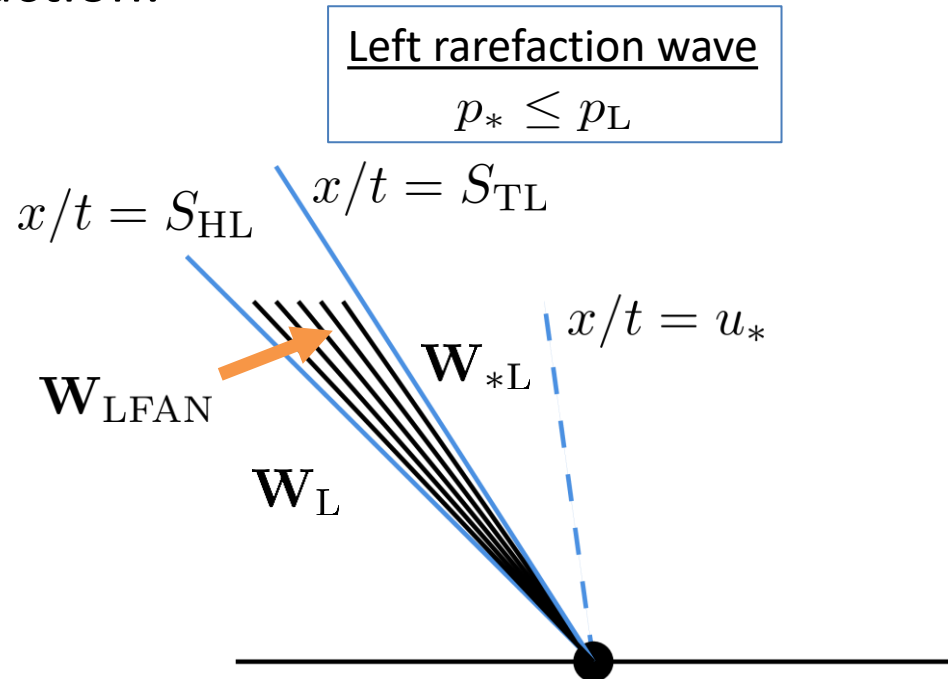
$$S_{TL} = u_* - a_{*L}$$

- The left head speed:

$$S_{HL} = u_L - a_L$$

- The values in the fan:

$$\mathbf{W}_{\text{LFAN}} = \begin{pmatrix} \rho \\ u \\ p \end{pmatrix}_{\text{LFAN}} = \begin{pmatrix} \rho_L \left[\frac{2}{\gamma+1} + \frac{(\gamma-1)}{(\gamma+1)a_L} \left(u_L - \frac{x}{t} \right) \right]^{\frac{2}{\gamma-1}} \\ \frac{2}{\gamma+1} \left[a_L + \frac{(\gamma-1)}{2} u_L + \frac{x}{t} \right] \\ p_L \left[\frac{2}{\gamma+1} + \frac{(\gamma-1)}{(\gamma+1)a_L} \left(u_L - \frac{x}{t} \right) \right]^{\frac{2\gamma}{\gamma-1}} \end{pmatrix}$$





Solution for all Possible States – Right Shock Wave

- Same formula as for the left shock wave but:
 - Need to replace left L with right R .
 - A sign in the right shock speed formula is different:

$$S_R = u_R \oplus a_R \left[\frac{(\gamma + 1)}{2\gamma} \frac{p_*}{p_R} + \frac{(\gamma - 1)}{2\gamma} \right]^{\frac{1}{2}}$$

- Same formula as for the left rarefaction wave but:

- Need to replace left L with right R.
- Some signs in the formula are different:

- The left tail speed: $S_{TR} = u_* \oplus a_{*R}$
- The left head speed: $S_{HR} = u_R \oplus a_R$
- The values in the fan:

$$\mathbf{W}_{\text{RFAN}} = \begin{pmatrix} \rho \\ u \\ p \end{pmatrix}_{\text{RFAN}} = \begin{pmatrix} \rho_R \left[\frac{2}{\gamma+1} \ominus \frac{(\gamma-1)}{(\gamma+1)a_R} \left(u_R - \frac{x}{t} \right) \right]^{\frac{2}{\gamma-1}} \\ \frac{2}{\gamma+1} \left[\ominus a_R + \frac{(\gamma-1)}{2} u_R + \frac{x}{t} \right] \\ p_R \left[\frac{2}{\gamma+1} \ominus \frac{(\gamma-1)}{(\gamma+1)a_R} \left(u_R - \frac{x}{t} \right) \right]^{\frac{2\gamma}{\gamma-1}} \end{pmatrix}$$



