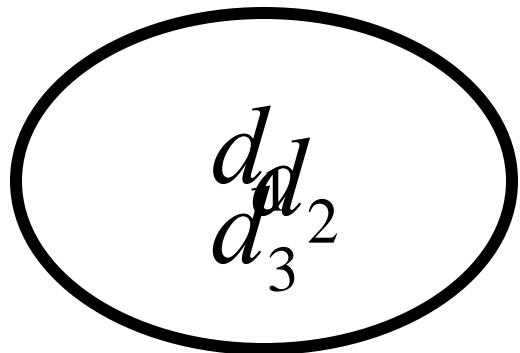


Radon Transforms and SSA

Mauricio D Sacchi

SAIG University of Alberta

Noise attenuation with “TRANSFORM-BASED” methods



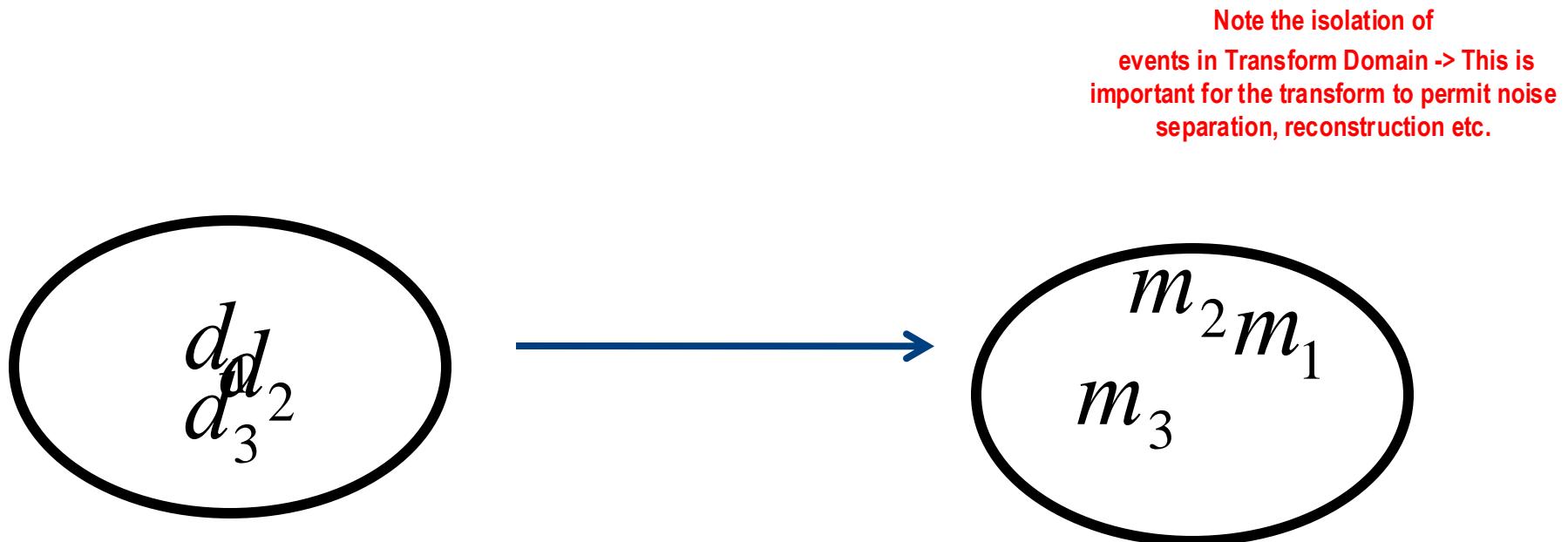
[Q] I want to remove d_3

$$d = d_1 + d_2 + d_3$$

How?

One possibility is to use PhotosShop. However, it is too time consuming for seismic data processing...

Noise attenuation with “TRANSFORM-BASED” methods

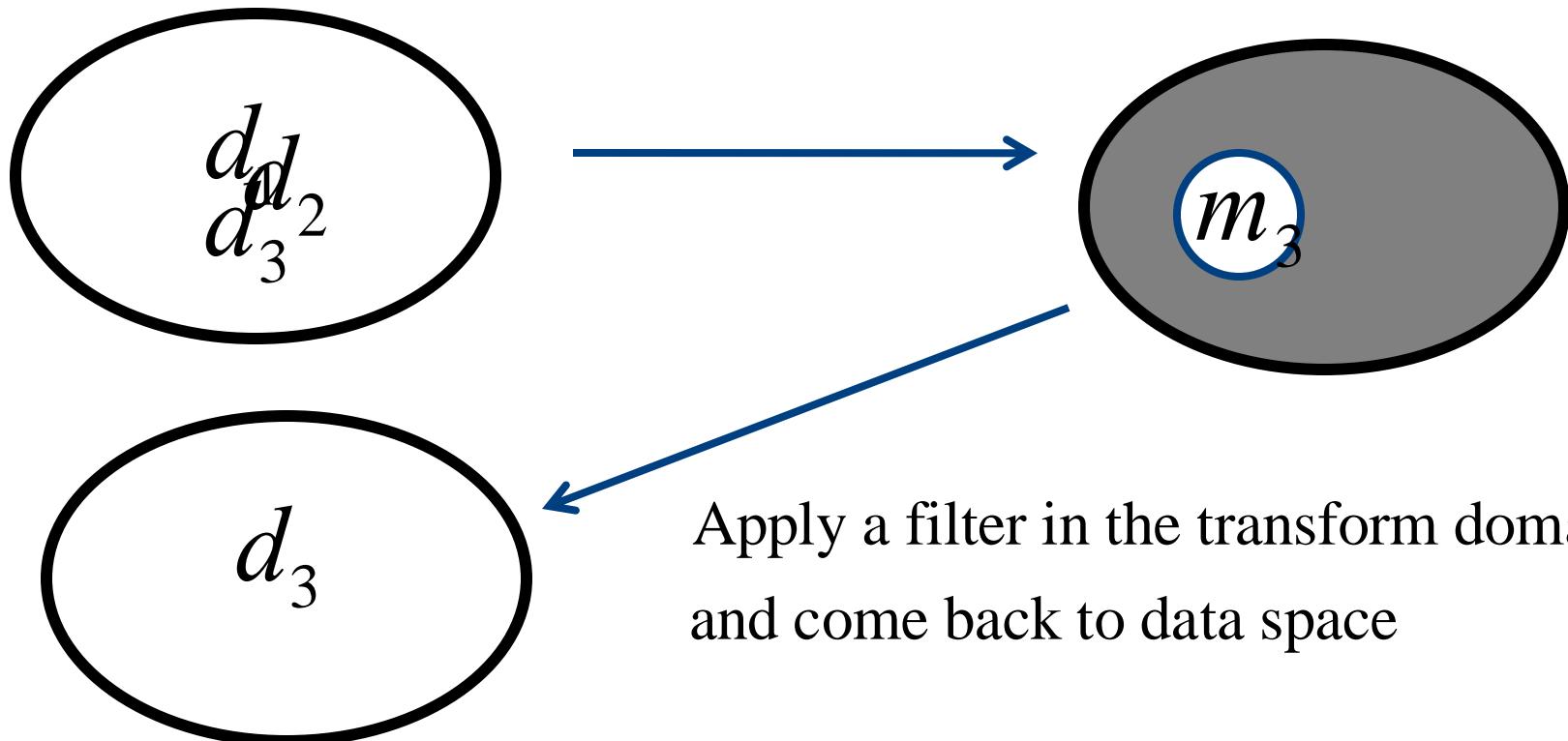


[A] Use a transform

$$d = d_1 + d_2 + d_3 \rightarrow m = m_1 + m_2 + m_3$$

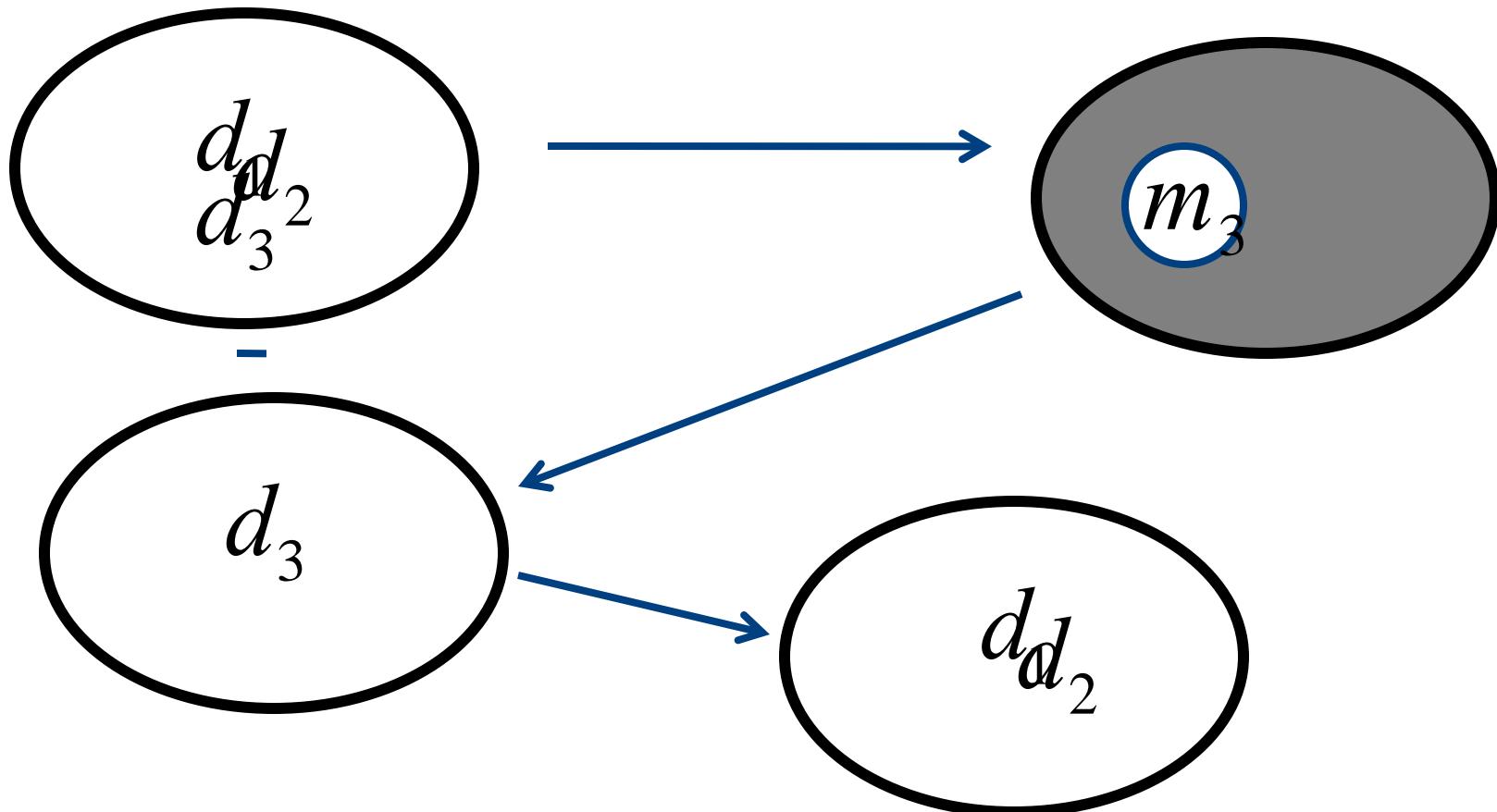
Noise attenuation with “TRANSFORM-BASED” methods

Note the isolation of events in Transform Domain -> This is important for the transform to permit noise separation, reconstruction etc.



Noise attenuation with “TRANSFORM-BASED” methods

Note the isolation of events in Transform Domain -> This is important for the transform to permit noise separation, reconstruction etc.



Radon transforms

- Search for coherent waveforms along particular travel-time curves
- Three common implementations: linear, parabolic, hyperbolic
- Resolution is severely affected by amplitude variations and aperture
 - Resolution = Ability to distinguish events with similar moveout
- HR implementations can cope with aperture problems but not with AVO.
- AVO could be modeled but might be expensive and hard to parameterize for processing large batches of data.
- Radon transforms are often used as non-local operators (Operator spans the full aperture of the gather) This is a problem as it entails the assumption of high degree of matching of the signal to the transform.
 - Solution: Going local - but again hard to parameterize

- Companies often use the name
- HIGH RESOLUTION RADON → SPARSE RADON TRANSFOMR
- Is a RADON TRANFORM with FOCUSING of coefficients

- ***Time Variant and Time Invariant Transforms***
 - Parabolic and Linear → Time Invariant
 - Hyperbolic → Time Variant
 - In general, Time Invariant transforms lead to fast algorithms with operators with Toeplitz form.

Some refs.

- **High Resolution Hyperbolic RT** [Time]
Thorson and Claerbout (GEOP, 1985)
- **Least-squares parabolic RT** [Freq]
Dan Hampson (CSEG, 1986)
- **Sparse Parabolic RT** (high-resolution) [Freq]
Sacchi and Ulrych – (GEOP, 1995)
- **De-aliased Parabolic RT** (high-resolution) [Freq]
Phillippe Hermann et al – (SEG, 2000)
- **Latest views of the sparse Radon transform** [Time and Freq]
Trad, Ulrych and Sacchi (Geop 2003)

Radon Integration Traveltime Paths

time invariant curves

$$t = t + qh \text{ Linear}$$
$$t = t + qh^2 \text{ Parabolic}$$
$$t = \sqrt{t^2 + qh^2} \text{ hyperbolic}$$
$$t = \sqrt{t^2 + q(h - h_0)^2} \text{ hyperbolic apex shifted}$$

This is also a post-stack
migration algorithm

Fourier shift theorem (**Important**)

Fourier delay theorem is used to design f-domain Transforms:

$$f(t - t_0) \leftrightarrow F(w) \exp(-iwt_0)$$

Time and freq. domain forms for the PRT

$$m(t, q) = \mathring{\mathbf{a}}_h d(t = t + qh^2, h)$$

$$M(w, q) = \mathring{\mathbf{a}}_h D(w, h) e^{i w q h^2}$$

These are the transforms that we do use...

$$d(t, h) = \hat{\mathbf{a}}_q m(t = t - qh^2, q)$$

$$D(w, h) = \hat{\mathbf{a}}_q M(w, q) e^{-iwh^2}$$

Forward/Adjoint or Forward/Transpose transform pair

$$d(t, h) = \mathcal{\hat{A}}_q m(t = t - qh^2, q)$$

Forward or
Synthesis

$$D(w, h) = \mathcal{\hat{A}}_q M(w, q) e^{-iwh^2}$$

$$\hat{m}(t, q) = \mathcal{\hat{A}}_h d(t = t + qh, h^2)$$

Adjoint

$$\hat{M}(w, q) = \mathcal{\hat{A}}_h D(w, h) e^{iwh^2}$$

The hat is important (discuss it)

Key points about Radon

- The adjoint is not the inverse (unless the operator is orthogonal)
- Time or Freq. Let's write the problem in Matrix or Operator form

$$\mathbf{d} = \mathbf{Lm}$$

We write our data \mathbf{d} as the sum of events in Radon space \mathbf{m}

$$\mathbf{d} = \mathbf{Lm} + \mathbf{n}$$

We also add noise and modeling errors

Find the Radon Panel via the solution of an inverse problem

$$J = \| \mathbf{d} - \mathbf{Lm} \|^2_2 + m \mathbf{R}(m)$$

Time domain RT:

Solve all at once

Freq. domain RT:

