Regression methods in waveform modeling: a comparative study

Michael Pürrer AEI Potsdam-Golm UNC, October 15, 2020

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Regression methods in waveform modeling: a comparative study

Yoshinta Setyawati^{1,2} (ii), Michael Pürrer³ and Frank Ohme^{1,2} (iii)
Published 3 March 2020 • © 2020 The Author(s). Published by IOP Publishing Ltd
Classical and Quantum Gravity, Volume 37, Number 7

https://arxiv.org/abs/1909.10986





Structure of this talk

- Introduction to building models of GWs from compact binaries
- Regression methods
- Setup of this study
- Results
- Conclusion



Introduction to building models of GWs from compact binaries



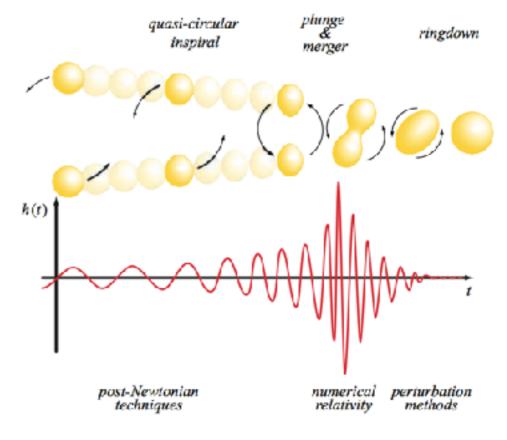
What do we need to model?

• A waveform model is a parametrized function of the waveform polarizations $h_{+,\times}(t; \overrightarrow{\lambda})$ or complex modes $h_{lm}(t; \overrightarrow{\lambda})$

$$h_{+} - ih_{\times} = \sum_{l,m} h_{lm}(t; \overrightarrow{\lambda})^{-2} Y_{lm}(\theta, \phi)$$

 Need to model the inspiral, merger and ringdown stages in binary black hole coalescence.

BBH coalescence



[Baumgarte & Shapiro, Numerical Relativity]

GW detectors record GW strain:

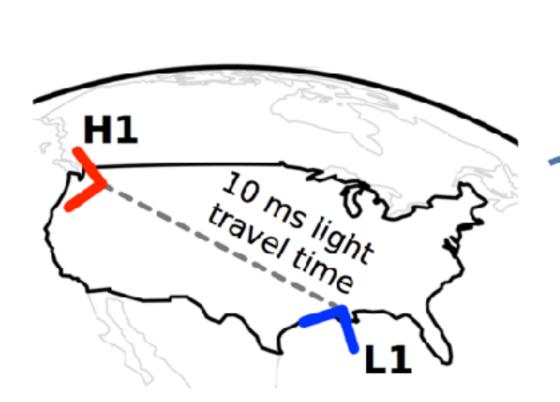
$$h(t; \overrightarrow{\theta}) = h_{+}(t; \overrightarrow{\lambda}) F_{+}(\hat{n}, \psi) + h_{\times}(t; \overrightarrow{\lambda}) F_{\times}(\hat{n}, \psi)$$

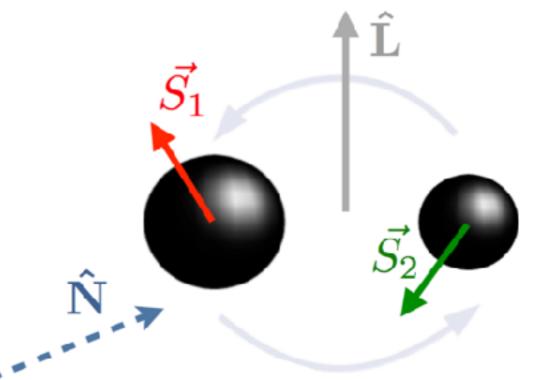


Model parameters

Intrinsic parameters:

masses, spins, eccentricity, tidal deformability





Extrinsic parameters:

time, sky position, distance, orientation, reference phase

Credit: LIGO/Virgo



Types of waveform models

Analytic:

- post-Newtonian
- (uncalibrated) **Effective-one body** (EOB) models
- Numerical relativity (NR):
 - Numerical solution of Einstein's equations for binary systems
- "Data driven":
 - Semi-analytical models combining analytical information with tuning to numerical relativity (NR) simulations

Phenom(enological) models EOBNR models

- Surrogate / reduced order models

Model directly NR or a semi-analytical model



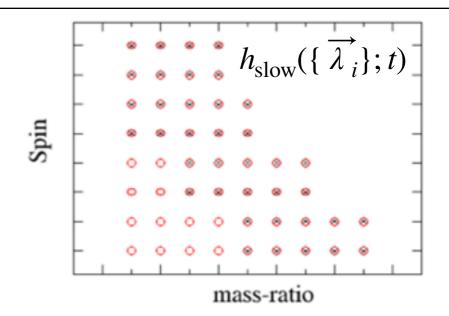
• Start with a *slow* model (e.g. EOB, NR) $h_{\text{slow}}(\overrightarrow{\lambda};t)$



- Start with a *slow* model (e.g. EOB, NR) $h_{slow}(\overrightarrow{\lambda};t)$
- Build fast and accurate surrogate model or ROM:
 - **surrogate**: substitute, proxy, replacement
 - reduced order model: truncated orthonormal basis expansion

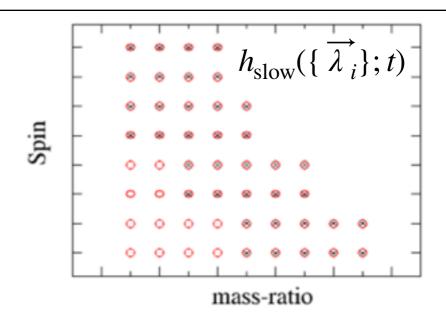


- Start with a *slow* model (e.g. EOB, NR) $h_{\text{slow}}(\overrightarrow{\lambda};t)$
- Build fast and accurate surrogate model or ROM:
 - Build **training set**:from **slow model** $h_{\text{slow}}(\{\overrightarrow{\lambda}_i\};t)$



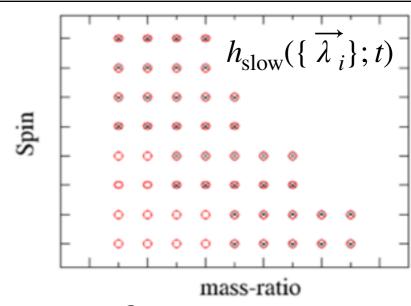


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- Build fast and accurate surrogate model or ROM:
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 - Decompose waveform into data pieces and orthonormal bases