Complete Solution

- Let's define: $\mathbf{W}_{*i}^{\text{sho}} = [\rho_{*i}^{\text{sho}}, u_*, p_*]$, $i=L,R$
 $\mathbf{W}_{*i}^{\text{rar}} = [\rho_{*i}^{\text{rar}}, u_*, p_*]$

- The complete solution is:

$$\text{RP}(\mathbf{W}_L, \mathbf{W}_R, x/t) = \left\{ \begin{array}{l} \left. \begin{array}{l} \mathbf{W}_{*L}^{\text{sho}} \quad S_L \leq \frac{x}{t} \leq u_* \\ \mathbf{W}_L \quad \frac{x}{t} \leq S_L \\ \mathbf{W}_L \quad \frac{x}{t} \leq S_{HL} \\ \mathbf{W}_{LFAN} \quad S_{HL} \leq \frac{x}{t} \leq S_{TL} \\ \mathbf{W}_{*L}^{\text{rar}} \quad S_{TL} \leq \frac{x}{t} \leq u_* \end{array} \right\} \quad \left. \begin{array}{l} p_* > p_L \\ p_* \leq p_L \end{array} \right\} \quad \frac{x}{t} \leq u_* \\ \left. \begin{array}{l} \mathbf{W}_{*R}^{\text{sho}} \quad u_* \leq \frac{x}{t} \leq S_R \\ \mathbf{W}_R \quad \frac{x}{t} \geq S_R \\ \mathbf{W}_R \quad \frac{x}{t} \geq S_{HR} \\ \mathbf{W}_{RFAN} \quad S_{TR} \leq \frac{x}{t} \leq S_{HR} \\ \mathbf{W}_{*R}^{\text{rar}} \quad u_* \leq \frac{x}{t} \leq S_{TR} \end{array} \right\} \quad \left. \begin{array}{l} p_* > p_R \\ p_* \leq p_R \end{array} \right\} \quad \frac{x}{t} \geq u_* \end{array} \right.$$

- Note that this solution assumed no vacuum (i.e. $\rho > 0$) and extended solution may be found for that case.

- Initial-Boundary Value Problem (IBVP) over $[0, t_{\text{final}}] \times [0, L]$:

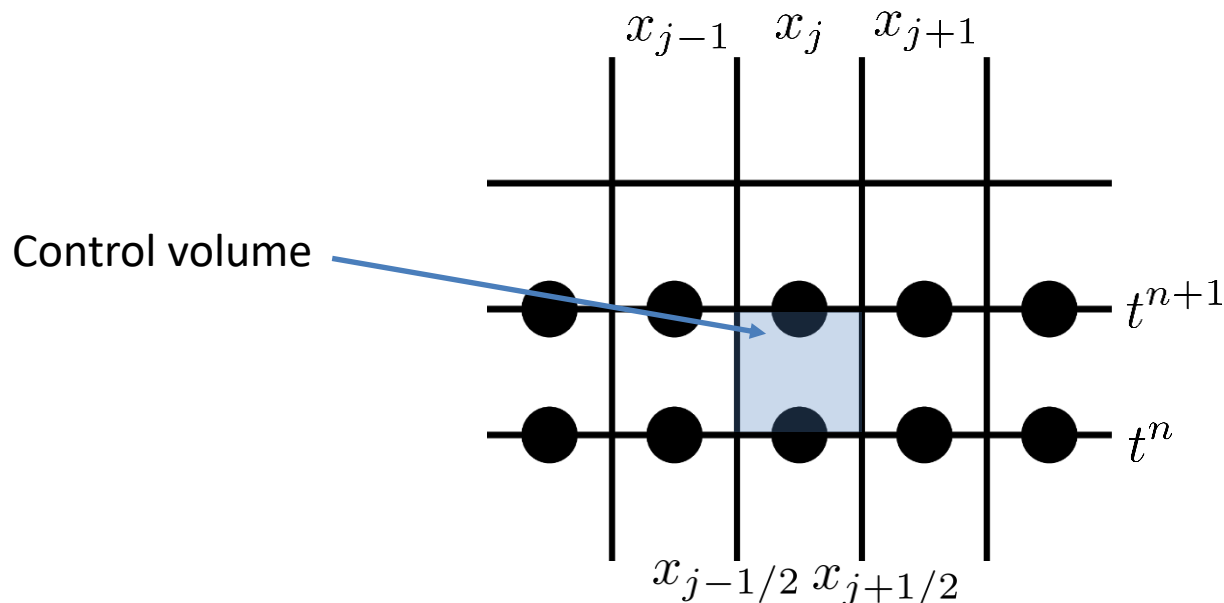
$$\text{PDE} : \mathbf{U}_t + \mathbf{F}(\mathbf{U})_x = 0$$

$$\text{IC} : \mathbf{U}(x, 0) = \mathring{\mathbf{U}}(x)$$

$$\text{BC} : \mathbf{U}(0, t) = \mathbf{U}_l(t)$$

$$\text{BC} : \mathbf{U}(L, t) = \mathbf{U}_r(t)$$

- Discretization of $[0, L]$ with cells centered in x_j for $j = 1, \dots, m$



- The initial condition is approximated with a constant piecewise function such that:


$$\mathring{\mathbf{U}}(x) = \mathring{\mathbf{U}}(x_j), \quad x_{j-1/2} < x < x_{j+1/2}, \quad j = 1, \dots, m$$

