

# CrIS FOV 5 Anomaly and the Nonlinearity Correction

\*\*\*\* DRAFT \*\*\*\*

H. E. Motteler, L. L. Strow

UMBC Atmospheric Spectroscopy Lab  
Joint Center for Earth Systems Technology

September 29, 2016

# introduction

- ▶ Bob Knuteson (most recently) and others have shown FOV5 at  $668.125\text{ cm}^{-1}$  can have an anomalously warm brightness temperature relative to the other FOVs when the  $900\text{ cm}^{-1}$  window channel is cold
- ▶ we show a corresponding anomaly from the DC level integral of the nonlinearity correction, and for that correction, which may explain the brightness temp anomaly

# methods

- ▶ the nonlinearity correction applied to FOV 5 is significantly less than the corrections applied to other FOVs
- ▶ however the correction applied to FOV 5 relative to the other FOVs is slightly greater for spectra with a cold vs a warm window region
- ▶ this is because the DC level integral for FOV 5 relative to the other FOVs is slightly greater for spectra with a cold vs a warm window region

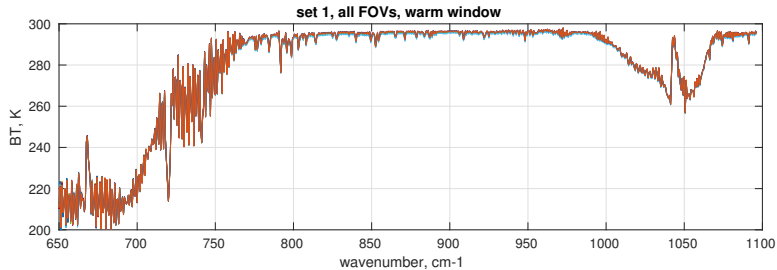
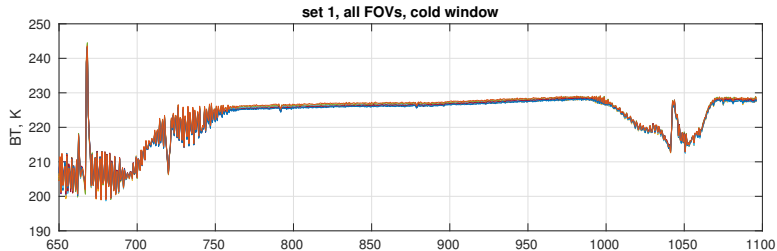
## methods

- ▶ we find homogeneous scenes for two cases: warm and cold at  $900\text{ cm}^{-1}$ , as follows
- ▶ LW brightness temperature spectra were compared for all FOVs for FOR 15 and 16 from 1–3 Jan 2016. If the max RMS difference between any pair of FOVs was less than 1K and the mean over all FOVs at  $900\text{ cm}^{-1}$  was less than 230K, the FOR was saved as a “cold FOR”
- ▶ similarly, if the max RMS difference between any pair of FOVs was less than 1K and the mean over all FOVs at  $900\text{ cm}^{-1}$  was greater than 270K, the FOR was saved as a “warm FOR”.

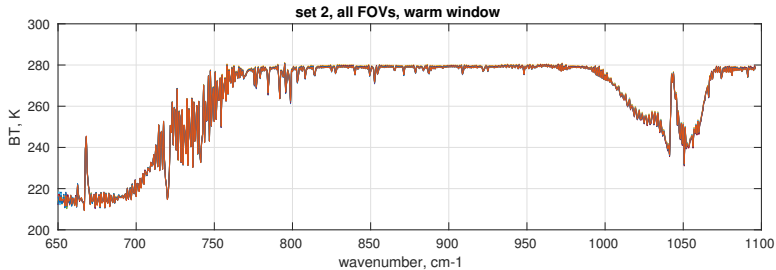
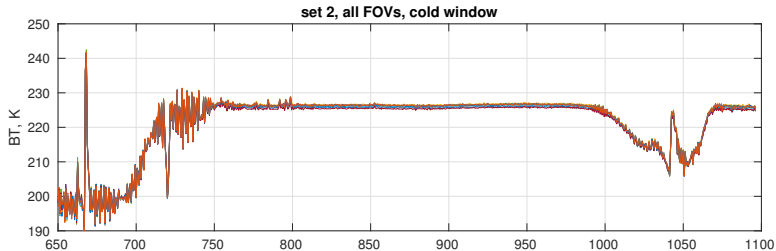
## methods

