

Investigating Flux Densities and Polarization of Methanol Masers in the Orion Nebula



Maneth Perera

Illinois Math and Science Academy

Advisor: Dr. Esteban Araya

Collaborators: Gabriel Sojka, Kenneth

VanZuiden

Western Illinois University



NASA, Hubble Space Telescope

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- ACCESS



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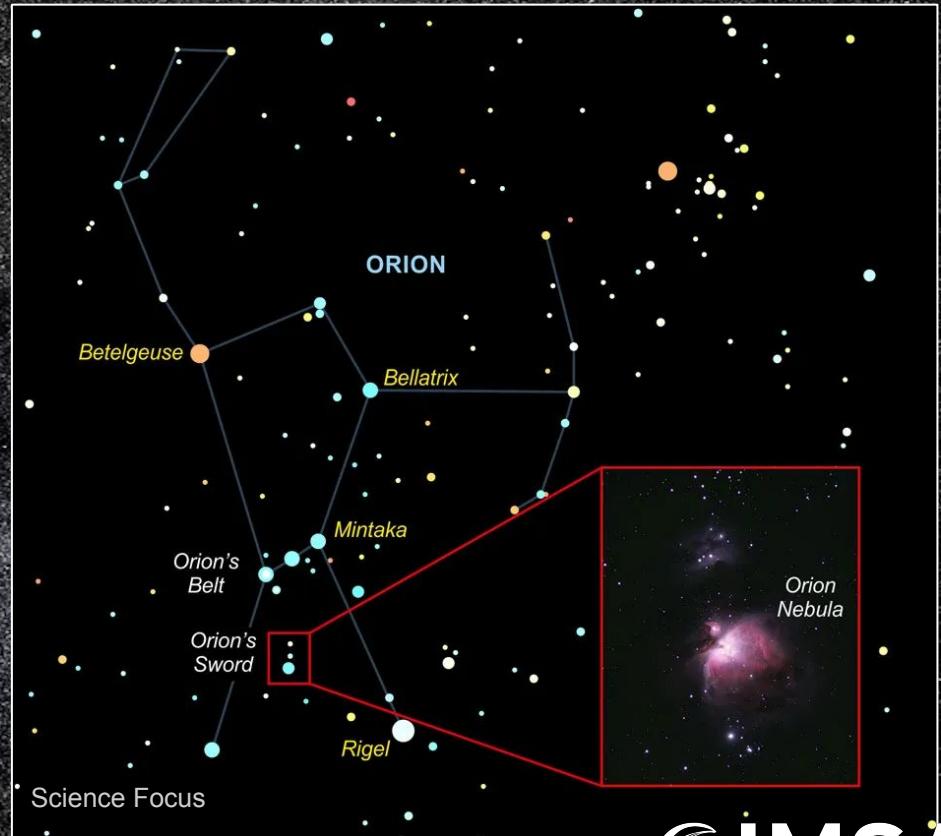
Background

- VLA (Very Large Array)
 - Radio Telescope
 - New Mexico, U.S.
- VOLA (VLA Orion-A Large Survey)
- 20 Epochs
- From April 3, 2022
- To July 1, 2022



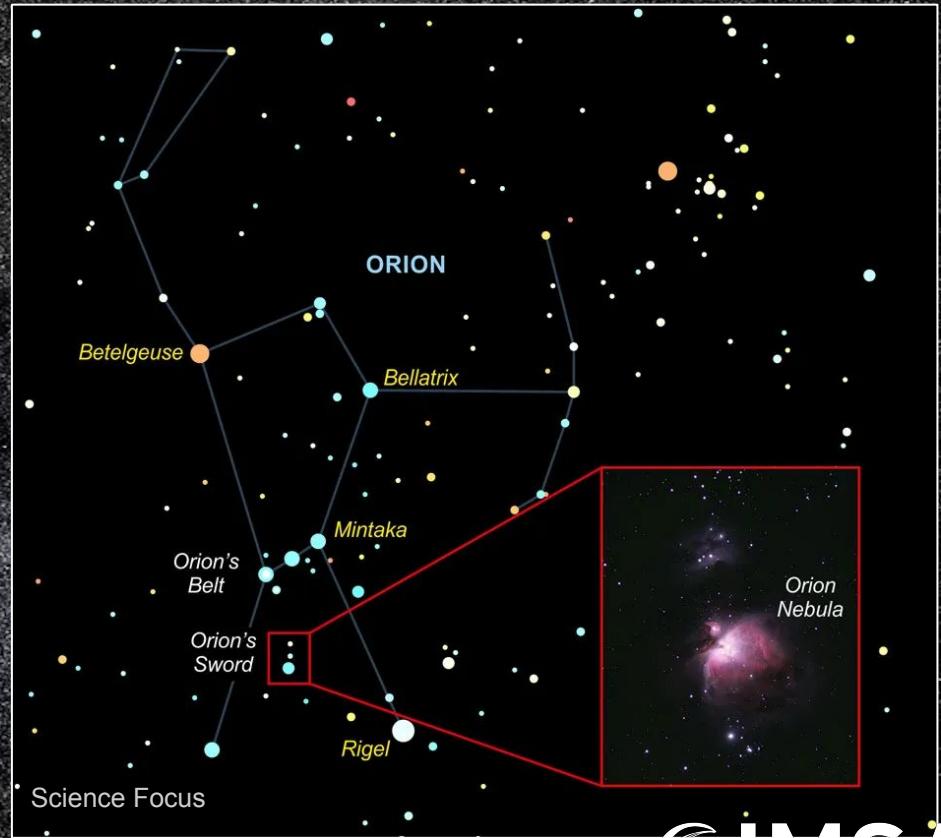
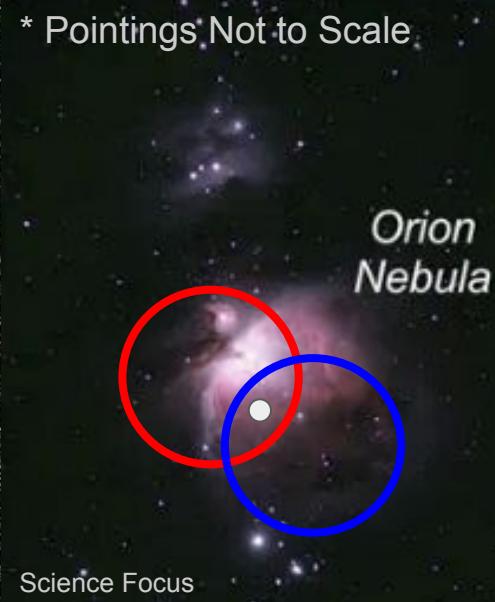
Background

