## 1 Evolution

We start from the acceleration:

$$\ddot{r} = -\frac{GM}{r^2} + \frac{\Lambda}{3}r\tag{1}$$

First we rewrite  $\ddot{r}$  in terms of  $x = \ln(a)$ :

$$\frac{d^2}{dt^2} = \frac{d}{dt} \left[ \frac{da}{dt} \frac{d}{da} \right] 
= \frac{d^2a}{dt^2} \frac{d}{da} + \left[ \frac{da}{dt} \right]^2 \frac{d^2}{da^2} 
= \ddot{a} \frac{d}{dx} \frac{dx}{da} + \dot{a}^2 \frac{d}{da} \frac{d}{dx} \frac{dx}{da} 
= \frac{\ddot{a}}{a} \frac{d}{dx} + \dot{a}^2 \left( \frac{d^2x}{da^2} \frac{d}{dx} + \frac{dx}{da} \frac{d^2}{dx^2} \frac{dx}{da} \right) 
= \left( \frac{\ddot{a}}{a} - \frac{\dot{a}^2}{a^2} \right) \frac{d}{dx} + \frac{\dot{a}^2}{a^2} \frac{d^2}{dx^2}.$$
(2)

Now applying this to r and using that  $\frac{\dot{a}^2}{a^2}=H^2$  and the notation  $'=\frac{d}{dx}$  we have

$$\left(\frac{\ddot{a}}{a} - H^2\right)r' + H^2r'' = -\frac{GM}{r^2} + \frac{\Lambda}{3}r\tag{3}$$

or

$$H^2r^{\prime\prime} = -\frac{GM}{r^2} + \frac{\Lambda}{3}r - \left(\frac{\ddot{a}}{a} - H^2\right)r^\prime. \tag{4}$$

Now we find the mass in terms of the initial overdensity  $\delta_i$ , it is found through

$$M = \frac{4\pi}{3}\rho_m r^3$$

where  $\rho_m = \rho_{c,0}\Omega_{m,0}(1+\delta_i)(\frac{r_i}{ra_i})^3$  and  $\rho_{c,0} = \frac{3H_0^2}{8\pi G}$ , which gives

$$M = \frac{H_0^2}{2G} \left(\frac{r_i}{a_i}\right)^2 \Omega_{m,0} (1 + \delta_i). \tag{5}$$

Inserting eq. (5), the unit less  $\tilde{r} = \frac{ra_i}{r_i}$ , which implies that  $\tilde{r}' = \frac{r'a_i}{r_i}$  and  $\tilde{r}'' = \frac{r''a_i}{r_i}$ , and  $\Omega_{\Lambda} = \frac{\Lambda}{3H_0^2}$  we get the equation

$$H^2\tilde{r}'' = H_0^2 \left( -\frac{\Omega_{m,0}(1+\delta_i)}{\tilde{2r}^2} + \Omega_{\Lambda}\tilde{r} \right) - \left( \frac{\ddot{a}}{a} - H^2 \right) \tilde{r}'.$$
 (6)

We now find

$$\ddot{a} = \frac{d\dot{a}}{dt} = \frac{d}{dt} \left( aH_0^2 \sqrt{\Omega_{m,0} a^{-3} + \Omega_{\Lambda}} \right)$$

$$= a \left( H^2 - \frac{3}{2} H_0^2 \Omega_{m,0} a^{-3} \right), \tag{7}$$

which means that

$$\frac{\ddot{a}}{a} - H^2 = -\frac{3}{2} \frac{H_0^2 \Omega_{m,0}}{a^3}.$$
 (8)

Finally inserting (8) into (6) and dividing by  $H^2 = H_0^2(\Omega_{m,0} + \Omega_{\Lambda})$  for a flat background we get

$$\tilde{r}'' = \left(\Omega_{\Lambda}\tilde{r} - \frac{\Omega_{m,0}(1+\delta_i)}{2\tilde{r}^2} + \frac{3\Omega_{m,0}\tilde{r}'}{2a^3}\right) \left[\Omega_{m,0}a^{-3} + \Omega_{\Lambda}\right]$$
(9)

## 2 Virialization

We start with the potentials

$$U_m = -\frac{3GM^2}{5r} \tag{10}$$

and

$$U_{\Lambda} = -\frac{\Lambda M}{6}r^2 \tag{11}$$

for the gravitational and DE potential respectively. Now we use the relation for a virialized system with kinetic energy T and potential energy U

$$T = \frac{1}{2}r\frac{\partial U}{\partial r}$$

$$= \frac{1}{2}r\left(\frac{3GM^2}{5r^2} - \frac{\Lambda M}{3}r\right)$$

$$= \frac{3GM^2}{10r} - \frac{\Lambda M}{6}r^2.$$
(12)

Now using that  $T = \frac{1}{2}M\dot{r}^2 = \frac{1}{2}H^2Mr'^2$  we get

$$H^2 r'^2 = \frac{3GM}{5r} - \frac{\Lambda}{3}r^2. \tag{13}$$

Now we can insert  $\Omega_{\Lambda} = \frac{\Lambda}{3H_0^2}$  and eq. (5) for the mass. We also use the variable  $\tilde{r}$  to get

$$H^{2}\tilde{r}^{2} = H_{0}^{2} \left( \frac{3}{10} \frac{\Omega_{m,0}(1+\delta_{i})}{\tilde{r}} - \Omega_{\Lambda}\tilde{r}^{2} \right). \tag{14}$$

Which we can write as

$$\tilde{r}^{2} = \left(\frac{3\Omega_{m,0}(1+\delta_i)}{10\tilde{r}} - \Omega_{\Lambda}\tilde{r}^2\right) \left(\Omega_{m,0}a^{-3} + \Omega_{\Lambda}\right). \tag{15}$$