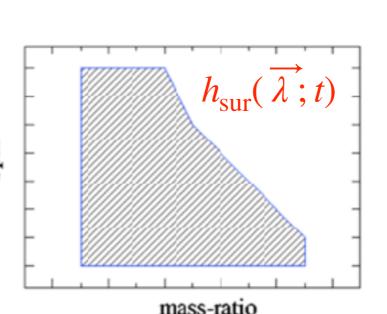




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 - Decompose waveform into data pieces and orthonormal bases
 - Interpolate coefficients over parameter space

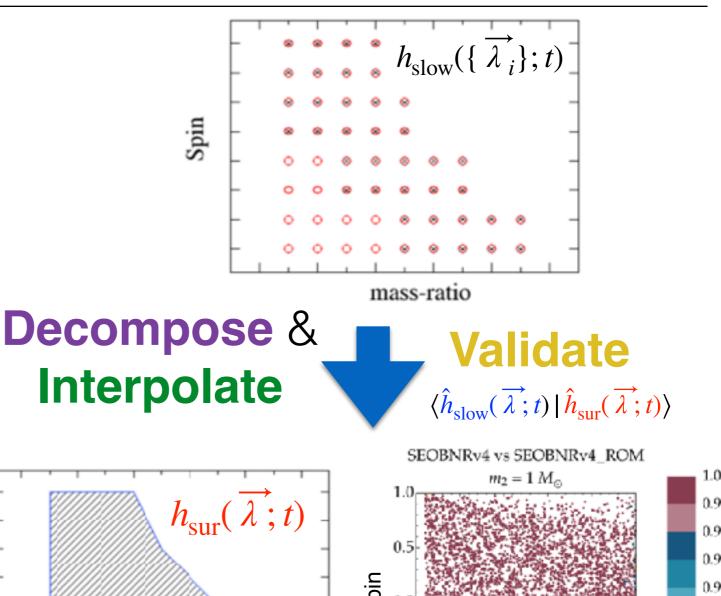


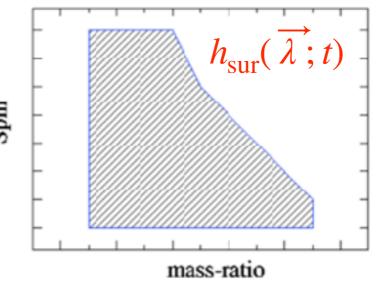
Decompose & Interpolate

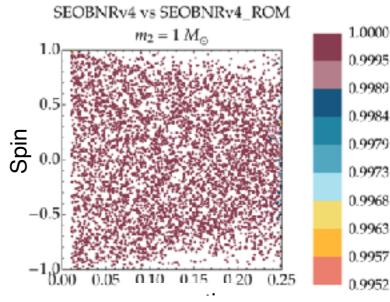




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 - Validate the surrogate

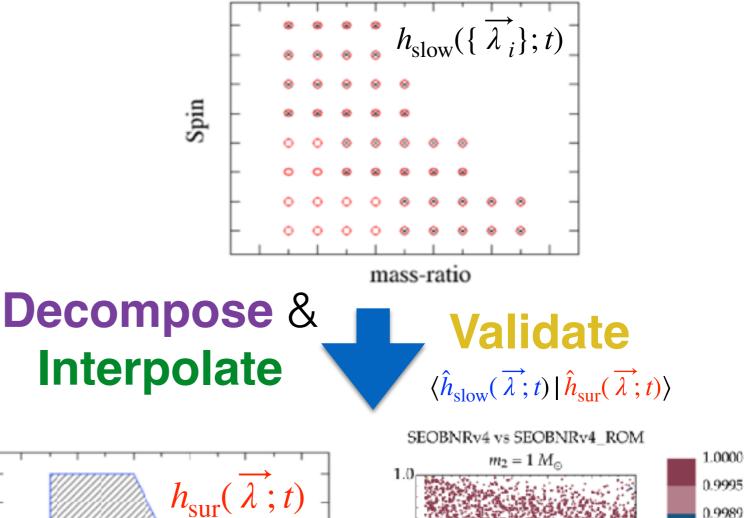


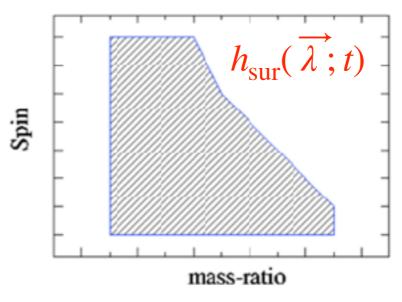


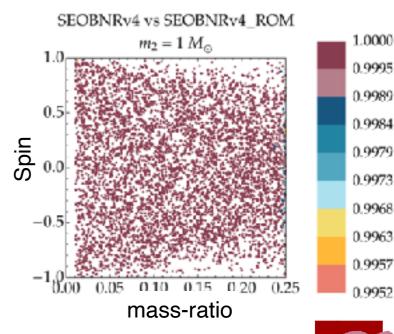




- Start with a slow model (e.g. EOB, NR) $h_{\text{slow}}(\overline{\lambda};t)$
- Build fast and accurate surrogate model or ROM:
 - Build training set:from **slow** model $h_{\text{slow}}(\{\overrightarrow{\lambda}_i\};t)$
 - **Decompose** waveform into data pieces and orthonormal bases
 - Interpolate coefficients over parameter space
 - Validate the surrogate
 - **Speedup** ~ O(1000)









0.9973

0.9952

Reduced order / surrogate models

- Usually decompose waveform data pieces in a reduced orthonormal basis
 - SVD or greedy basis construction
 - Work with coefficients
 - Empirical interpolation method: can transform basis such that coefficients are waveform data piece at a certain time or frequency

$$I_N[X](t; \vec{\lambda}) = \sum_{i=1}^N c_i(\vec{\lambda})e^i(t) = \sum_{j=1}^N X(T_j; \vec{\lambda})b^j(t)$$