■ Orion Nebula



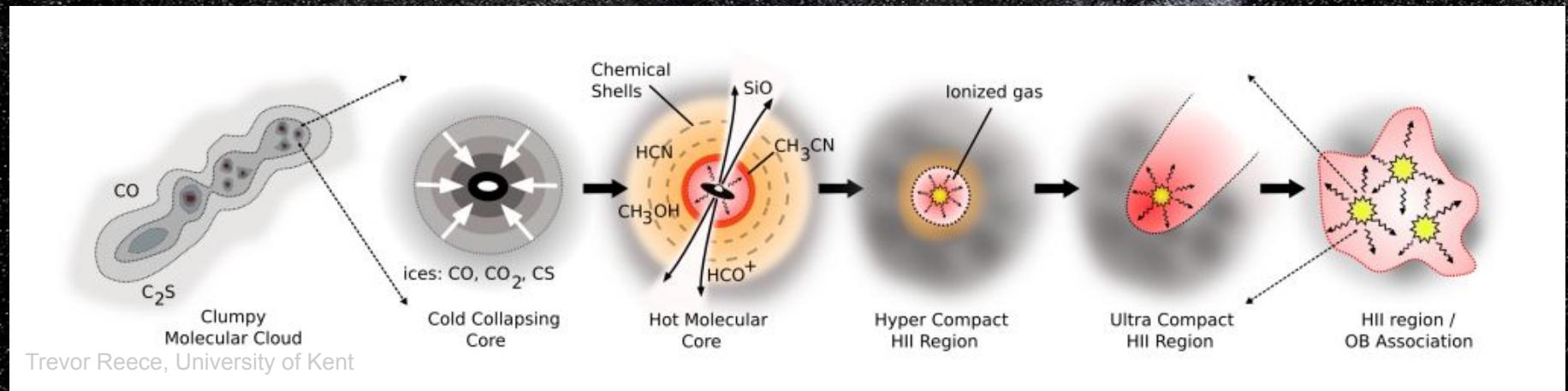
Background

- Orion Nebula
- Pointings



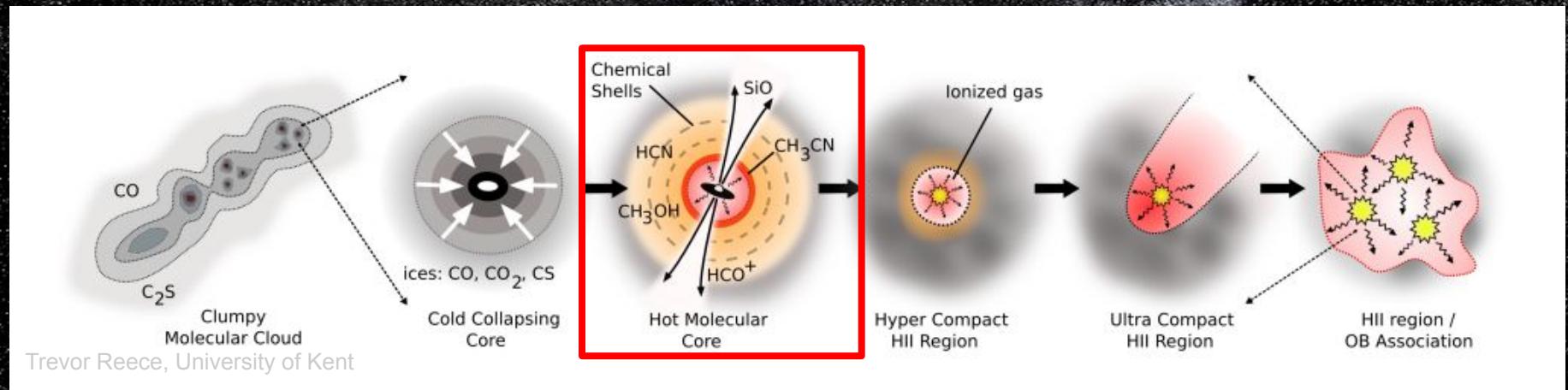
Background

■ High Mass Star Formation



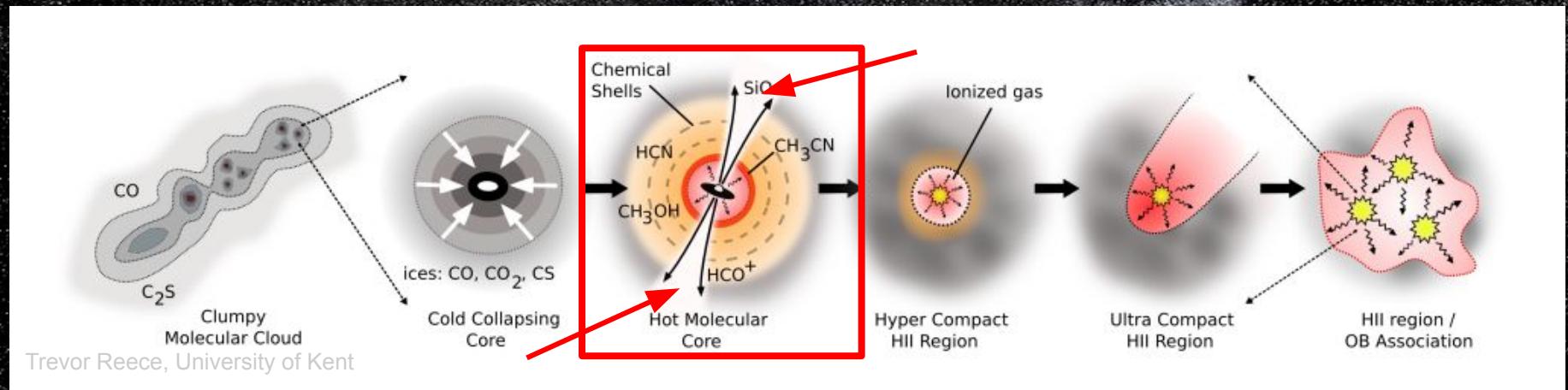
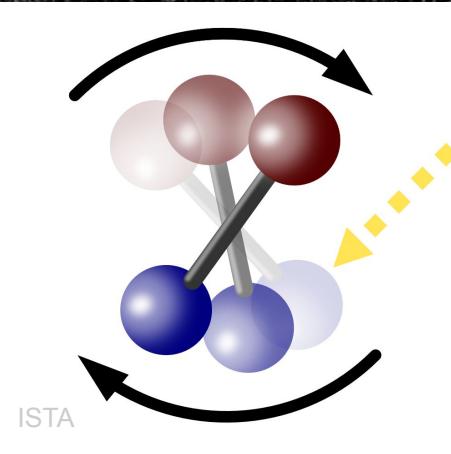
Background

- High Mass Star Formation
 - Molecular Outflow



Background

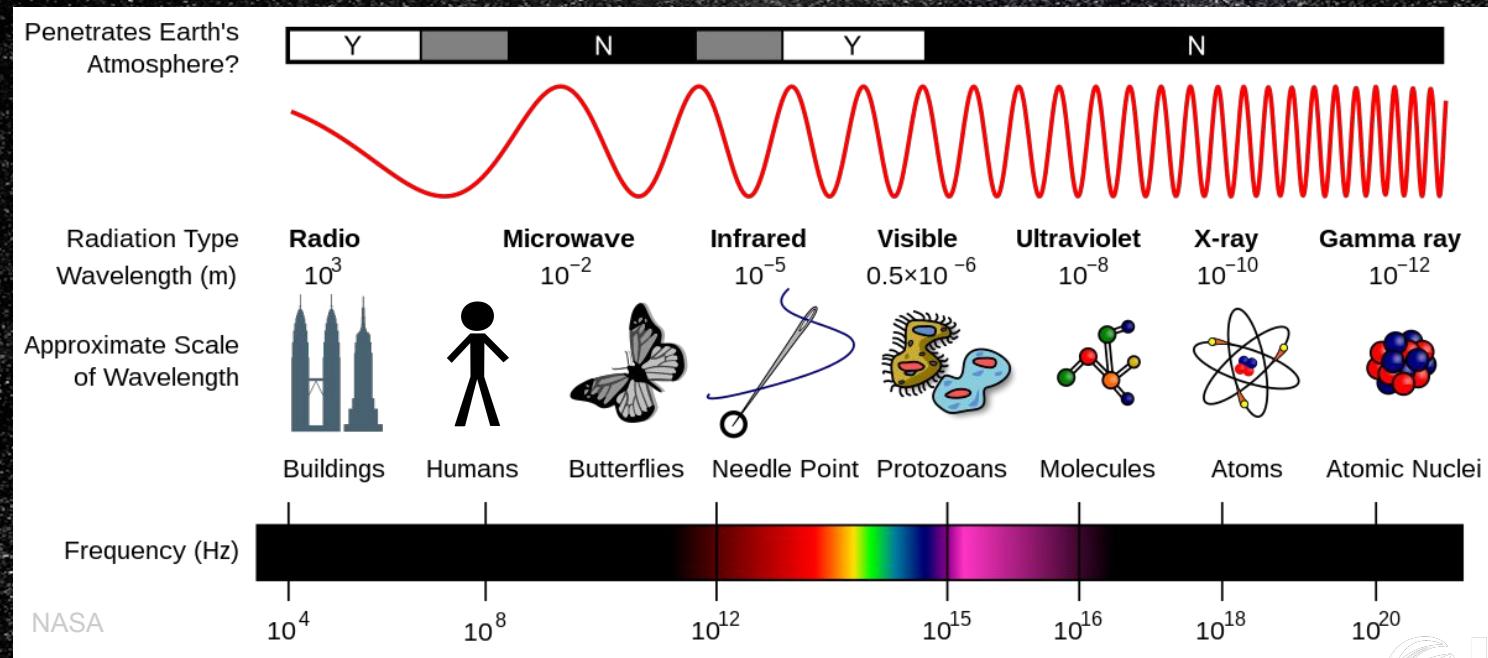
- High Mass Star Formation
 - Molecular Outflow
 - Maser



Background

- Microwave Radiation
- Flux Density

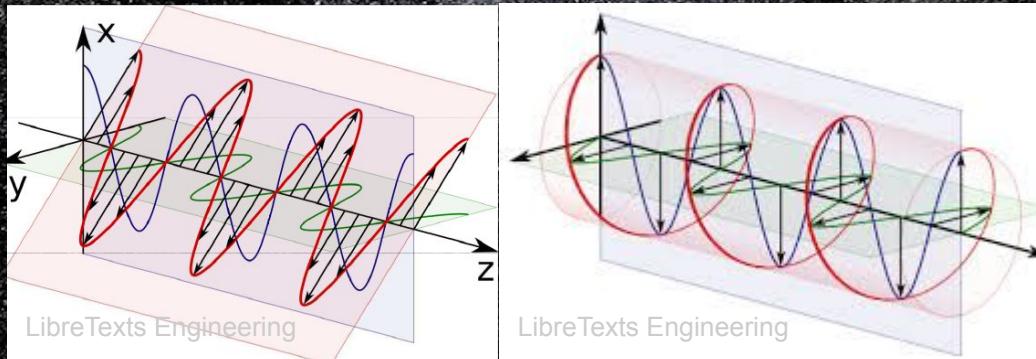
$$1 \text{ Jy} = 10^{-26} \frac{W}{m^2 Hz}$$



TMSA

Background

- Light Wave Polarization
 - Stokes Intensity - IQUV
 - I - Overall Signal Intensity
 - Q, U - Linear Polarization, 45 degree increments
 - V - Circular Polarization, Left/Right-Hand