Solve one small IP for each frequency

Some nice features of freq domain RT

- Toeplitz structure

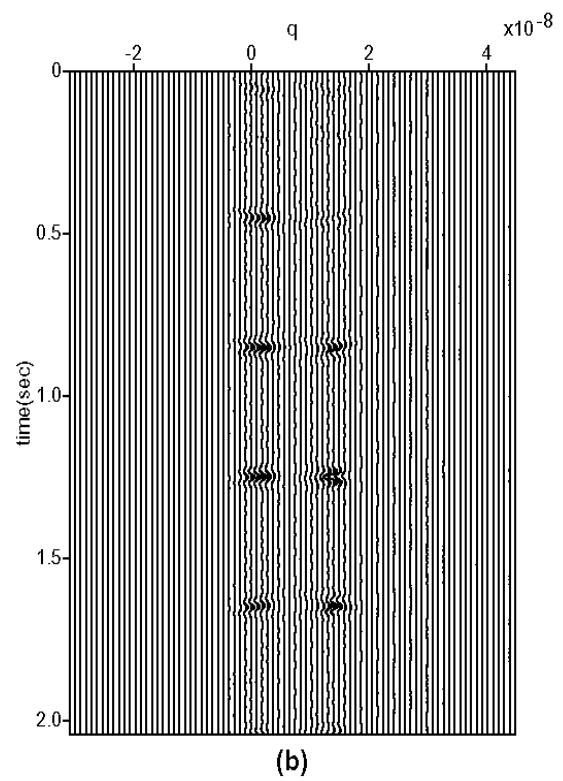
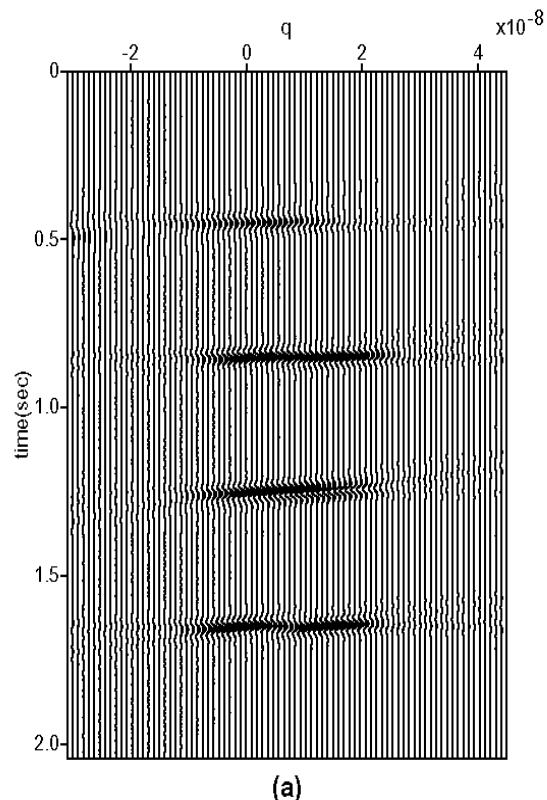
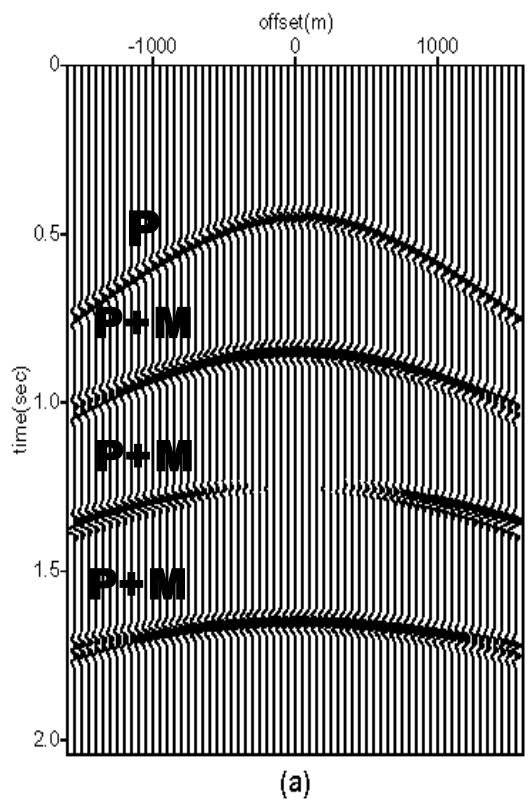
$$J(\omega) = \| \mathbf{d}(\omega) - \mathbf{L}(\omega)\mathbf{m}(\omega) \|_2^2 + \mu \| \mathbf{m}(\omega) \|_2^2$$

$$\nabla J(\omega) = 0 \rightarrow$$

$$\mathbf{m}(\omega) = (\mathbf{L}(\omega)^H \mathbf{L}(\omega) + \mu \mathbf{I})^{-1} \mathbf{L}(\omega)^H \mathbf{d}(\omega)$$

- Toeplitz if and only if q -axis is regularly sampled. This leads to fast algorithms.
- Valid for the linear and parabolic cases (time-invariant) and for the time-invariant hyperbolic transform proposed by Mosher and Foster (Geophysics 1992)

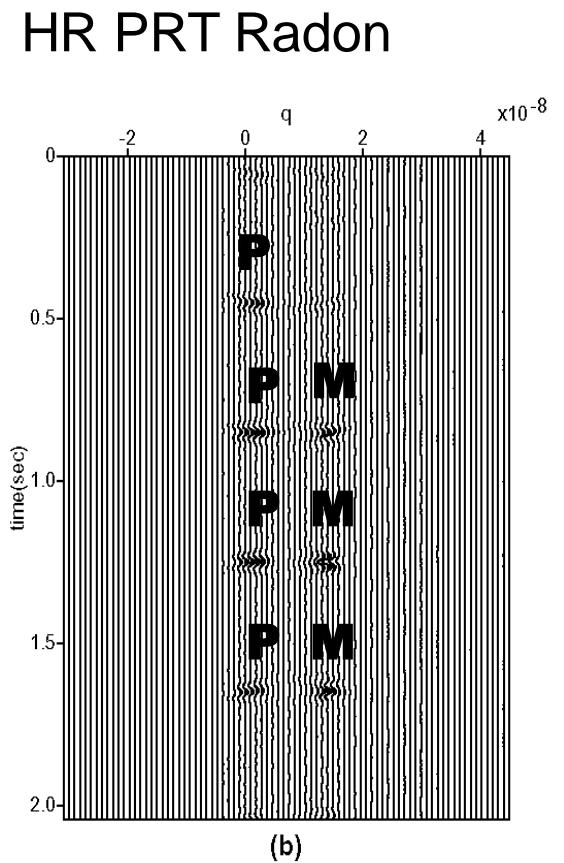
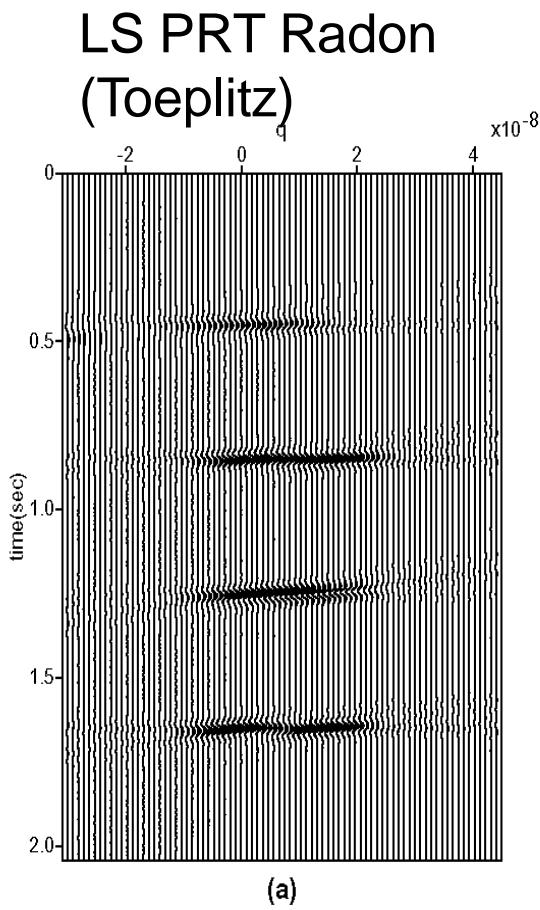
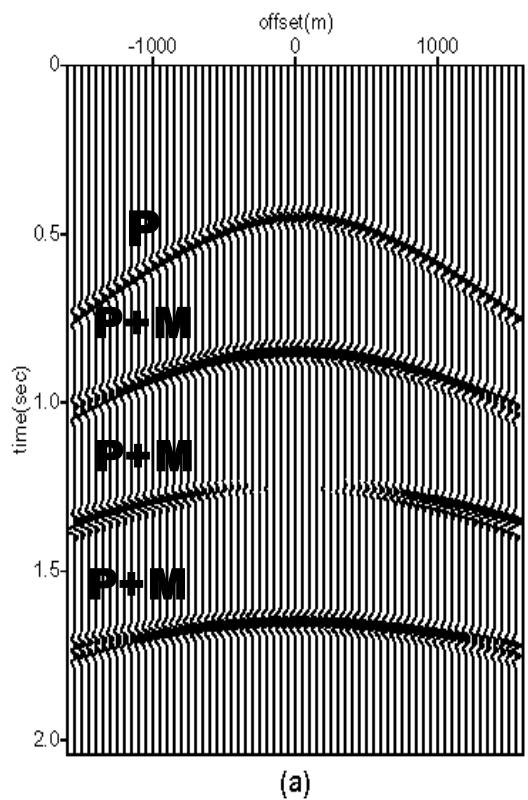
Improving resolution



smooth model

sparse model

Improving resolution



smooth model

sparse model

Flavors of Freq. domain RT (Sparse or Hi-Res Radon)

$$J(\omega) = \| \mathbf{d}(\omega) - \mathbf{L}(\omega)\mathbf{m}(\omega) \|_2^2 + \mu R(\mathbf{m}(\omega)) \|_2^2$$

$$\nabla J(\omega) = 0 \rightarrow$$

$$\mathbf{m}(\omega) = (\mathbf{L}(\omega)^H \mathbf{L}(\omega) + \mu \mathbf{Q}(\mathbf{m}))^{-1} \mathbf{L}(\omega)^H \mathbf{d}(\omega)$$

$$Q_{ii}(m(\omega)) \propto \frac{1}{|m(\omega)_i|}$$

All high res methods must use weights with that form.
Why?

Flavors of Freq. domain RT

IRLS (Sacchi and Ulrych, GEO95)

$$\mathbf{m}^k(\mathcal{W}) = (\mathbf{L}(\mathcal{W})^H \mathbf{L}(\mathcal{W}) + m\mathbf{Q}(\mathbf{m}^{k-1}(\mathcal{W})))^{-1} \mathbf{L}(\mathcal{W})^H \mathbf{d}(\mathcal{W})$$

CGG - P. Herrmann et al., 02

$$\mathbf{m}(\mathcal{W}_k) = (\mathbf{L}(\mathcal{W}_k)^H \mathbf{L}(\mathcal{W}_k) + m\mathbf{Q}(\mathbf{m}(\mathcal{W}_{k-1})))^{-1} \mathbf{L}(\mathcal{W}_k)^H \mathbf{d}(\mathcal{W}_k)$$

Flavors of Freq. domain RT (I. Moore)

2 - passes

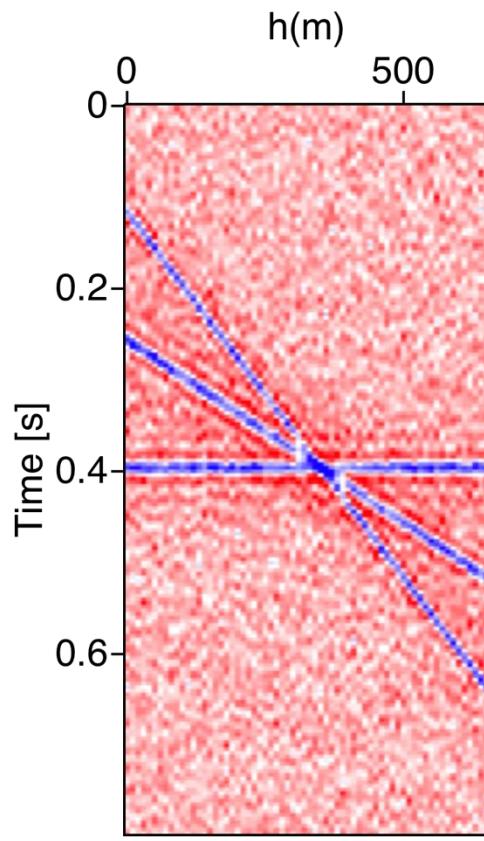
$$1) \mathbf{m}(W) = (\mathbf{L}(W)^H \mathbf{L}(W) + m\mathbf{I})^{-1} \mathbf{L}(W)^H \mathbf{d}(W)$$

$$2) Q^{-1} \uplus Semblance(m)$$

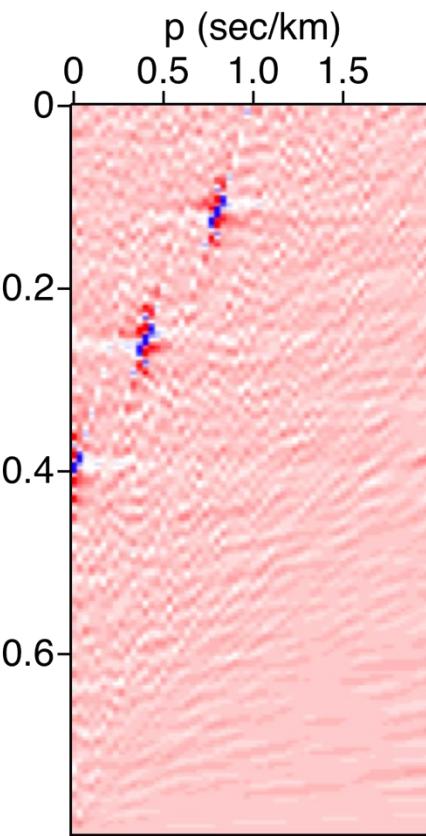
$$3) \mathbf{m}(W_k) = (\mathbf{L}(W_k)^H \mathbf{L}(W_k) + m\mathbf{Q}(S(W)))^{-1} \mathbf{L}(W_k)^H \mathbf{d}(W_k)$$

Comparison of Least-squares (Smooth) versus Sparse (high-res) Radon transform → IRLS