Reduced order / surrogate models

- Usually decompose waveform data pieces in a reduced orthonormal basis
 - SVD or greedy basis construction
 - Work with coefficients
 - Empirical interpolation method: can transform basis such that coefficients are waveform data piece at a certain time or frequency
- Use sophisticated regression methods:
 - Greedy step-wise forward polynomial fits
 - Tensor-product spline interpolation (regular grid)
 - Gaussian process regression

```
[Field+13, MP 14, 15, Bohé+...MP 17, Lackey+16, Doctor+17 (MP) Lackey+...MP 19, Blackman+15,+17,+17, Varma+19,+19, Cotesta+20 (MP)]
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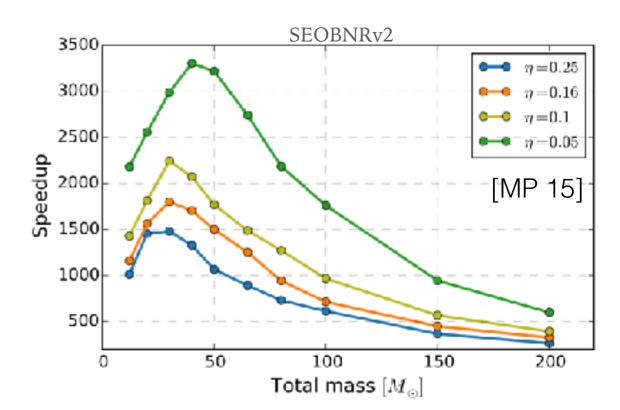
Reduced order / surrogate models

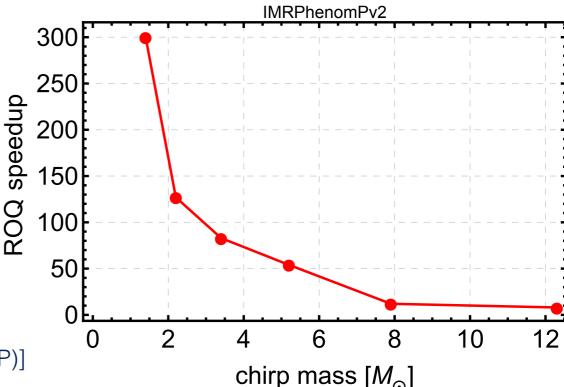
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[Field+13, MP 14, 15, Bohé+...MP 17, Lackey+16, Doctor+17 (MP) Lackey+...MP 19, Blackman+15,+17,+17, Varma+19,+19, Cotesta+20 (MP)]





Regression methods

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Overview of regression methods

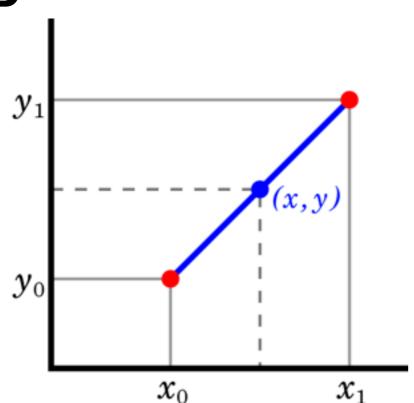
- Polynomial interpolation & polynomial fits
 - Linear interpolation
 - Tensor product interpolation (TPI)
 - Greedy multivariate polynomial fit (GMVP)
- Radial basis functions (RBF)
- Gaussian process regression (GPR)
- Artificial neural networks (ANNs)



Linear interpolation

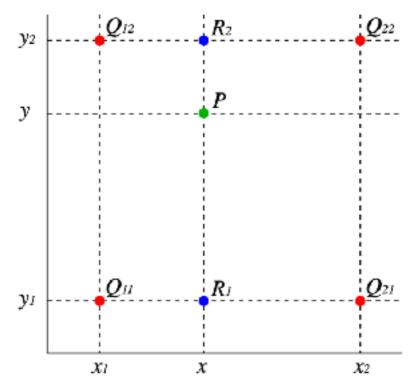
 Multivariate linear interpolation on a regular grid using the regular grid interpolator (RGI) in scipy





$$y = y_0 + (x - x_0) \frac{y_1 - y_0}{x_1 - x_0}$$

2D



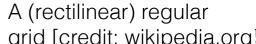
- 1) Interpolate in x-direction: -> R₁, R₂
- 2) Interpolate in y-direction: -> P



Tensor product interpolation (TPI)

 On a regular grid can use the same univariate interpolation method (e.g. splines, spectral interpolation) in each dimension.

$$I[X](t_{\mathrm{i}};ec{\lambda}) = \sum_{j_{1},...,j_{d}} a_{j_{1},...,j_{d}} \left(\Psi_{j_{1}}\otimes\ldots\otimes\Psi_{j_{d}}
ight)(ec{\lambda})_{\mathrm{f}}$$



- Splines are piecewise polynomials with continuity conditions. grid [credit: wikipedia.org]