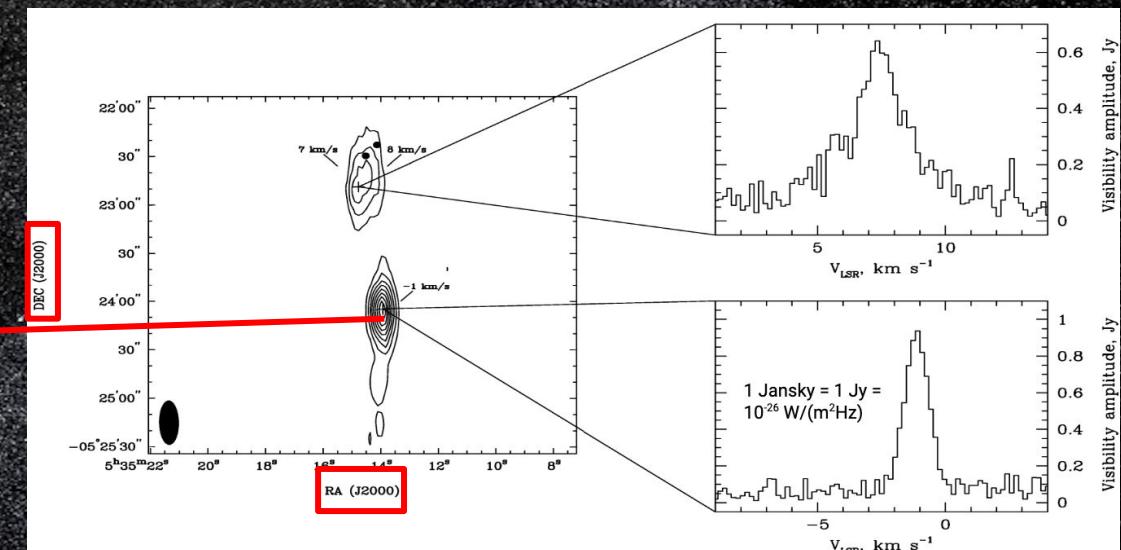


Background

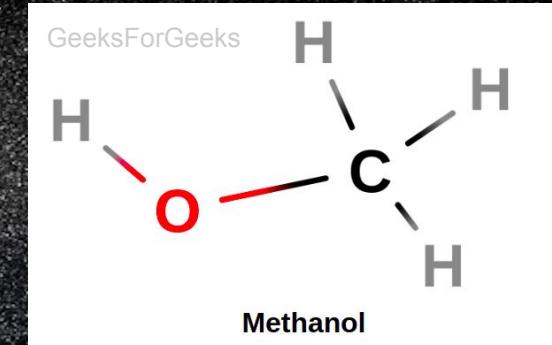
- Methanol (CH_3OH) Masers in Orion



Voronkov et al, (2005, MNRAS 362, 995)

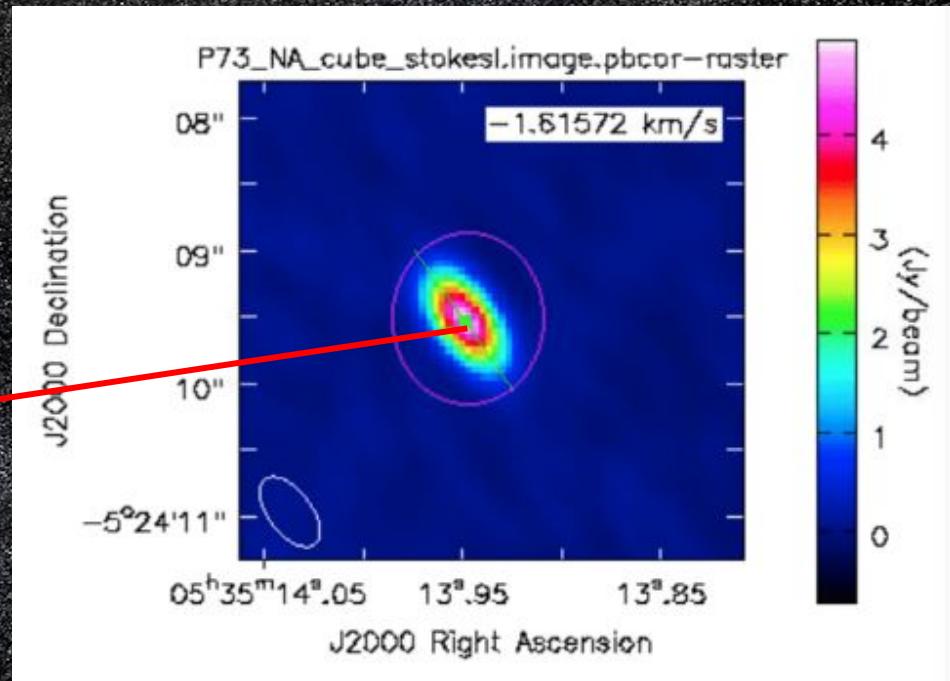
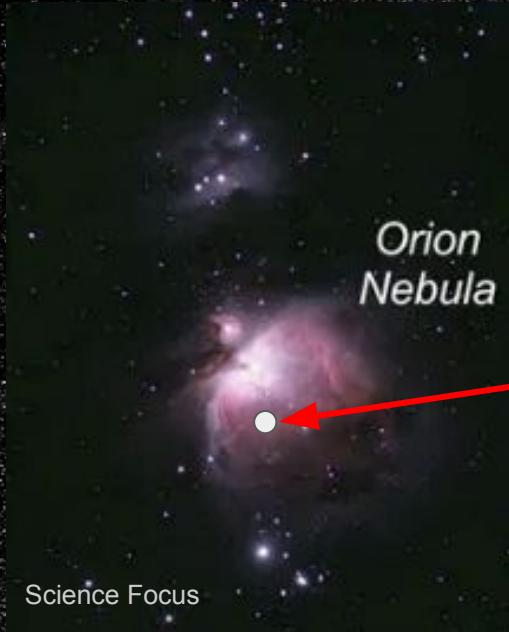


- Right Ascension, Declination => x, y in the Night Sky



Background

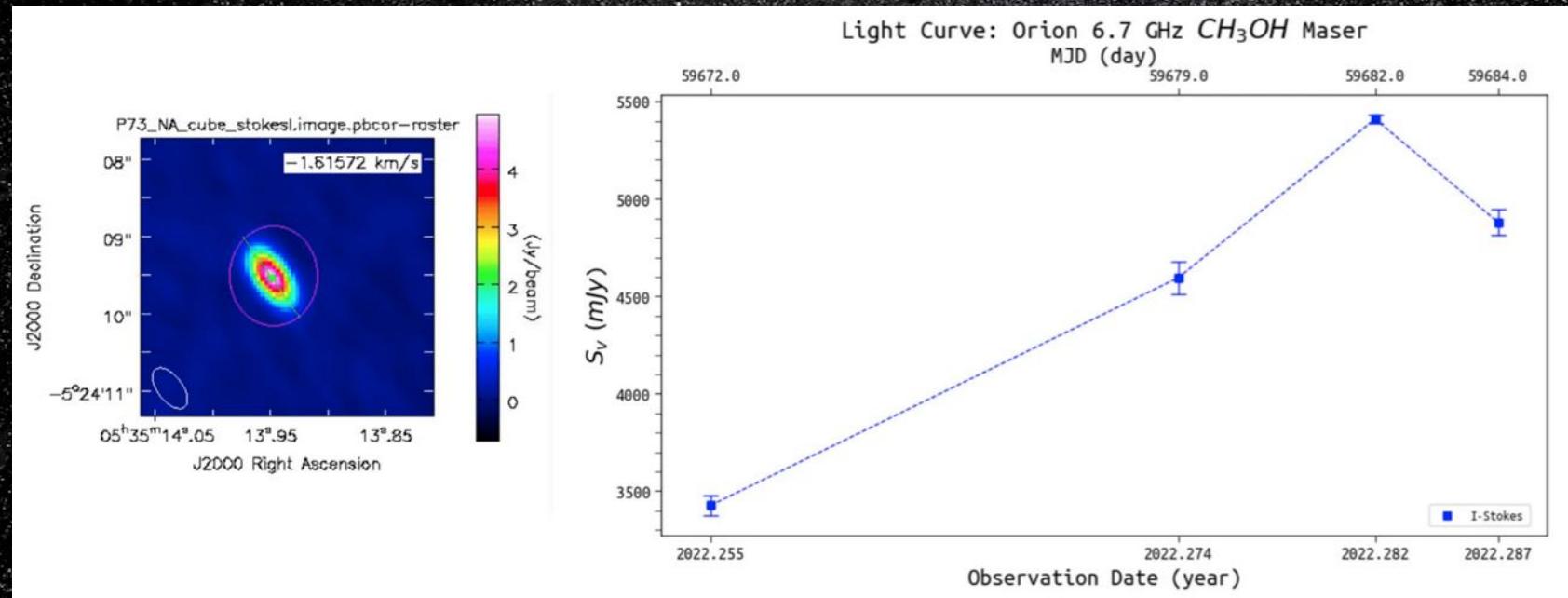
■ Methanol (CH_3OH) Masers in Orion



Sojka et al. (2024; AAS 244, 302.06)