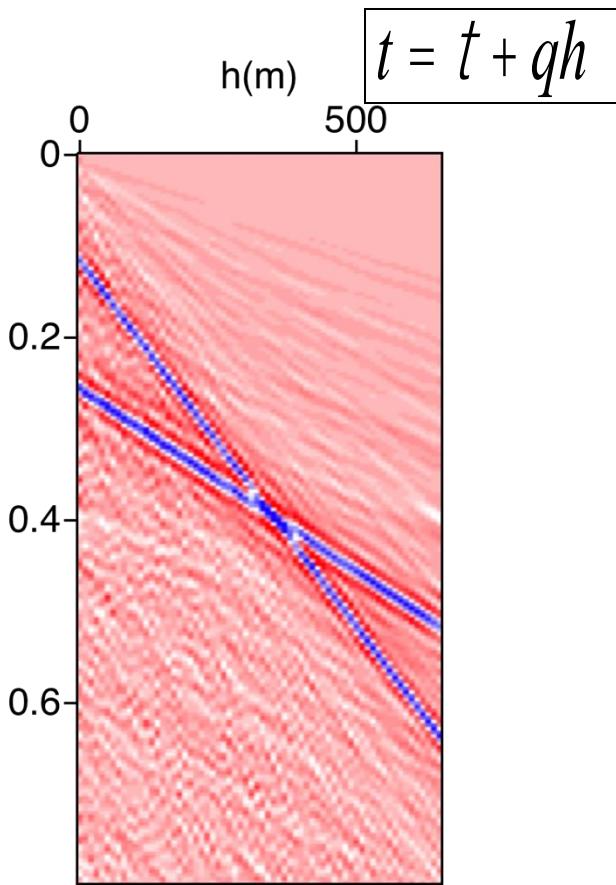
Linear Radon Transform



Data

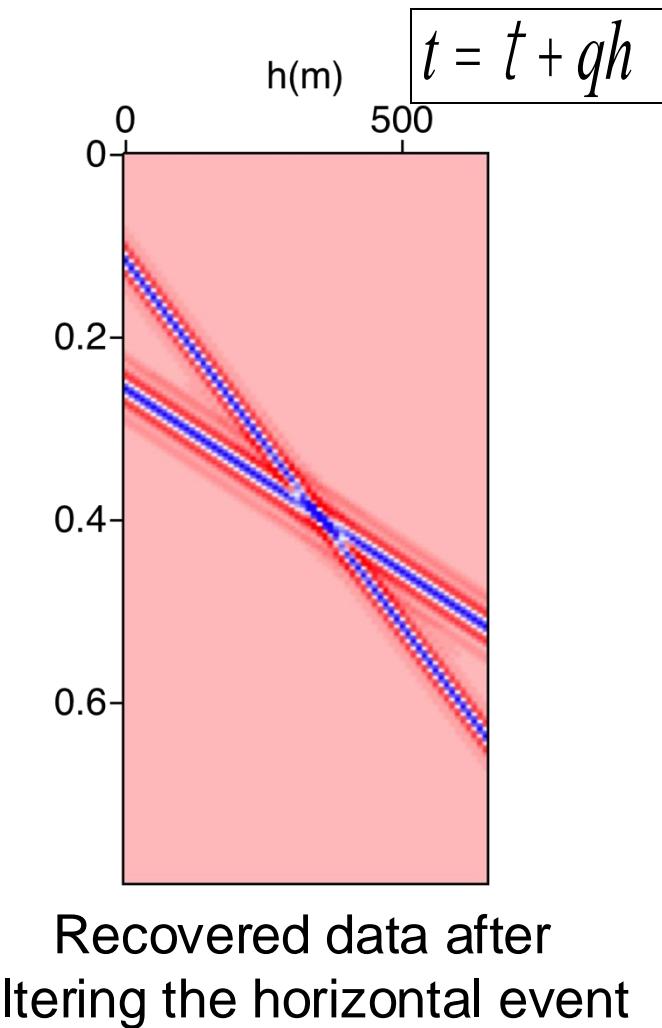
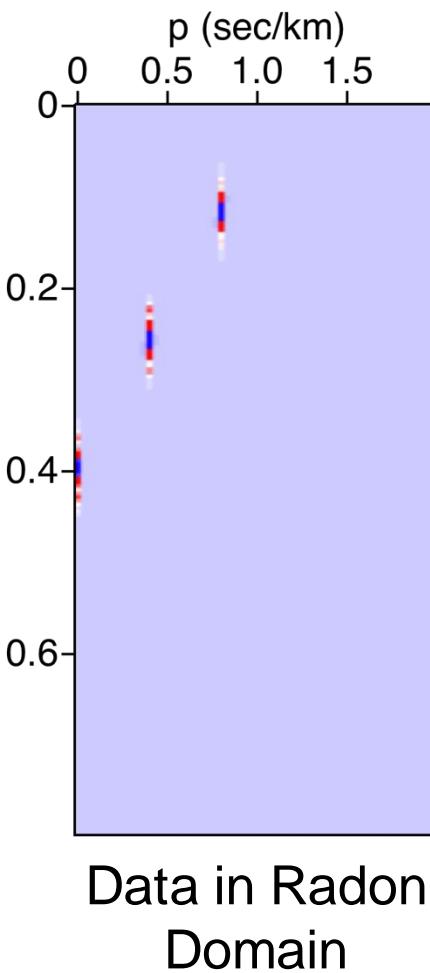
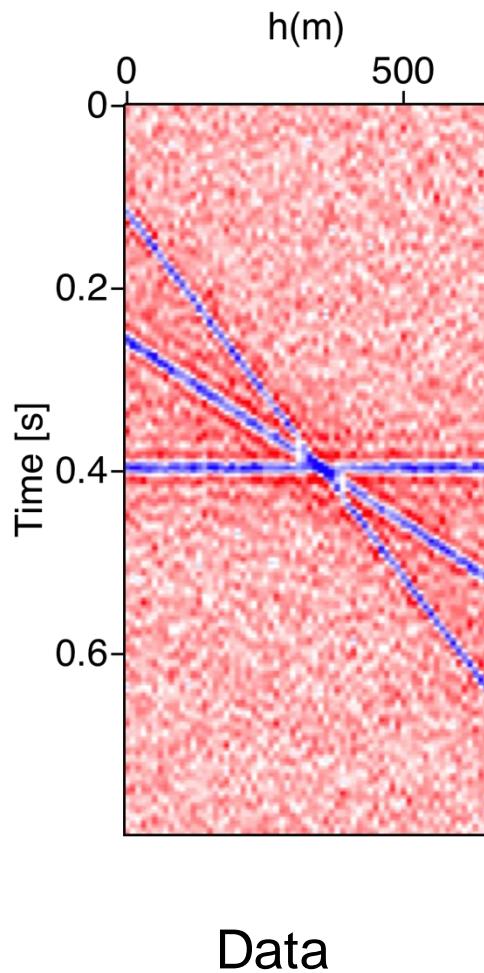


Data in Radon
Domain

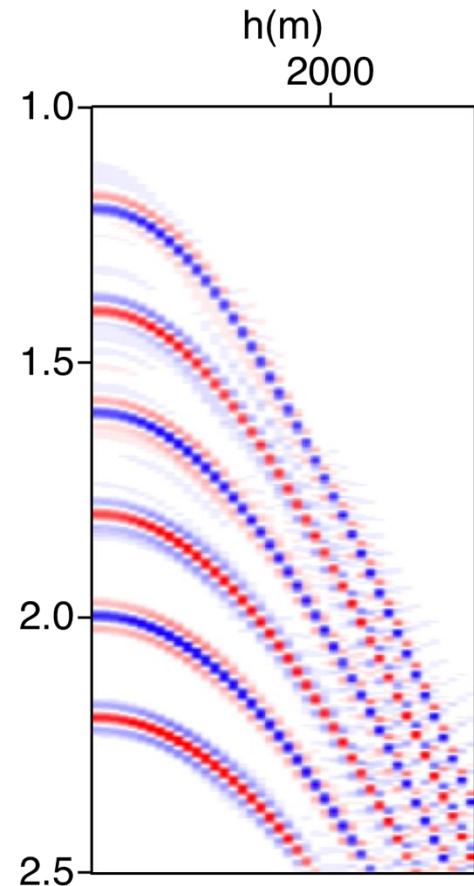
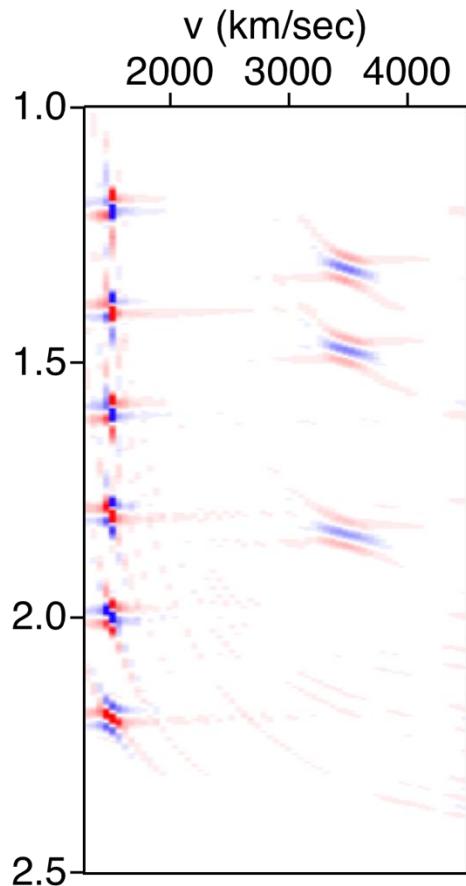
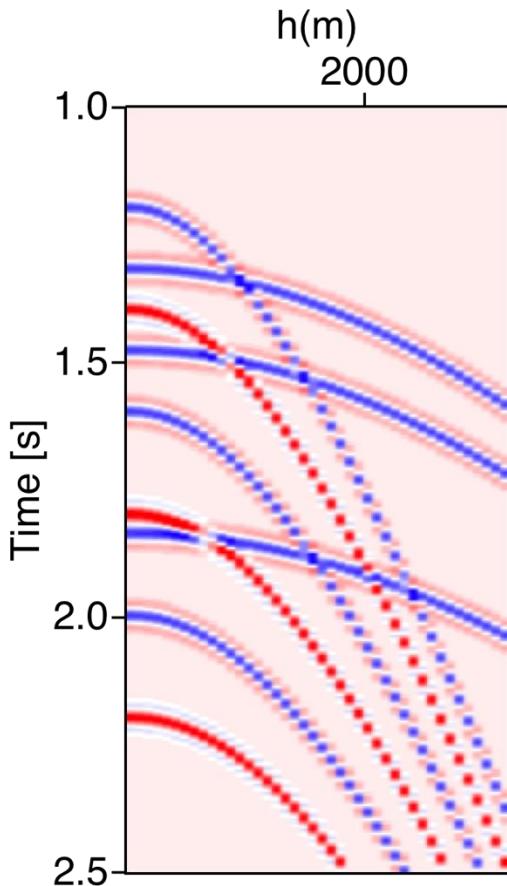


Recovered data after
filtering the horizontal event

Linear Radon Transform

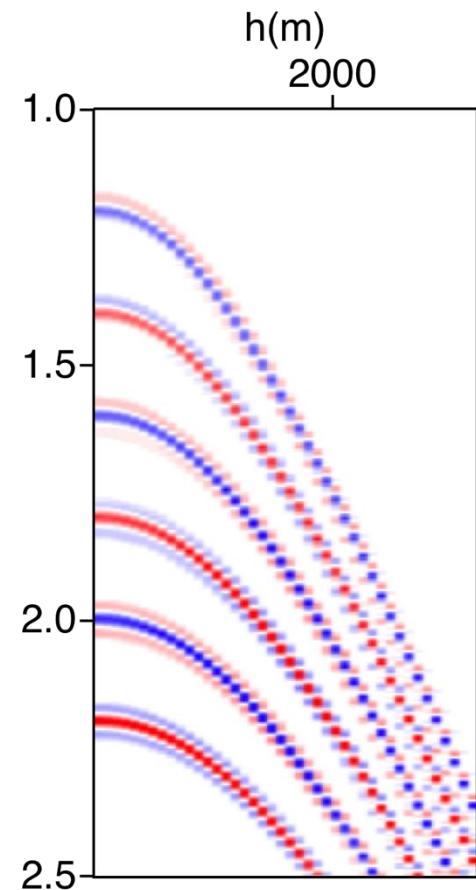
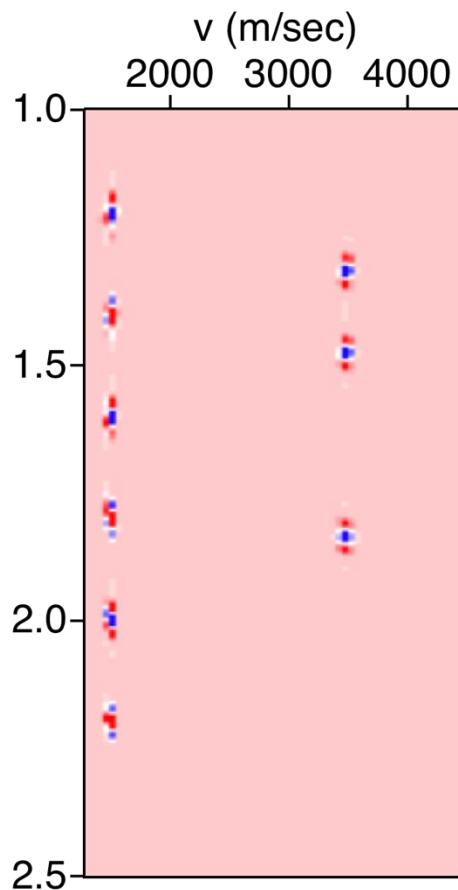
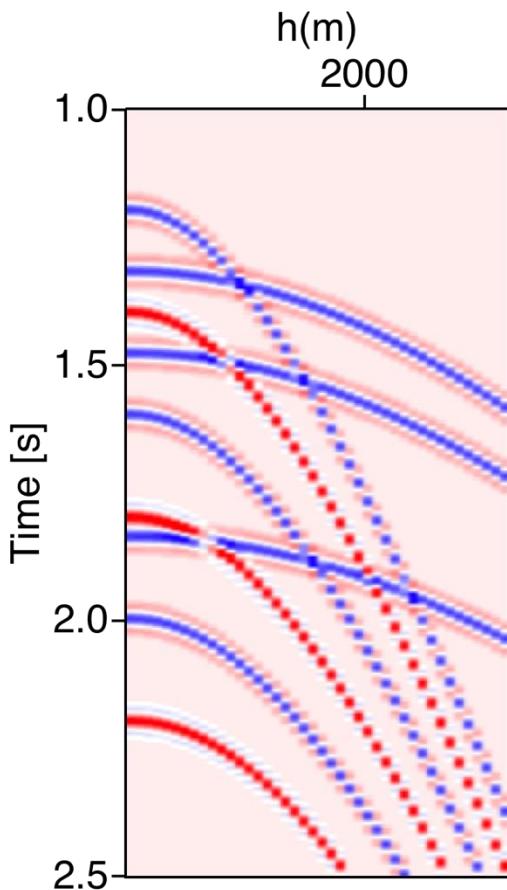


Hyperbolic Radon Transform



$$t = \sqrt{t^2 + qh^2}$$

Hyperbolic Radon Transform



$$t = \sqrt{t^2 + qh^2}$$

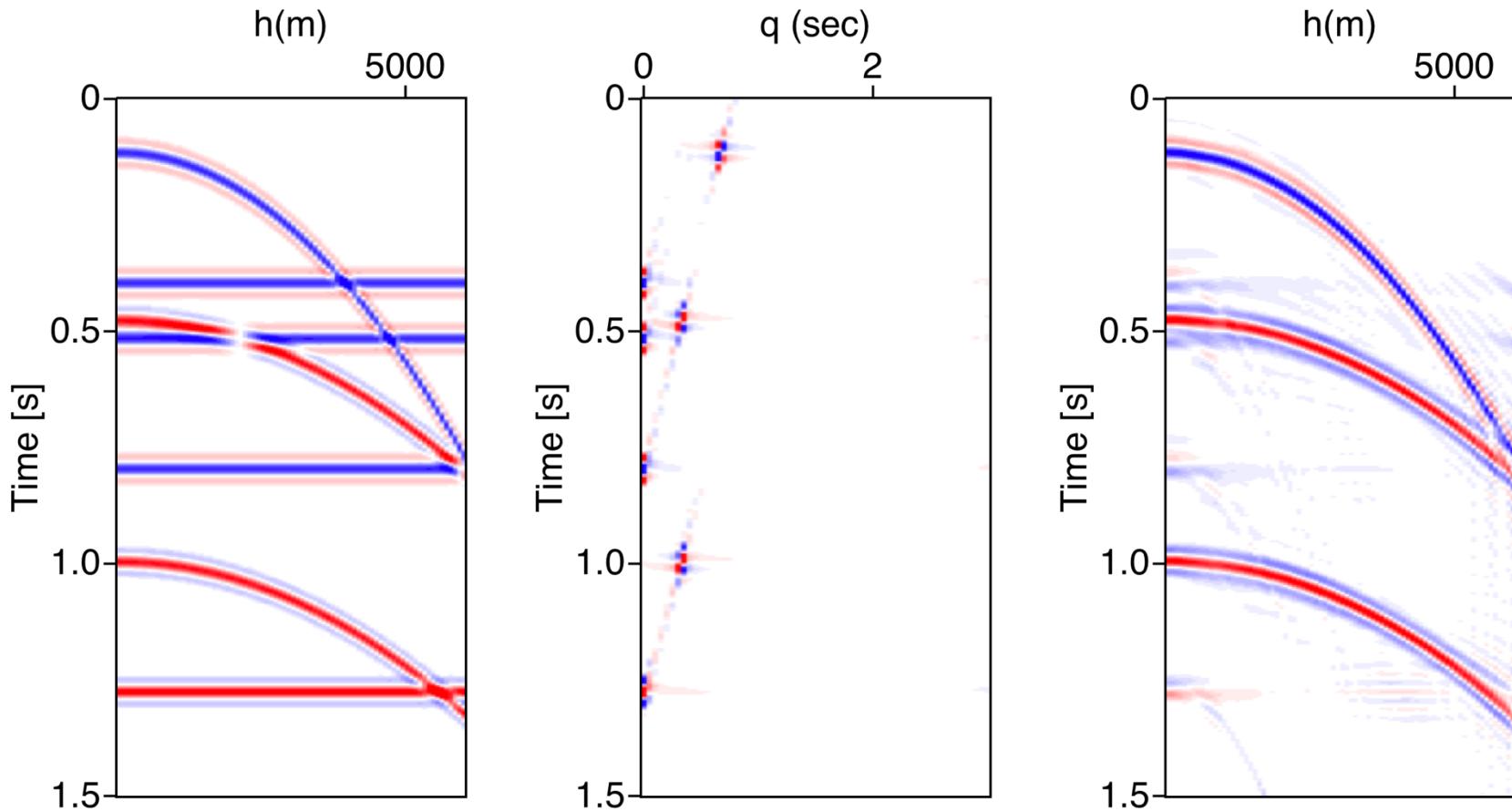
Parabolic Radon Transform: A note on q

- q is residual moveout at far offset (in secs)

