## 勉強したことのまとめ

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# 07/08/2024

#### 1.1 光の測地線方程式

粒子の場を表す方程式である、クラインゴルドン方程式

$$\left(\frac{1}{c^2}\frac{\partial^2}{\partial t^2} - \nabla^2 + \left(\frac{mc}{\hbar}\right)^2\right)\phi(\boldsymbol{x},t) = 0$$

質量のない光では

$$0 = \left(\frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \nabla^2\right) \phi(\mathbf{x}, t)$$
$$= \Box \phi(\mathbf{x}, t)$$

これを一般の時空に拡張すると

$$g^{ij}\nabla_i\nabla_j\phi=0$$

ここで、

$$\phi = Ce^{i\frac{s}{\epsilon}}$$

と書くと

$$\begin{split} \nabla_{j}\phi &= \nabla\left(Ce^{\frac{is}{\epsilon}}\right) \\ &= (\nabla_{j}C)\,e^{\frac{is}{\epsilon}} + \frac{iC}{\epsilon}e^{\frac{is}{\epsilon}}\,(\nabla_{j}S) \\ \nabla_{i}\nabla_{j}\phi &= \nabla_{i}\left((\nabla_{j}C)\,e^{\frac{is}{\epsilon}} + \frac{iC}{\epsilon}e^{\frac{is}{\epsilon}}\,(\nabla_{j}S)\right) \\ &= (\nabla_{i}\nabla_{j}C)\,e^{\frac{is}{\epsilon}} + (\nabla_{i}S)\,(\nabla_{j}C)\,\frac{2i}{\epsilon}e^{\frac{is}{\epsilon}} + (\nabla_{i}\nabla_{j}S)\,\frac{iC}{\epsilon}e^{\frac{is}{\epsilon}} - (\nabla_{i}S)\,(\nabla_{j}S)\,\frac{1}{\epsilon^{2}}e^{\frac{2is}{\epsilon}} \\ g^{ij}\nabla_{i}\nabla_{j}\phi &= g^{ij}\,(\nabla_{i}\nabla_{j}C)\,e^{\frac{is}{\epsilon}} + g^{ij}\,(\nabla_{i}S)\,(\nabla_{j}C)\,\frac{2i}{\epsilon}e^{\frac{is}{\epsilon}} + g^{ij}\,(\nabla_{i}\nabla_{j}S)\,\frac{iC}{\epsilon}e^{\frac{is}{\epsilon}} - g^{ij}\,(\nabla_{i}S)\,(\nabla_{j}S)\,\frac{1}{\epsilon^{2}}e^{\frac{2is}{\epsilon}} \\ &= (\Box C)\,e^{\frac{is}{\epsilon}} + (\nabla_{i}S)\,(\nabla^{i}C)\,\frac{2i}{\epsilon}e^{\frac{is}{\epsilon}} + (\Box S)\,\frac{iC}{\epsilon}e^{\frac{is}{\epsilon}} - (\nabla_{i}S)\,(\nabla^{i}S)\,\frac{1}{\epsilon^{2}}e^{\frac{2is}{\epsilon}} \\ &= \left(\Box C\right)e^{\frac{is}{\epsilon}} + (\nabla_{i}S)\,(\nabla^{i}C)\,\frac{2i}{\epsilon}e^{\frac{is}{\epsilon}} + (\Box S)\,\frac{iC}{\epsilon}e^{\frac{is}{\epsilon}} - (\nabla_{i}S)\,(\nabla^{i}S)\,\frac{1}{\epsilon^{2}}e^{\frac{2is}{\epsilon}} \\ &= \left(\Box C\right)e^{\frac{is}{\epsilon}} + (\nabla_{i}S)\,(\nabla^{i}C)\,\frac{2i}{\epsilon}e^{\frac{is}{\epsilon}} + (\Box S)\,\frac{iC}{\epsilon}e^{\frac{is}{\epsilon}} - (\nabla_{i}S)\,(\nabla^{i}S)\,\frac{1}{\epsilon^{2}}e^{\frac{2is}{\epsilon}} \\ &= \left(\Box C\right)e^{\frac{is}{\epsilon}} + (\nabla_{i}S)\,(\nabla^{i}C)\,\frac{2i}{\epsilon}e^{\frac{is}{\epsilon}} + (\Box S)\,\frac{iC}{\epsilon}e^{\frac{is}{\epsilon}} - (\nabla_{i}S)\,(\nabla^{i}S)\,\frac{1}{\epsilon^{2}}e^{\frac{2is}{\epsilon}} \\ &= \left(\Box C\right)e^{\frac{is}{\epsilon}} + (\nabla_{i}S)\,(\nabla^{i}C)\,\frac{2i}{\epsilon}e^{\frac{is}{\epsilon}} + (\Box S)\,\frac{iC}{\epsilon}e^{\frac{is}{\epsilon}} - (\nabla_{i}S)\,(\nabla^{i}S)\,\frac{1}{\epsilon^{2}}e^{\frac{2is}{\epsilon}} \\ &= \left(\Box C\right)e^{\frac{is}{\epsilon}} + (\nabla_{i}S)\,(\nabla^{i}C)\,\frac{2i}{\epsilon}e^{\frac{is}{\epsilon}} + (\Box S)\,\frac{iC}{\epsilon}e^{\frac{is}{\epsilon}} - (\nabla_{i}S)\,(\nabla^{i}S)\,\frac{1}{\epsilon^{2}}e^{\frac{2is}{\epsilon}} \\ &= \left(\Box C\right)e^{\frac{is}{\epsilon}} + (\nabla_{i}S)\,(\nabla^{i}C)\,\frac{2i}{\epsilon}e^{\frac{is}{\epsilon}} + (\Box S)\,\frac{iC}{\epsilon}e^{\frac{is}{\epsilon}} - (\nabla_{i}S)\,(\nabla^{i}S)\,\frac{1}{\epsilon^{2}}e^{\frac{2is}{\epsilon}} \\ &= \left(\Box C\right)e^{\frac{is}{\epsilon}} + (\nabla_{i}S)\,(\nabla^{i}C)\,(\nabla^{i}S)$$

