





Implications from the Terrestrial Formation of Jarosite - Alunite Minerals for Weathering on Mars

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"El Capitan" at Meridiani Planum shows fine parallel lamination in its upper part. It also contains scattered sphere-shaped objects (ranging from 1 to 2 mm in size) rich in jarosite and hematite. (©NASA/JPL/Cornell).

Compositions of Rocks and Soils Pathfinder Samples, Ares Vallis, July 1997 (after Economou et al., 2003)

	Rocks	Soils	Average Earth (Taylor, 1964)
No analyses	(5)	(6)	
SiO ₂	49.7	41.6	51.3
Al ₂ O ₃	11.5	10.2	16.6
Fe ₂ O ₃	16.9	22.1	12.3
MnO	0.4	0.3	0.2
MgO	3.7	7.4	7.5
CaO	6.4	6.0	9.4
Na ₂ O	4.0	2.8	2.6
K ₂ O	0.9	0.6	1.0
TiO ₂	0.6	0.8	1.5
Cr ₂ O ₃	0.1	0.3	0.03
P ₂ O ₅	0.5	0.8	0.3
SO ₃	2.8	6.1	0.06
CI	0.6	0.9	0.006
H ₂ O	<u>1.9</u>	_	
	100		

Opportunity Landing Area and Rover Path

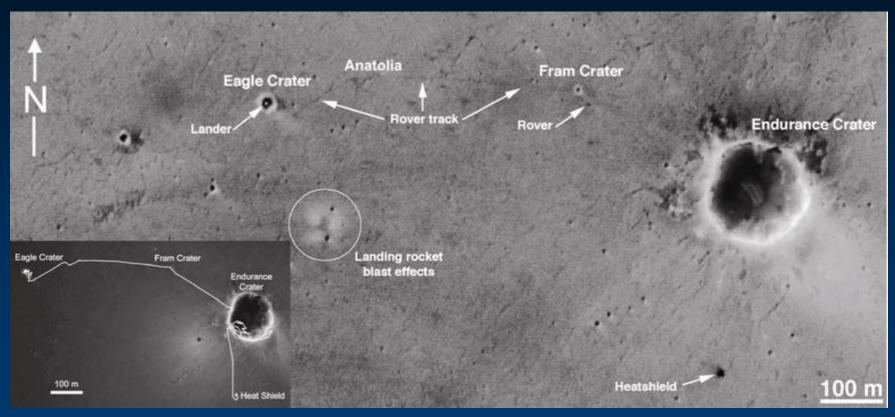


Image credit: NASA/JPL/MSSS

'Burns' formation at Burns cliff, SE rim Endurance crater

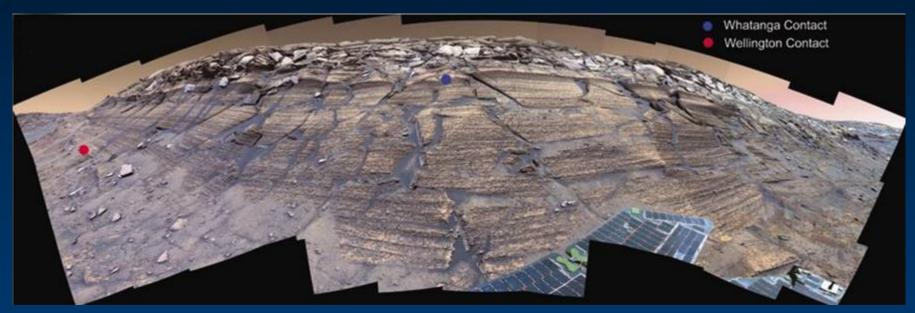


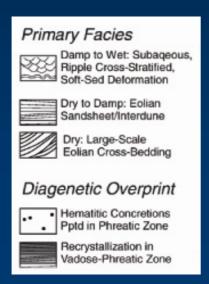
Image credit: NASA/JPL/MSSS

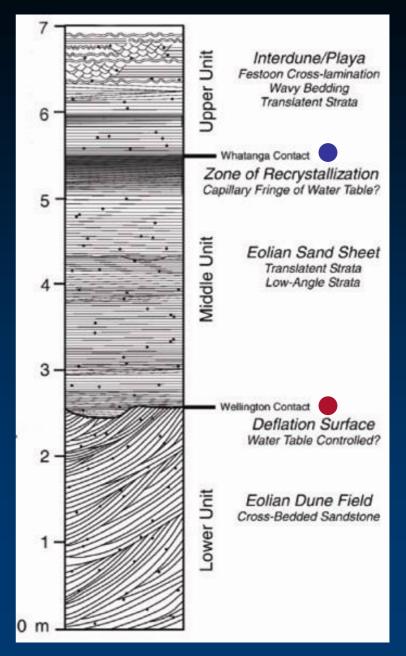
'Burns' formation at Burns cliff, SE rim Endurance crater



