

# ccast and noaa relative fov response

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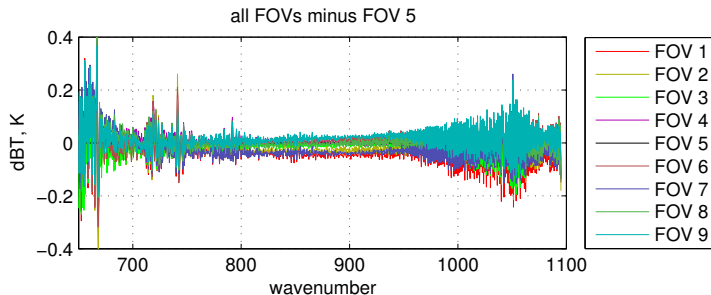
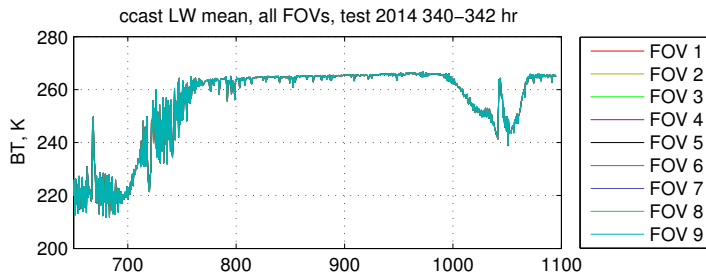
UMBC Atmospheric Spectroscopy Lab  
Joint Center for Earth Systems Technology

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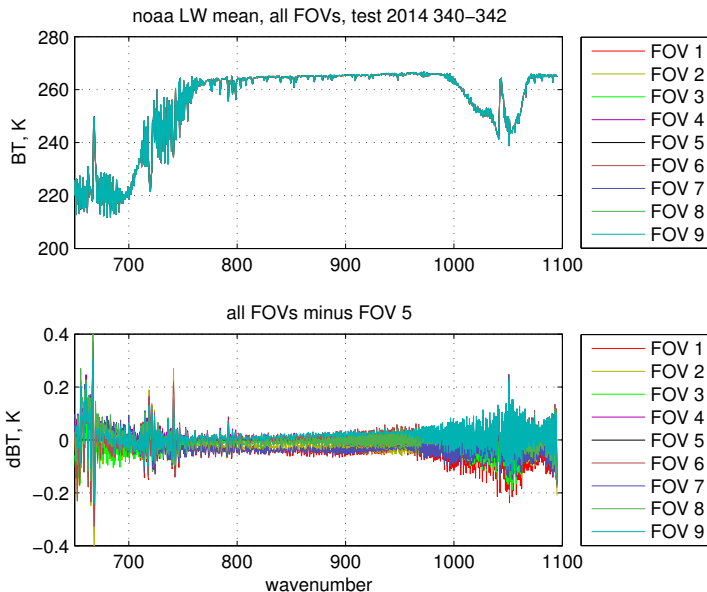
## test methods

- ▶ start with CCAST and NOAA high res data from 6–8 Dec 2014
- ▶ take the average and standard deviation of FOR 15 and 16 independently for each FOV, and compare these values with the values for FOV 5
- ▶ results shown here are for 32,186 CCAST and 32,120 NOAA descending FORs
- ▶ as a precaution, FORs where any LW channel was greater than 320K were discarded
- ▶ the intent is to show variation among FOVs, as might arise from varying nonlinearity or artifacts of the self-apodization correction

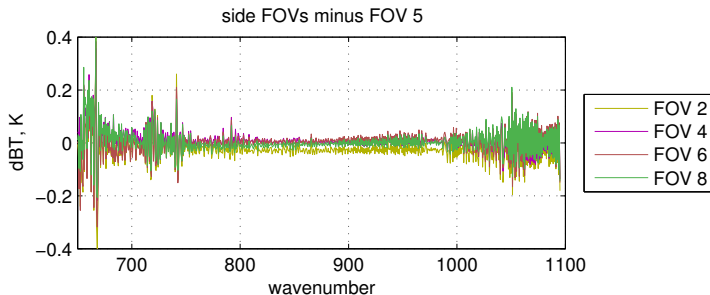
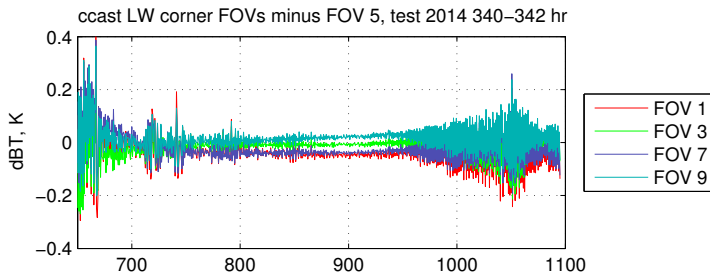
# ccast LW mean