 - Can be written in terms of B-spline basis functions of degree k with a "knot vector" t (connection points of polynomials)
 - Solve a linear system for the spline coefficients $s = \sum s_i B_{i,k,t}(x)$
- If data is smooth Chebyshev interpolation is a good option.
- We use TPI cubic splines provided by the Cython package at https://github.com/mpuerrer/TPI



Polynomial fits

- A multiple linear regression model where the independent variables form a polynomial
- Example in 1D:
 - Assume we are given data points $(x_i, y_i)_{i=1}^N$
 - Polynomial ansatz: $f(ec{x}) = c_0 x^k + c_1 x^{k-1} + \cdots + c_{k-1} x + c_k$
 - If ansatz has as many d.o.f. as data points, can solve linear system

$$\begin{pmatrix} x_1^k & x_1^{k-1} & \cdots & x_1 & 1 \\ x_2^k & x_2^{k-1} & \cdots & x_2 & 1 \\ \vdots & & \ddots & & \vdots \\ x_N^k & x_N^{k-1} & \cdots & x_N & 1 \end{pmatrix} \begin{pmatrix} c_0 \\ c_1 \\ \vdots \\ c_k \end{pmatrix} = \begin{pmatrix} y_1 \\ y_2 \\ \vdots \\ y_N \end{pmatrix}$$

- In general, the linear system may be over- or under-determined such that no unique solution would exist
- Use discrete **least squares fit** to minimize the error $\sum_{j=1}^N |f(x_j) y_j|^2$
- We use the fitting function polyfitnd() from https://bitbucket.org/chadgalley/rompy



Greedy multivariate polynomial fit (GMVP)

- Procedure (London & Fauchon-Jones, CQG 36, 2019):
 - Assume we have data at $\vec{x}_j = \{x_j^1, x_j^2, \dots, x_j^d\}$
 - Ansatz $f(\vec{x}) \approx \sum \mu_k \, \phi_k(\vec{x})$
 - Basis functions $\phi_k(\bar{x})$ are chosen to be
 - multivariate polynomials of maximal degree D Select terms from set $\phi_k(\vec{x}) \in \left\{ (x^1)^{\alpha_1} (x^2)^{\alpha_2} \dots (x^n)^{\alpha_d}, \sum_{i=1}^n \alpha_i \leq D \right\}$
 - Find least-squares solution
 - Iteratively add terms using a greedy algorithm to minimize the error $\epsilon^2 = \frac{\sum_j \left[f(\vec{x}_j) - \sum_k \mu_k \phi_k(\vec{x}_j)\right]^2}{\sum_i \left[f(\vec{x}_j)\right]^2}$
- A similar method has been used in the construction of NR-surrogate models Blackman+15,+17,+17, Varma+19,+19



Radial basis functions (RBF)

- RBF Method (we use scipy.interpolate.Rbf())
 - Assume we have data $\{(\vec{x}_i, y_i) \in \mathbb{R}^d \times \mathbb{R} | i = 1, \dots, N\}$