Image credit: NASA/JPL/MSSS

Stratigraphy of 'Burns' formation



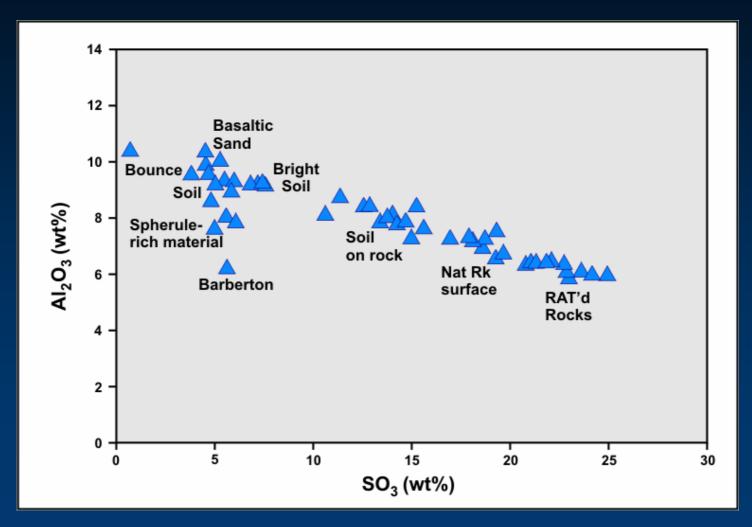


Average Mineralogy (vol %), Eagle crater Mini TES (Miniature Thermal Emission Spectroscopy)

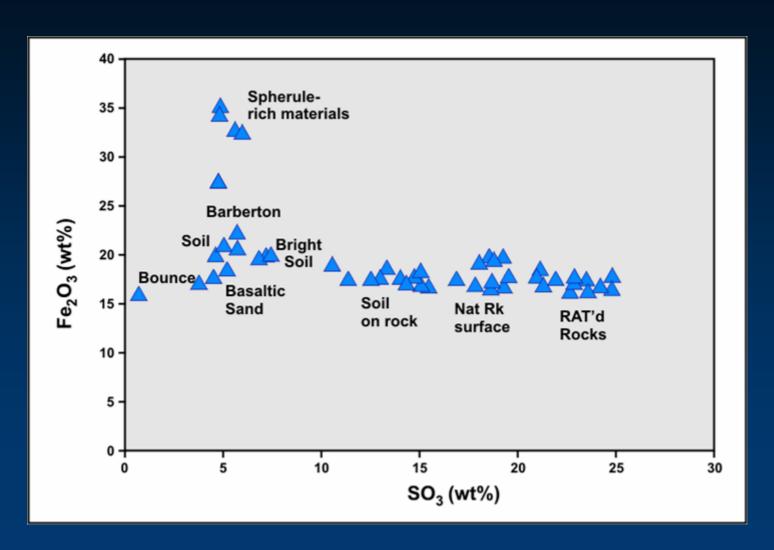
	Average (13)	Range
Sulfate	25	10 - 35
Hematite	24	0 - 45
Sheet silicates (e.g. micas/cla	y) 3	0 - 20
Glass	15	5 - 25
Oxide/hydroxide	6	0 - 20
Feldspar	12	0 - 45
Olivine	9	0 - 20
Pyroxene	6	0 - 15

(after Christiansen et al., 2004)

Meridiani Materials APXS (Alpha Particle X-ray Spectroscopy)



Meridiani Materials APXS (Alpha Particle X-ray Spectroscopy)



Features of Martian Materials

- High SO₃ content (present as Ca and Mg sulfates + jarosite)
- High CI and Br content
- Features in 'Burns' formation imply reworked sand and evaporative minerals derived from an evaporative basin (interdune depression or playa lake)

Jarosite – Alunite Minerals

$$AB_3 (XO_4)_2 (OH)_6$$

A = K, Pb, Ca, Na,
$$H_3O^+$$
, NH_4^+ , Sr, REE
B = Fe, Al, (Cu, Zn)
 $XO_4 = SO_4^{2-}$, PO_4^{3-} , AsO_4^{3-}

OH sites can contain halides

Common members of the Jarosite – Alunite family

Alunite $K Al_3(SO_4)_2(OH)_6$

Jarosite $K Fe_3(SO_4)_2(OH)_6$

Natroalunite Na $Al_3(SO_4)_2(OH)_6$

Natrojarosite Na $Fe_3(SO_4)_2(OH)_6$

Hinsdalite Pb Al₃(PO₄)(SO₄)(OH)₆

Plumbogummite Pb Al₃H(PO₄)₂(OH)₆

Features of Martian Materials

- Acidic pH ~ 2 jarosite
- Oxidizing Fe³⁺
- Sulfate rich environment

Jarosite – Alunite Formation on Earth

- in the supergene zone of many ore deposits
- in weathering profiles unrelated to mineralization, particularly in micaceous, sulfide-bearing sediments
- as an authigenic component of Quaternary saltlake sediments in Australia or as an authigenic precipitate from acid crater-lake waters
- as products of sulfate aeolinites in the playa-dune system of White Sands, New Mexico