Background

■ Methanol (CH_3OH) Masers in Orion

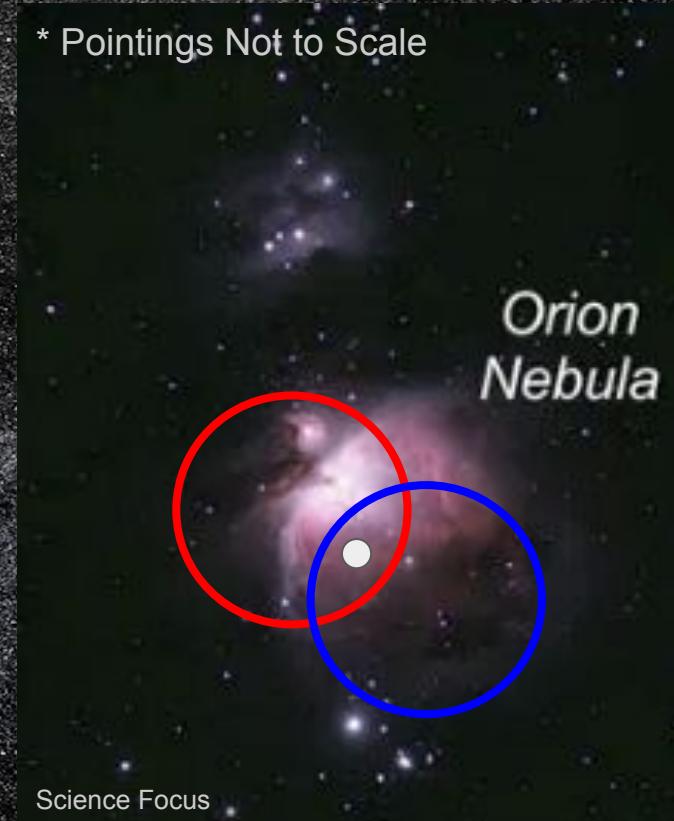


Sojka et al. (2024, AAS 244, 302.06)

Objectives

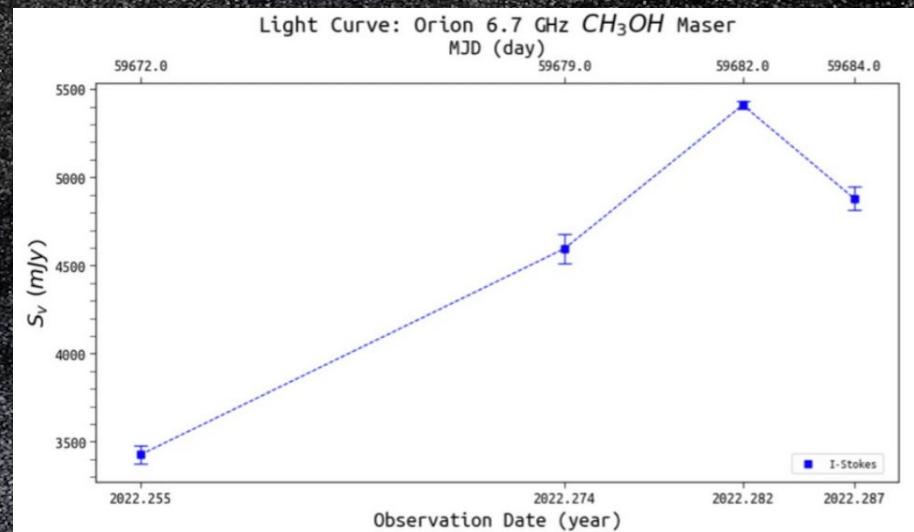
- Obj 1: Confirmation using pointings in the night sky

* Pointings Not to Scale



Objectives

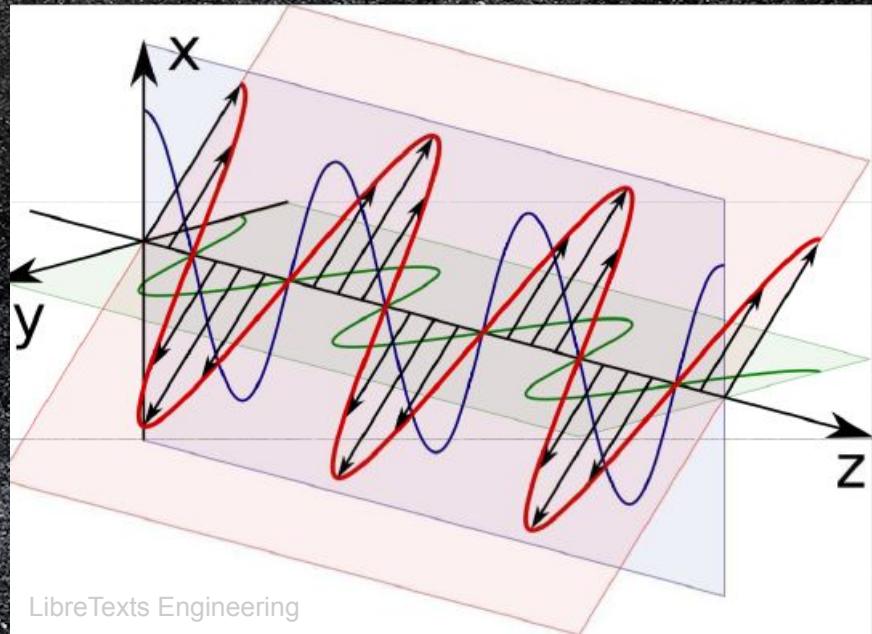
- Obj 1: Confirmation using pointings in the night sky
- Obj 2: Look for fast variability in the maser



Sojka et al. (2024, AAS 244, 302.06)

Objectives

- Obj 1: Confirmation using pointings in the night sky
- Obj 2: Look for fast variability in the maser
- Obj 3: Check for linear polarization (as well as circular polarization)



LibreTexts Engineering

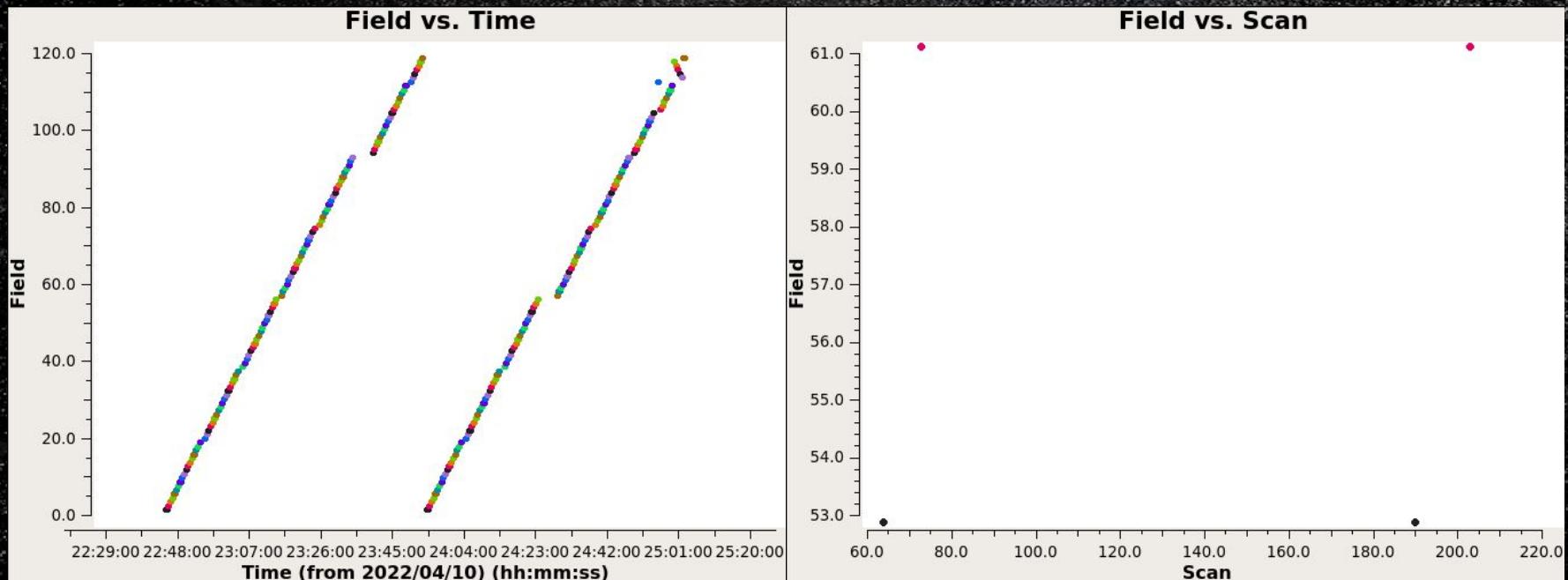
Methodology

- NSF ACCESS, Jetstream2, University of Indiana
- CASA (Common Astronomical Software Applications)



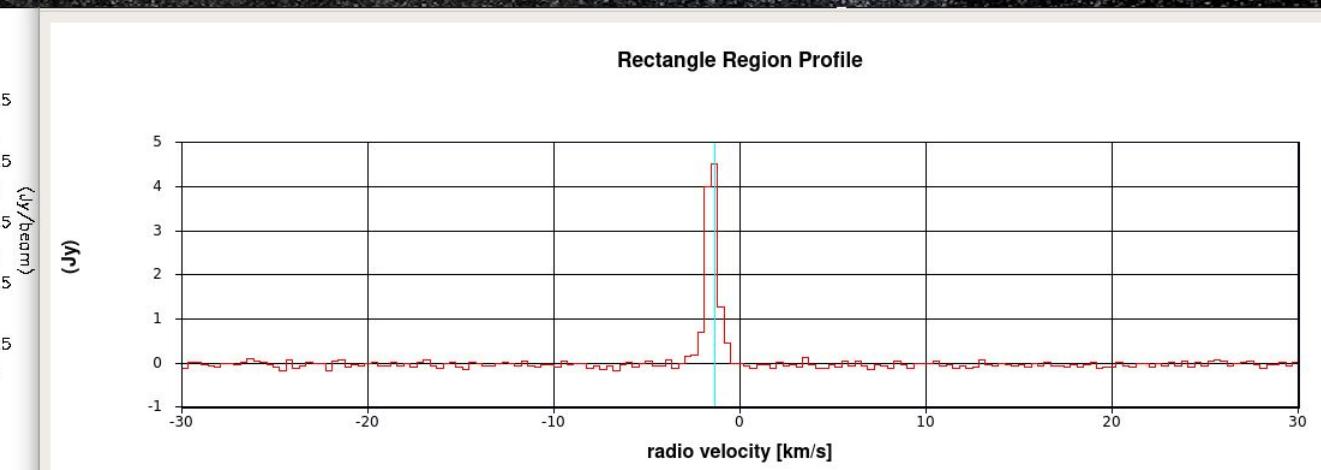
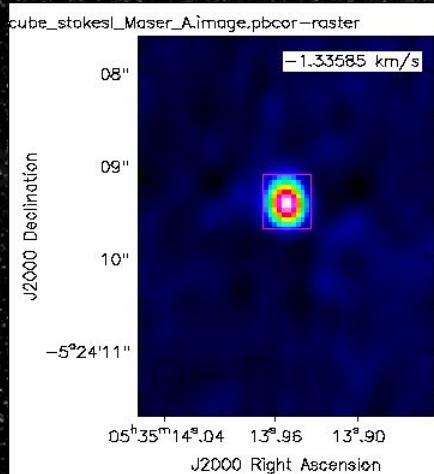
Methodology