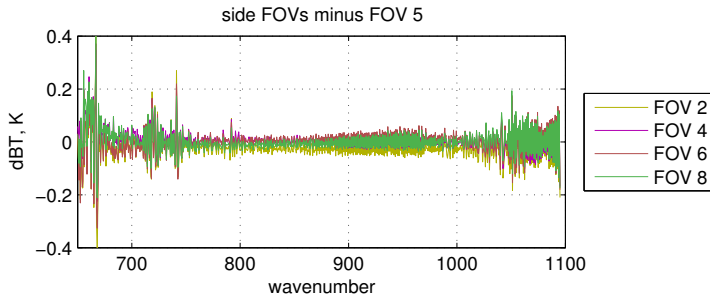
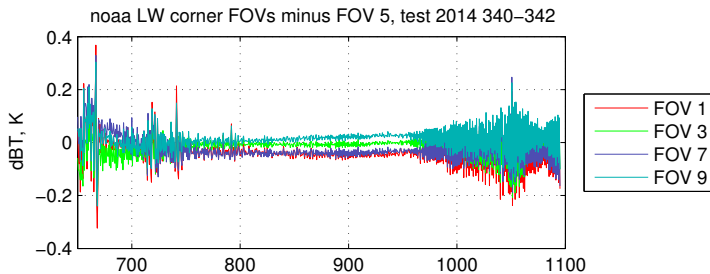
# noaa LW mean