$$t = \tau + q h^2 \rightarrow t = \tau + q \frac{h^2}{h_{\max}^2}$$

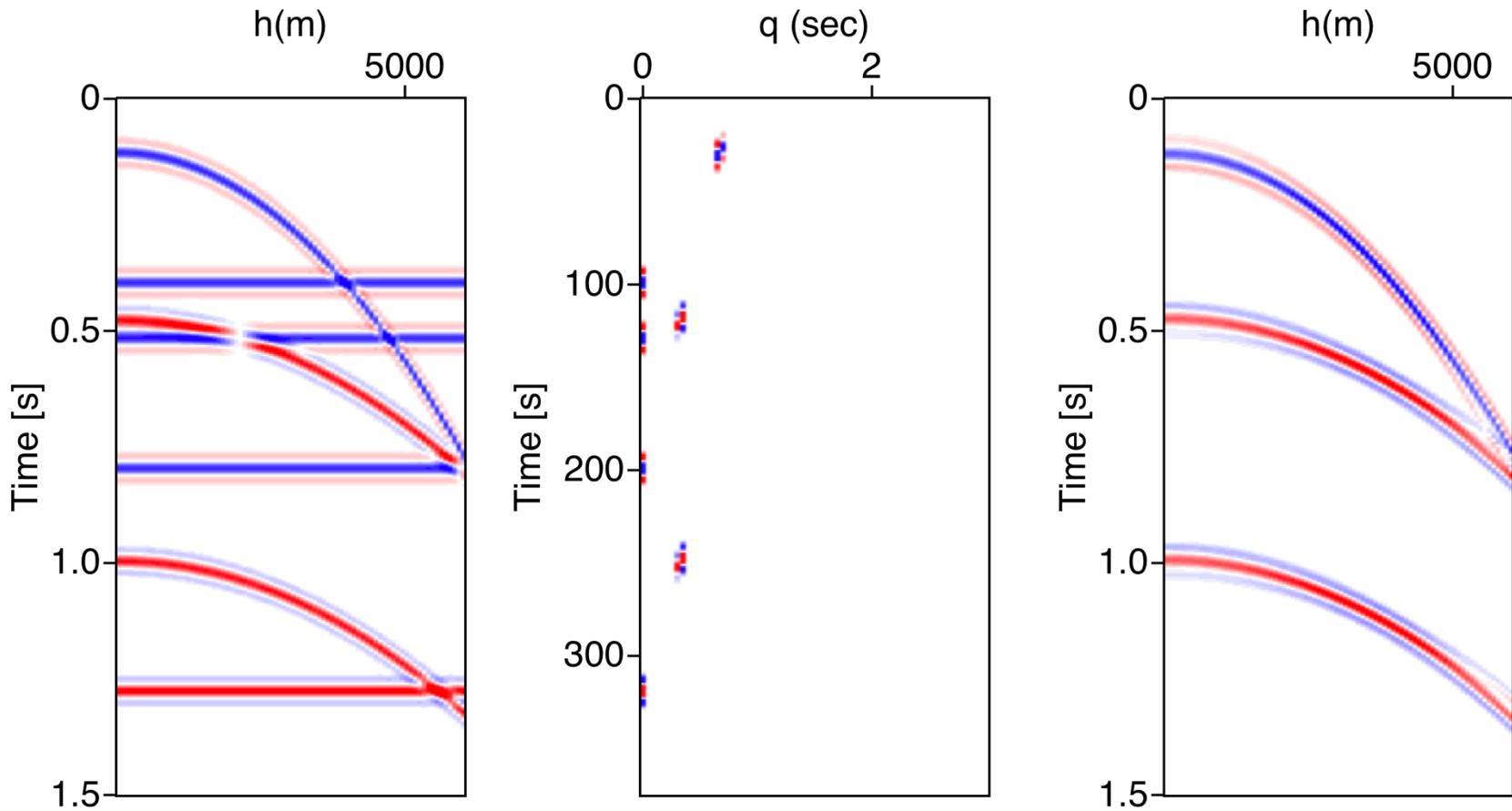
- Hyperbolic traveltimes after NMO correction can be approximated by parabolic traveltimes curves.

Parabolic Radon Transform



$$t = t + qh^2$$

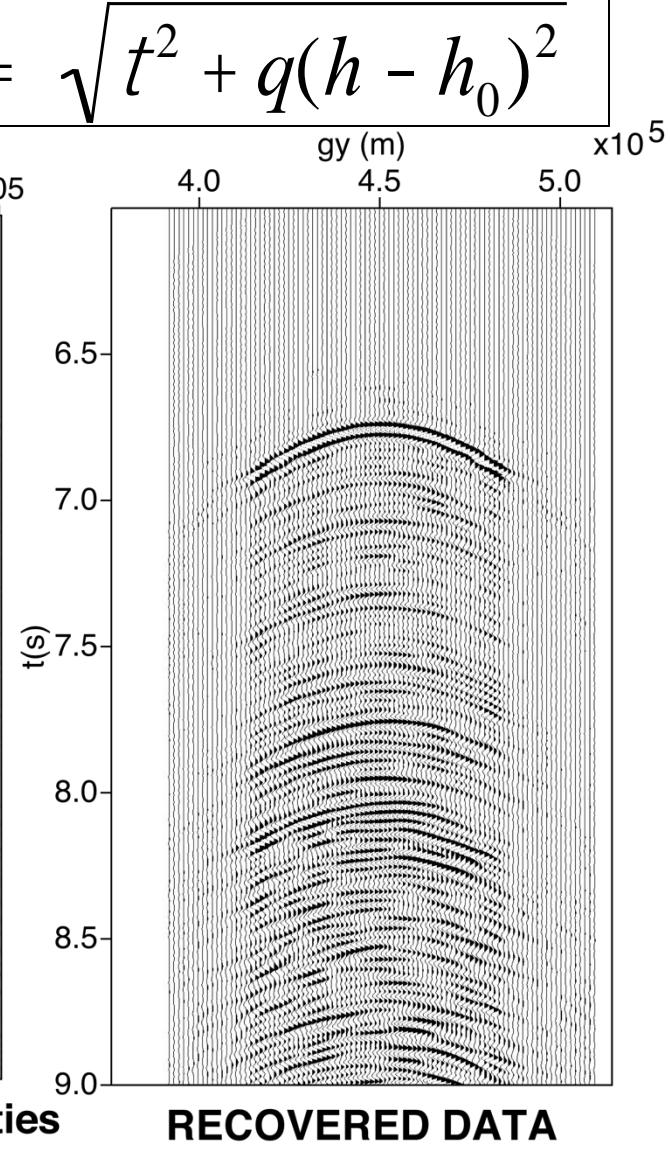
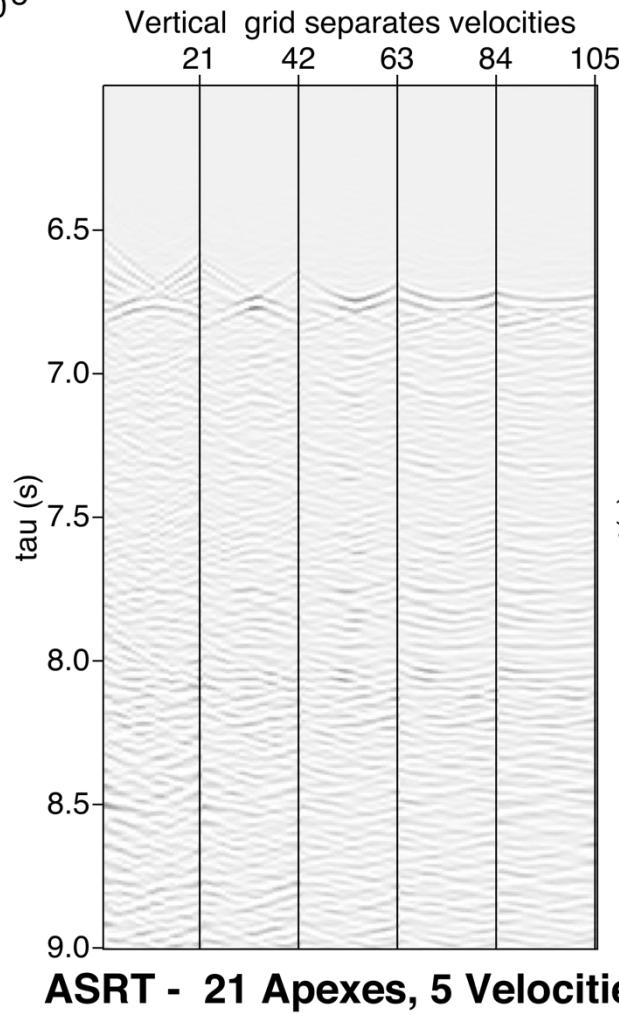
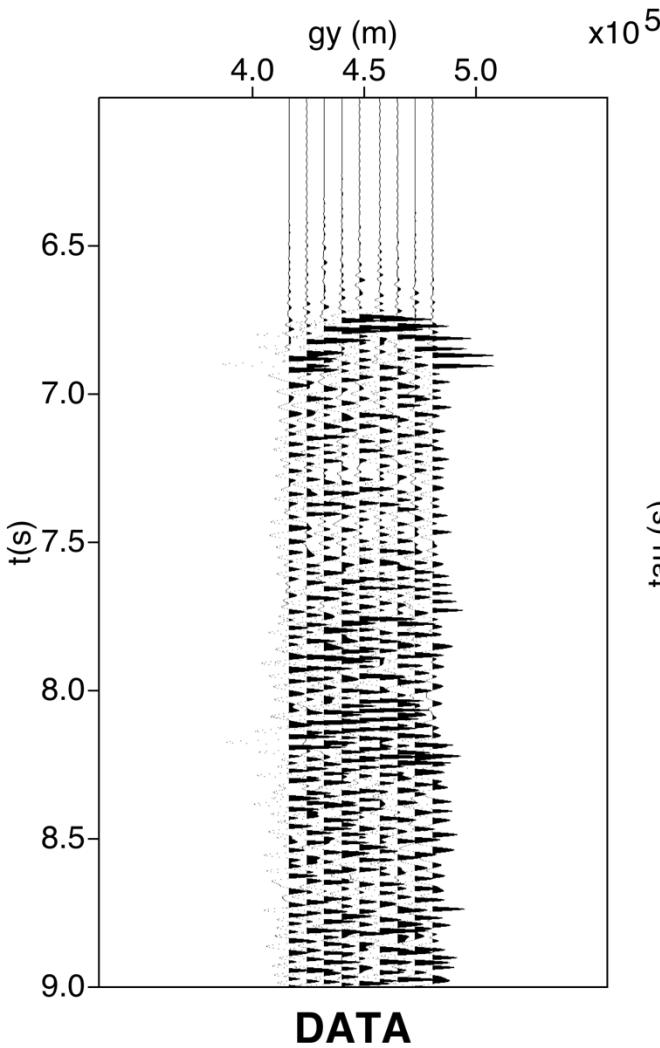
Parabolic Radon Transform