Jarosite – Alunite Formation on Earth

- as diagenetic component of Tertiary to Quaternary sediments in South Australia.
- as evaporite crusts precipitated from acid sulfate and iron-rich waters of Rio Tinto, Southern Spain.
- in acid sulfate soils occurs where pyrite in previously saturated soils is exposed to oxygen

Locations and Distribution of Jarosite from Different Geological and Climatic Settings throughout Australia



Supergene Jarosite Formation

$$FeS_2 + \frac{7}{2}O_2 + H_2O \rightarrow Fe^{2+} + 2SO_4^{2-} + 2H^+$$
Pyrite

2 Fe²⁺ +
$$\frac{1}{2}$$
 O₂ \longrightarrow Fe³⁺ + H₂O

$$4KAISi_3O_8 + 22H_2O \longrightarrow 4K^+ + 4OH^- + AI_4Si_4(OH)_8 + 8H_4SiO_4$$

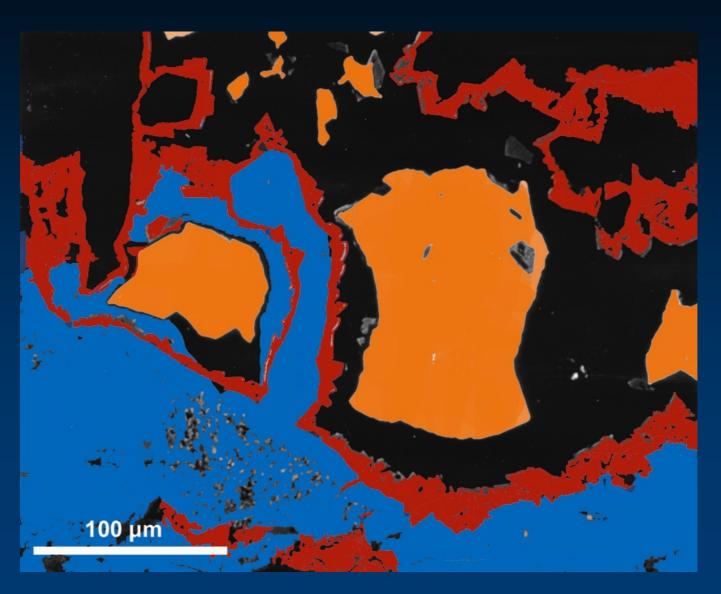
Feldspar Kaolinite

$$3Fe^{3+} + K^{+} + 2SO_4^{2-} + 6OH^{-} \rightarrow KFe_3(SO_4)_2(OH)_6$$

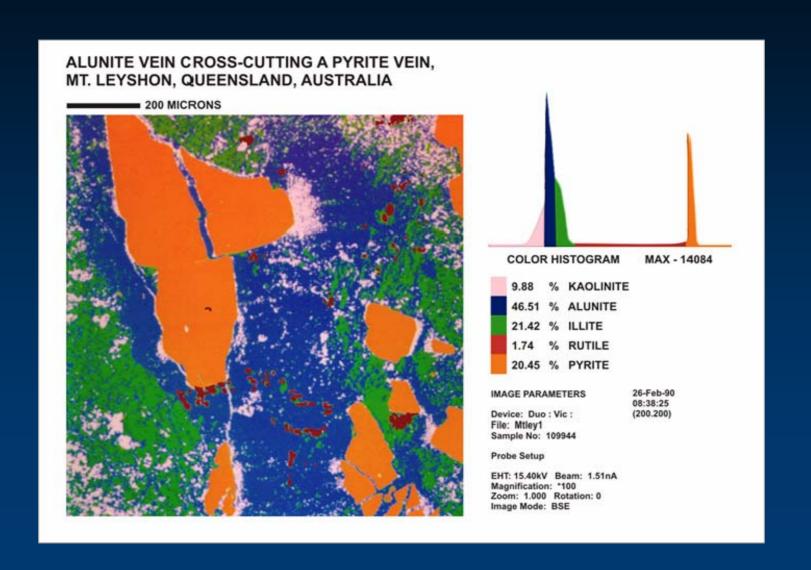
Jarosite Vein (yellow) within the Weathering Profile at the Mt Leyshon Gold Mine, Queensland