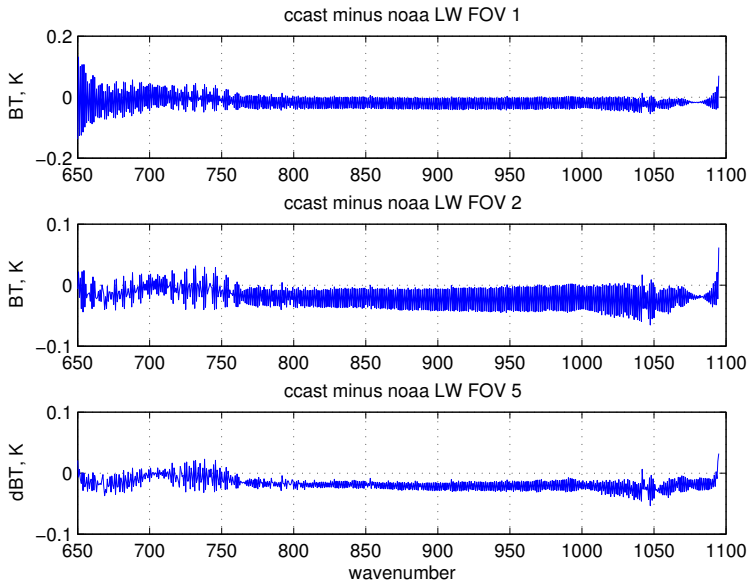
# ccast LW fov groups



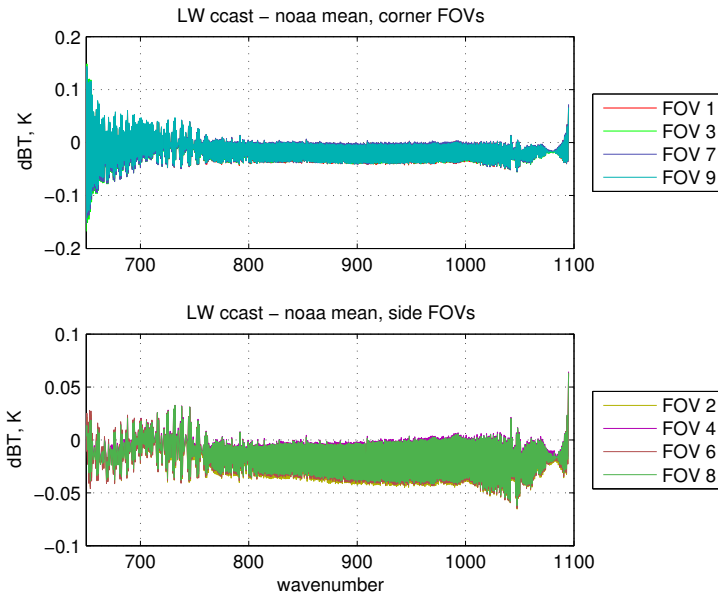
# noaa LW fov groups



# ccast minus noaa fovs 1, 2, and 5

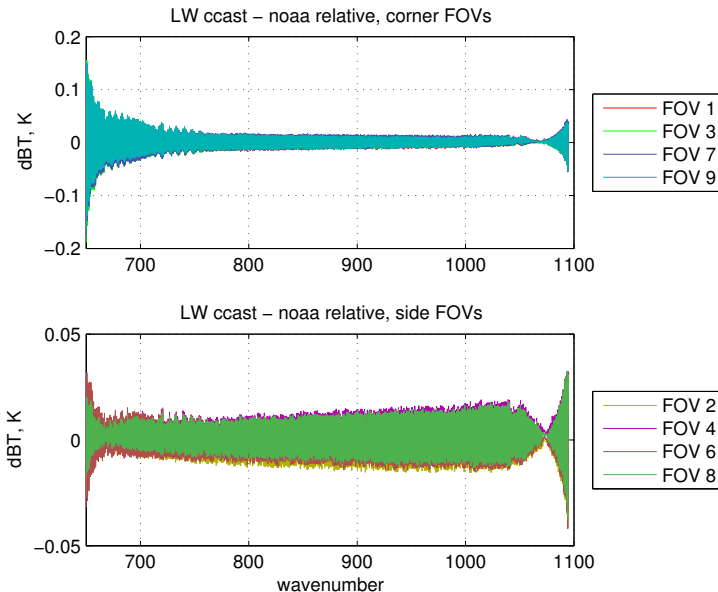


# ccast minus noaa fov groups





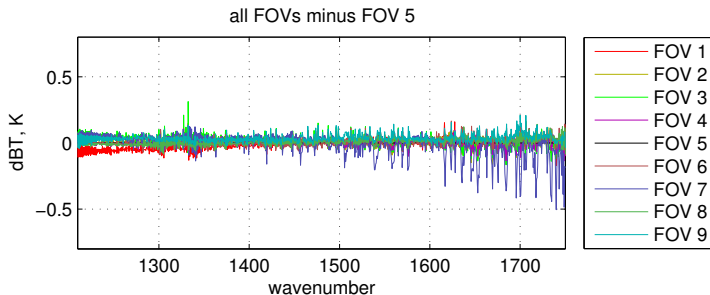
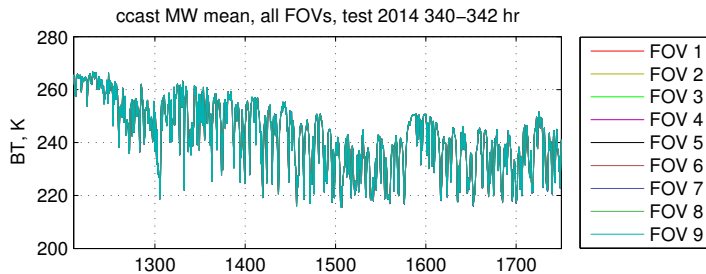
# ccast minus noaa relative fov groups



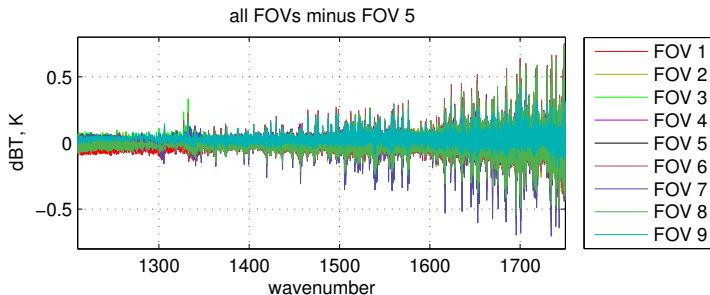
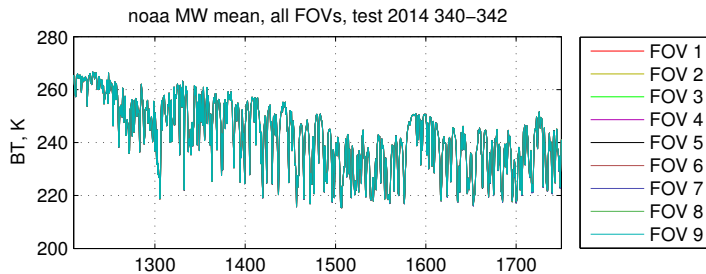
# LW discussion

- ▶ CCAST and NOAA are generally in good agreement
- ▶ in the LW, CCAST is around 0.02K colder than NOAA
- ▶ in the previous slide, “ccast minus noaa relative” is  $(\text{ccast all FOVs} - \text{FOV 5}) - (\text{noaa all FOVs} - \text{FOV 5})$
- ▶ the CCAST nonlinearity correction uses the UW a2 values
- ▶ the slightly greater difference for the corner FOVs at the low end of the band may be due to different processing filters

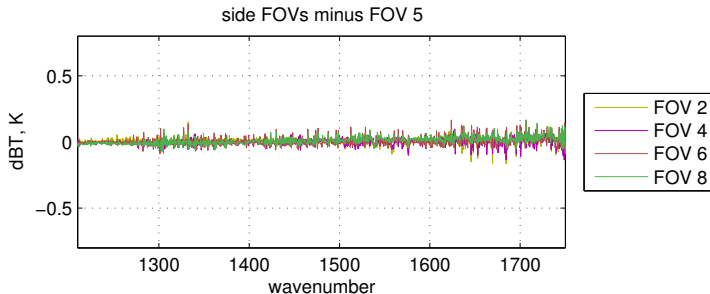
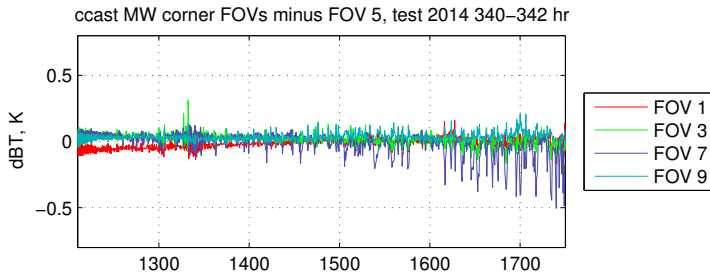
# ccast MW mean