■ Sky Scanning



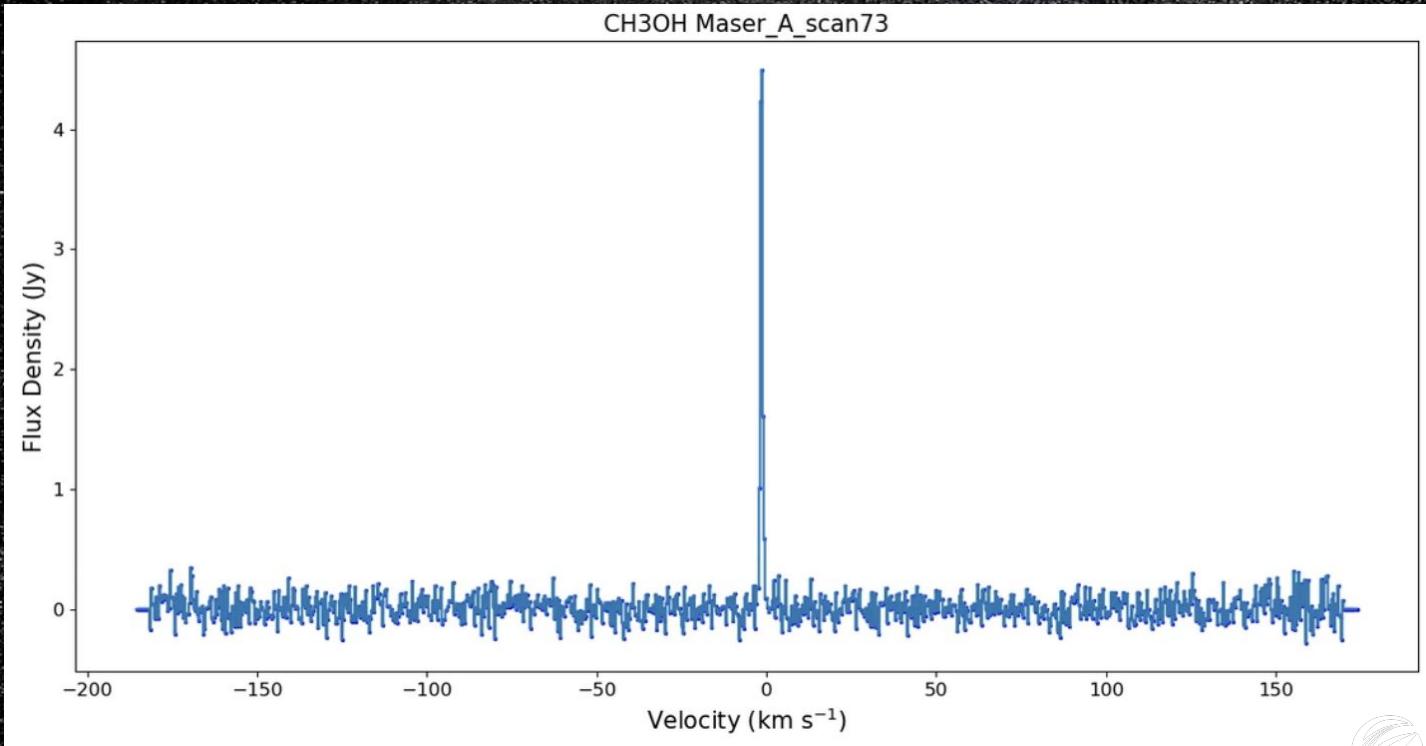
Methodology

- Spectral Imaging
- Spectral Lines



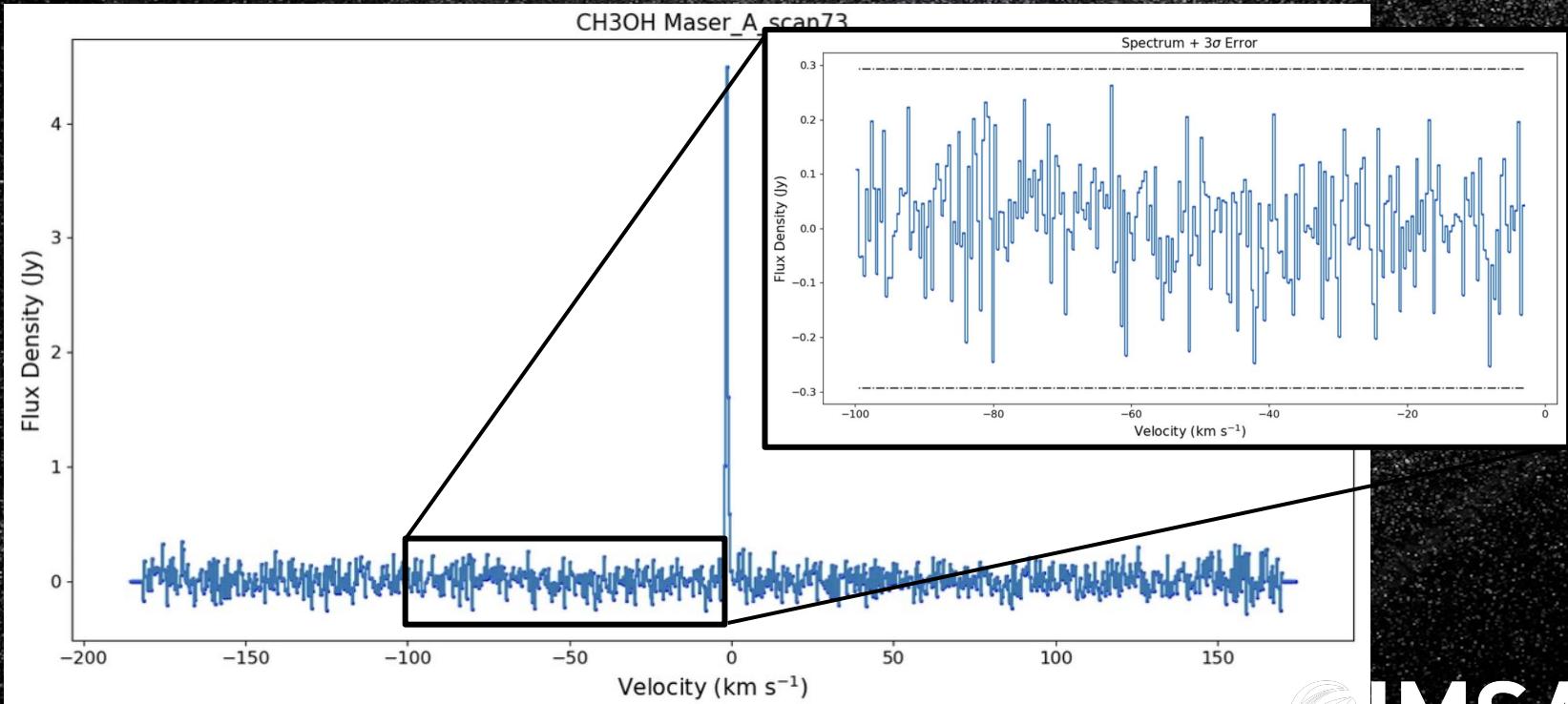
Methodology

■ Examples of Signal and Noise Graphs



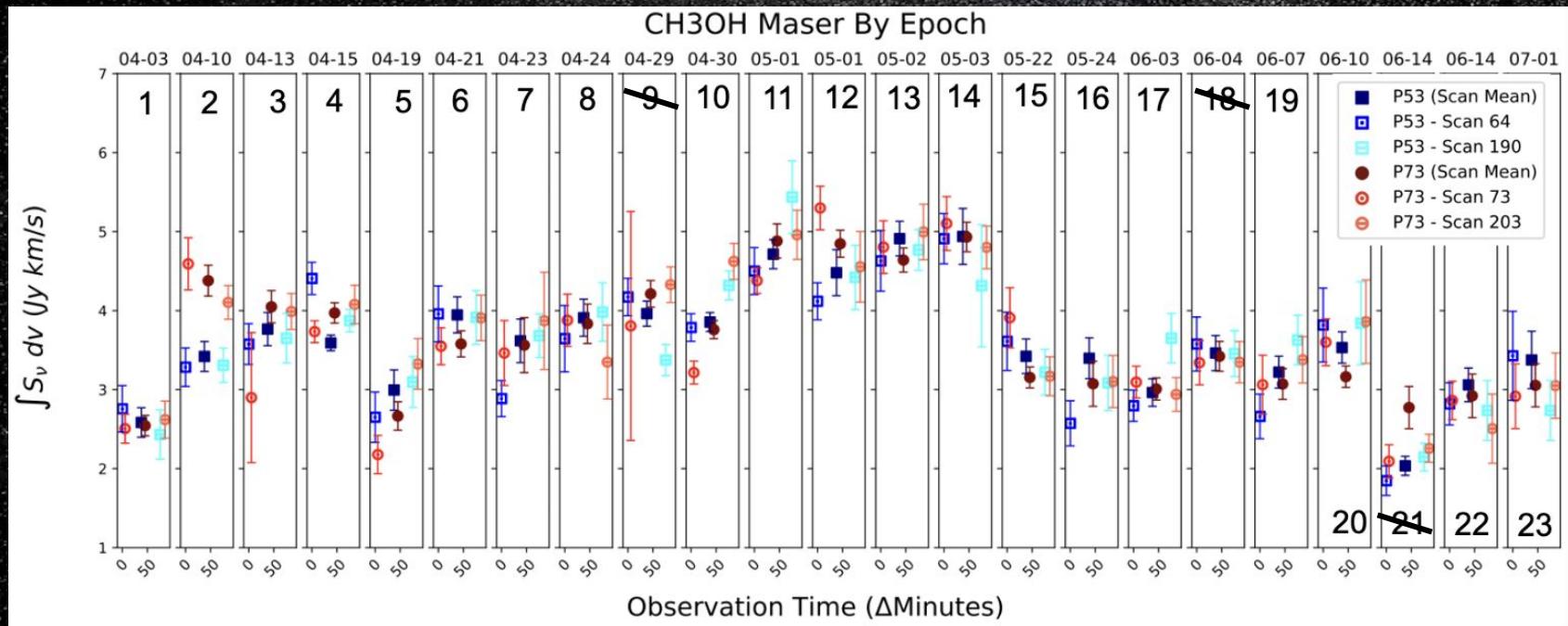
Methodology

■ Examples of Signal and Noise Graphs



Results and Discussion

■ Combined Light Curve of 23 epochs



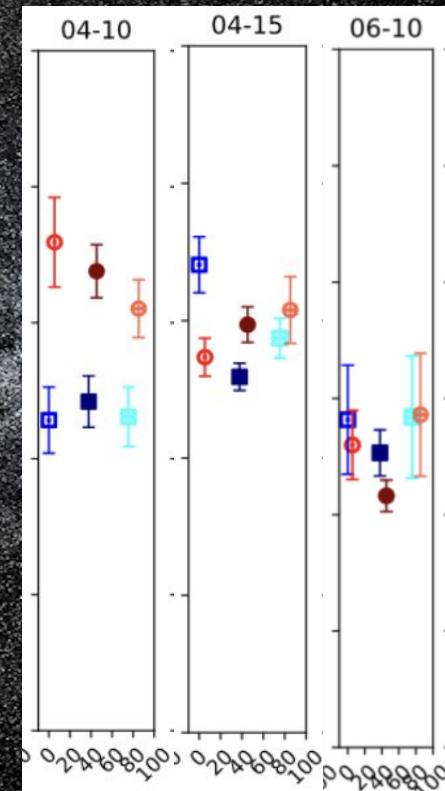
All Epochs

Objective 1

Epoch 2, Epoch 4

■ Overall Findings