 - Make ansatz that depends only on Euclidean distance

$$s(\vec{x}) = \sum_{i=1}^{N} w_i \varphi(r)$$
 $r = \|\vec{x} - \vec{x_i}\|$

- Need to solve linear system $\Phi(r)\vec{w} = \vec{Y}$
- Choose radial basis function (or kernel)

Need to ensure that matrix $\Phi(r)$ is non-singular

Common choice: multiquadric kernel function

$$\varphi(r) = \sqrt{1 + \left(\frac{r}{\varepsilon}\right)^2}$$



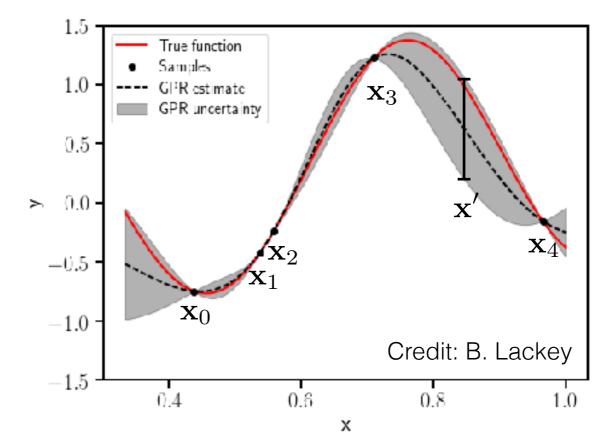
Gaussian process regression (GPR)

• GPR assumes the values of a function y_i at the points x_i are random variables that follow a multivariate Gaussian:

$$p(y_i|\mathbf{x}_i) = \mathcal{N}(0, k(\mathbf{x}_i, \mathbf{x}_j))$$

- Covariance function between two points $k(\mathbf{x}_i, \mathbf{x}_j)$
- Can estimate the functional value y' at x' given data (x_i, y_i)
- Mean: $E[p(y'|\mathbf{x}_i,\mathbf{x}',y_i)]$
- Uncertainty: $Var[p(y'|\mathbf{x}_i, \mathbf{x}', y_i)]$
- Example kernels:

$$k_{SE}(r) = \exp\left(rac{-r^2}{\ell^2}
ight)
onumber$$
 $k_M(r) = rac{2^{1-
u}}{\Gamma(
u)} \left(rac{\sqrt{2
u}r}{\ell}
ight)^
u K_
u \left(rac{\sqrt{2
u}r}{\ell}
ight)$





Gaussian process regression (GPR)

• Define a Gaussian process for zero mean $y(\vec{x}) \sim GP\Big(\mu(\vec{x}) = 0, k(\vec{x}, \vec{x}')\Big)$

$$y(\vec{x}) \sim GP\Big(\mu(\vec{x}) = 0, k(\vec{x}, \vec{x}')\Big)$$

Can write vector of training and test outputs as

$$\begin{bmatrix} \vec{y} \\ y_* \end{bmatrix} \sim \mathcal{N} \left(0, \begin{bmatrix} K(X, X) + \sigma_n^2 I & K(X, X_*) \\ K(X_*, X) & K(X_*, X_*) \end{bmatrix} \right)$$

• Find conditional probability $p(y_*|\vec{x}_i, \vec{x}_*, \vec{y}, \vec{\theta}) = \mathcal{N}(\bar{y}_*, \text{var}(y_*))$

$$\bar{y}_* = K(X_*, X)(K(X, X))_{ij}^{-1} y_j$$

$$\text{var}(y_*) = K(X_*, X_*) - K(X_*, X_i)(K(X, X))_{ij}^{-1} K(X_*, X_j)$$

$$K(x_i, x_j) = \sigma_f^2 k(x_i, x_j) + \sigma_n^2 \delta_{ij}$$

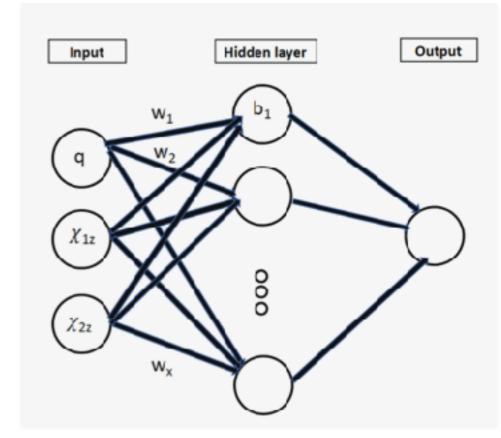
• Optimize over hyper-parameters $\sigma_f, \sigma_n, \{\ell_i\}_{i=1}^n$ by maximizing the marginal log-likelihood

$$\ln p(y_i|\vec{x}_i,\vec{\theta}) = -\frac{1}{2} \Big(y_i (K(X,X))_{ij}^{-1} y_j + \ln |K(X,X)| + N \ln 2\pi \Big)$$