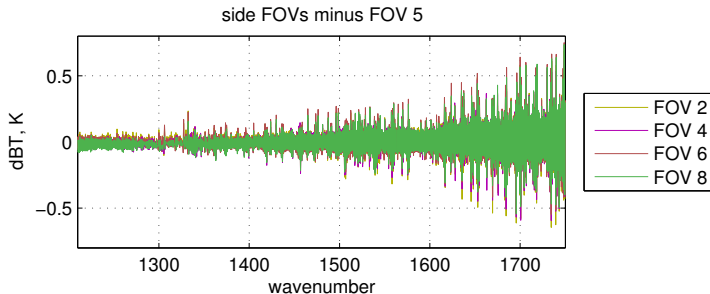
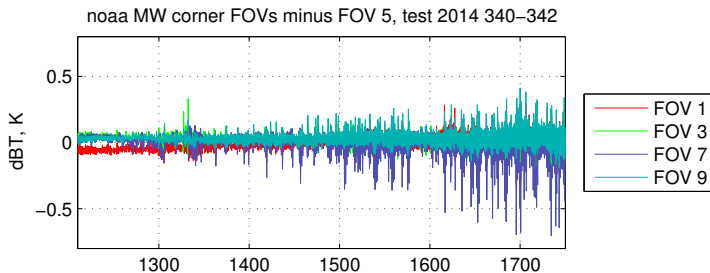
# noaa MW mean



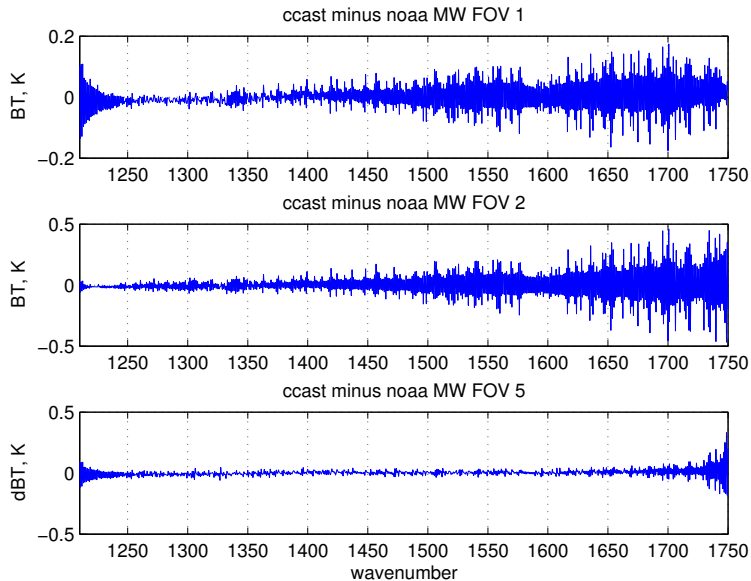
# ccast MW fov groups



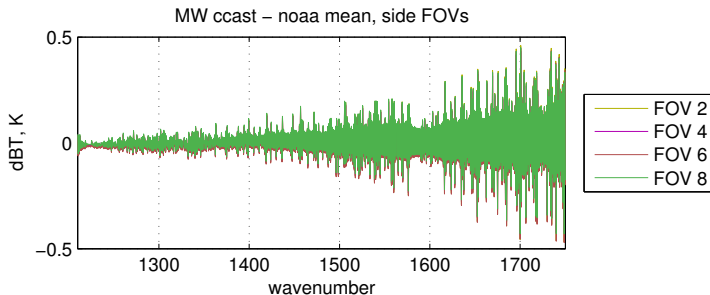
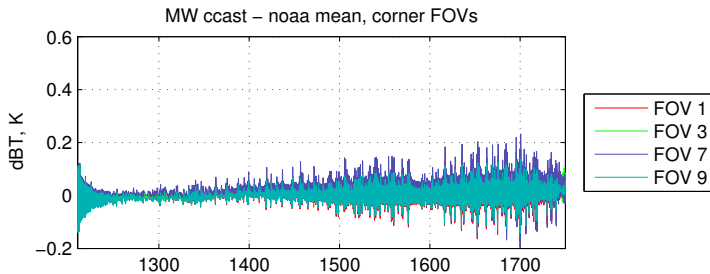
# noaa MW fov groups