ここで、

$$(\nabla_i S) \left( \nabla^i S \right) = 0$$

は光の測地線方程式を表している。

$$\nabla_i S = k_i$$

とおけば

$$\begin{split} 0 &= (\nabla_i S) \left( \nabla^i S \right) \\ &= k_i k^i \\ &= \nabla_j (k_i k^i) \\ &= \nabla_j (k_i g^{ij} k_l) \\ &= g^{il} (\nabla_j k_l) k_i + g^{il} (\nabla_j k_i) k_l \\ &= 2 (\nabla_j k_i) k^i \\ &= 2 g^{lj} (\nabla_j \nabla_i S) k^i \\ &= 2 (\nabla_i g^{lj} k_j) k^i \\ &= 2 \left( \nabla_i k^l \right) k^i \\ &= 2 \left( \frac{\partial}{\partial x^i} \left( \frac{dx^l}{d\lambda} \right) + \Gamma^l_{im} \left( \frac{dx^m}{d\lambda} \right) \right) \frac{dx^i}{d\lambda} \end{split}$$

光の測地線方程式

$$0 = \left(\frac{d^2x^l}{d\lambda^2}\right) + \Gamma^l_{\ im} \left(\frac{dx^i}{d\lambda}\frac{dx^m}{d\lambda}\right)$$

### 1.2 積分

$$\left(\frac{d\mu}{d\phi}\right)^{2} = 2MG(\mu)$$

$$\frac{d\mu}{d\phi} = -\sqrt{2MG(\mu)}$$

$$\begin{cases}
\mu_{1} = -\frac{Q + 2M - P}{4MP} \\
\mu_{2} = -\frac{1}{P} \\
\mu_{3} = -\frac{Q - 2M + P}{4MP}
\end{cases}$$
(1.2)

$$\int_0^{\phi_{\infty}} d\phi = \int_{\mu_2}^0 \frac{d\mu}{-\sqrt{2MG(\mu)}}$$
$$\phi_{\infty} = \frac{1}{\sqrt{2M}} \int_0^{\mu_2} d\mu \left( (\mu - \mu_1)(\mu - \mu_2)(\mu - \mu_3) \right)^{-\frac{1}{2}}$$