$$t = t + qh^2$$

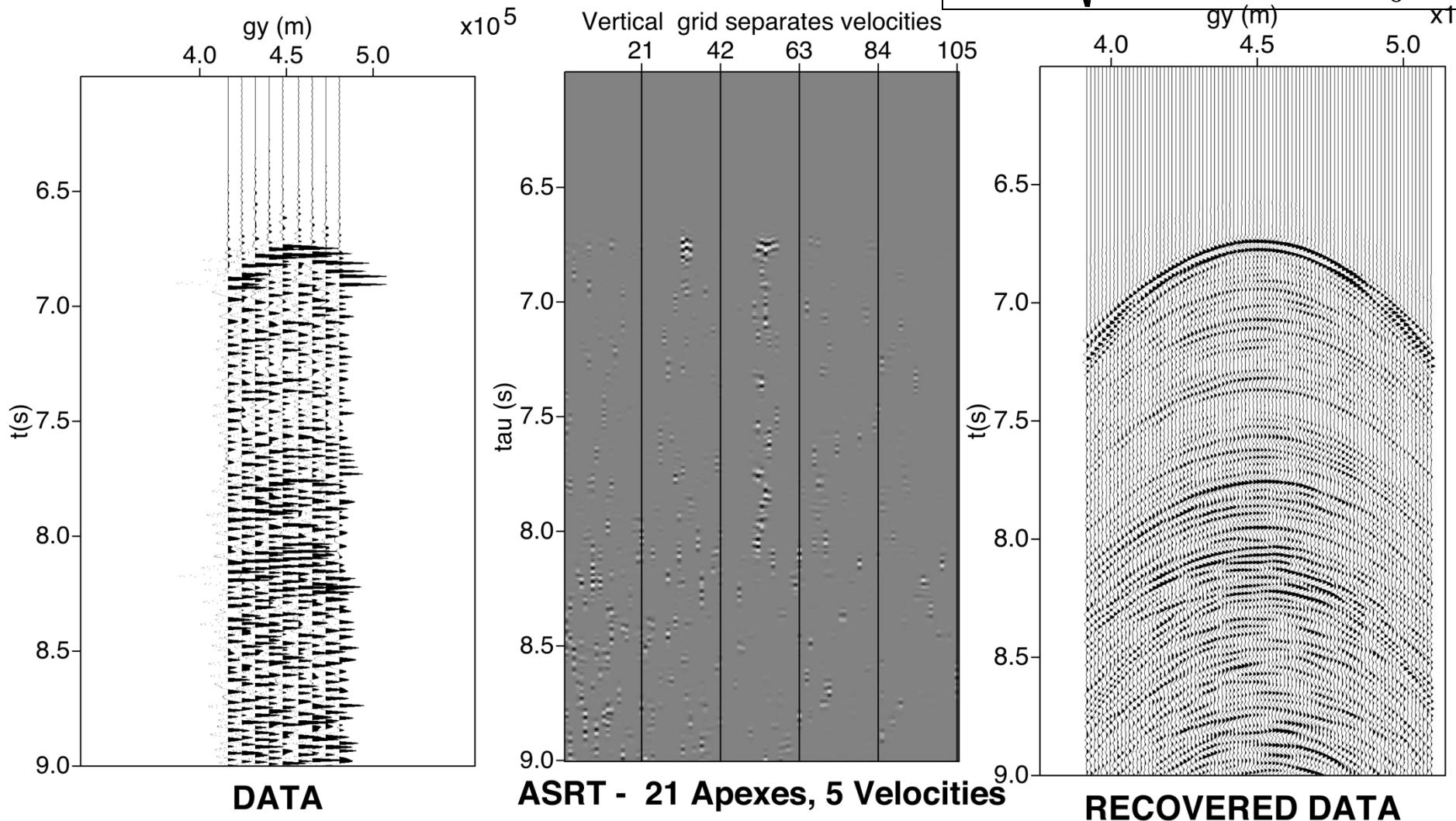
Hyperbolic Radon Transform with apex

$$t = \sqrt{t^2 + q(h - h_0)^2}$$



Hyperbolic Radon Transform with apex

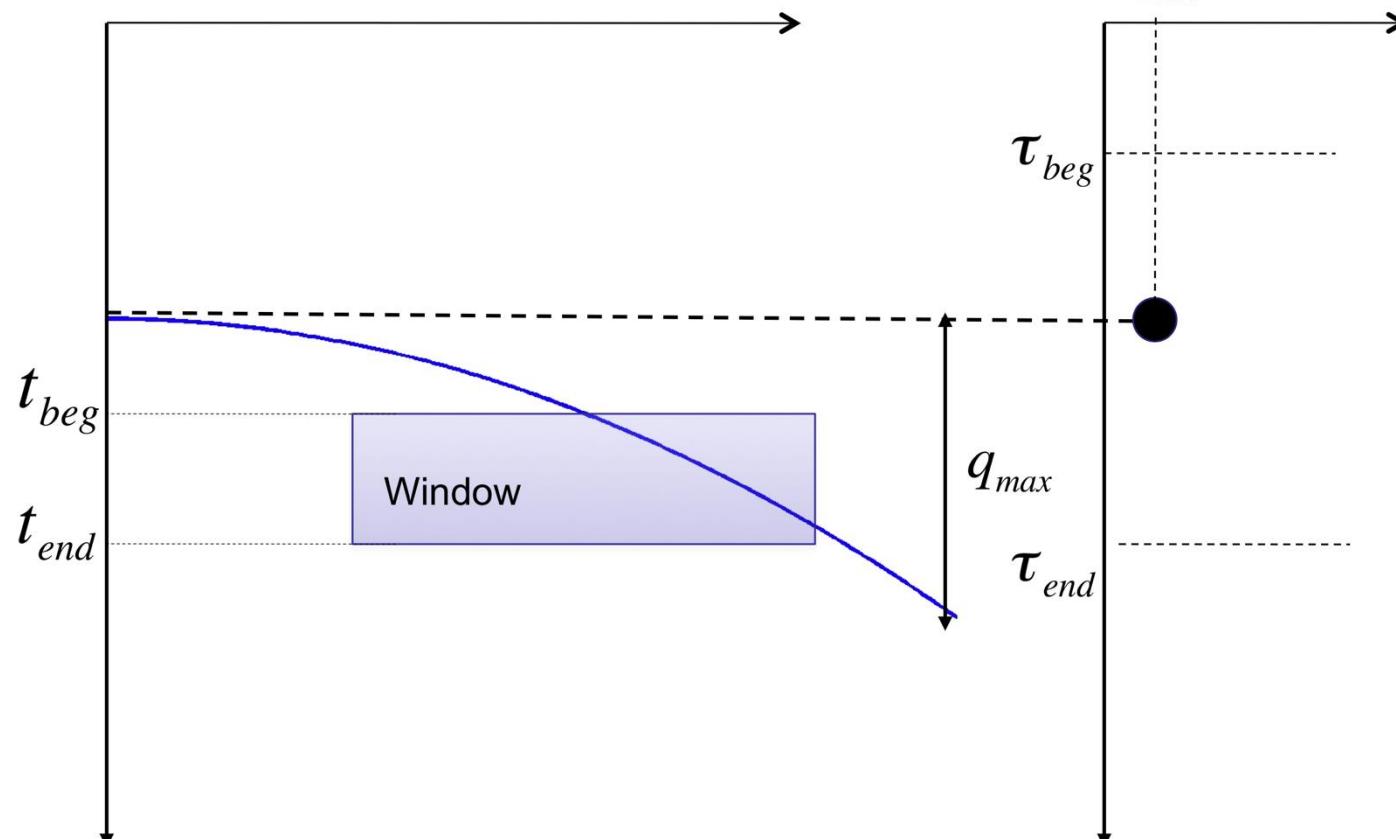
$$t = \sqrt{t^2 + q(h - h_0)^2}$$



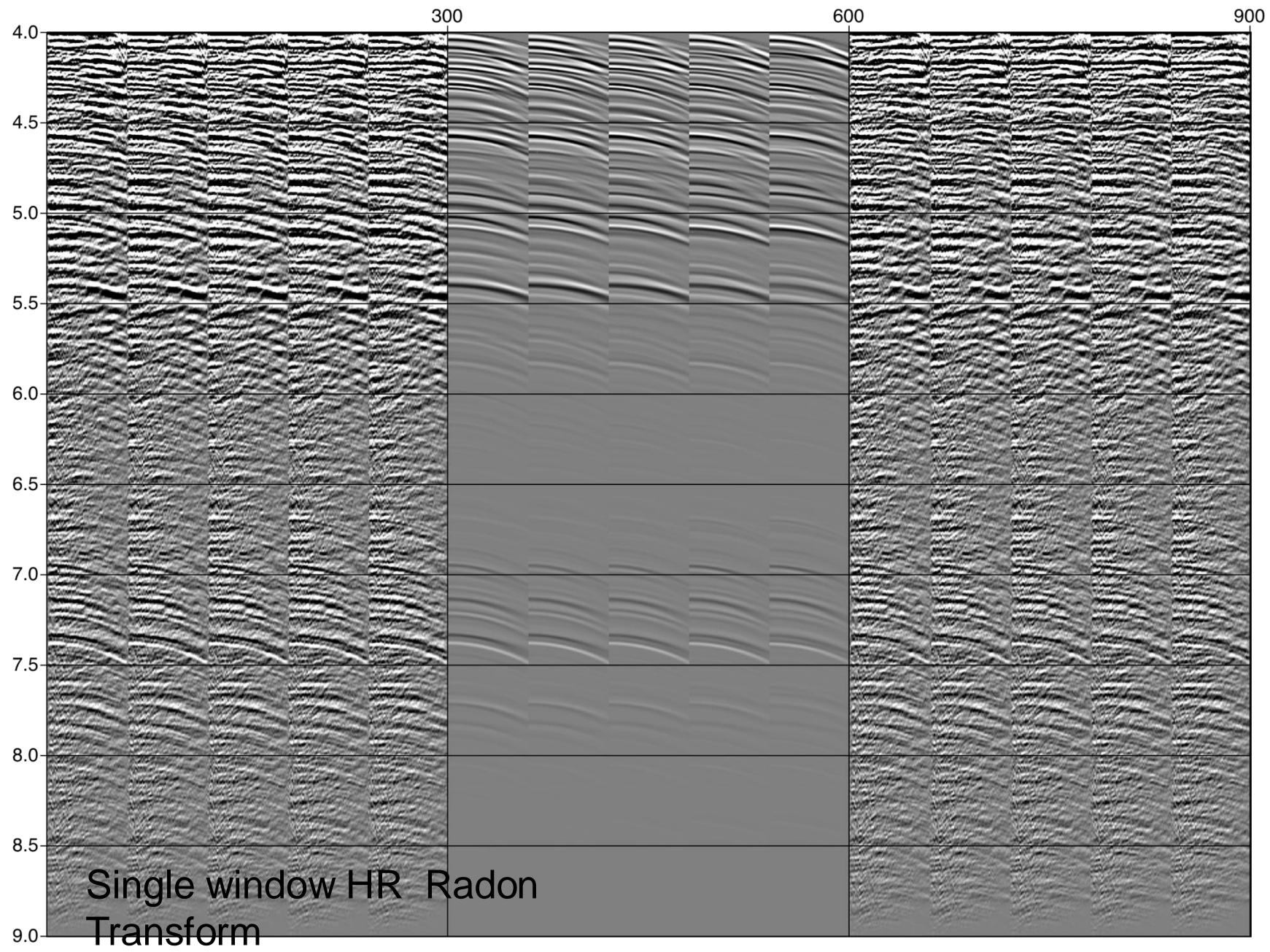
Warning: Sparsity works better in small windows

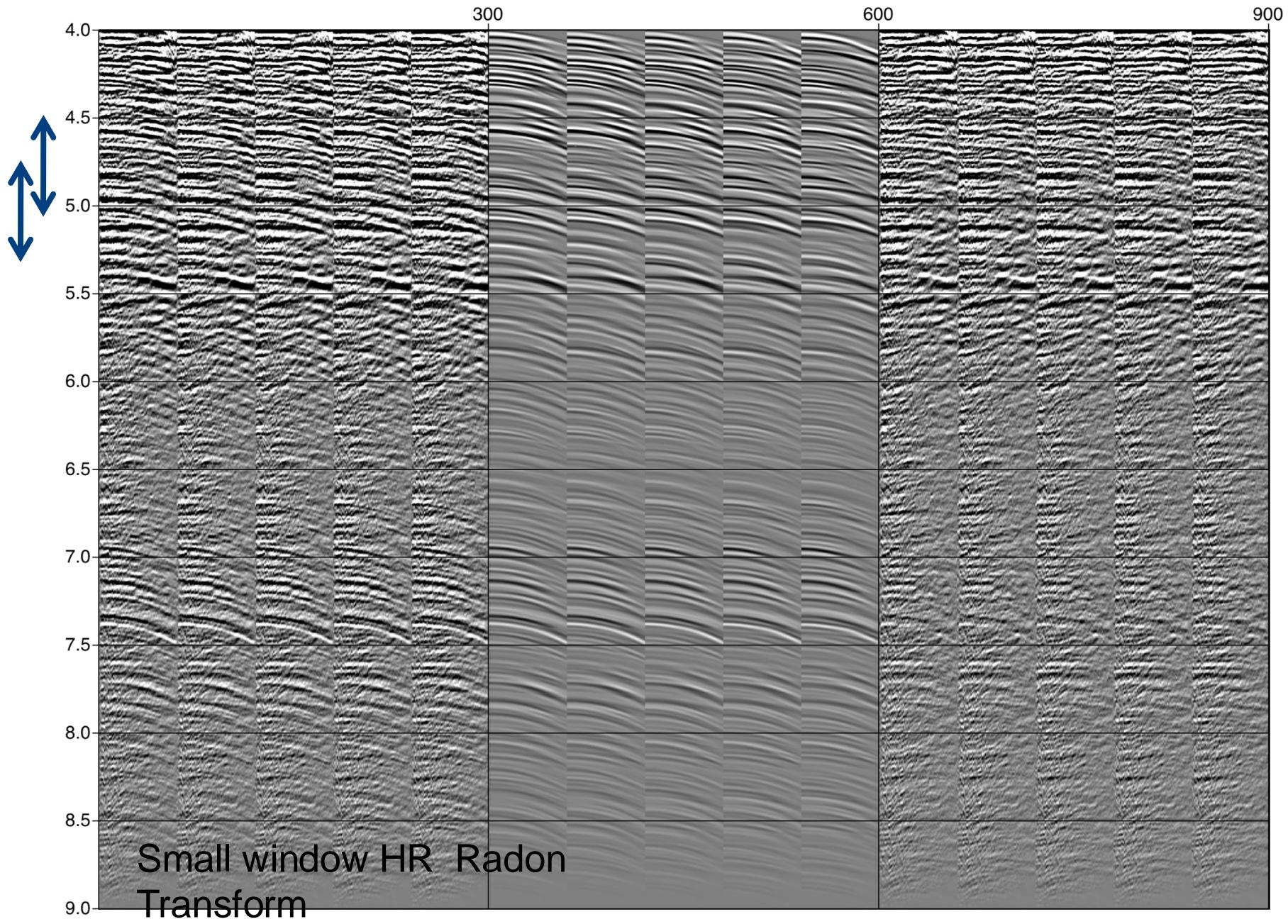
(similar problem in MED-Deconv, Sparse decon, etc)

Parameterization of
the operator: Parabolic case



The Radon window should be bigger than the data window





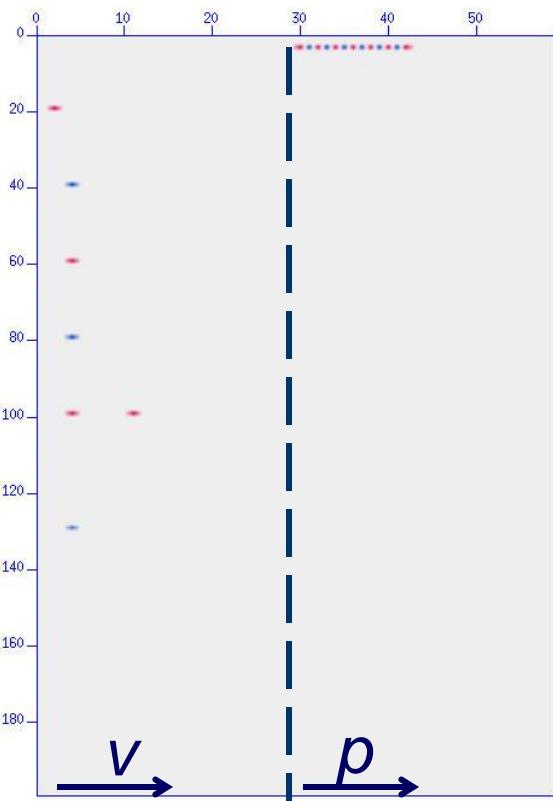
Hybrid Radon Transform

Like a dictionary of Radons

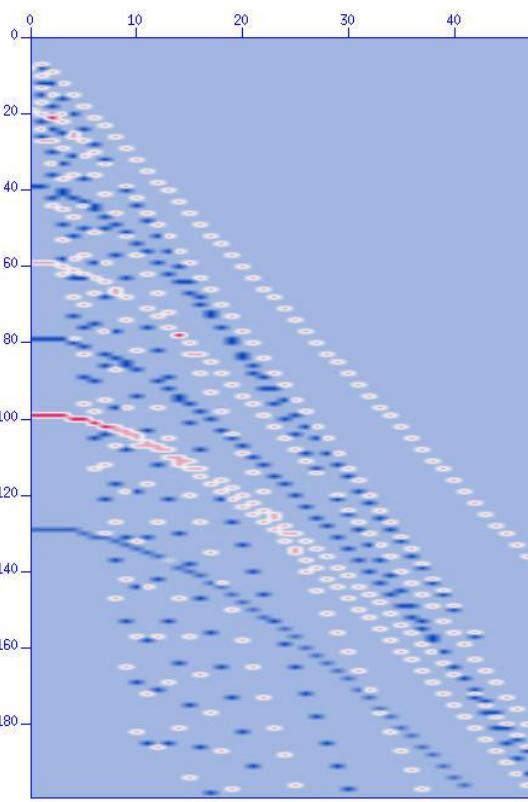
$$d = \begin{pmatrix} L_h & L_l \end{pmatrix} \begin{matrix} \overset{\text{am}}{\underset{\text{em}}{\underset{\text{C}}{\underset{\text{S}}{\underset{\text{E}}{\underset{\text{O}}{\div}}}}} \\ m_h \\ m_l \\ \emptyset \end{matrix} + n$$