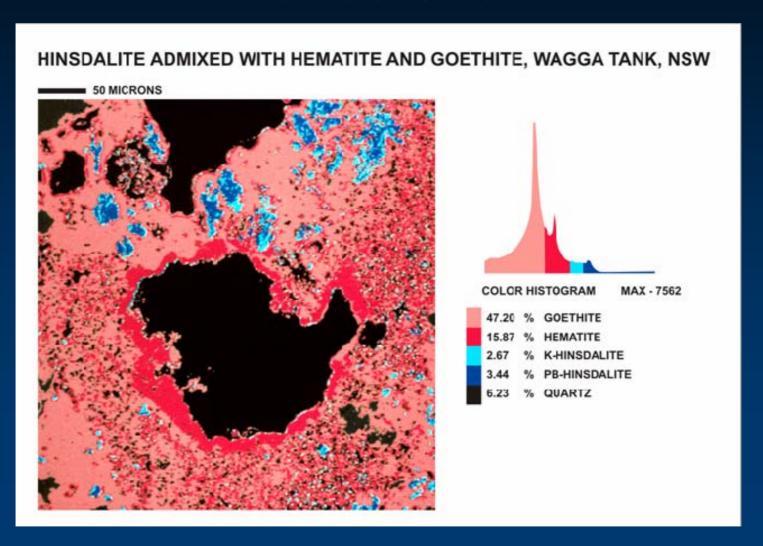
Pyrite, Jarosite and Alunite at Mt Leyshon showing Pyrite (orange) oxidising to Jarosite (red) and Alunite (blue)



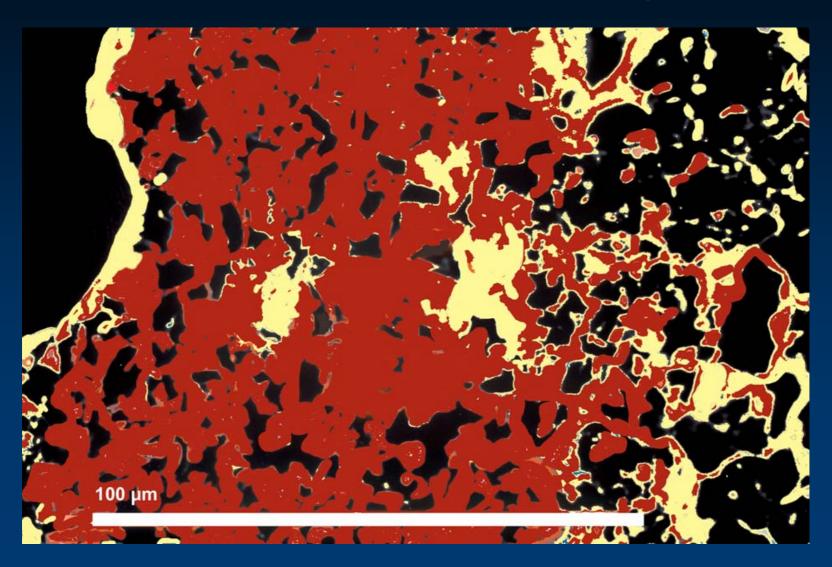
Alunite (blue) from Weathering of Pyrite, intimately associated with Kaolinite and Rutile, Mt Leyshon



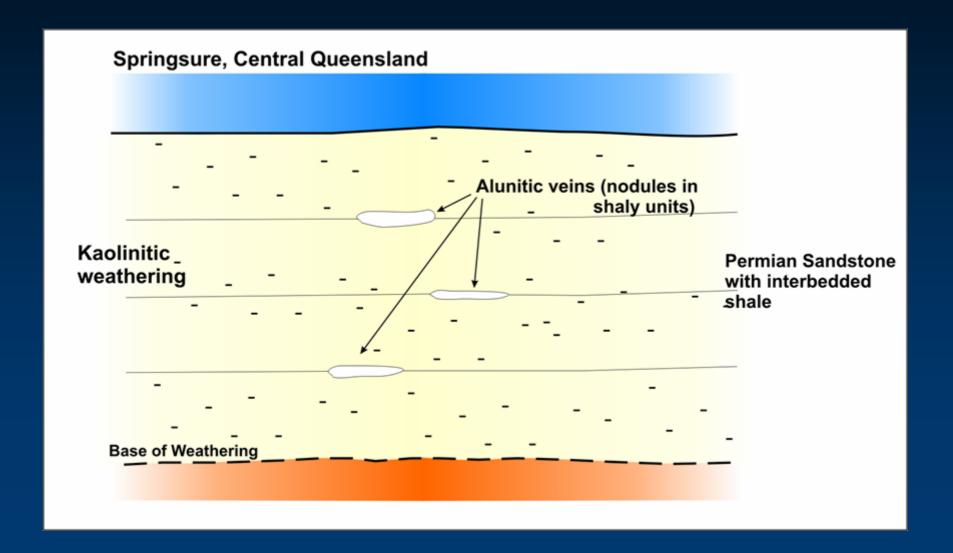
Pb and K rich Hinsdalite (blue) from Wagga Tank, NSW, intimately associated with Goethite (pink) and Hematite (red). The Hinsdalite shows compositional variations at the micron scale.