- We recall the one-dimensional integral form:

$$\int_{x_l}^{x_r} \mathbf{U}(x, t_2) dx = \int_{x_l}^{x_r} \mathbf{U}(x, t_1) dx + \int_{t_1}^{t_2} \mathbf{F}(\mathbf{U}(x_l, t)) dt - \int_{t_1}^{t_2} \mathbf{F}(\mathbf{U}(x_r, t)) dt$$

- We apply this form to each control volume $j = 1, \dots, m$:

$$\begin{aligned} \int_{x_{j-1/2}}^{x_{j+1/2}} \mathbf{U}(x, t^{n+1}) dx = \\ \int_{x_{j-1/2}}^{x_{j+1/2}} \mathbf{U}(x, t^n) dx + \int_{t^n}^{t^{n+1}} \mathbf{F}(\mathbf{U}(x_{j-1/2}, t)) dt - \int_{t^n}^{t^{n+1}} \mathbf{F}(\mathbf{U}(x_{j+1/2}, t)) dt \end{aligned}$$




$$\int_{x_{j-1/2}}^{x_{j+1/2}} \mathbf{U}(x, t^{n+1}) dx = \int_{x_{j-1/2}}^{x_{j+1/2}} \mathbf{U}(x, t^n) dx + \int_{t^n}^{t^{n+1}} \mathbf{F}(\mathbf{U}(x_{j-1/2}, t)) dt - \int_{t^n}^{t^{n+1}} \mathbf{F}(\mathbf{U}(x_{j+1/2}, t)) dt$$

- Let's define the cell averages:

$$\mathbf{U}_j^n = \frac{1}{\Delta x} \int_{x_{j-1/2}}^{x_{j+1/2}} \mathbf{U}(x, t^n) dx$$

$$\mathbf{U}_j^{n+1} = \frac{1}{\Delta x} \int_{x_{j-1/2}}^{x_{j+1/2}} \mathbf{U}(x, t^{n+1}) dx$$

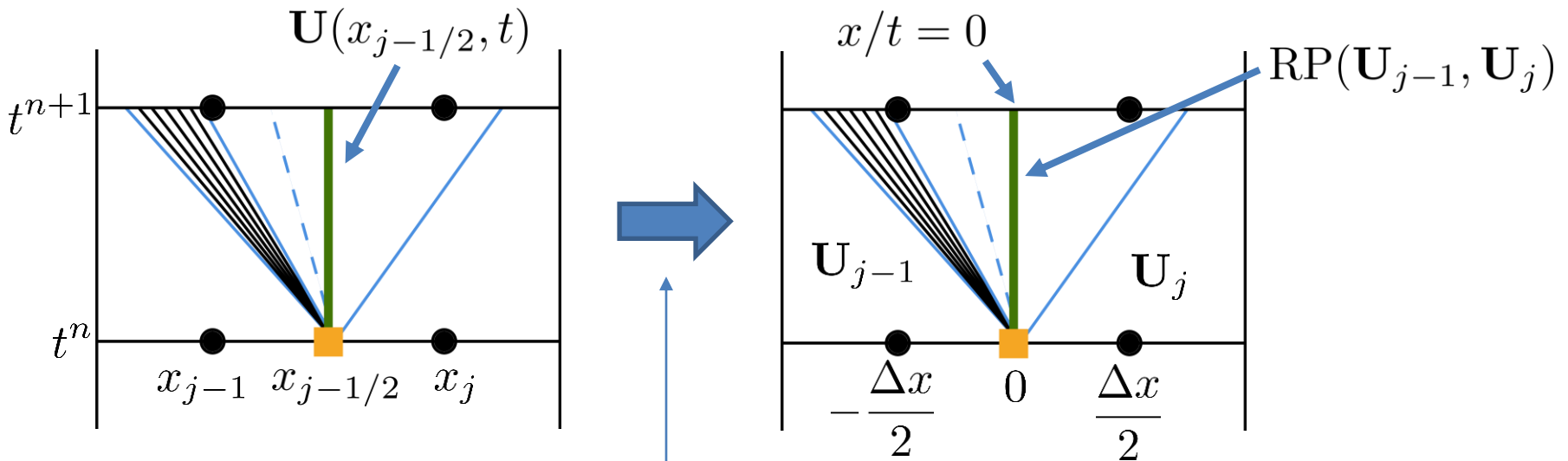
- Assuming a piecewise constant solution:



$$\mathbf{U}_j^{n+1} \Delta x = \mathbf{U}_j^n \Delta x + \int_{t^n}^{t^{n+1}} \mathbf{F}(\mathbf{U}(x_{j-1/2}, t)) dt - \int_{t^n}^{t^{n+1}} \mathbf{F}(\mathbf{U}(x_{j+1/2}, t)) dt$$

