# PHY 107 Gravitation

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### **OUTLINE**

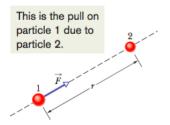
- Newton's law of gravitation
- Superposition
- Potential Energy

#### Newton's Law of Gravitation

Gravitation: Tendency of bodies to move toward one another **Newton's law of gravitation:** Every particle attracts any other particle with a gravitational force of magnitude

$$F = G \frac{m_1 m_2}{r^2}$$

$$G = 6.67 \ X \ 10^{-11} \text{Nm}^2 / \text{kg}^2$$



## Principle of Superposition

Given a group of particles, we find the net (or resultant) gravitational force on any one of them from the others by using the principle of superposition.

For n interacting particles, we can write the principle of superposition for the gravitational forces on particle 1 as:  $\overrightarrow{F_{1,pet}} = \overrightarrow{F_{12}} + \overrightarrow{F_{13}} + \overrightarrow{F_{14}} + \dots \overrightarrow{F_{1p}}$ 

$$\overrightarrow{F_{1,net}} = \overrightarrow{F_{12}} + \overrightarrow{F_{13}} + \overrightarrow{F_{14}} + \dots \overrightarrow{F_{1n}}$$

#### Newton's Law of Gravitation

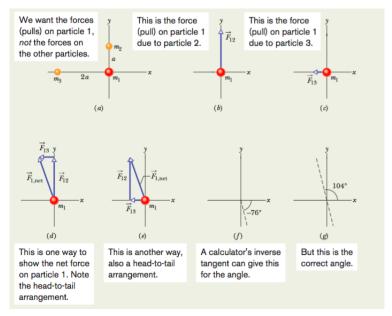
**Example** Figure shows an arrangement of three particles, particle 1 of mass  $m_1 = 6.0$  kg and particles 2 and 3 of mass  $m_2 = 4.0$  kg, and distance a = 2.0 cm. What is the net gravitational force  $F_{1,net}$  on particle 1 due to the other particles?

#### **Solution:**

$$F_{12} = \frac{Gm_1m_2}{a^2}$$

$$F_{13} = \frac{Gm_1m_3}{(2a)^2}$$

#### Newton's Law of Gravitation



## Gravitational Potential Energy

Let's broaden our view...

U=0 for 
$$r = \infty$$

**Potential Energy** is negative for any finite separation and becomes progressively more negative as the particles move closer together.

#### GPE of the two-particle system:

 $U = -\frac{GMm}{r}$  (gravitational potential energy)

(Check the book for proof)

GPE: a property of the two particle system (rather than of either particle alone).

What would be the GPE for a three particle system?

Potential Energy and Force:  $F = -\frac{dU}{dr}$ 

## Escape speed

Consider a projectile of mass m leaving the surface of a planet with escape speed v:

$$v = \sqrt{\frac{2GM}{R}}$$
 (Can be derived by COE)

M: mass of the planet

R: radius of the planet

## Example

#### Asteroid falling from space, mechanical energy

An asteroid, headed directly toward Earth, has a speed of 12 km/s relative to the planet when the asteroid is 10 Earth radii from Earth's center. Neglecting the effects of Earth's atmosphere on the asteroid, find the asteroid's speed  $v_f$  when it reaches Earth's surface.

#### Solution:

COE: 
$$K_f + U_f = K_i + U_i$$
  
 $0.5mv_f^2 - \frac{GMm}{R_E} = 0.5mv_i^2 - \frac{GMm}{10R_E}$   
We solve for  $v_f$ 

## Problems of importance

#### Reference book (Extended 9th edition)

Newton's Law of Gravitation: 3

Gravitation and Principle of Superposition: 6

#### Reference

Fundamentals of Physics by Halliday and Resnik