$$= Lm + n$$

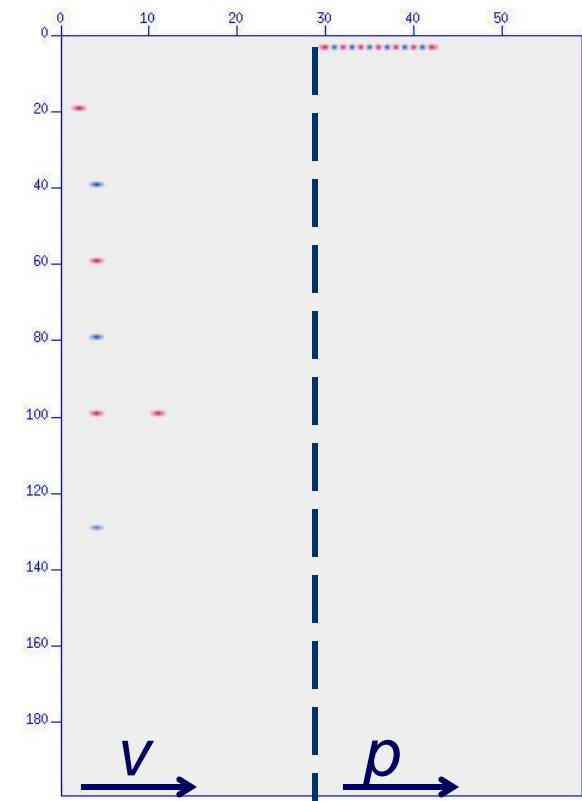
Hybrid Radon Transform



m

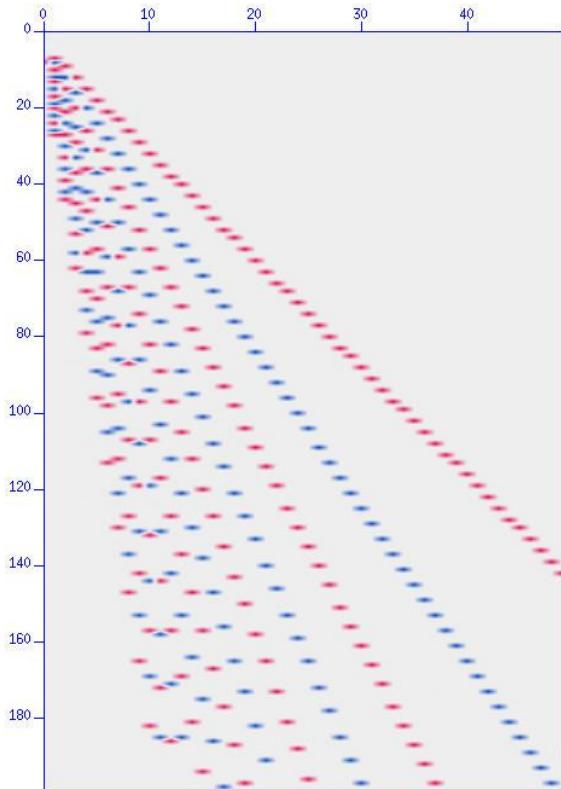


d

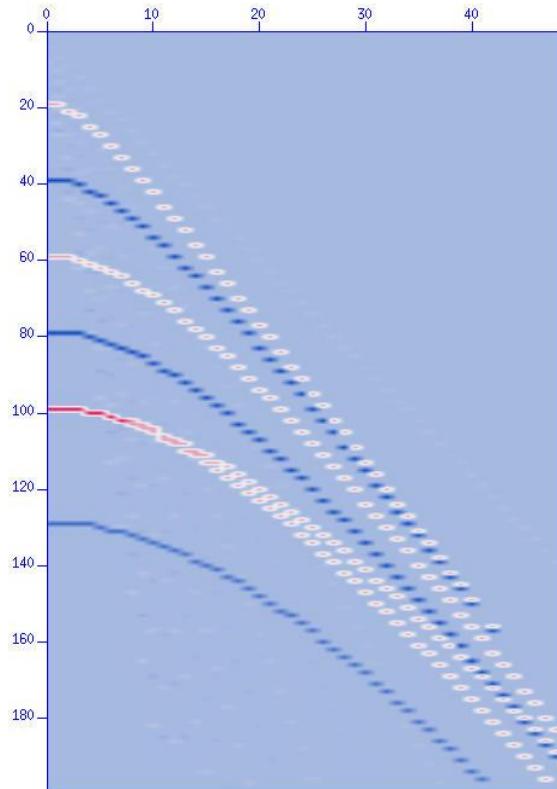


*Inverted m
Sparse Inversion*

Hybrid Radon Transform

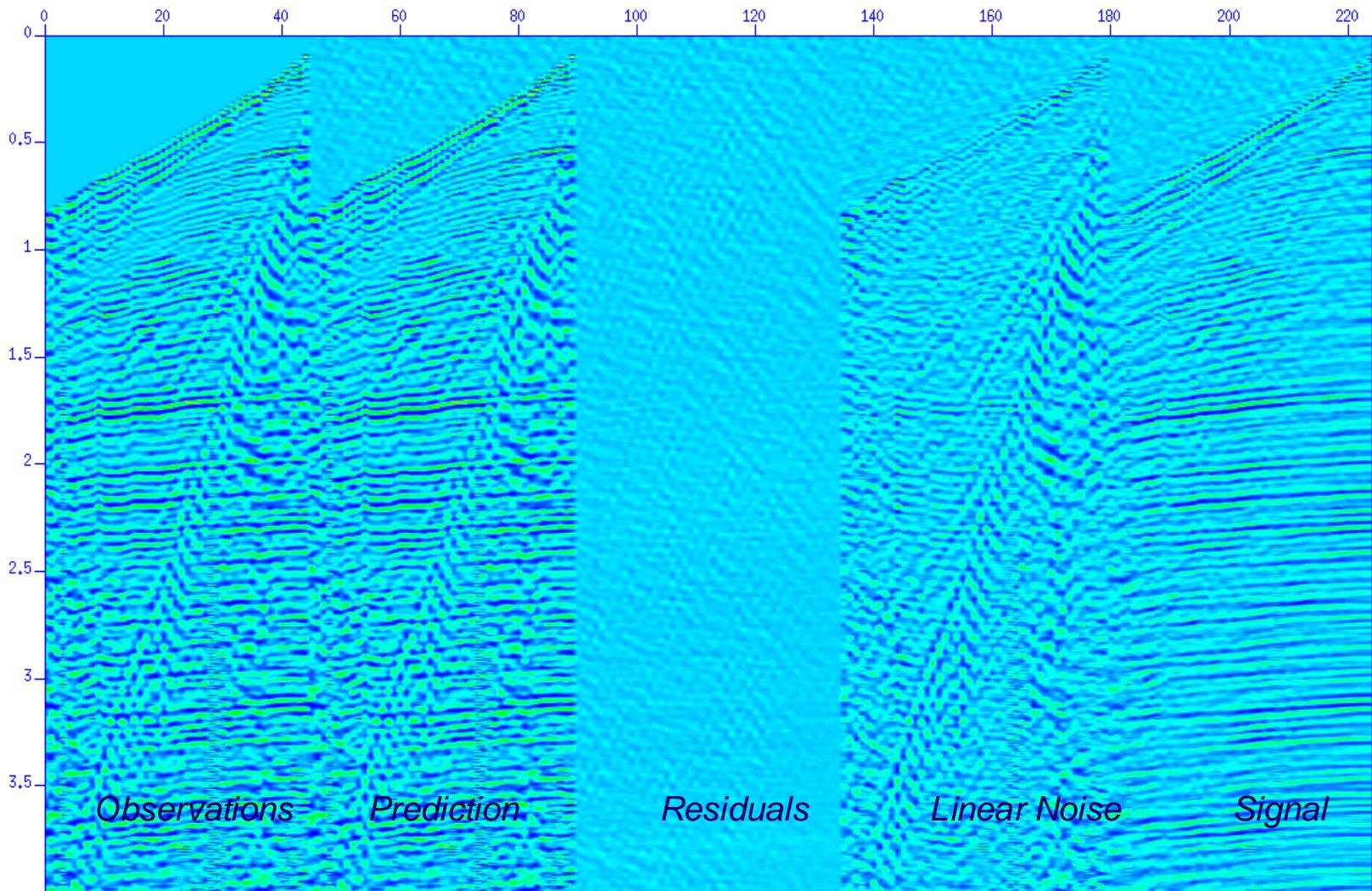


Recovered Model
of linear events



Recovered model of
hyperbolic events

Hybrid Radon



The Linear Event Model: FX Decon and Cazdow / SSA filters

Mauricio Sacchi

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Deconvolution in F-X

- Why f-x?
- Linear events in t-x are predictable in f-x

A Model for FX data (One event)

- Time Domain Delay \cup Frequency Domain Linear Phase

$$s(x, t) = a(t - xp)$$



$$S(x, \omega) = A(\omega) e^{-i\omega px}$$

A Model for FX data (One event)

- Time Domain Delay \Leftrightarrow Frequency Domain Linear Phase:

$$S(x, \omega) = A(\omega) e^{-i\omega p x}$$

$$x = (j - 1).\Delta x, \quad \omega = \omega_l$$

$$S_{jl} = A_l e^{-i\omega_l p(j-1)\Delta x}$$

- Same model as a function of Wave-number:

$$k_l = \omega_l p$$

$$S_{jl} = A_l e^{-ik_l(j-1)\Delta x}$$

A Model for FX data (One event)

- We now consider the problem as one of predicting data along the spatial coordinate for fix temporal frequency

$$S_j = A e^{-ik(j-1)\Delta x}, \quad j = 1, \dots, N$$

- Take for instance a linear event in t - x with the above representation in f - x , we write:

$$\begin{aligned} S_{j-1} &= A e^{-ik(j-2)\Delta x} \\ &= A e^{-ik(j-1)\Delta x} e^{ik\Delta x} \\ &= S_j e^{ik\Delta x} \end{aligned}$$

A Model for FX data (One event)

- From last expression, we can derive a recursive relation between the signal recorder at receiver j as a function of the signal at receiver $j-1$

$$\begin{aligned} S_j &= S_{j-1} e^{-ik\Delta x} \\ &= aS_{j-1} \end{aligned}$$

- The last simple expression is the basis for f - x prediction/deconvolution and snr enhancement
- If the signal can be predicted, then the non-predictable part must be the noise

Canales, L. L., 1984, Random noise reduction: 54th Annual International Meeting, SEG, Expanded Abstracts, Session:S10.1.

Gulunay, N., 1986, FX decon and complex Wiener prediction filter: 56th Annual International Meeting, SEG, Expanded Abstracts, Session:POS2.10.

Gulunay, N., 2000, Noncausal spatial prediction filtering for random noise reduction on 3-D poststack data: Geophysics, 65, 1641-1653.

A Model for FX data (One event)

- From last expression, we can derive a recursive relation between the signal recorder at receiver j and a function of the signal at receiver $j-1$

$$S_j = aS_{j-1} + N_j, \quad j = 1 \dots nx$$

- Determine a from available data the estimate noise

$$\begin{aligned} S_2 &= aS_1 + N_2 & \ddot{u} \\ S_3 &= aS_2 + N_3 & \ddot{i} \\ S_4 &= aS_3 + N_4 & \ddot{j} \\ \cdot & & \ddot{y} \\ \cdot & & \ddot{b} \\ S_N &= aS_{N-1} + N_{nx} \end{aligned}$$

System of equations from where one can estimate a and the noise (Least Squares!)

A Model for FX data (NP events)

- We can prove that if the data consists of NP linear events

$$s(x, t) = \sum_{n=1}^{NP} a(t - p_n x)$$

↔

$$S(x, \omega) = \sum_{n=1}^{NP} A(\omega) e^{-i\omega p_n x}$$

$$x = (j-1).\Delta x, \quad \omega = \omega_l$$

$$S_{jl} = \sum_{n=1}^{NP} A_l e^{-i\omega_l p_n (j-1)\Delta x}$$

A Model for FX data (NP events)

- We can prove that if the data consists of NP linear events

$$\begin{aligned} S_j &= \sum_{n=1}^{NP} A e^{-K_n(j-1)\Delta x} \\ &= a_1 S_{j-1} + a_2 S_{j-2} + a_3 S_{j-3} + \cdots + a_{NP} S_{j-NP} \end{aligned}$$

Recursion of order NP

- The coefficients of the recursion are also called prediction error coefficients - They are related to the wavenumber of each linear event

A Model for FX data (Adding noise)

- We now consider also noise:

$$S_j = a_1 S_{j-1} + a_2 S_{j-2} + a_3 S_{j-3} + \square + a_{LF} S_{j-LF} + N_j$$

- Suppose $LF=3$ and $nx=24$ channels

$$S_4 = a_1 S_3 + a_2 S_2 + a_3 S_1 + N_4$$

$$S_5 = a_1 S_4 + a_2 S_3 + a_3 S_2 + N_5$$

$$S_6 = a_1 S_5 + a_2 S_4 + a_3 S_3 + N_6$$

...

$$S_{24} = a_1 S_{23} + a_2 S_{22} + a_3 S_{21} + N_{24}$$

LF (not NP) coefficients

??

Notice that I have added noise

A Model for FX data (Using the available information)

Ignoring non-recorded data (Using available information
only !)

$$S_4 = a_1 S_3 + a_2 S_2 + a_3 S_1 + N_4$$

$$S_5 = a_1 S_4 + a_2 S_3 + a_3 S_2 + N_5$$

$$S_6 = a_1 S_5 + a_2 S_4 + a_3 S_3 + N_6$$

...

$$S_{24} = a_1 S_{23} + a_2 S_{22} + a_3 S_{21} + N_{24}$$

P

$$\begin{array}{ccccccccc} \alpha & S_4 & \ddot{\alpha} & \alpha & S_3 & S_2 & S_1 & \ddot{\alpha} & \alpha & N_4 & \ddot{\alpha} & \ddot{\alpha} \\ \zeta & S_5 & \div & \zeta & S_4 & S_3 & S_2 & \div & \alpha & a_1 & \ddot{\alpha} & \zeta \\ \zeta & S_6 & \div & = & \zeta & S_5 & S_4 & \div & \zeta & a_2 & \div & \zeta \\ \zeta & \square & \div & \zeta & \square & \square & \square & \div & \zeta & a_3 & \div & \zeta \\ \zeta & \div & \zeta & \div & \zeta & \div & \div & \div & \zeta & \div & \div & \div \\ \text{e} & N_{24} & \emptyset & \text{e} & S_{23} & S_{22} & S_{21} & \emptyset & \text{e} & N_{24} & \emptyset & \emptyset \end{array} \quad \mathbf{s} = \mathbf{M}\mathbf{a} + \mathbf{n}$$

A Model for FX data (Assuming data outside aperture=0)

Considering non-recorded data = 0

$$S_2 = a_1 S_1 + N_2$$

$$S_3 = a_1 S_2 + a_2 S_1 + N_3$$

$$S_4 = a_1 S_3 + a_2 S_2 + a_3 S_1 + N_4$$

$$S_5 = a_1 S_4 + a_2 S_3 + a_3 S_2 + N_5$$

$$S_6 = a_1 S_5 + a_2 S_4 + a_3 S_3 + N_6$$

□

...

$$S_{24} = a_1 S_{23} + a_2 S_{22} + a_3 S_{21} + N_{24}$$

$$0 = a_1 S_{24} + a_2 S_{23} + a_3 S_{22} + N_{25}$$

$$0 = a_2 S_{24} + a_3 S_{23} + N_{26}$$

$$0 = a_3 S_{24} + N_{27}$$

A Model for FX data (Assuming data outside aperture=0)

Considering non-recorded data = 0

$$\begin{array}{ccccccccc}
 \mathbb{A} S_2 & \ddot{0} & \mathbb{A} S_1 & 0 & 0 & \ddot{0} & \mathbb{A} N_2 & \ddot{0} & \ddot{U} \\
 \zeta S_3 & \div & \zeta S_2 & S_1 & 0 & \div & \zeta N_3 & \div & \ddot{I} \\
 \zeta S_4 & \div & \zeta S_3 & S_2 & S_1 & \div & \zeta N_4 & \div & \ddot{I} \\
 \zeta S_5 & \div & \zeta S_4 & S_3 & S_2 & \div \mathbb{A} a_1 & \ddot{0} & \zeta N_5 & \div \\
 \zeta \square & \div = & \zeta S_5 & S_4 & S_3 & \div \zeta a_2 & \div + & \zeta \square & \div \text{ } \bar{y} \\
 \zeta S_{24} & \div & \zeta \square & \square & \square & \div \zeta a_3 & \div \emptyset & \zeta N_{24} & \div \\
 \zeta 0 & \div & \zeta S_{24} & S_{23} & S_{22} & \div & \zeta N_{25} & \div & \ddot{I} \\
 \zeta 0 & \div & \zeta 0 & S_{24} & S_{23} & \div & \zeta N_{26} & \div & \ddot{I} \\
 \mathbb{E} 0 & \div & \mathbb{E} 0 & 0 & S_{24} \emptyset & \div & \mathbb{E} N_{27} \emptyset & \div & \mathbb{B}
 \end{array}$$

s = Ca + n

A Model for FX data (NP events)

Two formulations: $\mathbf{s} = \mathbf{M}\mathbf{a} + \mathbf{n}$ (Transient free case)

$\mathbf{s} = \mathbf{C}\mathbf{a} + \mathbf{n}$ (Toeplitz case)

Two Least Squares Solutions:

$$\hat{\mathbf{a}}_1 = (\mathbf{M}^H \mathbf{M})^{-1} \mathbf{M}^H \mathbf{s}$$

$\mathbf{M}^H \mathbf{M}$ is a Non - Toeplitz correlation data matrix

$$\hat{\mathbf{a}}_2 = (\mathbf{C}^H \mathbf{C})^{-1} \mathbf{C}^H \mathbf{s}$$

$\mathbf{C}^H \mathbf{C}$ is the Toeplitz correlation data matrix

A Model for FX data (Algorithm)

Algorithm to recover the noise

1 - Original data in $t - x \rightarrow f - x$

2 -

For each frequency :

Find \hat{a} that solves $s = Ma + n$,

Use \hat{a} to predict the noise $\hat{n} = s - M\hat{a}$

End for

3 - $f - x \rightarrow t - x$

4 - Subtract noise from original data

A Model for FX data (Trade-off parameters)

- NP does not define the filter length LF
(Number of coefficients of the prediction error operator \mathbf{a} .)
 $LP=LF$ only if the data is a superposition of NP linear events and $noise=0$
- In general, events in the data are approximated by linear events. In addition, data are contaminated with random noise, then we choose $LF > NP$
- Similar to the deconvolution problem, we can also add a stability factor:

$$\hat{\mathbf{a}} = (\mathbf{M}^H \mathbf{M} + m\mathbf{I})^{-1} \mathbf{M}^H \mathbf{s}$$

A Model for FX data (Trade-off parameters)

- LF and M control the degree of noise that is modeled
- Long operators will attempt to model the signal and the noise and therefore, the noise estimate will be contaminated by the signal
- Short filters will only partially model the noise
- FX DECON needs small overlapping windows.
Why?.. Because assumes windows with constant dip data..... and a small number of dips. OK!!

Why the name F-X Deconvolution ?

$$S_j = a_1 S_{j-1} + a_2 S_{j-2} + a_3 S_{j-3} + \square + a_{LF} S_{j-LF} + N_j$$

or

$$S_j - a_1 S_{j-1} - a_2 S_{j-2} - a_3 S_{j-3} - \square - a_{LF} S_{j-LF} = N_j \quad \left| \begin{array}{l} f_0 = 1 \\ f_k = -a_k, k = 1, \dots, LF \end{array} \right.$$

▷

$$f_0 S_j + f_1 S_{j-1} + f_2 S_{j-2} + f_3 S_{j-3} + \square + f_{LF} S_{j-LF} = N_j$$

$$f * S = N$$

f predicts white noise - in spiking deconvolution f also
predicts white noise (the reflectivity)

Prediction error filter - A wave number view

S : spatial signal with one or more dominant wavenumber k

N : spatial white noise (same energy at all wavenumbers)

$$f * S = N \vdash$$

f : Will have to filter the dominant signals and let the noise pass
(Notch filter?)