ここで、

$$\chi = \sqrt{\mu - \mu_1}$$
$$d\chi = \frac{1}{2}(\mu - \mu_1)^{-\frac{1}{2}}d\mu$$

とおくと

$$\phi_{\infty} = \frac{1}{\sqrt{2M}} \int_{\sqrt{-\mu_1}}^{\sqrt{\mu_2 - \mu_1}} \frac{2\chi d\chi}{\sqrt{\chi^2(\chi^2 + \mu_1 - \mu_2)(\chi^2 + \mu - \mu_3)}}$$

ここで、

$$\chi = \sqrt{\mu_2 - \mu_1} \eta$$
$$d\chi = \sqrt{\mu_2 - \mu_1} d\eta$$

とおくと

$$\phi_{\infty} = \frac{2}{\sqrt{2M}} \int_{\sqrt{\frac{\mu_1}{\mu_1 - \mu_2}}}^{1} \frac{\sqrt{\mu_2 - \mu_1} d\eta}{\sqrt{-(\mu_2 - \mu_1)(1 - \eta^2)((\mu_2 - \mu_1)\eta^2 - \mu_1 + \mu_3)}}$$

ここで、

$$\eta = \sin x$$
$$d\eta = dx \cos x$$

とおくと

$$\phi_{\infty} = \frac{2}{\sqrt{2M}} \int_{\sin^{-1}}^{\frac{\pi}{2}} \frac{\cos x dx}{\sqrt{\cos^{2} x} \sqrt{-(\mu_{2} - \mu_{1}) \sin^{2} x ((\mu_{2} - \mu_{1}) \eta^{2} - \mu_{1} + \mu_{3})}}$$

$$= \frac{2}{\sqrt{2M}} \int_{\sin^{-1}}^{\frac{\pi}{2}} \frac{dx}{\sqrt{\mu_{3} - \mu_{1}} \sqrt{-\frac{\mu_{2} - \mu_{1}}{\mu_{3} - \mu_{1}} \sin^{2} x + 1)}}$$

$$\begin{cases}
\mu_3 - \mu_1 = \frac{Q}{2MP} \\
\mu_2 - \mu_1 = \frac{Q + 6M - P}{4MP}
\end{cases}$$
(1.3)

$$\begin{cases} \frac{\mu_2 - \mu_1}{\mu_3 - \mu_1} = \frac{Q + 6M - P}{2Q} \\ \frac{\mu_1}{\mu_1 - \mu_2} = \frac{Q + 2M - P}{Q + 6M - P} \end{cases}$$
(1.4)

$$\phi_{\infty} = \frac{2}{\sqrt{2M}} \int_{\sin^{-1}}^{\frac{\pi}{2}} \frac{dx}{\sqrt{\frac{Q}{Q+6M-P}}} \frac{dx}{\sqrt{\frac{Q}{2MP}} \sqrt{1 - \frac{Q+6M-P}{2Q} \sin^2 x}}$$

ここで、

$$\kappa = \sqrt{\frac{Q + 6M - P}{2Q}}$$
 
$$\zeta = \sin^{-1} \sqrt{\frac{Q + 2M - P}{Q + 6M - P}}$$

さらに

$$K(\kappa) = \int_0^{\frac{\pi}{2}} \frac{dx}{\sqrt{1 - \kappa^2 \sin^2 x}}$$
$$F(\zeta, \kappa) = \int_0^{\zeta} \frac{dx}{\sqrt{1 - \kappa^2 \sin^2 x}}$$

とおくと

## $\phi_\infty$ についての式

$$\phi_{\infty} = 2\sqrt{\frac{P}{Q}} \int_{\zeta}^{\frac{\pi}{2}} \frac{dx}{\sqrt{1 - \kappa^2 \sin^2 x}}$$
$$= 2\sqrt{\frac{P}{Q}} \left( K(\kappa) - F(\zeta, \kappa) \right)$$