- ▶ we found hundreds of warm FORs but only 7 cold FORs over the 3-day period. The cold FORs were then further ordered by their temperature at  $668.125\text{ cm}^{-1}$ , and we took the warmest three (at 243, 242, and 241K) and paired these with warm sets with similar temperatures at  $668.125\text{ cm}^{-1}$

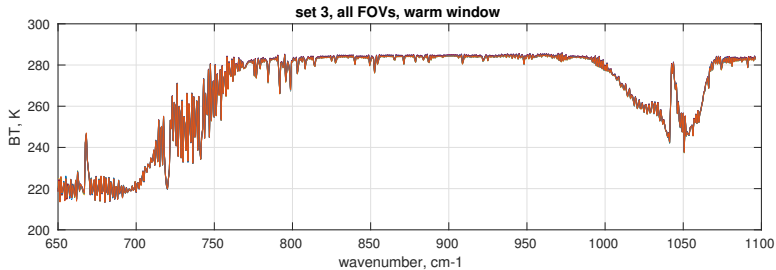
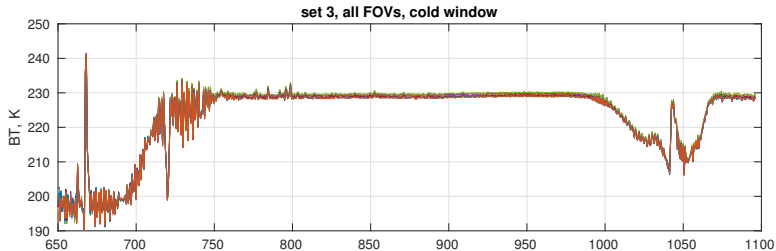
# test set 1



## test set 2

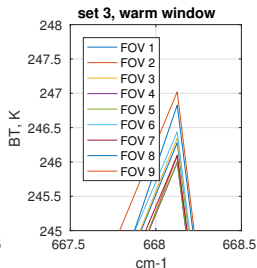
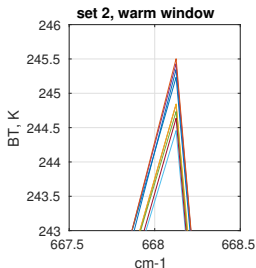
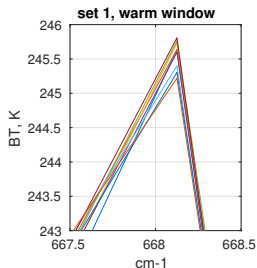
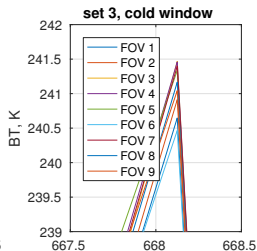
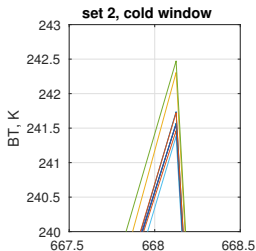
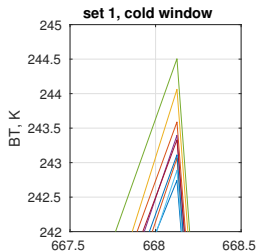


## test set 3





# 668 cm<sup>-1</sup> zoom



## nonlinearity correction

- ▶ let  $/$  be pointwise division, and

$$r_{\text{in}}^s = r_{\text{in}} / f_{\text{N}}$$

$$r_{\text{sp}}^s = r_{\text{sp}} / f_{\text{N}}$$

- ▶ the DC level is given by

$$v_{\text{dc}} = v_{\text{inst}} + \frac{2 \cdot \sum_{i=1}^n |r_{\text{in}}^s - r_{\text{sp}}^s|}{c_{\text{m}} \cdot c_{\text{a}} \cdot c_{\text{p}} \cdot d \cdot n}$$