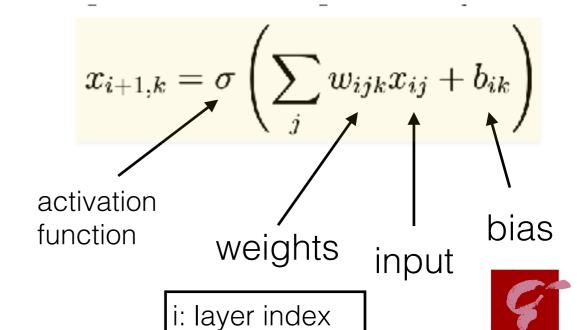


Artificial neural networks (ANNs)

- ANNs are a machine learning algorithm used e.g. for pattern recognition, classification.
- We use a multi-layer-perceptron with 2 hidden layers between input and output
 - Each layer acts as an *affine* transformation on its inputs
- Activation functions are used to introduced nonlinearity
 - We use the Rectified Linear Unit (ReLU) $\sigma(z) = \max(z,0)$



A single-layer ANN for 3d interpolation



j: neuron index

Overview of regression methods

Methods	Advantages	Disadvantages	Training time
Linear (RGI)	standard scipy	needs regular grid	$\mathcal{O}(N)$
TPI	robust and high accuracy	needs regular grid	$\mathcal{O}(N^k)$
GMVP	irregular grid fast execution time	complex	#basis function #error tolerance
Polynomial fit	irregular grid simple and fast	Runge's phenomenon only univariate in scipy	$\mathcal{O}(N)$ and #polynomial degree
RBF	scipy irregular grid	high computational complexity	$\mathcal{O}(N^3)$
GPR	irregular grid can predict uncertainty	depends on the choice of kernel and hyperparameters complex	$\mathcal{O}(N^3)$
ANN	irregular grid flexible architecture choices	complex	#neurons #hidden layers



Setup of this study



Key waveform quantities

- We transform the *inertial* frame waveform modes to the minimally rotating co-precessing frame and align the waveforms at a particular time (t = -2000 M)
- We further define a frame that follows the orbital motion and compute the modes in this co-orbital frame.
- In this study we don't build a full waveform model, instead we just use two key quantities at selected times

$$\phi(t) := \frac{1}{4} \left(\arg \left[\bar{\bar{h}}_{copr}^{2,-2}(t) \right] - \arg \left[\bar{\bar{h}}_{copr}^{2,2}(t) \right] \right)$$

$$A(t) := \text{Re } \bar{h}_{+}^{2,2} = \frac{1}{2} \text{Re } \left(\bar{h}_{\text{coorb}}^{2,2}(t) + \bar{h}_{\text{coorb}}^{2,-2^*}(t) \right)$$



Key waveform quantities: example

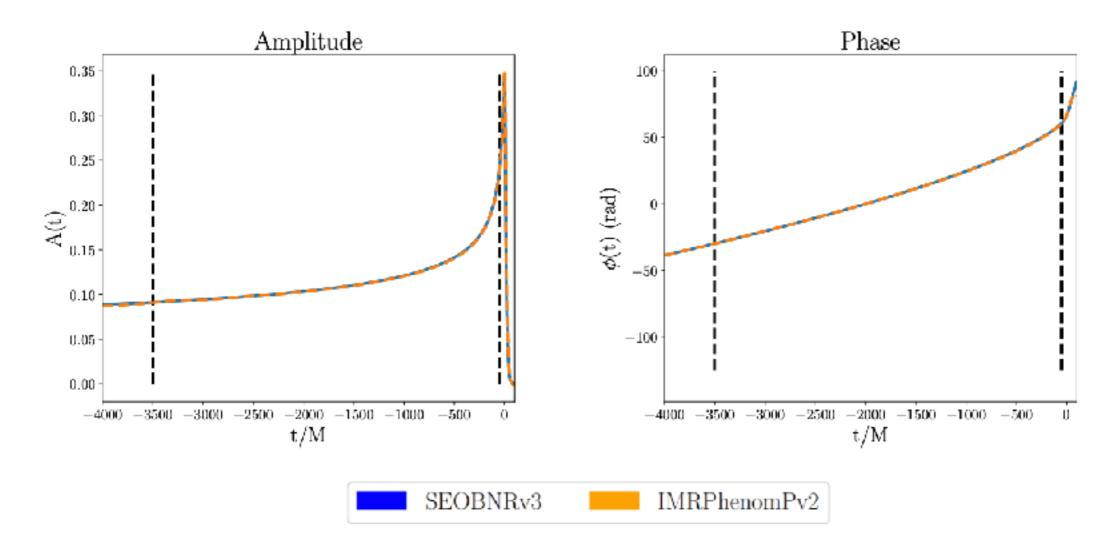


Figure 1. The key quantities of the GW signal of a precessing BBH, here illustrated for a binary with $(q, \chi_{1x}, \chi_{1y}, \chi_{1z}, \chi_{2x}, \chi_{2y}, \chi_{2z}) = (1.99, 0.51, 0.04, 0.03, 0.01, 0.6, 0.1)$. Left: the dimensionless amplitude A(t). Right: the phase $\phi(t)$ (in unit radian). The black dashed lines show the points in time-space, where we perform different interpolation methods (t=-3500M and t=-50M).