# 28/08/2024

#### 2.1 $P \gg M$

$$\begin{cases} Q \approx P \left( 1 + 2 \left( \frac{M}{P} \right) - 8 \left( \frac{M}{P} \right)^2 \right) \\ \kappa^2 \approx 4 \left( \frac{M}{P} \right) \\ \zeta \approx \frac{\pi}{4} - \frac{1}{2} \left( \frac{M}{P} \right) \end{cases}$$

以上を考慮して $\phi_{\infty}$ を計算する

$$\begin{split} \phi_{\infty} &= 2\sqrt{\frac{P}{Q}} \int_{\zeta}^{\frac{\pi}{2}} \frac{dx}{\sqrt{1 - \kappa^2 \sin^2 x}} \\ &\approx 2\sqrt{\frac{P}{Q}} \int_{\zeta}^{\frac{\pi}{2}} \left(1 + \frac{1}{2}\kappa^2 \sin^2 x\right) dx \\ &= 2\sqrt{\frac{P}{Q}} \int_{\zeta}^{\frac{\pi}{2}} \left(1 + \frac{1}{2}\kappa^2 - \frac{1}{4}\kappa^2 \cos 2x\right) dx \\ &\approx 2\left(1 - \frac{M}{P}\right) \int_{\frac{\pi}{4} - \frac{M}{2P}}^{\frac{\pi}{2}} \left(1 + 2\frac{M}{P} - \frac{M}{P} \cos 2x\right) dx \\ &= 2\left(1 - \frac{M}{P}\right) \left[x + \frac{M}{P}x - \frac{M}{2P} \sin 2x\right]_{\frac{\pi}{4} - \frac{M}{2P}}^{\frac{\pi}{2}} \\ &= \frac{\pi}{2} + \frac{2M}{P} \end{split}$$

### 2.2 $P \approx 3M(1+\epsilon)$

$$\begin{cases} Q \approx 3M + 5M\epsilon \\ \kappa \approx 1 - \frac{2}{3}\epsilon \\ \zeta \approx \sin^{-1}\sqrt{\frac{1}{3}} + \frac{\sqrt{3}}{9\cos\left(\sin^{-1}\sqrt{\frac{1}{3}}\right)}\epsilon \\ \sqrt{\frac{P}{Q}} \approx 1 - \frac{1}{3}\epsilon \\ \sqrt{1 - \kappa^2} \approx \sqrt{\frac{4}{3}\epsilon} \\ b \approx b_c \left(1 + \frac{3}{2}\epsilon^2\right) \end{cases}$$

#### $\kappa \approx 1$ での公式

$$K(\kappa) = \ln\frac{4}{\kappa'} + \left(\frac{1}{2}\right)^2 \left(\ln\frac{4}{\kappa'} - \frac{2}{1\cdot 2}\right) \kappa'^2 + \left(\frac{1\cdot 3}{2\cdot 4}\right)^2 \left(\ln\frac{4}{\kappa'} - \frac{2}{1\cdot 2} - \frac{2}{3\cdot 4}\right) \kappa'^4 + \dots$$

$$F(\zeta, \kappa) = \frac{2}{\pi} \mathbf{K'} \ln \tan\left(\frac{\zeta}{2} + \frac{\pi}{4}\right) - \frac{\tan\zeta}{\cos\zeta} \left(a'_0 - \frac{2}{3}a'_1 \tan^2\zeta + \dots\right)$$

$$\begin{cases} a'_0 = \frac{2}{\pi} \mathbf{K'} - 1\\ a'_n = a'_{n-1} - \left[\frac{(2n-1)!!}{2^n n!}\right]^2 \kappa'^{2n} \end{cases}$$

$$\mathbf{K'} = \frac{\pi}{2} \left(1 + \left(\frac{1}{2}\right)^2 \kappa'^2 + \dots\right)$$

$$\kappa' = \sqrt{1 - \kappa^2}$$

以上を考慮してbを計算する

$$K(\kappa) \approx \ln \frac{4}{\kappa'}$$
  
 $F(\zeta, \kappa) \approx \ln \tan \left(\frac{\zeta}{2} + \frac{\pi}{4}\right)$ 

$$\phi_{\infty} = 2\sqrt{\frac{P}{Q}} \left( K(\kappa) - F(\zeta, \kappa) \right)$$