Fxy Filters

- 2D Kernels - same idea but not clear to what kind of signal model one requires to justify 2D prediction error filters
- 1D Prediction can be performed with Forward or backward operators or a combination of both

Predicting from the left :

$$S_j = a_1 S_{j-1} + a_2 S_{j-2} + a_3 S_{j-3} + \square + a_{LF} S_{j-LF} + N_j$$



Predicting from the right :

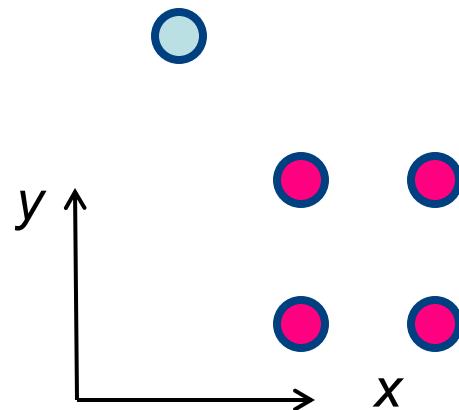
$$S_j = b_1 S_{j+1} + b_2 S_{j+2} + b_3 S_{j+3} + \square + b_{LF} S_{j+LF} + N_j$$



Fxy Filters

2D Prediction:

$$S_{ij} = a_{11}S_{i-1j-1} + a_{12}S_{i-1j-2} + a_{12}S_{i-2j-1} + a_{22}S_{i-2j-2} + \square + N_{ij}$$



Solution: could use Block Matrices and Block Solvers or operate on the flight with the Method of Conjugate Gradients

SSA Singular Spectrum Analysis or Cadzow Filter

- SSA: an alternative to FX deconvolution ?
- **++**
 - Easy to code and to understand
 - Easy to extend to ND
 - Possibility of creating noise attenuation methods that simultaneously operate in frequency and space
- **-+**
 - Memory?

Note:

SSA arises from people working in dynamical systems as a method for rank reduction of 1D time series. Cadzow decomposition is a more general approach to image/signal enhancement that leads to an algorithm very similar to SSA when used to enhance complex sinusoids.

References

- S. Trickett, 2008, F-xy Cadzow noise suppression, CSPG CSEG Convention
- Broomhead, D.S., and King, G., 1986: Extracting qualitative dynamics from experimental data, *Physica D*, **20**, 217-236.
- Vautard, R., and Ghil, M., 1989: Singular spectrum analysis in nonlinear dynamics, with applications to paleoclimatic time series, *Physica D*, **35**, 395-424.
- V Oropeza and M D Sacchi, 2011, Simultaneous seismic data de-noising and reconstruction via Multichannel Singular Spectrum Analysis (MSSA), *Geophysics*, **76**, V25-V32.

Theory (an easy tour of SSA)

Assume we are in FX working with data in x for fixed f

$$\mathbf{M} = \begin{matrix} \ddot{x} & S_1 & S_2 & S_3 & S_4 & \ddot{0} \\ \dot{\zeta} & & & & & \div \\ \zeta & S_2 & S_3 & S_4 & S_5 & \div \\ \dot{\zeta} & S_3 & S_4 & S_5 & S_6 & \div \\ \ddot{\zeta} & & & & & \div \\ \ddot{e} & S_4 & S_5 & S_6 & S_7 & \emptyset \end{matrix}$$

Hankel Matrix of a signal of length 7 with a view of length M=4 (Trajectory Matrix)

$$\mathbf{M} = \begin{matrix} & \ddot{\alpha} s_1 & s_2 & s_3 & s_4 & \ddot{o} \\ \ddot{\alpha} s_1 & & s_2 & s_3 & s_4 & \vdots \\ \zeta s_2 & & s_3 & s_4 & s_5 & \vdots \\ \zeta s_3 & & s_4 & s_5 & s_6 & \vdots \\ \zeta s_4 & & s_5 & s_6 & s_7 & \emptyset \end{matrix}$$

*Hankel Matrix of a signal of length 7 with a view of length
 $M=4$ (Trajectory Matrix)*

Average Along anti-diagonals

$$A(\mathbf{M}) = \frac{\partial S_1}{\partial} \div \frac{\partial S_2}{\partial} \div \frac{\partial S_3}{\partial} \div \frac{\partial S_4}{\partial} \div \frac{\partial S_5}{\partial} \div \frac{\partial S_6}{\partial} \div \frac{\partial S_7}{\partial} \emptyset$$

Linear event in TX = Complex Sinusoid in FX

$$s(t, x) = w(t - px)$$

$$S(w, x) = W(w)e^{-iwx}$$

$$S(\omega, x) = W(\omega) e^{-i\omega p x}$$

$$x = n \Delta x$$

$$S_n = W e^{-i\alpha n}$$

$$S_n = P S_{n-1}$$

Prediction operator of length 1 is able to propagate signal from channel $n-1$ to channel n .

Single linear event in FX

$$S_n = PS_{n-1}$$

$$\mathbf{M} = \begin{array}{cccccc} \mathbb{A} & S_1 & S_2 & S_3 & S_4 & \emptyset \\ \zeta & S_2 & S_3 & S_4 & S_5 & \div \\ \zeta & S_3 & S_4 & S_5 & S_6 & \div \\ \mathbb{E} & S_4 & S_5 & S_6 & S_7 & \emptyset \end{array}$$

Single linear event in $\mathcal{F}X$

$$S_n = PS_{n-1}$$

$$\mathbf{M} = \begin{matrix} \ddot{\alpha} s_1 & s_2 & s_3 & s_4 \emptyset \\ \dot{\zeta} s_2 & s_3 & s_4 & s_5 \div \\ \dot{\zeta} s_3 & s_4 & s_5 & s_6 \div \\ \dot{\epsilon} s_4 & s_5 & s_6 & s_7 \emptyset \end{matrix} = \begin{matrix} \ddot{\alpha} s_1 & Ps_1 & P^2 s_1 & P^3 s_1 \emptyset \\ \dot{\zeta} s_2 & Ps_2 & P^2 s_2 & P^3 s_2 \div \\ \dot{\zeta} s_3 & Ps_3 & P^2 s_3 & P^3 s_3 \div \\ \dot{\epsilon} s_4 & Ps_4 & P^2 s_4 & P^3 s_4 \emptyset \end{matrix}$$

\mathbf{M} is a Hankel matrix of $Rank = 1$

p linear events

- It can be shown that for a superposition of p complex sinusoids (p events of different slope in FX):

\mathbf{M} is a Hankel matrix of $\text{Rank} = p$

- We will use rank reduction to attenuate random noise
- The spectra of singular values can be used to determine the number of complex sinusoids (**plane waves**) in the window of analysis

SSA FX algorithm = Rank Reduction of M

$$s(t, x) \rightarrow s(\omega, x) \rightarrow \mathbf{s}(\omega)$$

Select p

Select $M(view)$

For $\omega = \omega_{\min}, \omega_{\max}$

$$\mathbf{M} = \text{Hankel}[\mathbf{s}(\omega)]$$

$$\mathbf{M} = \mathbf{U} \mathbf{S} \mathbf{V}^H \quad \text{SVD}$$

$$\mathbf{M}_p = \mathbf{U}_p \mathbf{U}_p^H \mathbf{M} \quad \text{SVD - Filtering to reduce rank}$$

$$\mathbf{s}_p(\omega) = A(\mathbf{M}_p) \quad \text{Sum along anti-diags}$$

End

$$\mathbf{s}_p(\omega) \rightarrow s_p(\omega, x) \rightarrow s_p(t, x)$$

The Cazdow/SSA filter

- Hankelization Operator
- Rank Reduction Operator

$$\begin{aligned}\mathbf{M}_p &= \mathbf{U}_p \mathbf{U}_p^H \mathbf{M} \\ &= R_p[\mathbf{M}]\end{aligned}$$