Waveforms and data sets

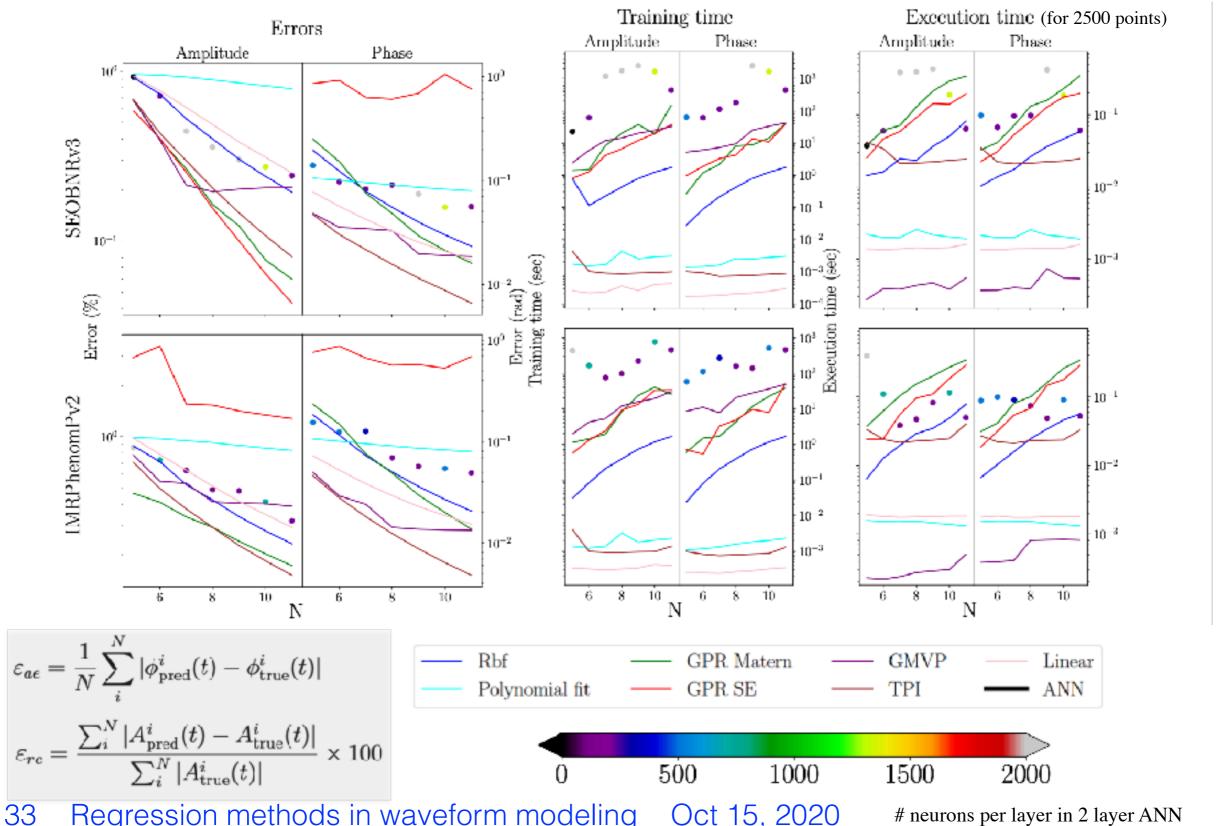
- Our choice of using waveform quantities at fixed times is motivated by the EIM modeling framework.
- We use two inspiral-merger-ringdown waveform models that describe GWs emitted from precessing black hole binaries:
 - **SEOBNRv3** (time domain waveform)
 - IMRPhenomPv2 (iFFT of Fourier domain waveform)
- We consider these data sets:
 - Dimensionality: 3D (q, aligned spin) or 7D (q, generic spins)
 - Early and late times: -3500M or -50M before merger
 - Number of training set points: 3D (5 11 per dimension), 7D (total up to 3000); 2500 random test points
 - Physical domain: 3D ($q \le 10, |\chi_i^z| \le 1$), 7D ($q \le 2, |\chi_i| \le 1/\sqrt{3}$)



Results



3D interpolation results at t=-3500M



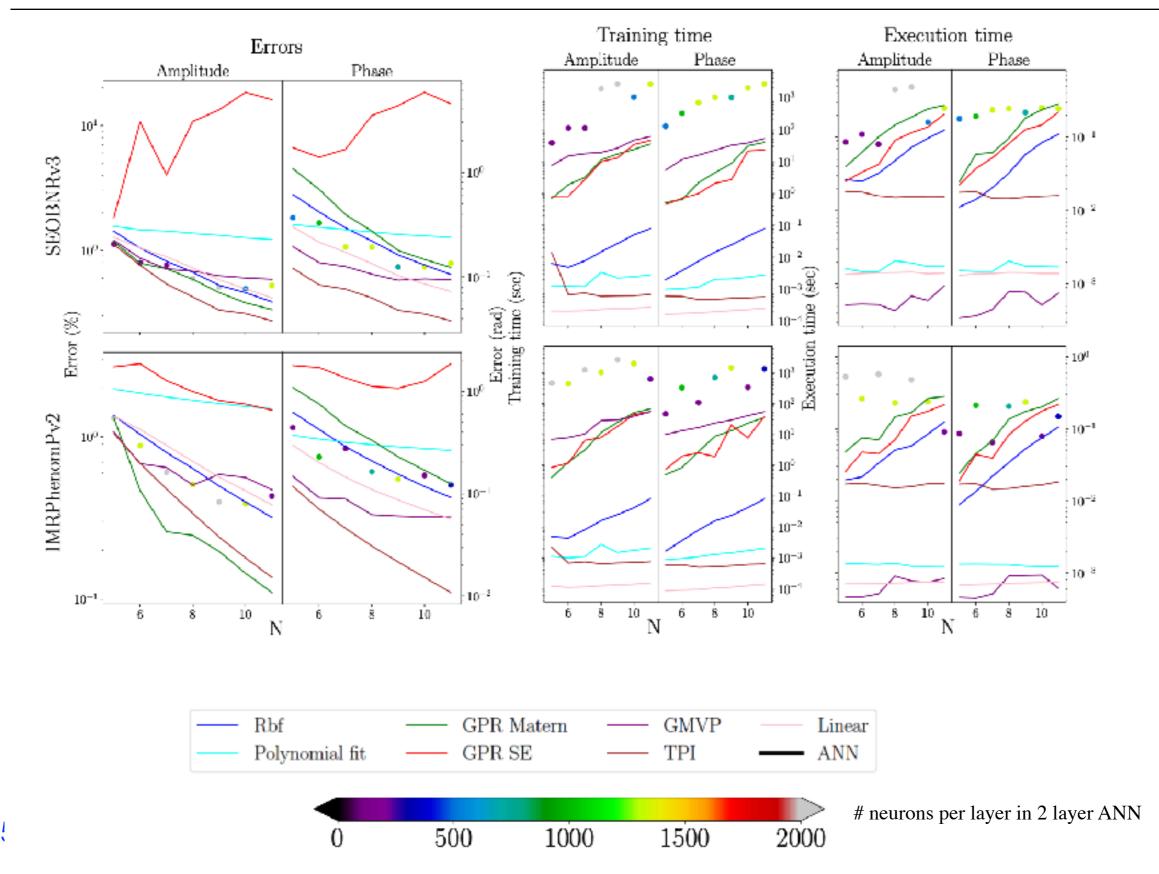


Observations for 3D results

- In general, errors decrease with the size of the training set.
- Errors are similar between the two waveform models, but:
 - GPR *amplitude error* for IMRPhenomPv2 is much higher for SE kernel compared to Matérn
 - Noise in the IMRPhenomPv2 data due to the iFFT
- Comments on the 2-layer ANNs:
 - We report *minimum error* over # neurons in [2, 2000]
 - Too many neurons lead to overfitting (dropout could help) and too few neurons lead to underfitting

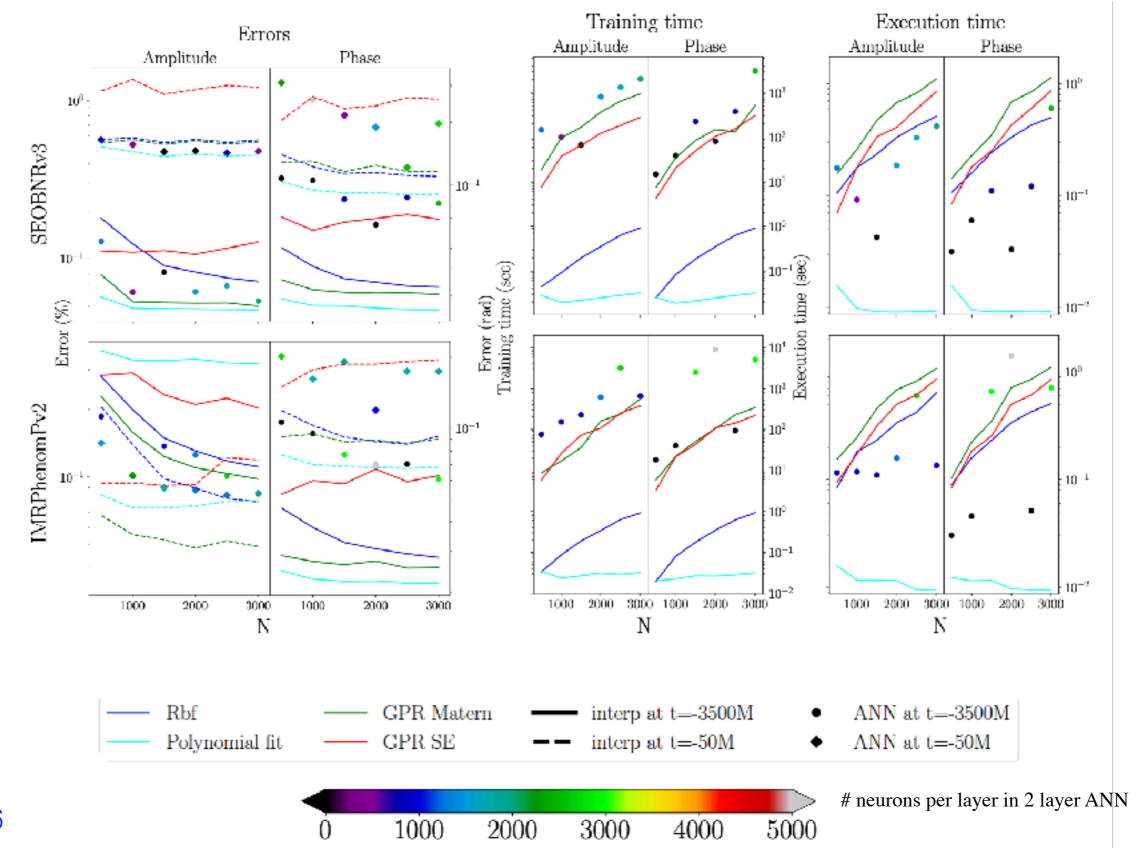


3D interpolation results at t=-50M





7D interpolation results





Observations for 7D results

Amplitude errors near merger / during inspiral:

- For SEOBNRv3, errors near merger are higher than in the inspiral.
- For IMRPhenomPv2, surprisingly, amplitude errors are *smaller* near merger.

• Phase errors:

- They are comparable for SEOBNRv3 and IMRPhenomPv2.

Execution time:

- for GPR and RBF it depends on the size of the TS,
- in contrast to ANNs.

Comparing errors between 3D and 7D:

- For the same parameter ranges, errors in 7D can be up to 100 times larger for A(t) and 15 times larger for φ(t).



Conclusions

