Homework 2

Physics 201

Due September 11th, 2020

Exercise 1. We will define $r_0 = R$ where R is the radius of the Earth, and $\Delta r = r$ where r is the distance above Earth.

$$F(r_0 + \Delta r) = GMm \cdot \frac{1}{(r_0 + \Delta r)^2}$$

Thus we can make the approximation

$$F(r_0 + \Delta r) = (F(r_0) + F'(r_0)\Delta r + \frac{1}{2}F''(r_0)\Delta r^2 + \dots)$$

So the first non-constant is $F'(r_0)\Delta r = -GMm\frac{2\Delta r}{r_0^3}$.

We will evaluate the constant term $F(r_0)=GMm\frac{1}{r_0^2}=\frac{(6.67\cdot 10^{-11})(5.972\cdot 10^{24})}{(6.3781\cdot 10^6)^2}\approx 9.7918$ which is close to the acceleration of gravity near the Earth's surface.

The first order correction is $F'(r_0)\Delta r = -GMm\frac{2\Delta r}{r_0^3} = -\frac{(6.67\cdot 10^{-11})(5.972\cdot 10^{24})(1)(1)(2)}{(6.3781\cdot 10^6)^3} \approx -0.00000307044736068$ which is absurdly small of course.

Exercise 2. With masses that are d apart they experience the same force,

$$F_1 = \frac{Gm_1m_2}{d^2}\hat{x} \implies a_1 = \frac{F_1}{m_1} = \frac{Gm_2}{d^2}\hat{x}$$

$$F_2 = -\frac{Gm_1m_2}{d^2}\hat{x} \implies a_2 = \frac{F}{m_2} = -\frac{Gm_1}{d^2}\hat{x}.$$

If $m_2 = -m_1$ then,

$$F_1 = -\frac{Gm_1^2}{d^2}\hat{x} \implies a_1 = \frac{F_1}{m_1} = -\frac{Gm_1}{d^2}\hat{x}$$

$$F_2 = \frac{Gm_1^2}{d^2}\hat{x} \implies a_2 = \frac{F}{m_2} = -\frac{Gm_1}{d^2}\hat{x}.$$

We can use Newton's second law to find their motion. First mass 1,

$$m_1 \ddot{x}_1(t) = -\frac{Gm_1^2}{d^2}$$

$$\ddot{x}_1(t) = -\frac{Gm_1}{d^2}$$

$$\dot{x}_1(t) = -\frac{Gm_1}{d^2}t + A$$

$$x_1(t) = -\frac{Gm_1}{2d^2}t^2 + At + B$$

$$x_1(0) = B = -\frac{d}{2} \implies$$

$$x_1(t) = -\frac{Gm_1}{2d^2}t^2 - \frac{d}{2}$$

Now mass 2,

$$m_{2}\ddot{x}_{2}(t) = \frac{Gm_{1}^{2}}{d^{2}}$$

$$-m_{1}\ddot{x}_{2}(t) = \frac{Gm_{1}^{2}}{d^{2}}$$

$$\ddot{x}_{2}(t) = -\frac{Gm_{1}}{d^{2}}$$

$$\dot{x}_{2}(t) = -\frac{Gm_{1}}{d^{2}}t + A$$

$$x_{2}(t) = -\frac{Gm_{1}}{2d^{2}}t^{2} + At + B$$

$$x_{2}(0) = B = \frac{d}{2} \Longrightarrow$$

$$x_{2}(t) = -\frac{Gm_{1}}{2d^{2}}t^{2} + \frac{d}{2}$$

Exercise 3. We can use the trigonometric identity of $\sin(2x) = \sin(x)\cos(x)$ to translate $U(x) = U_0 \frac{\sin(\frac{2x}{l})}{2}$ which implies that the minimum is when $\sin(\frac{2x}{l}) = -1$ which is when $\frac{2x}{l} = \frac{3\pi}{2} \implies$ the minimum is $\frac{x}{l} = \frac{3\pi}{4}$.