- ▶ the corrected radiances (scaled by  $f_{\text{N}}$ ) are

$$r_{\text{out}}^s = r_{\text{in}}^s \cdot (1 + 2 \cdot a_2 \cdot v_{\text{dc}})$$

# parameters

- ▶  $r_{in}$  is scene count spectra
- ▶  $r_{sp}$  is space-look count spectra
- ▶  $n$  is the number of decimated points
- ▶  $d$  is the decimation factor
- ▶  $c_m$  is modulation efficiency
- ▶  $c_p$  is PGA gain
- ▶  $c_a$  is A/D gain
- ▶  $v_{inst}$  instrument contribution to DC level
- ▶  $v_{dc}$  is estimated DC level
- ▶  $f_N$  is the numeric filter at the sensor grid
- ▶  $a_2$  are the correction parameters

# DC Level Integral

Values from the DC level integral

| FOV | set 1  |        | set 2  |        | set 3  |        |
|-----|--------|--------|--------|--------|--------|--------|
|     | cold   | warm   | cold   | warm   | cold   | warm   |
| 1   | 0.1190 | 0.3988 | 0.1157 | 0.3047 | 0.1231 | 0.3323 |
| 2   | 0.1269 | 0.4202 | 0.1233 | 0.3233 | 0.1314 | 0.3476 |
| 3   | 0.1188 | 0.3965 | 0.1173 | 0.3062 | 0.1251 | 0.3276 |
| 4   | 0.1258 | 0.4171 | 0.1232 | 0.3192 | 0.1313 | 0.3476 |
| 5   | 0.1353 | 0.4447 | 0.1312 | 0.3407 | 0.1409 | 0.3683 |
| 6   | 0.1296 | 0.4250 | 0.1266 | 0.3291 | 0.1337 | 0.3552 |
| 7   | 0.1343 | 0.4476 | 0.1292 | 0.3420 | 0.1387 | 0.3711 |
| 8   | 0.1244 | 0.4188 | 0.1223 | 0.3196 | 0.1286 | 0.3458 |
| 9   | 0.1155 | 0.3801 | 0.1117 | 0.2899 | 0.1170 | 0.3139 |

# correction factor

correction factor

| FOV | set 1  |        | set 2  |        | set 3  |        |
|-----|--------|--------|--------|--------|--------|--------|
|     | cold   | warm   | cold   | warm   | cold   | warm   |
| 1   | 1.0576 | 1.0684 | 1.0575 | 1.0648 | 1.0577 | 1.0659 |
| 2   | 1.0443 | 1.0527 | 1.0442 | 1.0500 | 1.0445 | 1.0507 |
| 3   | 1.0498 | 1.0588 | 1.0498 | 1.0558 | 1.0500 | 1.0565 |
| 4   | 1.0677 | 1.0804 | 1.0676 | 1.0761 | 1.0679 | 1.0774 |
| 5   | 1.0392 | 1.0475 | 1.0391 | 1.0447 | 1.0394 | 1.0455 |
| 6   | 1.0523 | 1.0620 | 1.0522 | 1.0589 | 1.0525 | 1.0597 |
| 7   | 1.0469 | 1.0560 | 1.0467 | 1.0530 | 1.0470 | 1.0538 |
| 8   | 1.0529 | 1.0631 | 1.0528 | 1.0597 | 1.0530 | 1.0606 |
| 9   | 1.0912 | 1.1073 | 1.0909 | 1.1018 | 1.0913 | 1.1033 |

## conclusions

- ▶ with the switch to extended res, we have seen a significant convergence in calibration algorithm performance
- ▶ the NOAA “SA-1 first” algorithm does slightly better when compared with reference truth convolved with responsivity, while the CCAST “ratio first” algorithm does slightly better when compared with reference truth convolved with a flat passband
- ▶ this may be because responsivity cancels out more completely in the ratio-first method
- ▶ because reference truth convolved with a flat passband is a more conventional and non instrument-specific standard, the ccast algorithm, or some similar ratio-first method, may be preferable