$$\approx 2 \left( 1 - \frac{\epsilon}{3} \right) \left( \ln \frac{4}{\kappa'} - \ln \tan \left( \frac{\zeta}{2} + \frac{\pi}{4} \right) \right)$$

$$= \left( 1 - \frac{\epsilon}{3} \right) \ln \left( \frac{\frac{12}{\epsilon}}{\tan^2 \left( \frac{\sin^{-1} \sqrt{\frac{1}{3}}}{2} + \frac{\pi}{4} \right)} \right)$$

$$\approx \ln \left( \frac{12}{\tan^2 \left( \frac{\sin^{-1} \sqrt{\frac{1}{3}}}{2} + \frac{\pi}{4} \right) \epsilon} \right) = \ln \frac{\alpha}{\epsilon}$$

ここで、

$$\phi_{\infty} = \frac{1}{2}\pi + \frac{1}{2}\mu$$

より、

$$\frac{\alpha}{\epsilon} = \exp\left(\frac{\pi}{2} + \frac{\mu}{2}\right)$$
$$\epsilon = \frac{\alpha}{\exp\left(\frac{\pi}{2} + \frac{\mu}{2}\right)}$$

したがって

$$b \approx b_c \left( 1 + \frac{3}{2} \epsilon^2 \right)$$

$$= b_c \left( 1 + \frac{3}{2} \left( \frac{\alpha}{\exp\left(\frac{\pi}{2} + \frac{\mu}{2}\right)} \right)^2 \right)$$

$$= b_c \left( 1 + \frac{3}{2} \alpha^2 e^{-\mu} e^{-\pi} \right)$$

$$\approx 5.19615M + 3.48228Me^{-\mu}$$

# 23/09/2024

### 3.1 3次式の解

 $b^2 = \frac{P^3}{P - 2M}$ 

この P を b について解きたい

$$\begin{cases} P = yM \\ b = xM \end{cases}$$

と置き、y を x について解く問題を考える。

$$0 = P^{3} - Pb^{2} + 2Mb^{2}$$

$$0 = y^{3}M^{3} - yx^{2}M^{3} + 2x^{2}M^{3}$$

$$0 = y^{3} - yx^{2} + 2x^{2}$$

$$y = u + v$$

とおくと

$$0 = (u+v)^3 - (u+v)x^2 + 2x^2$$
$$0 = (u^3 + v^3 + 2x^2) + (3uv - x^2)(u+v)$$
$$u+v \neq 0$$

と仮定すると

$$\begin{cases} u^3 + v^3 + 2x^2 = 0\\ 3uv - x^2 = 0 \end{cases}$$
 (3.1)

以上より、片方の符号を採用すると

$$\begin{cases} u^{3} = -x^{2} + x^{2} \sqrt{1 - \left(\frac{x}{3\sqrt{3}}\right)^{2}} \\ v^{3} = -x^{2} - x^{2} \sqrt{1 - \left(\frac{x}{3\sqrt{3}}\right)^{2}} \end{cases}$$
(3.2)

ここで (u,v) の 3 乗根の主値をそれぞれ  $(\alpha,\beta)$  とおくと

$$\alpha = \left(-x^2 + x^2 \sqrt{1 - \left(\frac{x}{3\sqrt{3}}\right)^2}\right)^{\frac{1}{3}}$$

$$\beta = \left(-x^2 - x^2 \sqrt{1 - \left(\frac{x}{3\sqrt{3}}\right)^2}\right)^{\frac{1}{3}}$$

$$\omega = \exp\left(\frac{2}{3}\pi i\right)$$

(1)  $D < 0 \text{ Tit } (x > 3\sqrt{3})$ 

$$(u,v) = \begin{cases} (\alpha,\beta) \\ (\omega\alpha,\omega^2\beta) \\ (\omega^2\alpha,\omega\beta) \end{cases}$$