➔
$$\mathbf{U}_j^{n+1} \Delta x = \mathbf{U}_j^n \Delta x + \int_{t^n}^{t^{n+1}} \mathbf{F}(\mathbf{U}(x_{j-1/2}, t)) dt - \int_{t^n}^{t^{n+1}} \mathbf{F}(\mathbf{U}(x_{j+1/2}, t)) dt$$

- Assuming no wave interaction at $x_{j-1/2}$ for $t \in [t^n, t^{n+1}]$ imply:
 - $\mathbf{U}(x_{j-1/2}, t) = \text{RP}(\mathbf{U}_{j-1}, \mathbf{U}_j, x/t = 0) = \text{RP}(\mathbf{U}_{j-1}, \mathbf{U}_j)$
 - $\mathbf{U}(x_{j+1/2}, t) = \text{RP}(\mathbf{U}_j, \mathbf{U}_{j+1}, x/t = 0) = \text{RP}(\mathbf{U}_j, \mathbf{U}_{j+1})$



Change of coordinates

$$\Rightarrow \mathbf{U}_j^{n+1} \Delta x = \mathbf{U}_j^n \Delta x + \int_{t^n}^{t^{n+1}} \mathbf{F}(\mathbf{U}(x_{j-1/2}, t)) dt - \int_{t^n}^{t^{n+1}} \mathbf{F}(\mathbf{U}(x_{j+1/2}, t)) dt$$

$$\Rightarrow \mathbf{U}_j^{n+1} \Delta x = \mathbf{U}_j^n \Delta x + \int_{t^n}^{t^{n+1}} \mathbf{F}(\text{RP}(\mathbf{U}_{j-1}^n, \mathbf{U}_j^n)) dt - \int_{t^n}^{t^{n+1}} \mathbf{F}(\text{RP}(\mathbf{U}_j^n, \mathbf{U}_{j+1}^n)) dt$$

- The integrant is piecewise constant:

$$\Rightarrow \mathbf{U}_j^{n+1} \Delta x = \mathbf{U}_j^n \Delta x + \mathbf{F}(\text{RP}(\mathbf{U}_{j-1}^n, \mathbf{U}_j^n)) \Delta t - \mathbf{F}(\text{RP}(\mathbf{U}_j^n, \mathbf{U}_{j+1}^n)) \Delta t$$

$$\Rightarrow \text{Godunov scheme : } \mathbf{U}_j^{n+1} = \mathbf{U}_j^n + \frac{\Delta t}{\Delta x} \left(\mathbf{F}_{j-1/2}^n - \mathbf{F}_{j+1/2}^n \right)$$

– with: $\mathbf{F}_{j-1/2}^n = \mathbf{F}(\text{RP}(\mathbf{U}_{j-1}^n, \mathbf{U}_j^n))$
 $\mathbf{F}_{j+1/2}^n = \mathbf{F}(\text{RP}(\mathbf{U}_j^n, \mathbf{U}_{j+1}^n))$

- The CFL condition for this Godunov scheme is:

$$\Delta t \leq \frac{\Delta x}{S_{\max}^n}$$

- where S_{\max}^n is the fastest wave at time t^n on the whole computational domain.

- For the Euler equations, a good estimate for the fastest wave is:

$$S_{\max}^n \approx \max_{1 \leq j \leq m} \{|u_j^n| + a_j^n\}$$

- Stability may be improved by using a CFL number lower than 1:

$$\Delta t \leq \text{CFL} \frac{\Delta x}{S_{\max}^n}$$

- Typically $\text{CFL} = 0.9$ is a practical choice.

- Exact Riemann solution and Godunov method:

- Domain: $0 \leq x \leq 1$
- Diaphragm at $x_0 = 0.3$
- 100 cells
- $\gamma = 1.4$
- CFL = 0.9
- Left initial condition:

$$\mathbf{W}_L = (\rho_L, u_L, p_L)^T$$

$$= (1, 0.75, 1)^T$$

- Right initial condition:

$$\mathbf{W}_R = (\rho_R, u_R, p_R)^T$$

$$= (0.125, 0, 0.1)^T$$

