Archive-name: space/constants

Last-modified: \$Date: 93/04/01 14:39:04 \$

This list was originally compiled by Dale Greer. Additions would be appreciated.

Numbers in parentheses are approximations that will serve for most blue-skying purposes.

Unix systems provide the 'units' program, useful in converting between different systems (metric/English, etc.)

7726 m/s (8000) -- Earth orbital velocity at 300 km altitude

3075 m/s (3000) -- Earth orbital velocity at 35786 km (geosync)

6371 km (6400) -- Mean radius of Earth

6378 km (6400) -- Equatorial radius of Earth

1738 km (1700) -- Mean radius of Moon

5.974e24 kg (6e24) -- Mass of Earth

7.348e22 kg (7e22) -- Mass of Moon

1.989e30 kg (2e30) -- Mass of Sun

3.986e14 m³/s² (4e14) -- Gravitational constant times mass of Earth

4.903e12 m³/s² (5e12) -- Gravitational constant times mass of Moon

1.327e20 m³/s² (13e19) -- Gravitational constant times mass of Sun

384401 km (4e5) -- Mean Earth-Moon distance

1.496e11 m (15e10) -- Mean Earth-Sun distance (Astronomical Unit)

1 megaton (MT) TNT = about 4.2e15 J or the energy equivalent of

about .05 kg (50 gm) of matter. Ref: J.R Williams, "The Energy Level

of Things", Air Force Special Weapons Center (ARDC), Kirtland Air

Force Base, New Mexico, 1963. Also see "The Effects of Nuclear

Weapons", compiled by S. Glasstone and P.J. Dolan, published by the

US Department of Defense (obtain from the GPO).

Where d is distance, v is velocity, a is acceleration, t is time.

Additional more specialized equations are available from:

ames.arc.nasa.gov:pub/SPACE/FAQ/MoreEquations

For constant acceleration

$$d = d0 + vt + .5at^2$$

$$v = v0 + at$$

$$v^2 = 2ad$$

Acceleration on a cylinder (space colony, etc.) of radius r and

rotation period t:

$$a = 4 pi**2 r / t^2$$

For circular Keplerian orbits where:

Vc = velocity of a circular orbit

Vesc = escape velocity

M = Total mass of orbiting and orbited bodies

G = Gravitational constant (defined below)

u = G * M (can be measured much more accurately than G or M)

$$K = -G * M / 2 / a$$

r = radius of orbit (measured from center of mass of system)

V = orbital velocity

P = orbital period

a = semimajor axis of orbit

$$Vesc = sqrt(2 * M * G / r) = sqrt(2) * Vc$$

$$V^2 = u/a$$

$$P = 2 pi/(Sqrt(u/a^3))$$

$$K = 1/2 V^{**}2 - G * M / r$$
 (conservation of energy)

The period of an eccentric orbit is the same as the period

of a circular orbit with the same semi-major axis.

Change in velocity required for a plane change of angle phi in a

circular orbit:

delta V = 2 sqrt(GM/r) sin (phi/2)

Energy to put mass m into a circular orbit (ignores rotational

velocity, which reduces the energy a bit).

GMm (1/Re - 1/2Rcirc)

Re = radius of the earth

Rcirc = radius of the circular orbit.

Classical rocket equation, where

dv= change in velocity

Isp = specific impulse of engine

Ve= exhaust velocity

x= reaction mass

m1= rocket mass excluding reaction mass

Ve= Isp * g

dv = Ve * ln((m1 + x) / m1)

= Ve * In((final mass) / (initial mass))

Relativistic rocket equation (constant acceleration)

t (unaccelerated) = c/a * sinh(a*t/c)

 $d = c^{**}2/a * (cosh(a^*t/c) - 1)$

v = c * tanh(a*t/c)

Relativistic rocket with exhaust velocity Ve and mass ratio MR:

at/c = Ve/c * In(MR), or

t (unaccelerated) = c/a * sinh(Ve/c * In(MR))

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d = c^{**}2/a * (cosh(Ve/C * ln(MR)) - 1)
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$$v = c * tanh(Ve/C * In(MR))$$

Converting from parallax to distance:

d (in parsecs) = 1 / p (in arc seconds)

d (in astronomical units) = 206265 / p

Miscellaneous

f=ma -- Force is mass times acceleration

w=fd -- Work (energy) is force times distance

Atmospheric density varies as exp(-mgz/kT) where z is altitude, m is

molecular weight in kg of air, g is local acceleration of gravity, T

is temperature, k is Bolztmann's constant. On Earth up to 100 km,

$$d = d0*exp(-z*1.42e-4)$$

where d is density, d0 is density at 0km, is approximately true, so

Atmospheric scale heightDry lapse rate

(in km at emission level) (K/km)

Earth 7.5 9.8

Mars 11 4.4

Venus 4.9 10.5

Titan 18 1.3

Jupiter 19 2.0

Saturn 37 0.7

Uranus 24 0.7

Neptune 21 0.8

Triton 8 1

Titius-Bode Law for approximating planetary distances:

 $R(n) = 0.4 + 0.3 * 2^N Astronomical Units (N = -infinity for$

Mercury, 0 for Venus, 1 for Earth, etc.) This fits fairly well except for Neptune. 6.62618e-34 J-s (7e-34) -- Planck's Constant "h" 1.054589e-34 J-s (1e-34) -- Planck's Constant / (2 * Pl), "h bar" 1.3807e-23 J/K(1.4e-23) - Boltzmann's Constant "k" 5.6697e-8 W/m^2/K (6e-8) -- Stephan-Boltzmann Constant "sigma" 6.673e-11 N m^2/kg^2 (7e-11) -- Newton's Gravitational Constant "G" 0.0029 m K (3e-3) -- Wien's Constant "sigma(W)" 3.827e26 W (4e26) -- Luminosity of Sun 1370 W / m² (1400) -- Solar Constant (intensity at 1 AU) 6.96e8 m (7e8) -- radius of Sun 1738 km (2e3) -- radius of Moon 299792458 m/s (3e8) -- speed of light in vacuum "c" 9.46053e15 m (1e16) -- light year 206264.806 AU (2e5) -- \ 3.2616 light years (3) -- --> parsec 3.0856e16 m (3e16) --/ Black Hole radius (also called Schwarzschild Radius): 2GM/c^2, where G is Newton's Grav Constant, M is mass of BH, c is speed of light Things to add (somebody look them up!) Basic rocketry numbers & equations

Aerodynamical stuff

Energy to put a pound into orbit or accelerate to interstellar velocities.

Non-circular cases?

NEXT: FAQ #7/15 - Astronomical Mnemonics