# ccast minus noaa fovs 1, 2, and 5

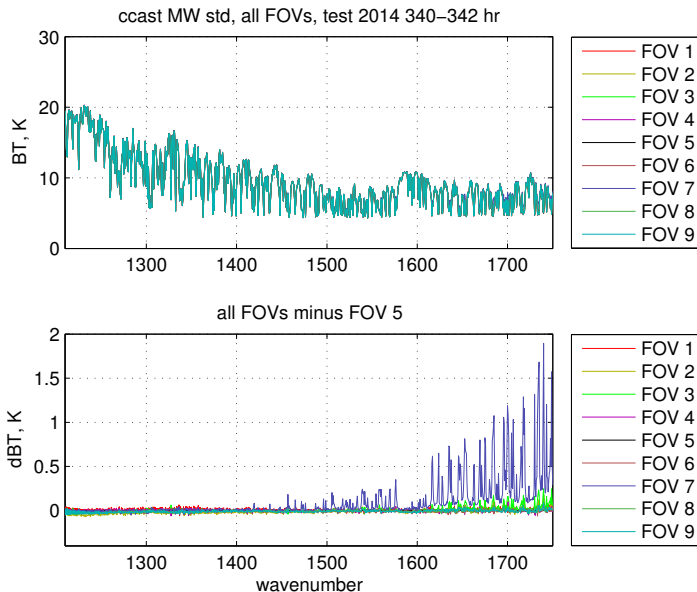


# ccast minus noaa fov groups





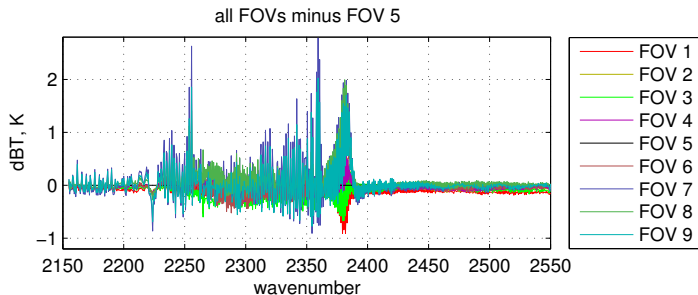
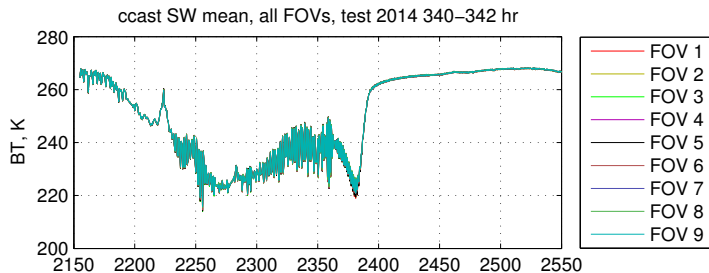
# ccast MW standard deviation



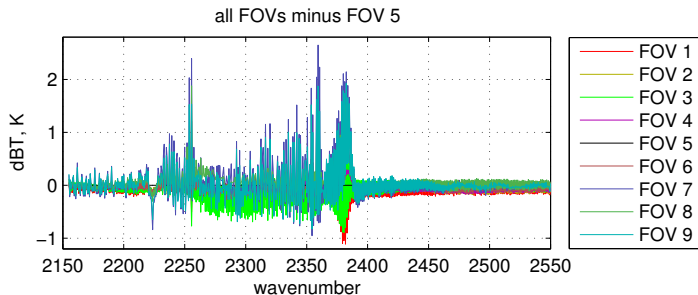
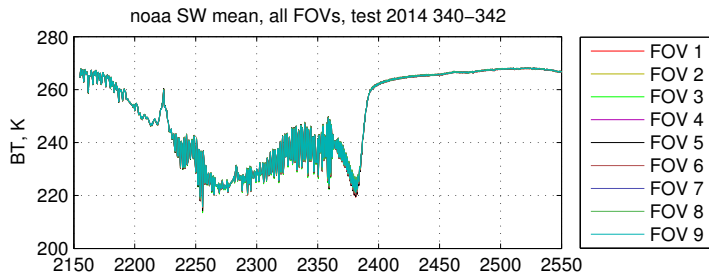
## MW discussion

- ▶ FOV 7 is the least linear, and only partially corrected for with the CCAST first order adjustment
- ▶ the NOAA variation in FOV response is much greater than CCAST
- ▶ this may be due to problems with the nonlinearity correction
- ▶ a normalized frequency domain representation of the numeric filter needs a scaling factor to match the original nonlinearity measurements. We used 1.6047 for LW, 0.9826 for MW, and 0.2046 for SW

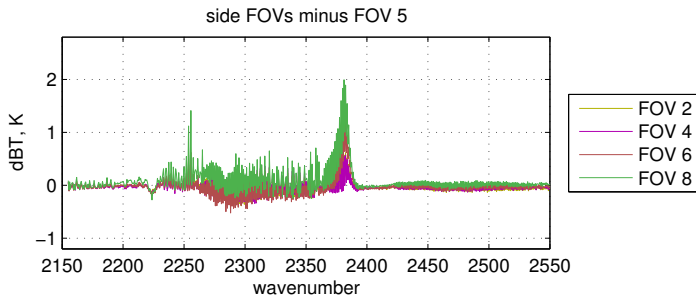
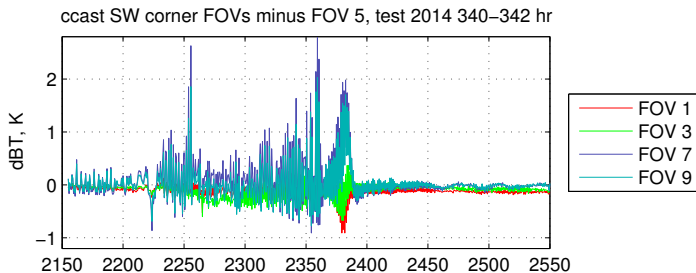
# ccast SW mean