Plumbojarosite (red) intimately associated with Goethite (cream), Currawang, NSW.

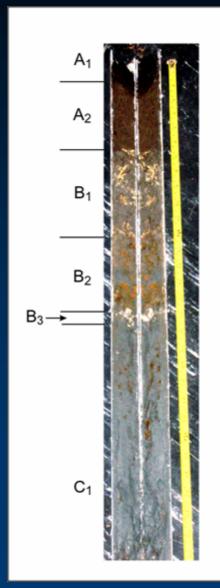


Jarosite / Alunite from Non – Mineralized Weathering Profiles



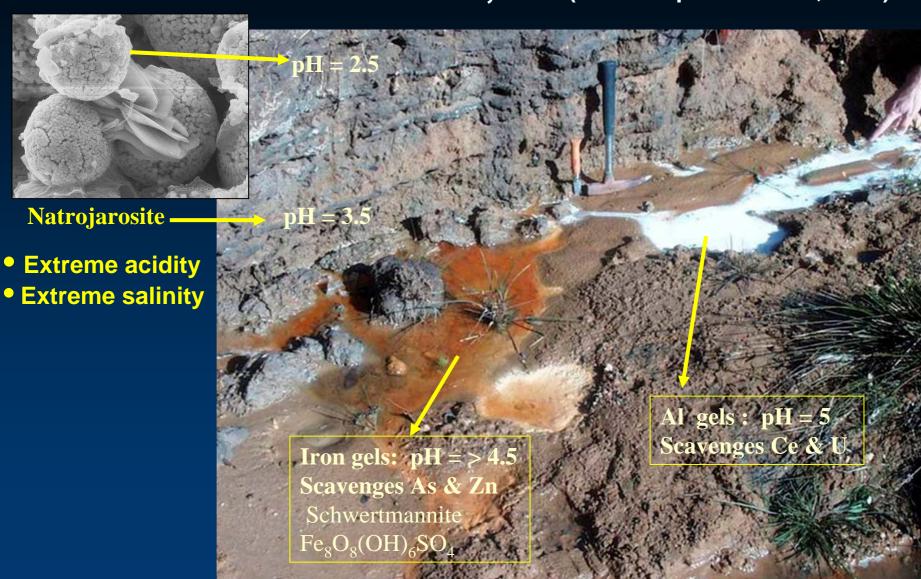
Jarosite from Acid Sulfate Soils

1.5m Vertical Profile, Mays Swamp, Kempsey area, NSW (after Isaacson et al., 2006)

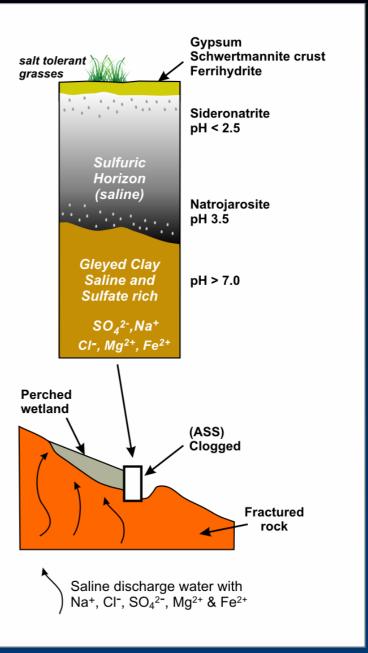


Horizon ID	Description		
A ₁	Black, rich in organic matter, peat like sediment		
A ₂	Lighter brown, alluvial like sediment, organic matter in the form of root cavities present		
Oxidised B ₁	Grey thick, estuarine sediment interlaid with prevalent yellow jarosite crystalline mottles associated with root cavities		
Oxidised B ₂	Grey thick, estuarine sediment interlaid with prevalent red/brown jarosite/goethite crystalline mottles associated with root cavities. Crystals up to 1cm prevalent		
Transition B ₃	Shell layer with thick grey estuarine sediment occurring between shell fragments. Gypsum immediately above shells		
Reduced C ₁	Blue/Grey thick estuarine sediment, red/brown goethite/hematite mottles occurring around limited root cavities		