(2) D = 0 では  $(x = 3\sqrt{3})$ 

$$(u,v) = \begin{cases} (\omega\alpha, \omega\beta) \\ (\alpha, \omega^2\beta) \\ (\omega^2\alpha, \beta) \end{cases}$$

(3) D > 0 では  $(x < 3\sqrt{3})$ 

$$(u,v) = \begin{cases} (\omega\alpha, \omega\beta) \\ (\alpha, \omega^2\beta) \\ (\omega^2\alpha, \beta) \end{cases}$$

# 02/12/2024

#### 4.1 Buchdahl 時空における光の軌道

Buchdahl 時空の計量は以下のように与えられる。

$$g_{ij} = \operatorname{diag}\left(-\frac{(1-f(r))^2}{(1+f(r))^2}, (1+f(r))^4, r^2(1+f(r))^4, r^2(1+f(r))^4 \sin^2\theta\right)$$
$$f(r) = \frac{a}{2\sqrt{1+kr^2}}$$

測地線を考えるには、赤道面  $\theta = \frac{\pi}{2}$  に注目すれば

$$S = \int \mathcal{L}d\tau$$

$$= \int \left( -\frac{(1 - f(r))^2}{(1 + f(r))^2} \dot{t}^2 + (1 + f(r))^4 \dot{r}^2 + r^2 (1 + f(r))^4 \dot{\phi}^2 \right) d\tau$$

この変分を計算すれば良い。 $\mathcal{L}$  は  $t,\phi$  をあらわには含まないので、オイラーラグランジュ方程式より

$$0 = \left(\frac{d\mathcal{L}}{d\dot{t}}\right) = \frac{-2(1 - f(r))^2}{(1 + f(r))^2}\dot{t}$$

$$\therefore E = \frac{(1 - f(r))^2}{(1 + f(r))^2}\dot{t}$$

$$0 = \left(\frac{d\mathcal{L}}{d\dot{\phi}}\right) = 2r^2(1 + f(r))^4\dot{\phi}$$

$$\therefore L = r^2(1 + f(r))^4\dot{\phi}$$

また、null の条件  $0 = g_{ij}\dot{x}^i\dot{x}^j$  より

$$0 = -\frac{(1 - f(r))^2}{(1 + f(r))^2} \dot{t}^2 + (1 + f(r))^4 \dot{r}^2 + r^2 (1 + f(r))^4 \dot{\phi}^2$$
$$\therefore \left(\frac{E}{L}\right)^2 = \frac{(1 - f(r))^2}{r^4 (1 + f(r))^6} \left(\frac{dr}{d\phi}\right)^2 + \frac{(1 - f(r))^2}{r^2 (1 + f(r))^6}$$

ここで  $u=\frac{1}{r},\ b=\frac{L}{E}$  とおくと

$$1 = r^4 \left(\frac{du}{dr}\right)^2$$

より

$$\left(\frac{1}{b}\right)^2 = \frac{(1 - f(1/u))^2}{(1 + f(1/u))^6} \left(\frac{du}{d\phi}\right)^2 + \frac{u^2(1 - f(1/u))^2}{(1 + f(1/u))^6}$$
$$\therefore \left(\frac{du}{d\phi}\right)^2 = G(u) := \left(\frac{1}{b}\right)^2 \frac{(1 + f(1/u))^6}{(1 - f(1/u))^2} - u^2$$

### 4.2 パラメータの制約

f(r) のとりうる範囲を確認しておく。

$$0 \le f(r) \le \frac{a}{2}$$

いくつかの条件から、パラメータの制約を考える

$$\rho(r) = \frac{24\pi k f(r)^5}{\pi a^4 (1 + f(r))^5}$$
$$p(r) = \frac{8k f(r)^6}{\pi a^4 (1 - f(r)^2)(1 + f(r))^4}$$