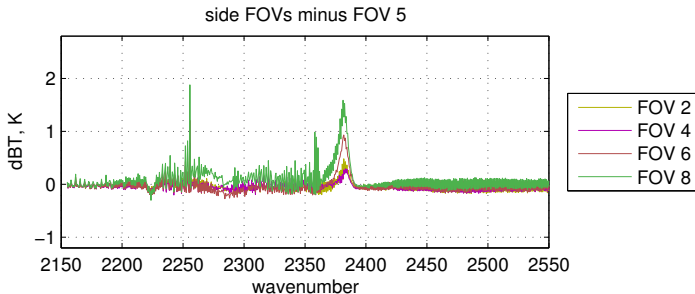
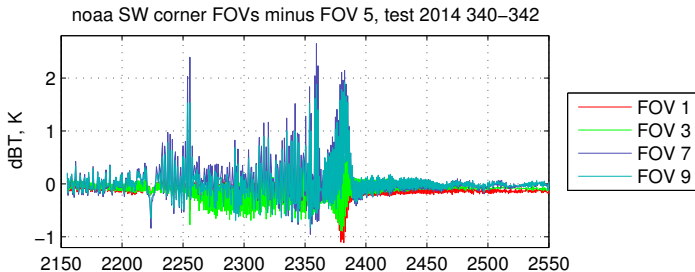
# noaa SW mean



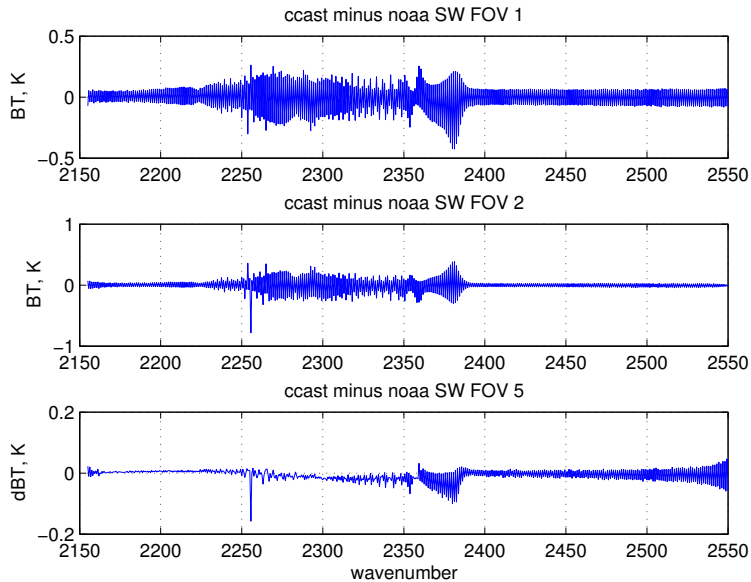
# ccast SW fov groups



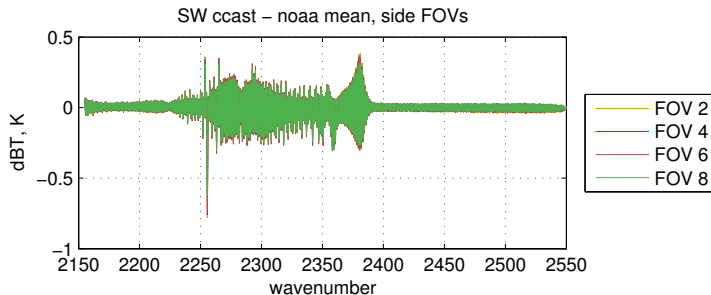
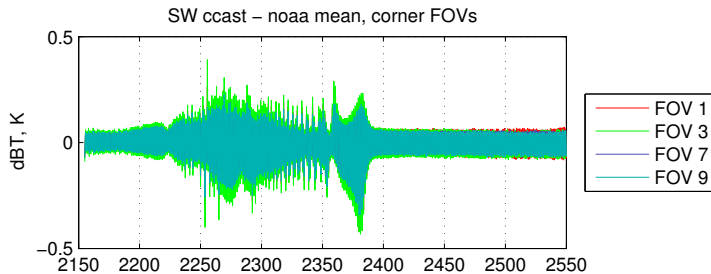
# noaa SW fov groups



# ccast minus noaa fovs 1, 2, and 5



# ccast minus noaa fov groups





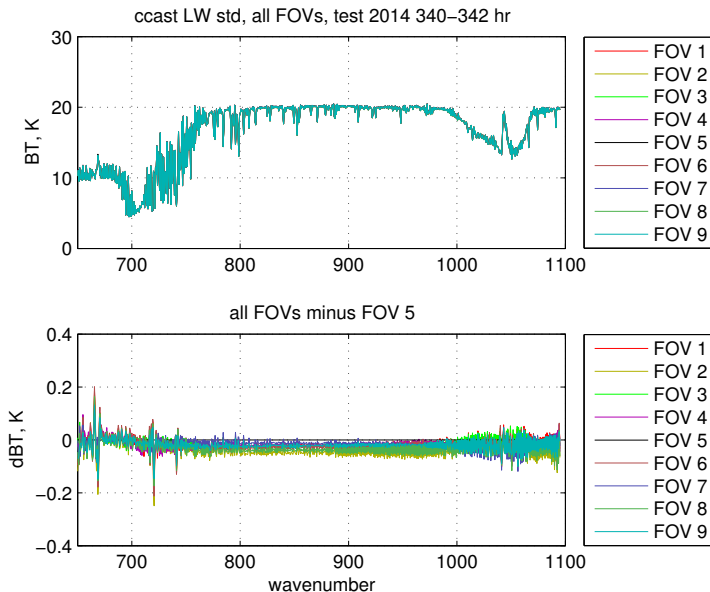
# SW discussion

- ▶ CCAST and NOAA are generally in good agreement
- ▶ residuals are significantly larger than for the LW band
- ▶ residuals and NOAA vs CCAST differences are generally greatest for the coldest lines and regions
- ▶ FOV 7 minus FOV 5 is significantly greater than for other FOVs at 2255 and 2359  $\text{cm}^{-1}$ , for both CCAST and NOAA