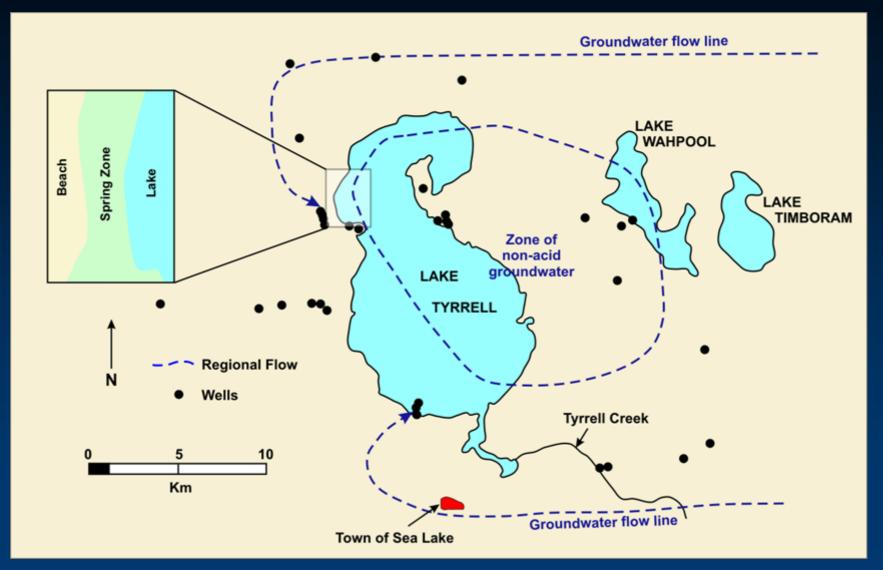
Jarosite from Acid Sulfate Soils Mt Torrens, SA (after Fitzpatrick et al., 2003)



Jarosite in Acid – Sulfate Soils

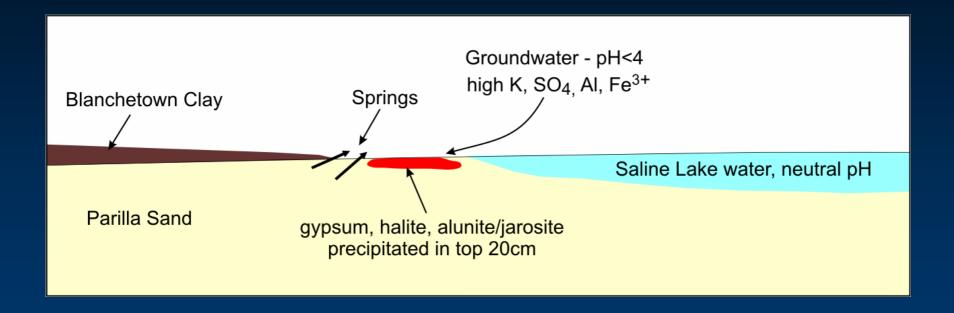


Jarosite Associated with Playa Lakes



Lake Tyrrell, Victoria (after Long et al., 1992)

Jarosite Associated with Playa Lakes

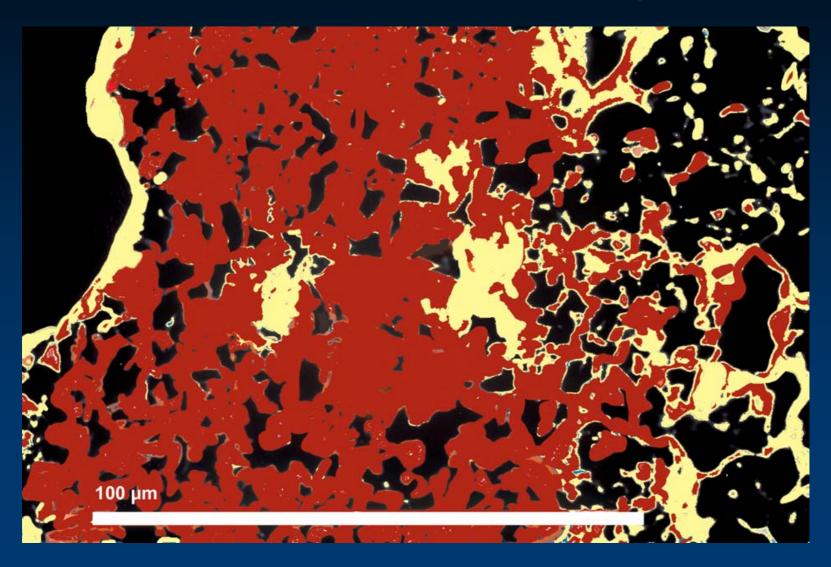


Lake Tyrrell, Victoria (after Long et al., 1992)