1) weak energy condition (WEC)

$$0 \le \rho(r)$$
$$0 \le \rho(r) + p(r)$$

解析すると

$$0 \le 1 + \frac{p(r)}{\rho(r)}$$

$$= 1 + \frac{8kf(r)^6}{\pi a^4 (1 - f(r)^2)(1 + f(r))^4} \times \frac{\pi a^4 (1 - f(r)^2)(1 + f(r))^4}{8kf(r)^6}$$

$$= 1 + \frac{f(r)}{3(1 - f(r))}$$

$$\therefore f(r) \le 1, \frac{3}{2} \le f(r)$$

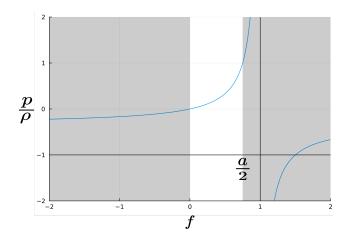


Figure 4.1: WEC

ここでのaの制約は

$$\frac{a}{2} \leq 1$$

2) strong energy condition (SEC)

$$0 \le \rho(r)$$
$$0 \le \rho(r) + 3p(r)$$

解析すると

$$0 \le 1 + \frac{3p(r)}{\rho(r)}$$
$$= 1 + \frac{f(r)}{1 - f(r)}$$
$$\therefore f(r) \le 1$$

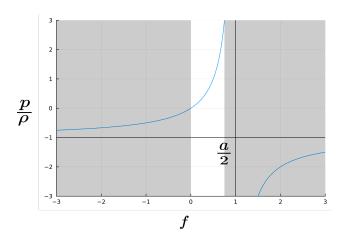


Figure 4.2: SEC

ここでのaの制約は

$$\frac{a}{2} \leq 1$$

3) dominant energy condition (DEC)

$$0 \le \rho(r)$$
$$0 \le \rho(r) + |p(r)|$$

解析すると

$$-1 \le \frac{p(r)}{\rho(r)} \le 1$$
$$-1 \le \frac{f(r)}{3(1 - f(r))} \le 1$$
$$\therefore f(r) \le \frac{3}{4}, \frac{3}{2} \le f(r)$$

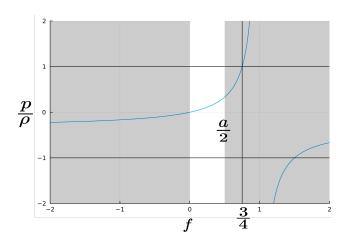


Figure 4.3: DEC

ここでの a の制約は

$$\frac{a}{2} \le \frac{3}{4}$$

以上より、パラメータの制約は

$$0 \le a \le \frac{3}{2}$$

# 16/12/2024

### 5.1 Buchdahl 時空の曲率

$$g_{ab} = \begin{pmatrix} -\frac{(1-f(r))^2}{(1+f(r))^2} \\ (1+f(r))^4 \end{pmatrix}$$

$$f(r) = \frac{a}{2\sqrt{1+kr^2}}$$

$$\frac{df}{dr} = -\frac{ar}{2(1+kr^2)^{\frac{3}{2}}} = -\frac{af(r)}{1+kr^2}$$

$$\Gamma^a_{bc} = \frac{1}{2}g^{ad}\left(\frac{\partial g_{db}}{\partial x^c} + \frac{\partial g_{dc}}{\partial x^b} - \frac{\partial g_{bc}}{\partial x^d}\right)$$

$$R^a_{bcd} = \frac{\partial \Gamma^a_{bd}}{\partial x^c} - \frac{\partial \Gamma^a_{bc}}{\partial x^d} + \Gamma^a_{ec}\Gamma^e_{bd} - \Gamma^a_{ed}\Gamma^a_{bc}$$

$$\frac{g_{tt}}{\partial x^r} = -\frac{af(r)}{1+kr^2} \cdot \frac{4(1-f(r))}{(1+f(r))^3}$$

$$\frac{g_{rr}}{\partial x^r} = -\frac{af(r)}{1+kr^2} \cdot 4(1+f(r))^3$$

$$\frac{g_{\phi\phi}}{\partial x^r} = 2r(1+f(r)) - \frac{af(r)}{1+kr^2} \cdot 4r^2(1+f(r))$$

 $\begin{pmatrix} a & b \\ c & d \end{pmatrix}$ 

$$\bar{w}_j = () \int$$

メモ用

1 = 0

$$\begin{cases} 1 = 0 \\ 1 = 0 \end{cases} \tag{5.1}$$

(1)