- Flux densities consistent between different pointings for all epochs except epochs 2, 4, and 20
- General consistency for 17/20 epochs => general maser consistency

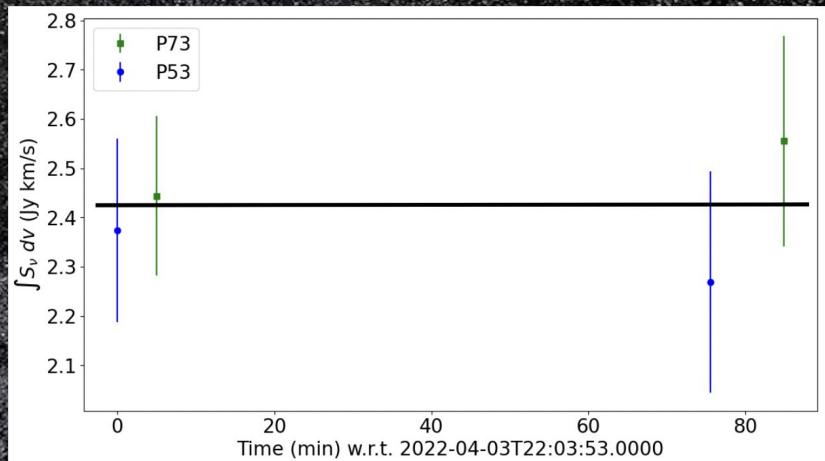


E2 E4 E20

Objective 2

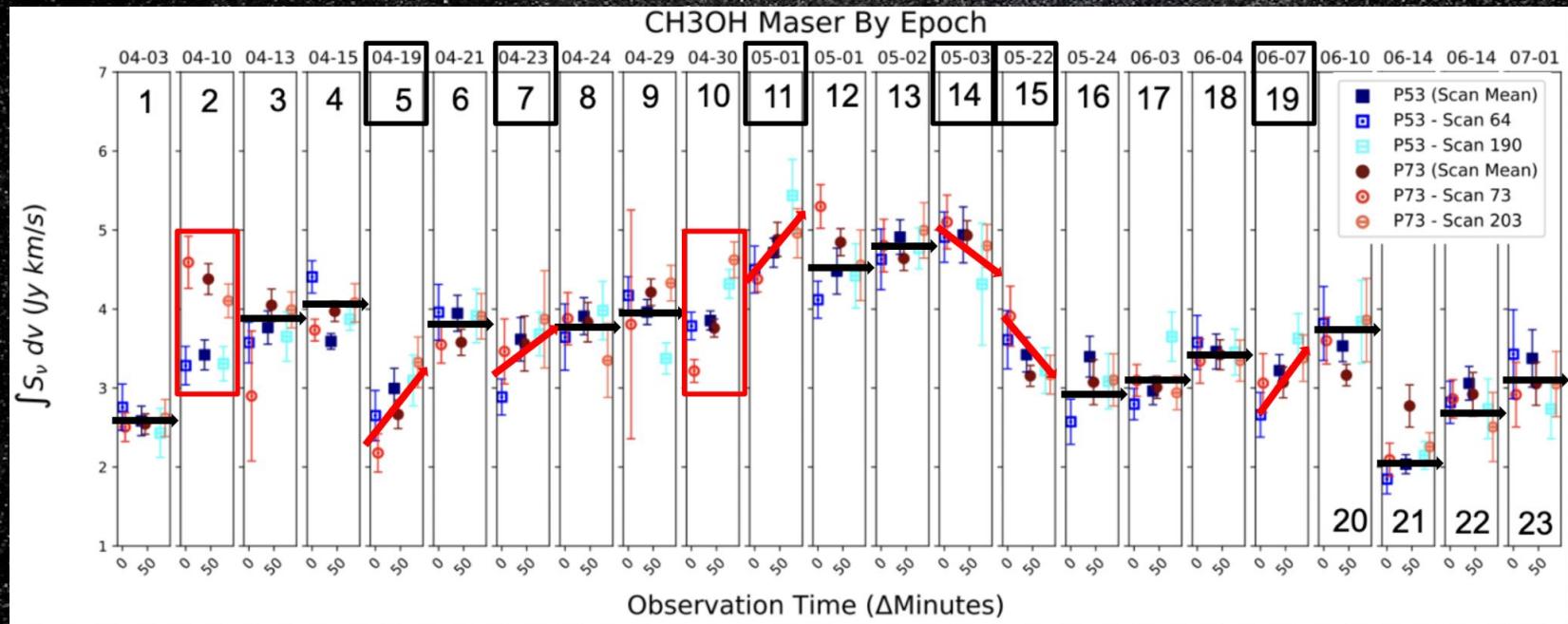
■ Overall Findings

- Possible evidence of fast variability in epochs with general trends
- Generally similar scan values for most epochs
- Offset scans in some epochs due to lack of telescope data or high noise