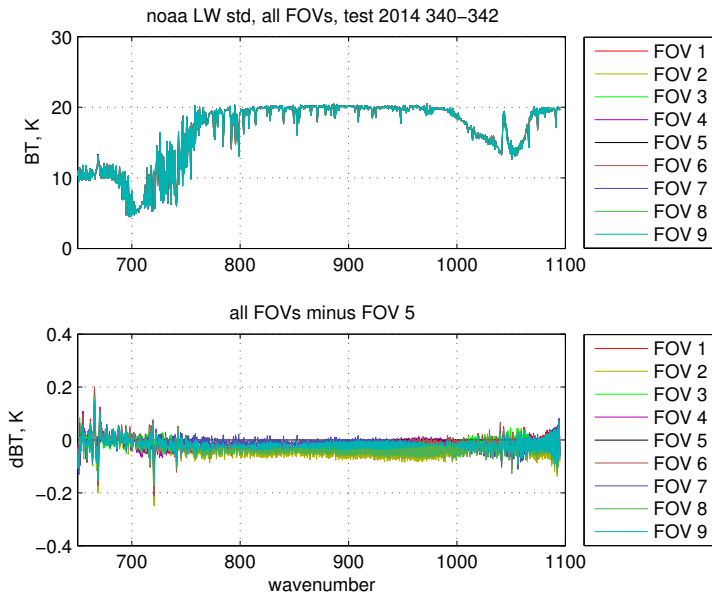
# conclusions

- ▶ there is significant convergence in the CCAST and NOAA processing
- ▶ variation due to nonlinearity, especially for the MW band, is significantly greater than some of the more subtle effects we have been considering recently
- ▶ note again that these results are relative to FOV 5 or are direct NOAA vs CCAST comparisons, and not comparisons with expected observed radiance from model data or radiance from other sounders
- ▶ supplementary slides
  - ▶ a comparison of standard deviations
  - ▶ CCAST ILS and calibration equations