Implications of Martian Jarosite Occurrences

 Fe Oxides (goethite) occur intimately with jarosite – Gusev crater

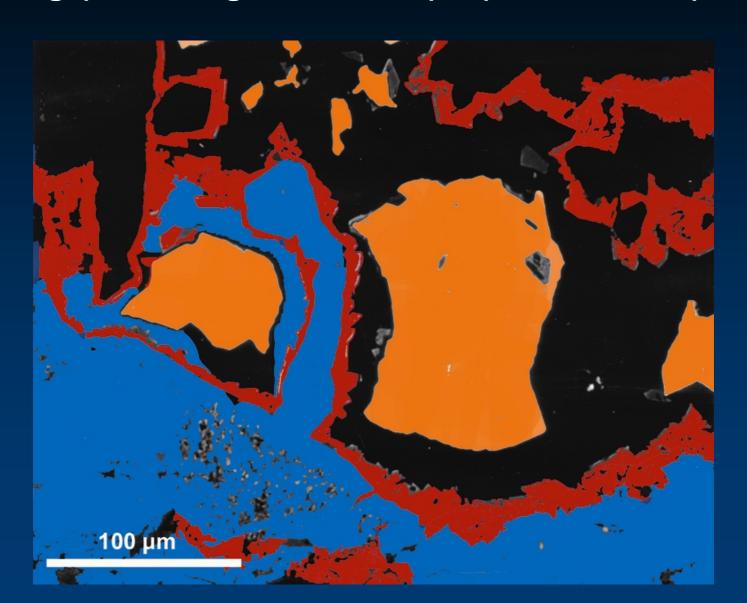
Plumbojarosite (red) intimately associated with Goethite (cream), Currawang, NSW.



Implications of Martian Jarosite Occurrences

- Fe Oxides (goethite) occur intimately with jarosite – Gusev crater
- Studies (~1930) suggest that goethite forms first
- Our studies imply that jarosite may form directly from pyrite

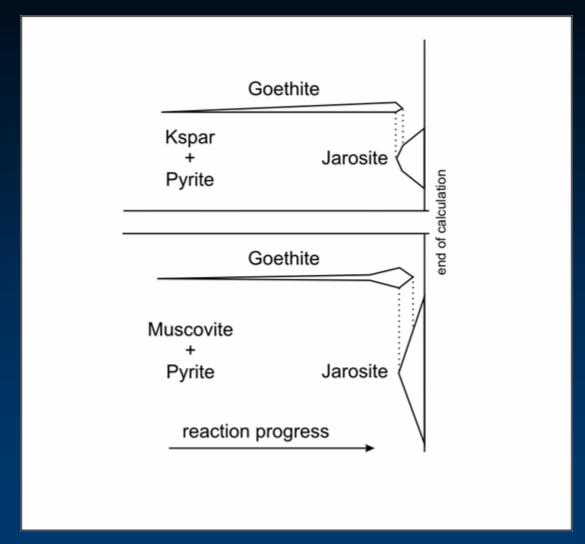
Pyrite, Jarosite and Alunite at Mt Leyshon showing Pyrite (orange) oxidising to Jarosite (red) and Alunite (blue)



Implications of Martian Jarosite Occurrences

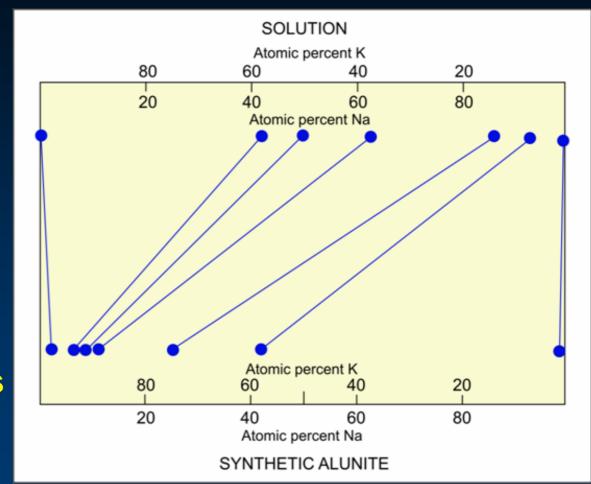
- Fe Oxides (goethite) occur intimately with jarosite – Gusev crater
- Studies (~1930) suggest that goethite forms first
- Our studies imply that jarosite may form directly from pyrite
- Thermodynamic studies show amounts of goethite formed during weathering relative to jarosite, is affected by the original rock

Thermodynamic studies of Jarosite Formation



Calculated paragenesis showing relative masses of minerals formed during the reaction of meteoric water with pyrite and K-feldspar or muscovite (after Bladh, 1982)