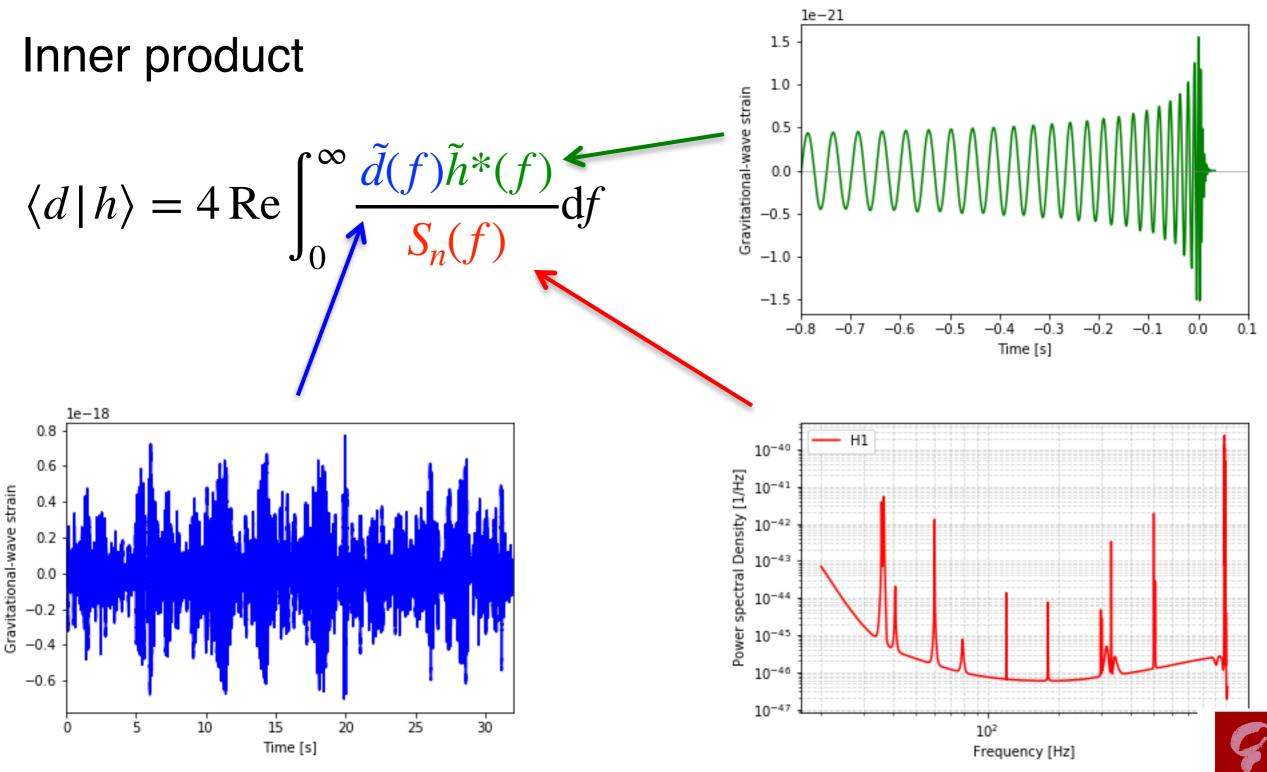
- A few take-aways from the regression results
- 3D:
 - TPI is very accurate and fast in training and evaluation
 - **GMVP** is also doing very well here
- 7D:
 - Polynomial fits are fast and accurate
 - GPRs with a Matérn kernel have good accuracy (with SE kernel errors are much higher)
 - Note: GMVP not included because implementation was too slow.
- Methods that have been used build GW surrogate models:
 - **TPI** (up to 5D)
 - Greedy polynomial fits (slightly different from GMVP) in 7D
 - **GPR** (up to 5D)



Thank you for your attention!

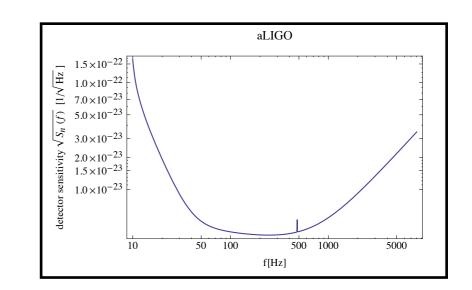


Matched filtering



Checking accuracy of waveforms

$$\langle h_1, h_2 \rangle = 4 \operatorname{Re} \int_{f_{\min}}^{f_{\max}} \frac{\tilde{h}_1(f)\tilde{h}_2(f)^*}{S_n(f)} df$$



SNR

$$\rho = ||h||$$

Match (overlap)

$$\mathcal{O}(h_1, h_2) = \frac{4}{\|h_1\| \|h_2\|} \max_{t_0} \left| \mathcal{F}^{-1} \left[\frac{\tilde{h}_1(f)\tilde{h}_2(f)^*}{S_n(f)} \right] (t_0) \right|$$

Mismatch (Faithfulness)

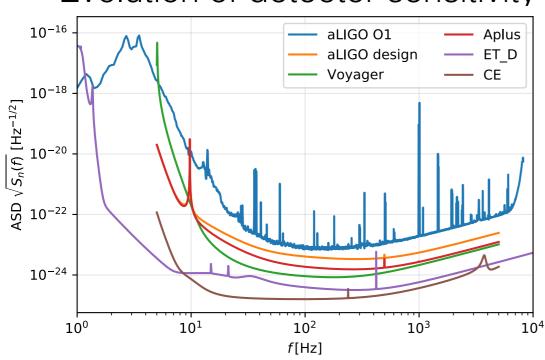
$$1-\mathcal{O}(h_1,h_2)$$

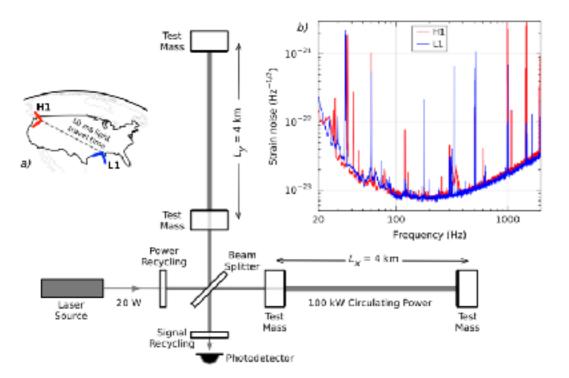
$$\mathcal{F}^{-1}[g(f)] = \int_{-\infty}^{\infty} g(f)e^{-2\pi i f t} df$$



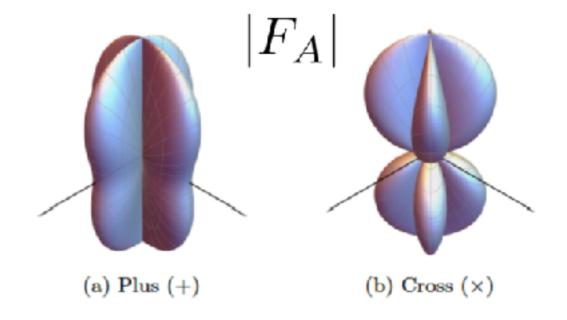
Interferometric GW detectors

Evolution of detector sensitivity





Angular response of IFO



$$h = h_A F^A \qquad F^A(\hat{n}, \psi) = D^{ab} e^A_{ab}$$
$$A = +, \times, \dots$$
$$D^{ab} = \frac{1}{2} \left(d_x^a d_x^b - d_y^a d_y^b \right)$$



Reduced Basis and Empirical interpolation method

$$\tilde{h}(f;\vec{\lambda}) pprox \sum_{i=1}^m c_i(\vec{\lambda}) e_i(f)$$
 greedy reduced basis $c_i(\vec{\lambda}) = \langle \tilde{h}(\cdot;\vec{\lambda}), e_i(\cdot) \rangle$

Empirical interpolation method finds "good" frequencies:

$$\sum_{i=1}^{m} c_i(\vec{\lambda}) e_i(F_j) \stackrel{!}{=} \tilde{h}(F_j; \vec{\lambda})$$
EI frequencies

• El interpolant:

$$\tilde{h}(f; \vec{\lambda}) \approx \sum_{i=1}^{m} B_j(f) \tilde{h}(F_j; \vec{\lambda}) \qquad B_j(f) = \sum_{i=1}^{m} e_i(f) (V^{-1})_{ij}$$
$$V_{ij} = e_i(F_j)$$

• Fit $\tilde{h}(F_i; \vec{\lambda})$ w.r.t. $\vec{\lambda}$ at each F_i using only data at greedy points