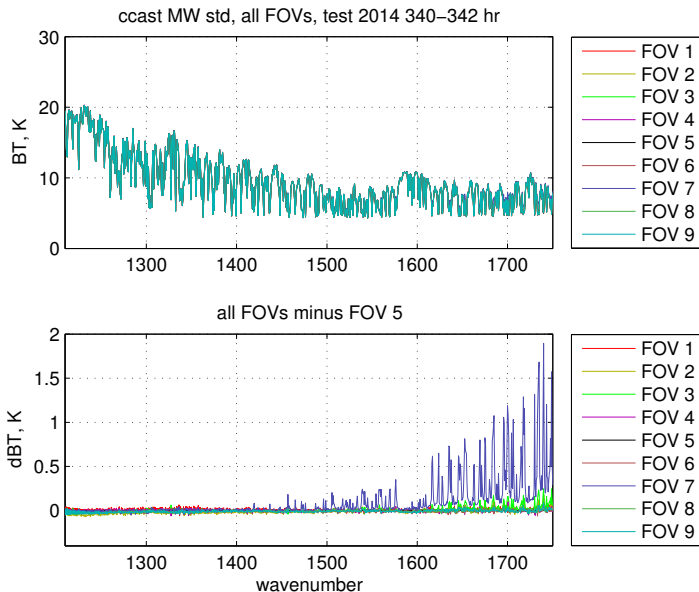
# ccast LW std



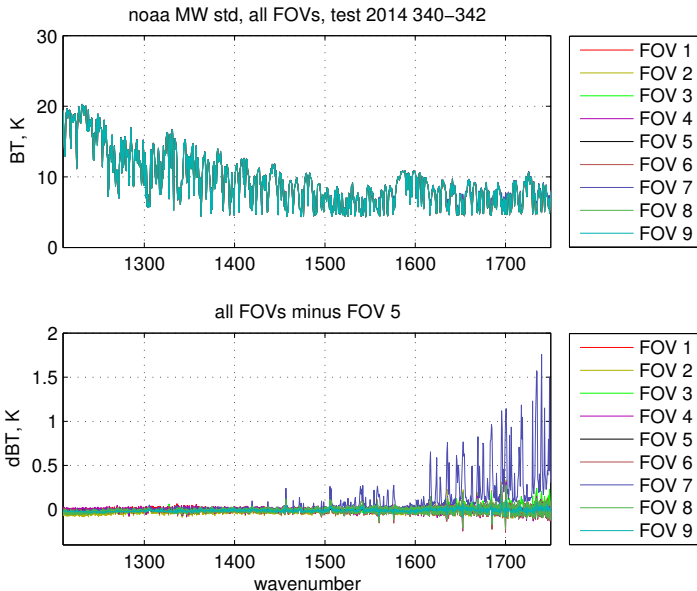
# noaa LW std



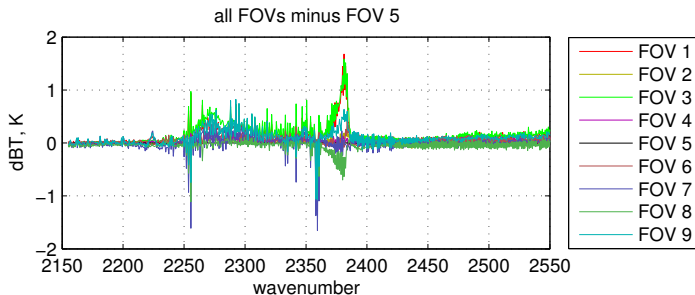
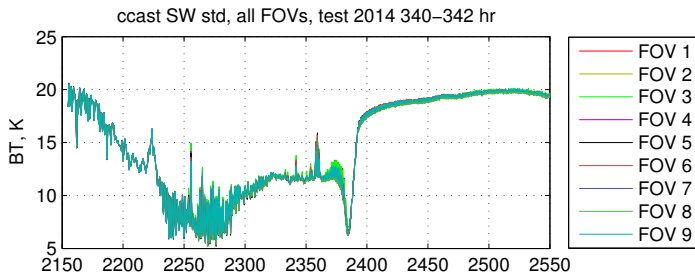
# ccast MW std



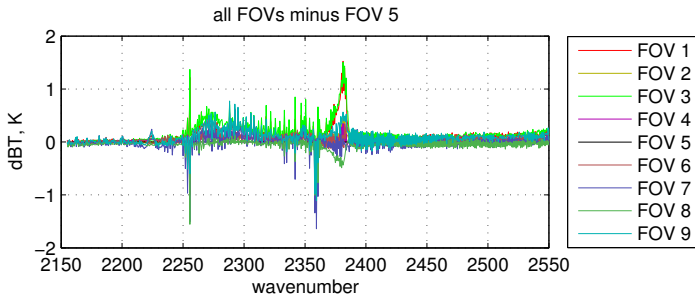
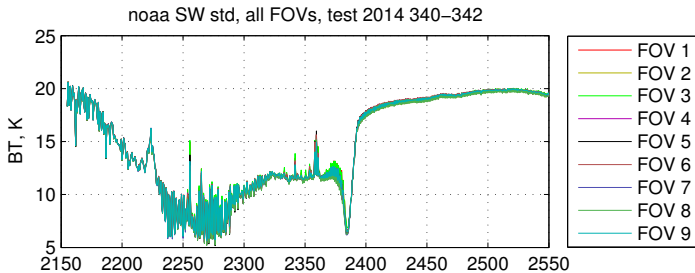
# noaa MW std



# ccast SW std



# noaa SW std





## calibration equation

The CCAST reference calibration equation is

$$r_{\text{OBS}} = F \cdot r_{\text{ICT}} \cdot f \cdot \text{SA}^{-1} \cdot f \cdot \frac{\text{ES} - \text{SP}}{\text{IT} - \text{SP}}$$

- ▶  $r_{\text{OBS}}$  is calibrated radiance at the user grid
- ▶  $F$  is Fourier interpolation from sensor to user grid
- ▶  $f$  is a raised-cosine bandpass filter
- ▶  $r_{\text{ICT}}$  is expected ICT radiance at the sensor grid
- ▶  $\text{SA}^{-1}$  is the inverse of the ILS matrix
- ▶ ES is earth-scene count spectra
- ▶ IT is calibration target count spectra
- ▶ SP is space-look count spectra

## calibration notes

- ▶ the IT and SP looks are averaged over several scans
- ▶ the UW nonlinearity correction is applied to count spectra before application of the calibration equation
- ▶ as part of the nonlinearity correction we divide the count spectra by the numeric filter at the sensor grid, but note this cancels out in the ratio  $(ES - SP)/(IT - SP)$
- ▶ the passband for  $f$  is the user grid. The wings are parameters currently set at 15, 20, and 22  $\text{cm}^{-1}$  for the LW, MW, and SW bands
- ▶  $f \cdot SA^{-1} \cdot f$  can be considered as a physically-based smoothing of the rows and columns of  $SA^{-1}$
- ▶  $F$  is a zero-filled double Fourier interpolation

## ccast ILS

the CrIS ILS for  $\text{FOV}_i$  can be represented as

$$\int_{\text{FOV}_i} w_i(\theta) \text{psinc}(2\pi d(\nu - \nu_0 \cos \theta)) d\theta$$

- ▶  $d$  is max OPD
- ▶  $\nu$  is frequency
- ▶  $\nu_0$  is reference or channel frequency
- ▶  $\text{psinc}(x) = \sin(x)/(\sin(x/n))$  for  $x \neq 0$ , 1 for  $x = 0$ , where  $n$  is the number of points in the sensor grid
- ▶  $\text{psinc}(2\pi d(\nu - \nu_0 \cos \theta))$  gives the ILS for a single ray at off-axis angle  $\theta$
- ▶ integration is over the intersection of on-axis arcs with  $\text{FOV}_i$ , with  $w_i(\theta)$  the length of an intersecting arc at off-axis angle  $\theta$