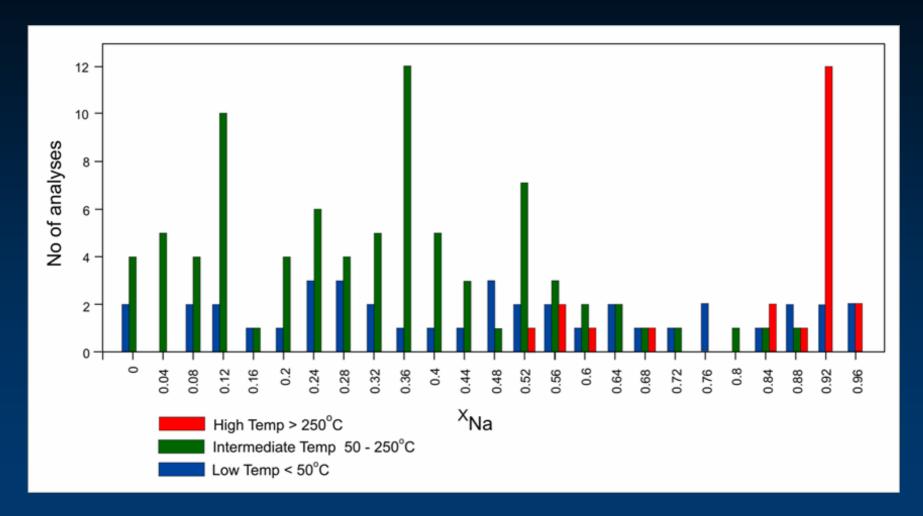
Experimental Studies of Alunite Formation 1 – Fluid Composition



K – rich alunites preferred

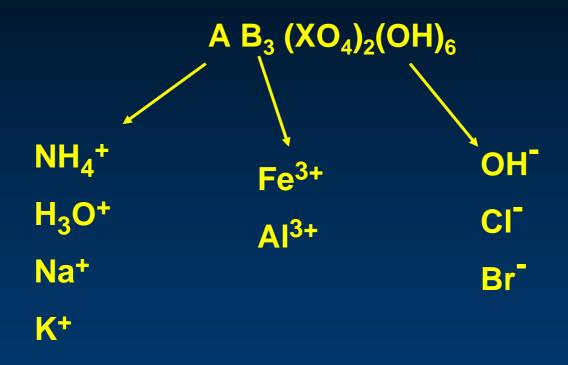
Relative atomic percentages of potassium and sodium in synthetic alunite and in solution prior to the precipitation of alunite (after Parker, 1962)

Experimental Studies of Alunite Formation 2 – Temperature



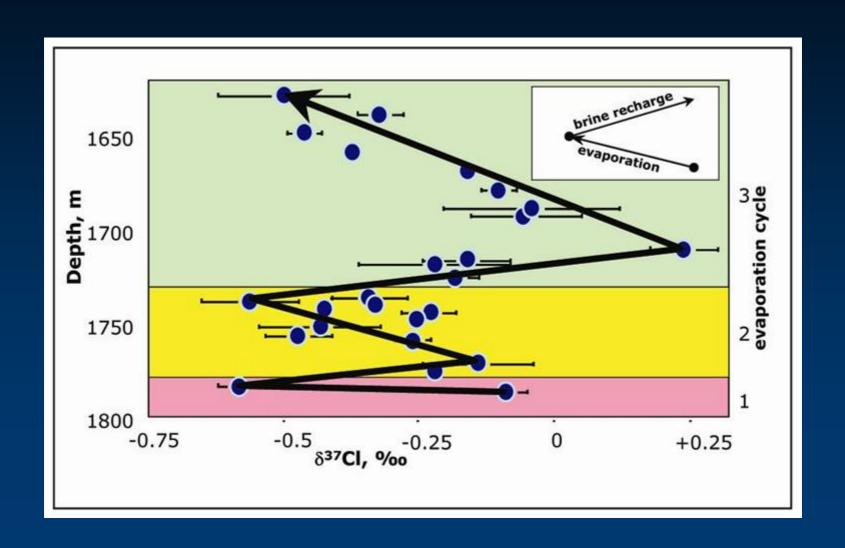
Compositional data on natural alunites (after Stoffregen and Cygan, 1990)

Jarosite – Alunite Minerals



What ions are actually in Martian Jarosite?

Chlorine Isotope Profile in a succession of Evaporite Beds



Jarosite – Alunite Minerals on Earth

- Jarosite-alunite minerals commonly found intimately associated with goethite
- Jarosite-alunite minerals form under acidic conditions
 and commonly under saline conditions → gypsum + halite
- Na-rich alunite favoured by higher temperature
- Na-rich alunite needs extremely Na rich fluid to form from solution
- More analog, experimental thermodynamic studies needed

Jarosite - Alunite Minerals on Mars

Questions to be answered by retrieval missions:

- Composition in A sites (Na⁺, K⁺, NH₄⁺, H₃O⁺). What does this tell us about the composition of the fluid from which it formed?
- Composition in B sites. Does this provide information about the fluid/rock interaction?
- Composition in (OH) sites. Are halides incorporated? If so, can they indicate about evaporation/recharge?
- O, H, S, CI isotopes can be useful in defining formation conditions







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