To determine the oscillatory motion, we must find the k constant, which comes from $U''(x_e)$, where $x_e = \frac{3\pi}{4}$.

$$U''(x_e) = \frac{d^2}{dt^2} \left(U_0 \frac{\sin\left(\frac{2x}{l}\right)}{2} \right) \Big|_{x_e} = -U_0 \frac{2\sin\left(\frac{2x_e}{l}\right)}{l^2} = \frac{2U_0}{l^2}.$$

$$T = 2\pi \sqrt{\frac{m}{U''(x_e)}} = 2\pi \sqrt{\frac{ml^2}{2U_0}} = \sqrt{2l\pi} \sqrt{\frac{m}{U_0}}$$

Exercise 4. a.

The positions that are allowed are 2,4, and 5 because they are below the total energy line, and nothing can be above due to no negative kinetic energy. 4 is fastest since it is farthest from the

E line meaning the least potential and most kinetic. 4 is slowest since it is closest to the E line meaning the most potential and least kinetic.

b.

2 and 4 since there is an decrease to the local minimum, local fastest speed, and a increase away from the local minimum, a decrease in speed, within the vicinity of the positions.

c.

4 since there is a slow decent to the local minimum and a slow rise away as opposed to 2 where it is a steep decent and climb.

Exercise 5.

$$f(t) = f_0 \sum_{j=0}^{\infty} \frac{t^j}{j!}$$

$$= f_0 \left(1 + t + \frac{t^2}{2} + \frac{t^3}{6} + \dots \right)$$

$$\frac{df(t)}{dt} = f_0 \left(\frac{d}{dt} (1) + \frac{d}{dt} (t) + \frac{d}{dt} \left(\frac{t^2}{2} \right) + \frac{d}{dt} \left(\frac{t^3}{6} \right) + \dots \right)$$

$$= f_0 \left(0 + 1 + t + \frac{t^2}{2} \dots \right)$$

$$= f_0 \left(1 + t + \frac{t^2}{2} + \frac{t^3}{6} + \dots \right)$$

$$= f_0 \sum_{j=0}^{\infty} \frac{t^j}{j!}$$

$$= f(t)$$

Or we may do it more symbolically as opposed to intuitively,

$$f(t) = f_0 \sum_{j=0}^{\infty} \frac{t^j}{j!}$$

$$\frac{df(t)}{dt} = f_0 \frac{d}{dt} \sum_{j=0}^{\infty} \frac{t^j}{j!}$$

$$= f_0 \sum_{j=0}^{\infty} \frac{d}{dt} \frac{t^j}{j!}$$

$$= f_0 \sum_{j=0}^{\infty} j \frac{t^{j-1}}{j!}$$
Let $l = j - 1$

$$= f_0 \sum_{l=-1}^{\infty} (l+1) \frac{t^l}{(l+1)!} = f_0 \sum_{l=0}^{\infty} \frac{t^l}{l!}$$

$$= f(t)$$

Exercise 6. Following from section 1.4.1,

$$\dot{f}(t) = t^p \sum_{j=0}^{\infty} a_j (j+p) t^{j-1}$$

Since we have the ODE in the form $\dot{f}(t) - \alpha f(t)$ we can rewrite this as,

$$t^{p} \left[a_{0}pt^{-1} + \sum_{j=0}^{\infty} a_{j+1}(j+p+1)t^{j} \right] - \alpha t^{p} \sum_{j=0}^{\infty} a_{j}t^{j} = 0$$
$$t^{p} \left[a_{0}pt^{-1} + \sum_{j=0}^{\infty} (a_{j+1}(j+p+1) - \alpha a_{j})t^{j} \right] = 0$$

We can find the recursion relation,

$$a_{j+1} = \frac{\alpha a_j}{j+p+1}$$

and once again set p = 0, thus

$$a_{j+1} = \frac{\alpha a_j}{j+1}$$

with the first few terms being,

$$a_1 = \alpha a_0$$

$$a_2 = \alpha \frac{a_1}{2} = \alpha^2 \frac{a_0}{2}$$

$$a_3 = \alpha \frac{a_2}{3} = \alpha^3 \frac{a_0}{6}$$

Thus we find,

$$f(t) = a_0 \sum_{j=0}^{\infty} \frac{(\alpha t)^j}{j!}$$

and we can set $a_0 = f_0$, thus

$$f(t) = f_0 \sum_{j=0}^{\infty} \frac{(\alpha t)^j}{j!} = f_0 e^{\alpha x}$$