Objective 2

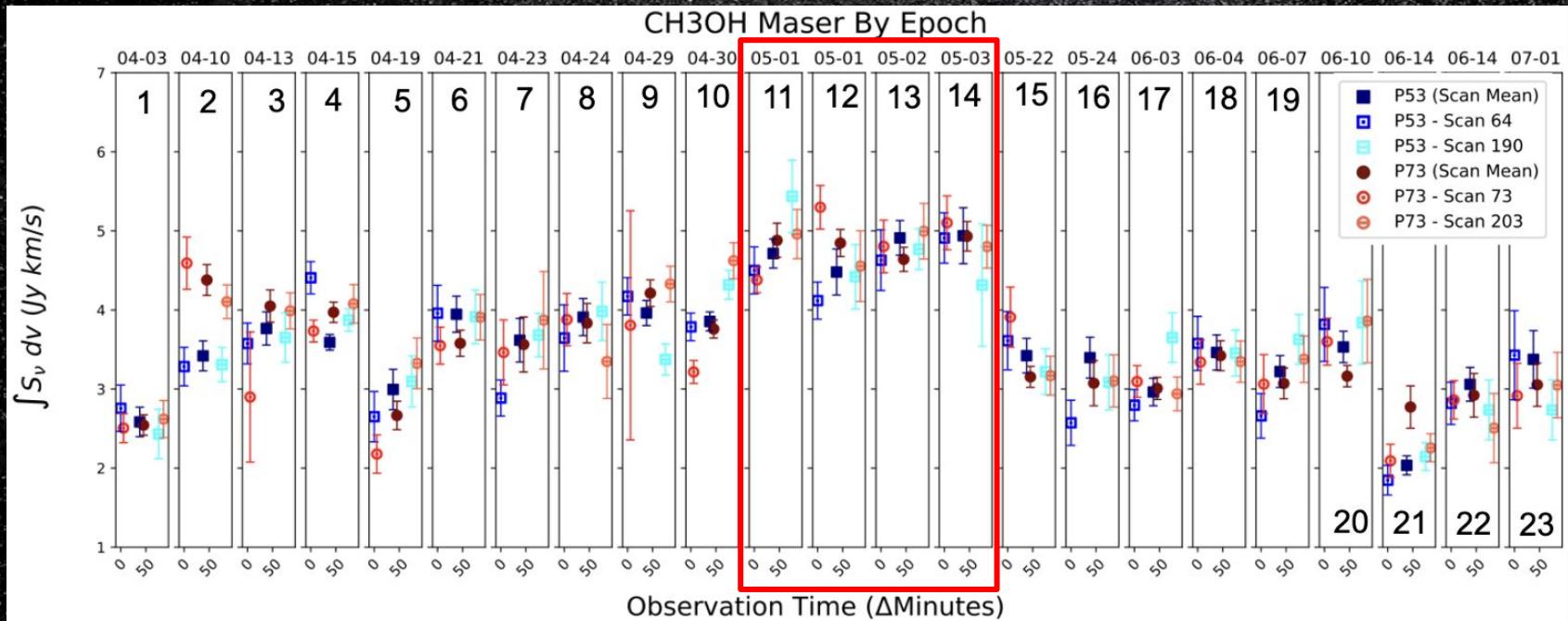
■ Fast Variability in the 20 Epochs



All Epochs

Discrepancy

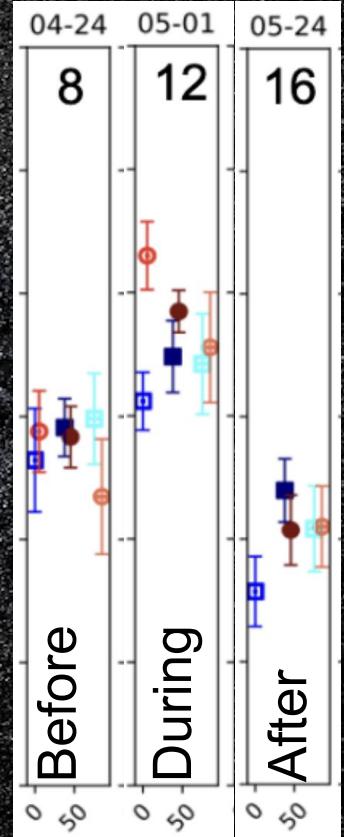
■ May 1st to 3rd Flare



Discrepancy

Epochs 8, 12, 16

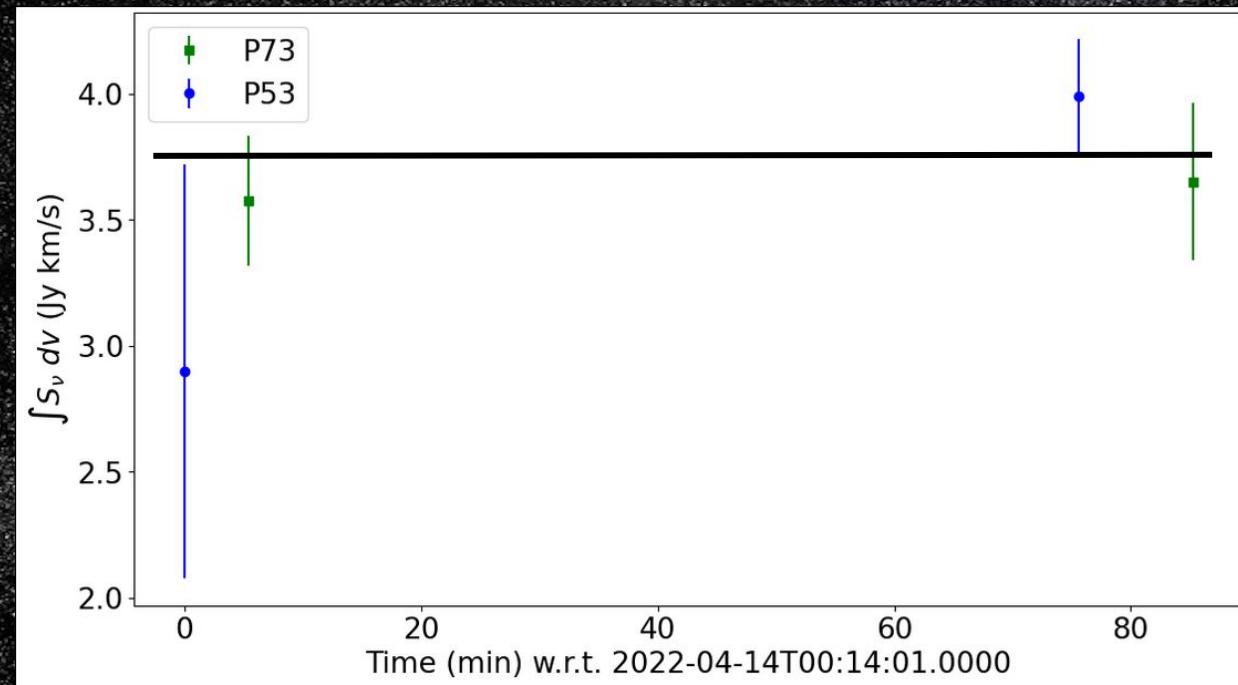
- May 1st to 3rd Flare
 - Significant increase in flux density in epochs 11 through 14
 - Possible Causes:
 - Nearby young stellar object's accretion disk grows, leading to a larger radiation field affecting maser
 - Shockwave front passes through maser



Discrepancy

■ Comparing Pointings

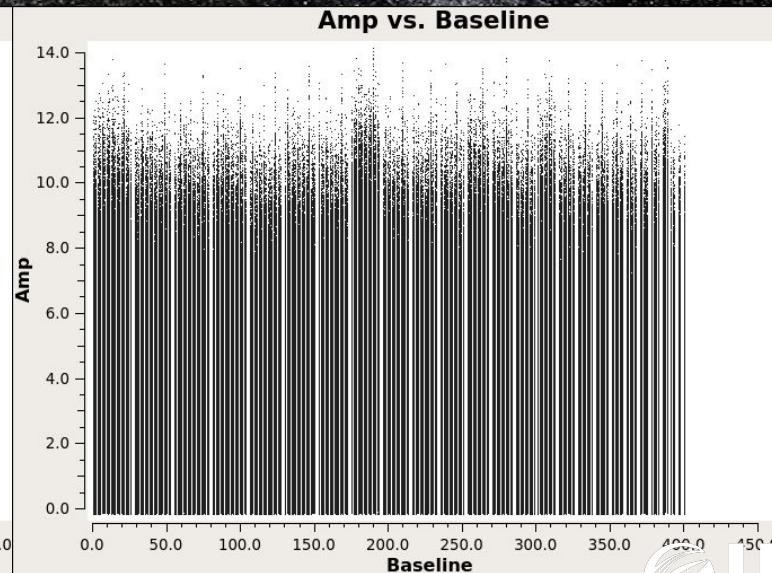
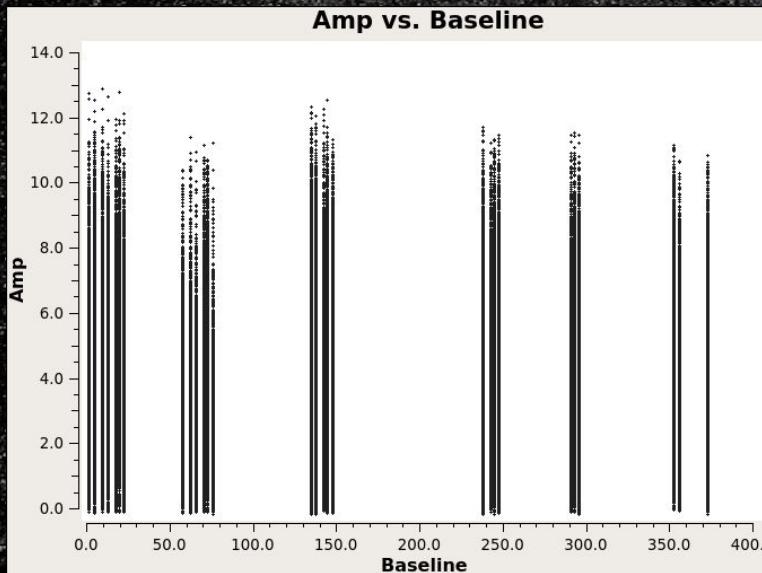
- Epoch 3 pointings are consistent, but one scan is offset



Discrepancy

■ Lack of data

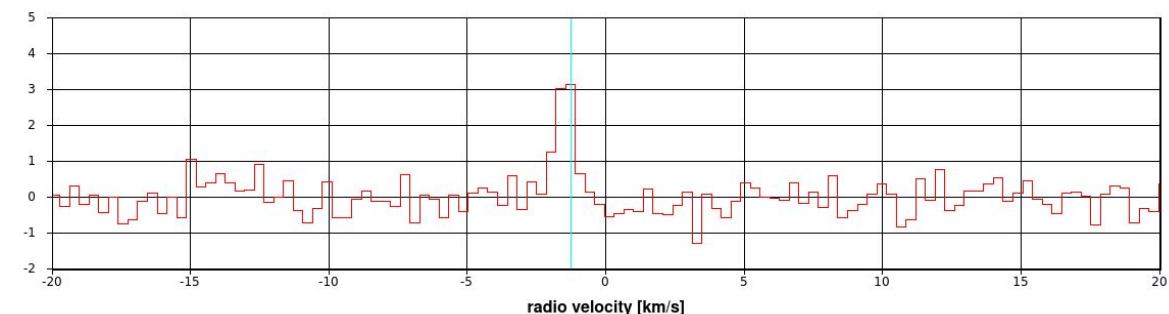
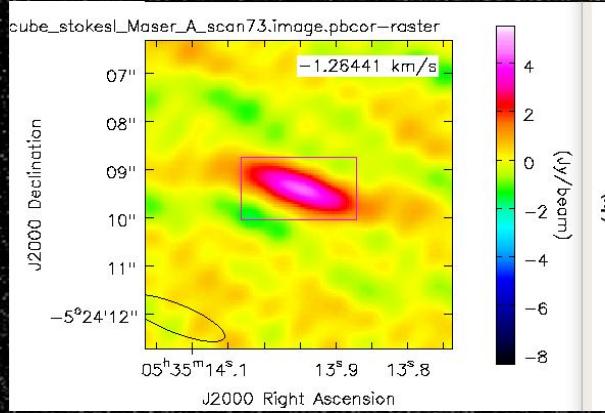
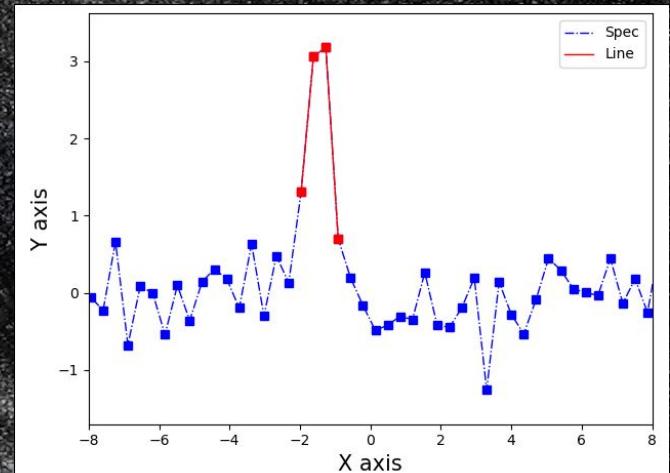
- Not all telescopes took a reading
- Left: Scan 73, Right: Scan 203



Discrepancy

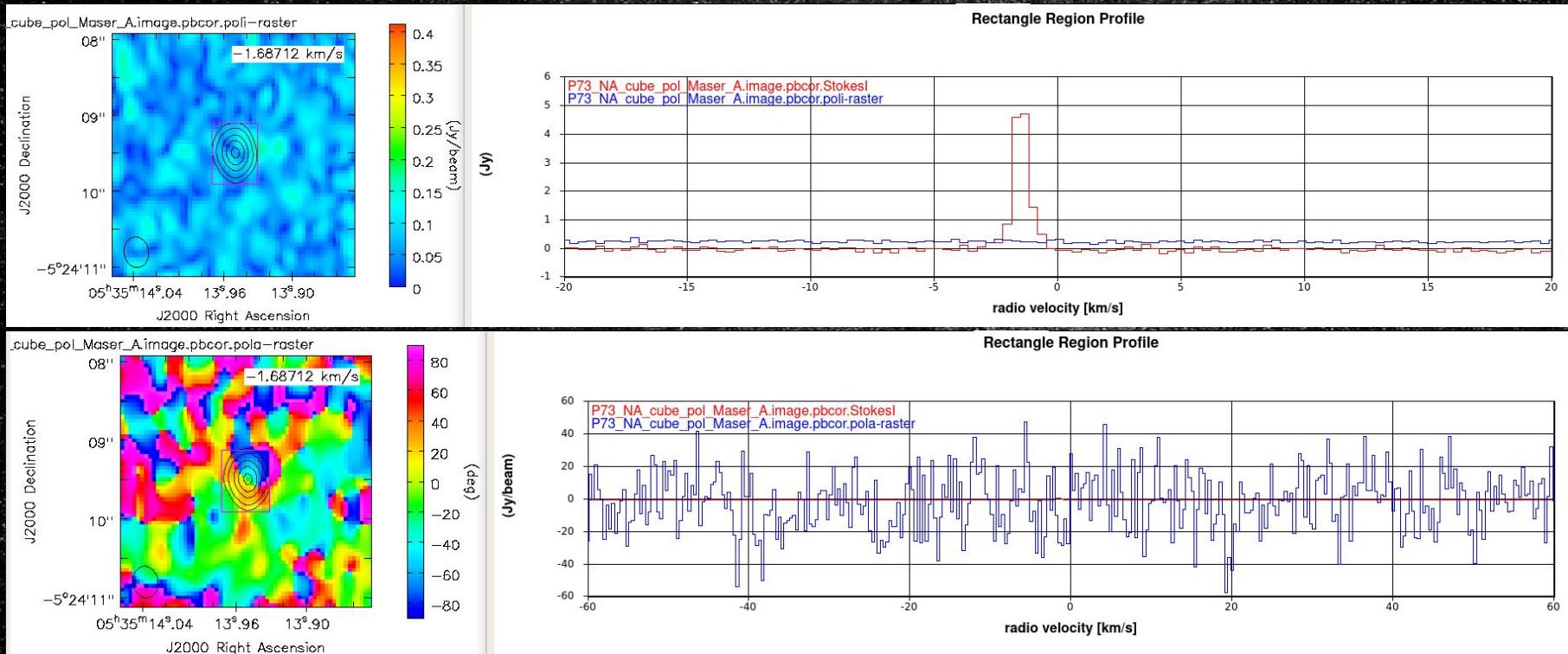
Epoch 3, Scan 73

- Scan 73
 - More noise
 - Less signal
 - Lower signal-to-noise ratio



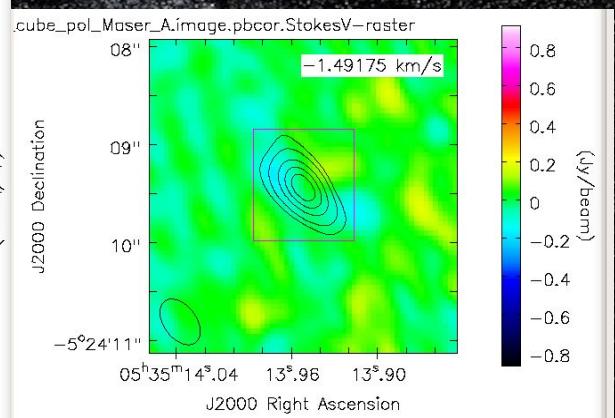
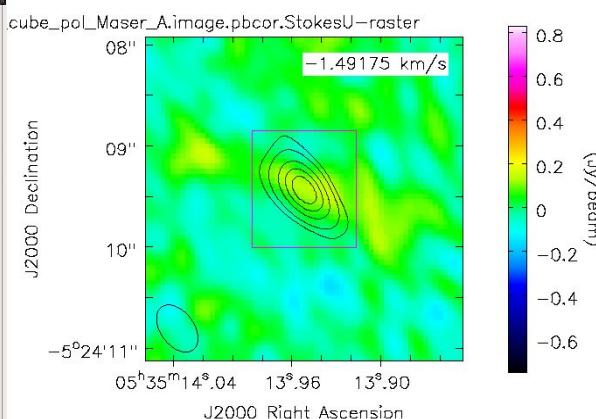
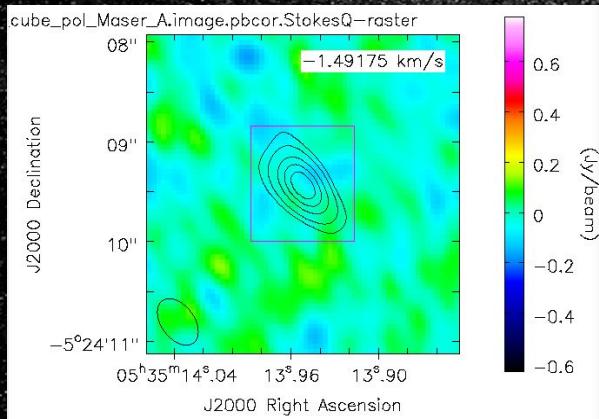
Objective 3

■ Lack of Polarization, Mix of Photonic Angles



Objective 3

- No dominant linear nor circular polarization found for any of the 20 epochs



Summary



NASA, Hubble Space Telescope

- Constructed light-curve to investigate short-term maser variability
- General consistency for pointing averages and scans for most epochs
- Some evidence of fast variability, no evidence of polarization

Acknowledgements

- Thank you to:
 - My advisor Dr. Araya for mentoring me
 - WIU Undergrad Mr. Sojka and WIU Grad Mr. VanZuiden for collaboration on the project
 - Dr. Dong, Dr. Trimm, and the rest of the SIR department

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