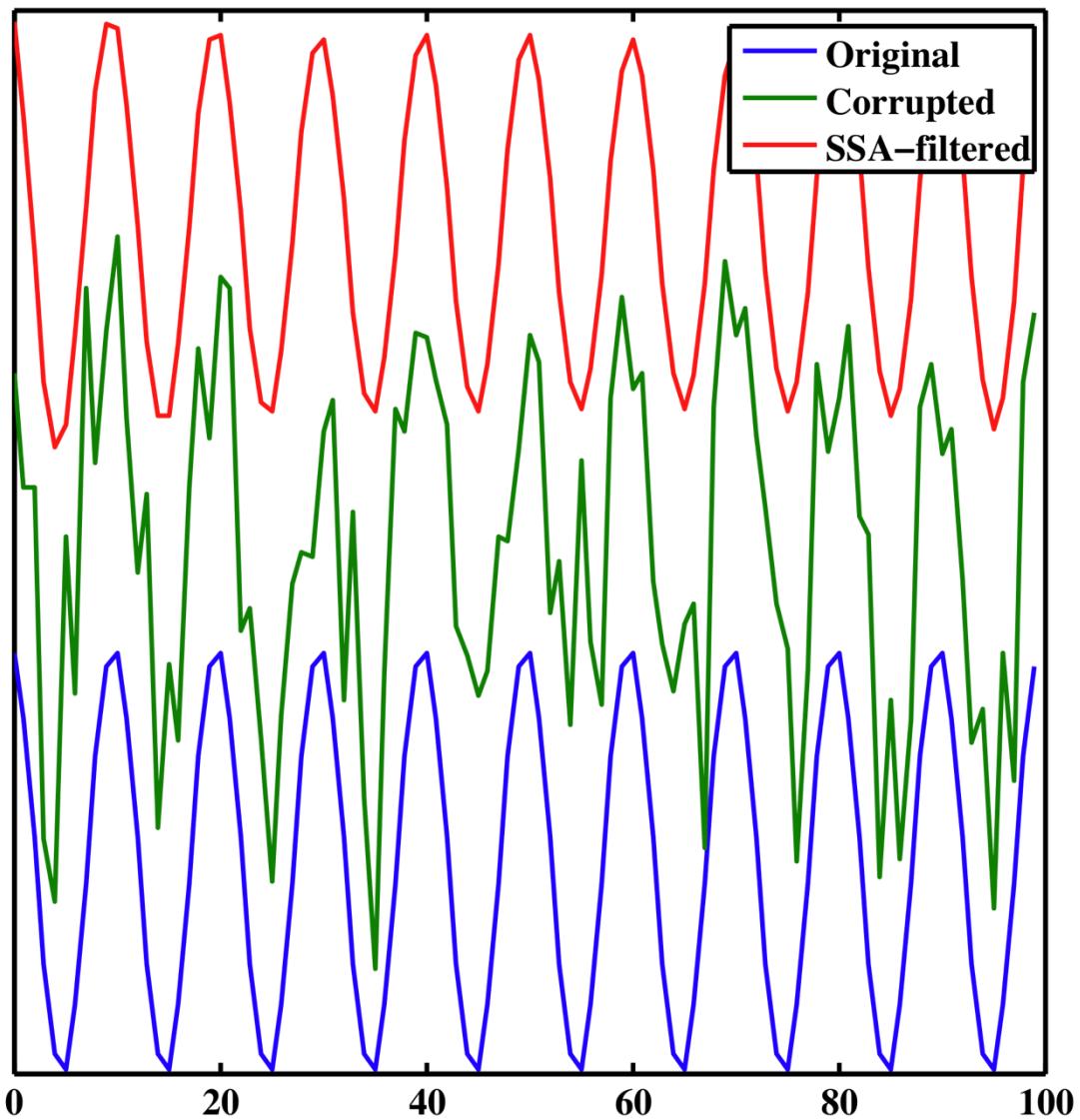
- Anti-diagonal Averaging Operator

$$\mathbf{s} = H[\mathbf{M}]$$

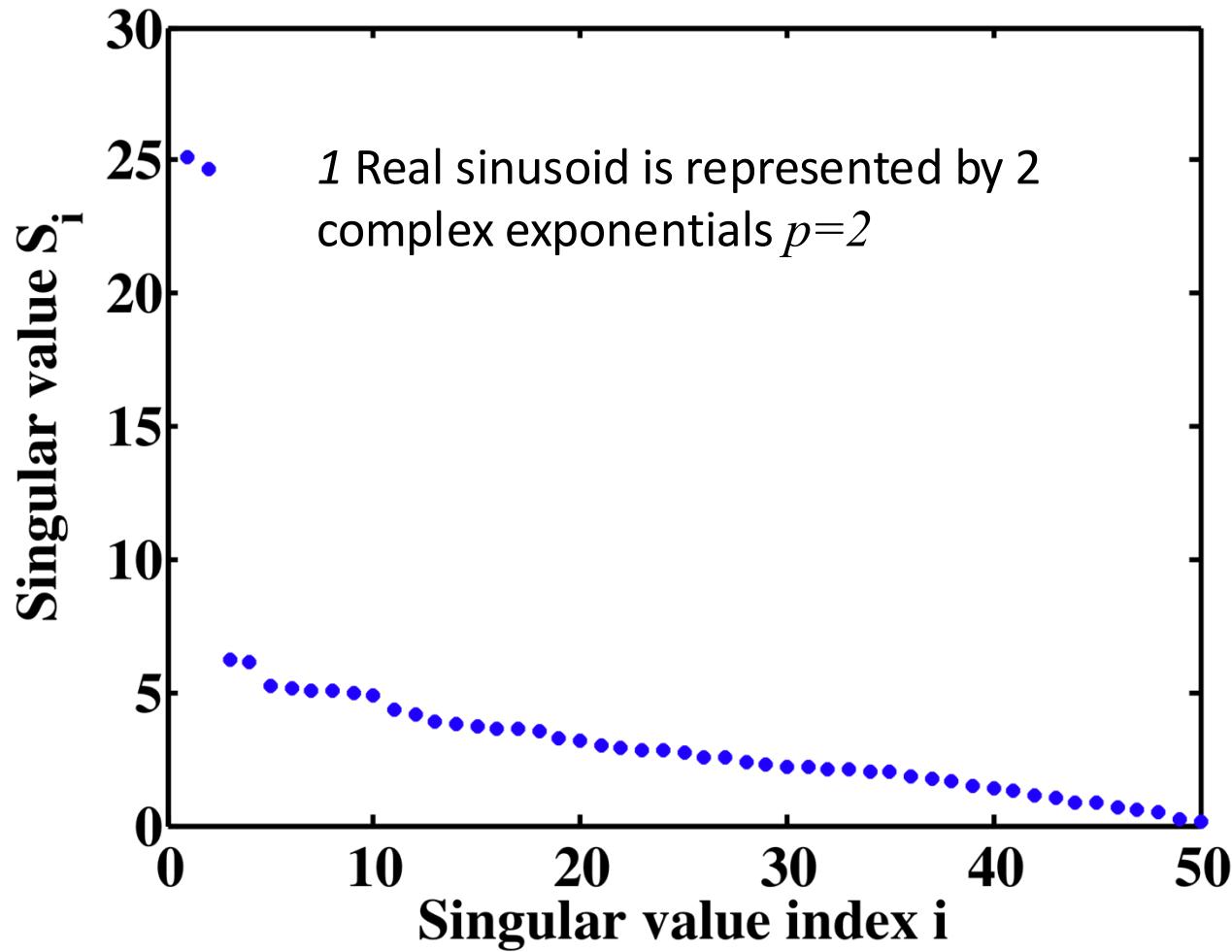
- Cazdow/SSA Filter

$$\hat{\mathbf{s}} = A[R_p[H[s]]]$$

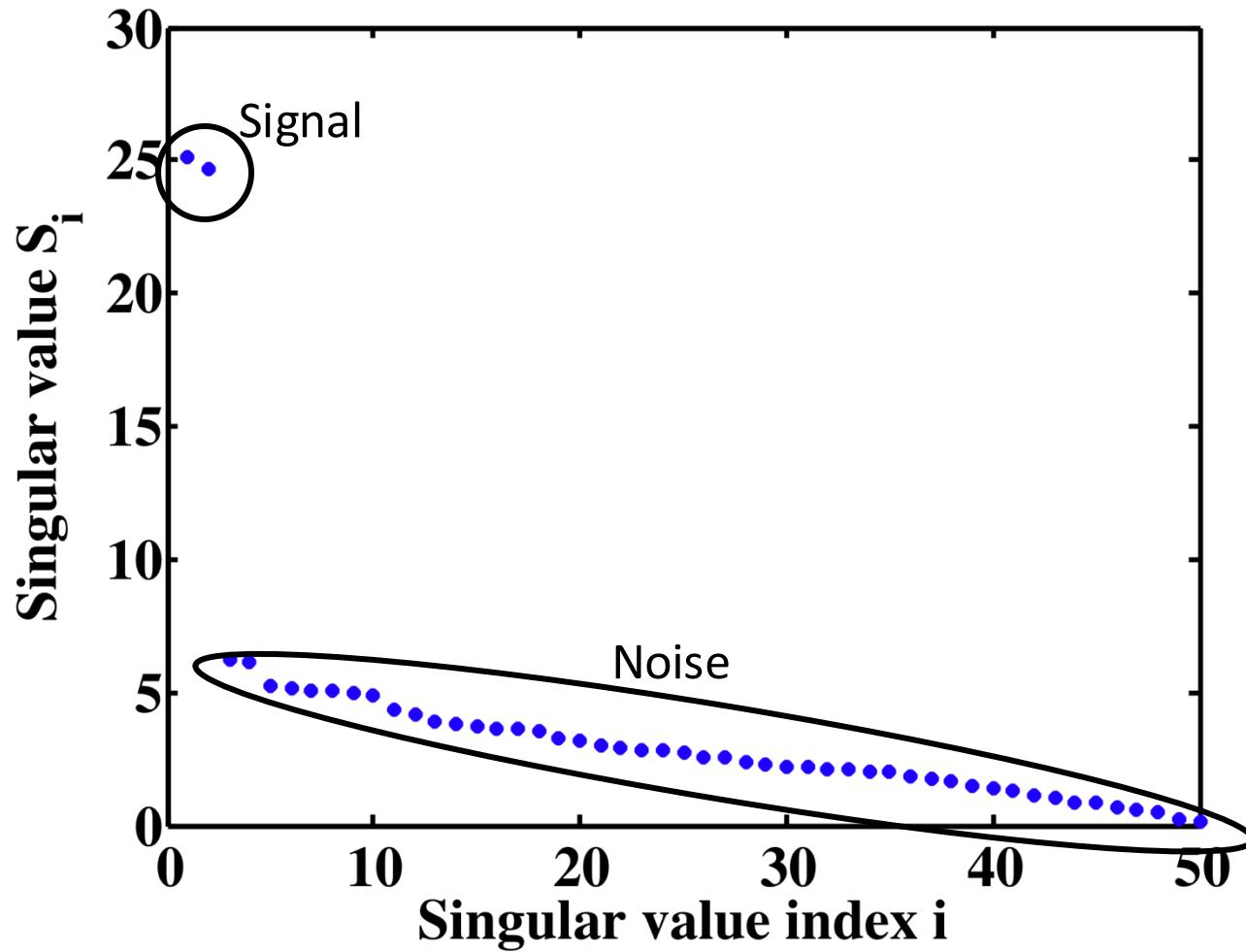
1D Example



1D example



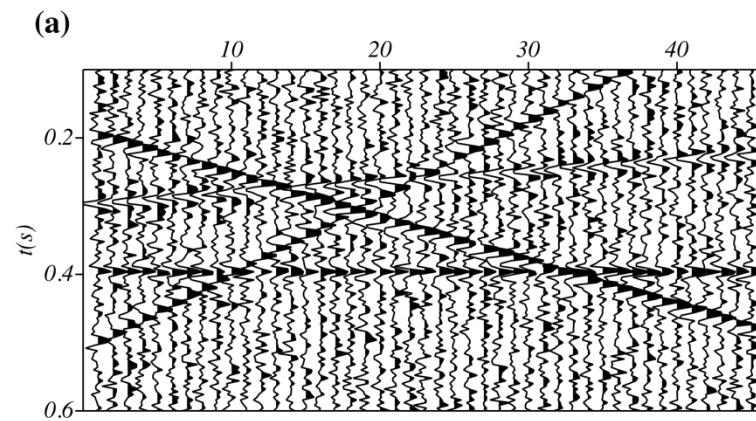
1D example



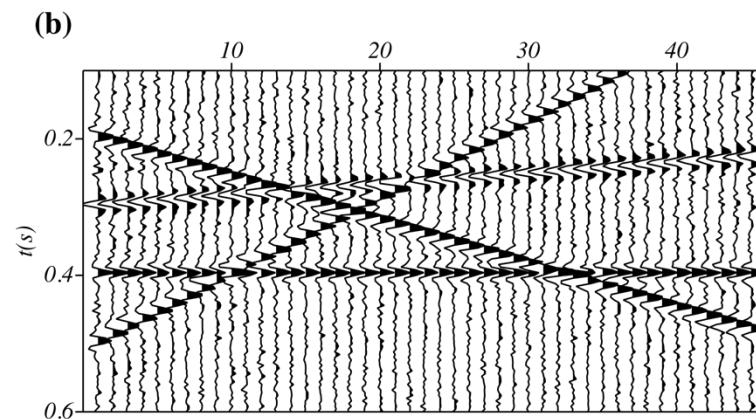
2D example (1D in FX)

- M (*view*) is selected to make \mathbf{M} (*Hankel Matrix*) close to a square matrix
- Rest of the algorithm is very similar to FX - decon:
 - Decide min-max frequencies for noise attenuation
 - Window patching/unpatching
 - Adjust filter length (p in ssa) by monitoring residual panel
 - Use frequency domain symmetries to save time

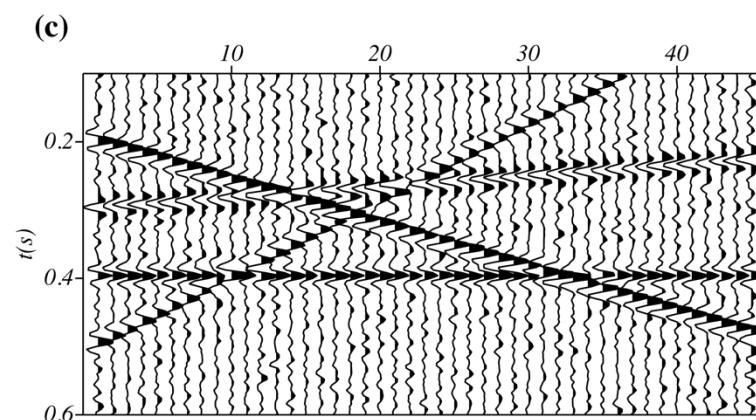
Data



FX SSA

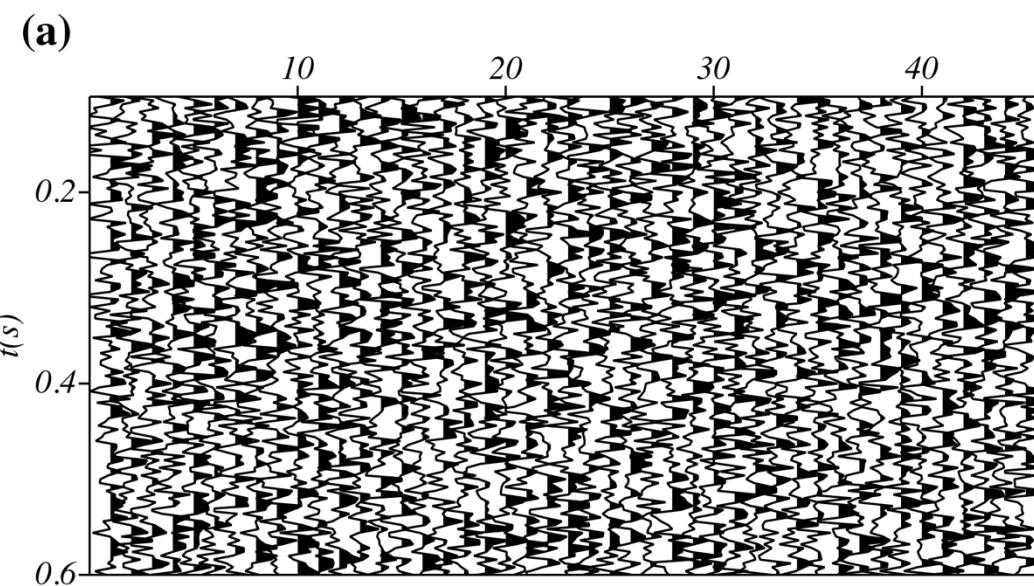


FX Decon

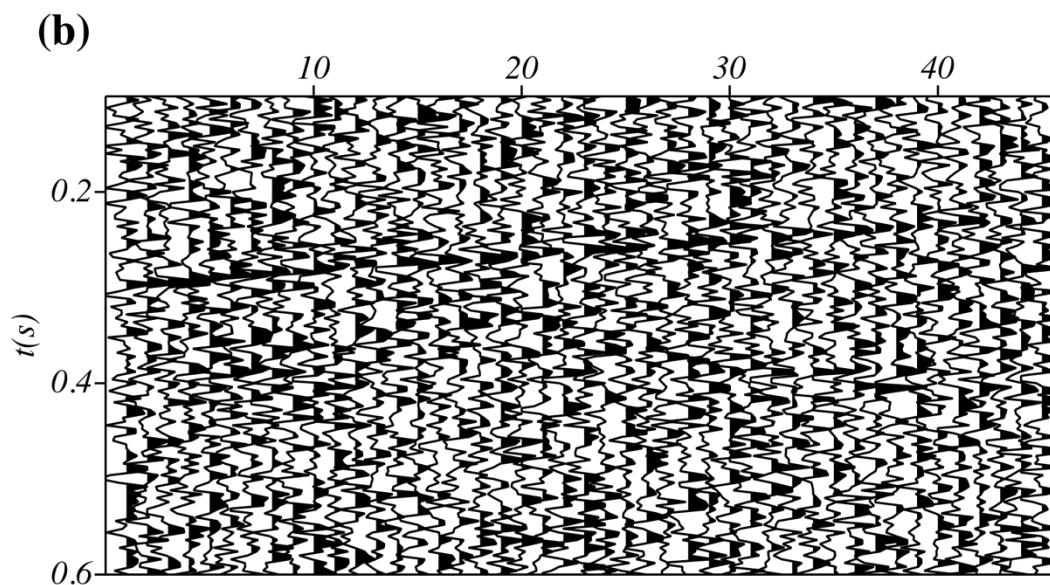


Error Panels

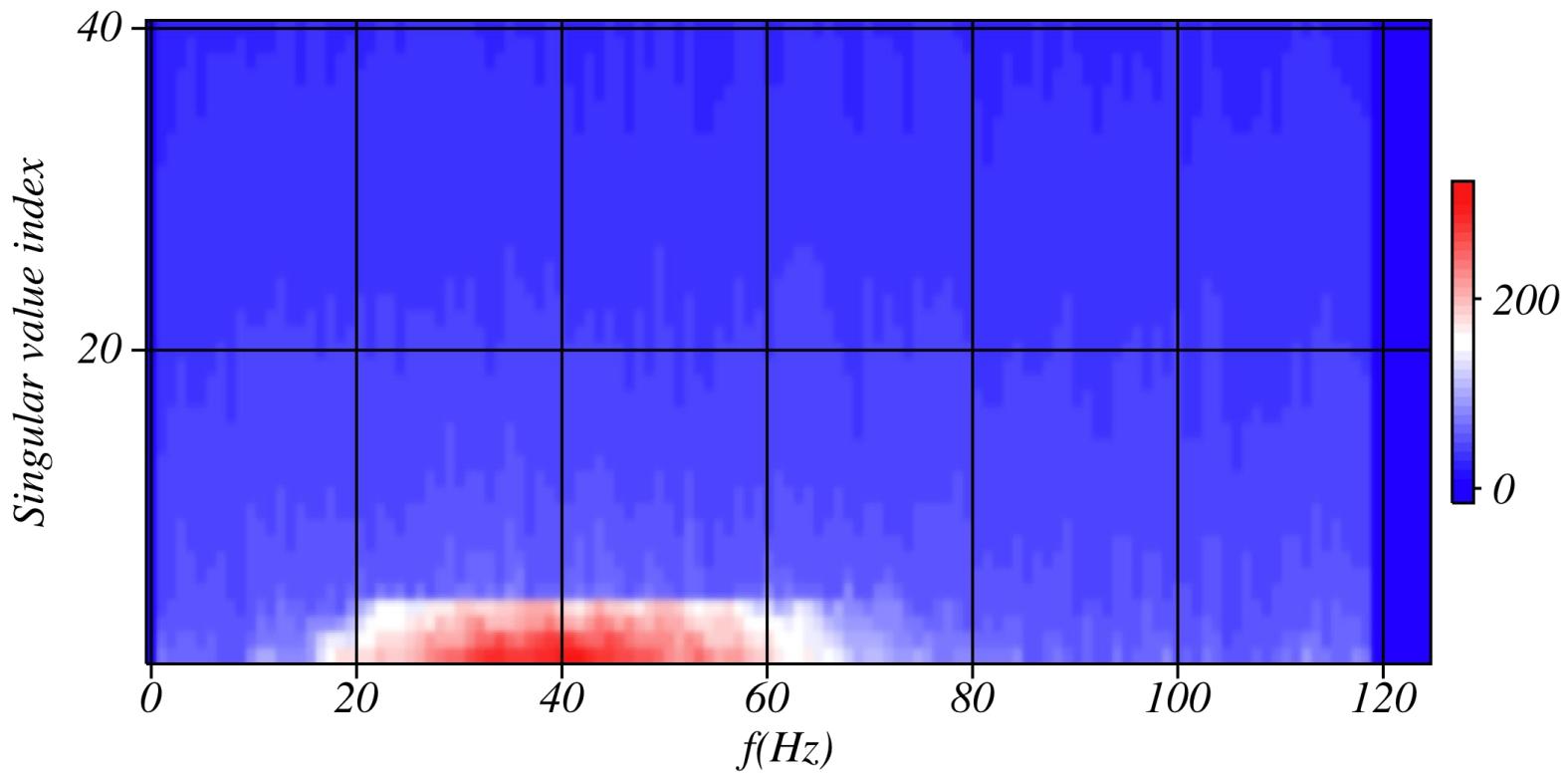
FX-
SSA



FX-Decon

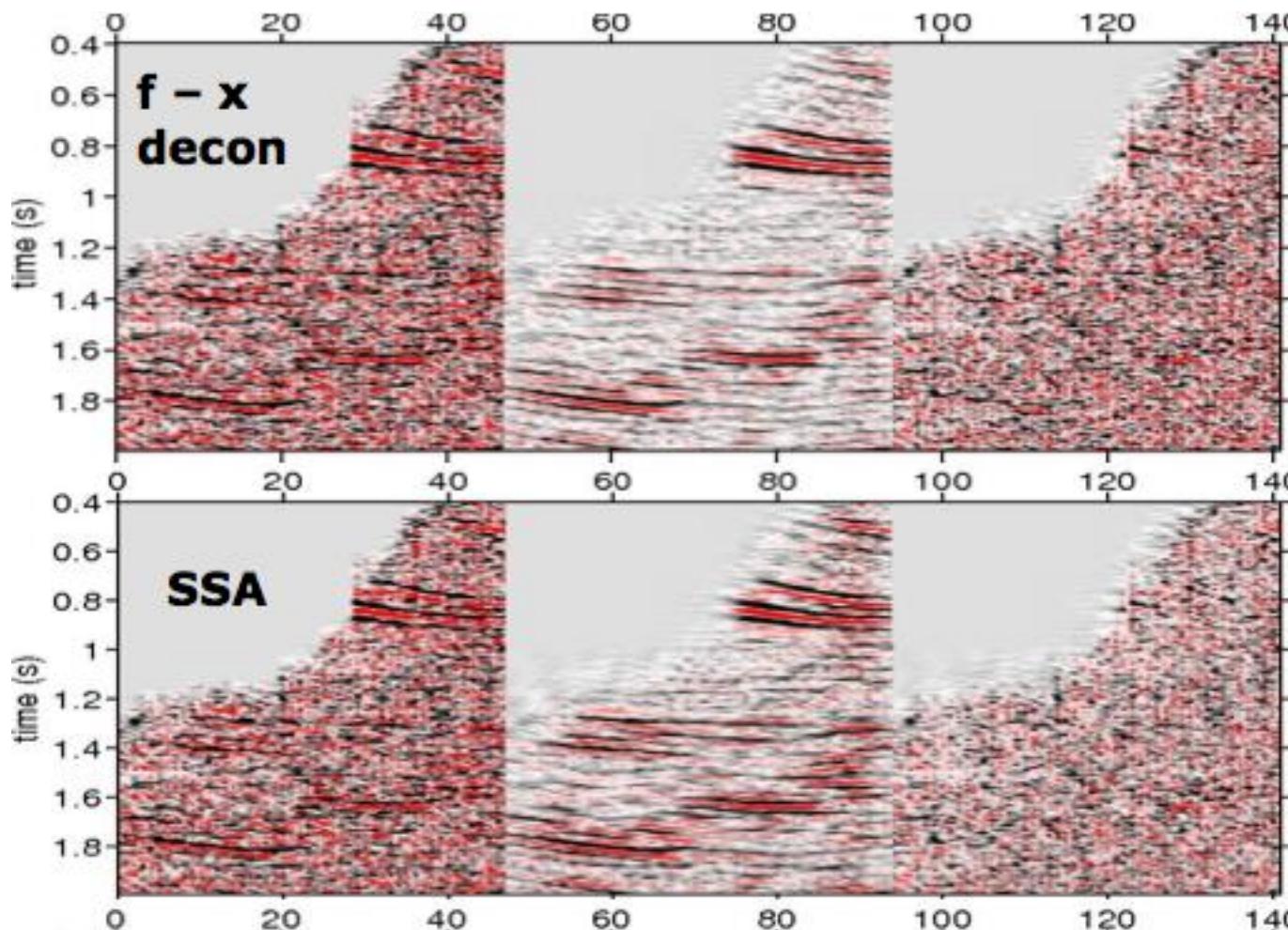


Spectrum of singular values vs. freq.

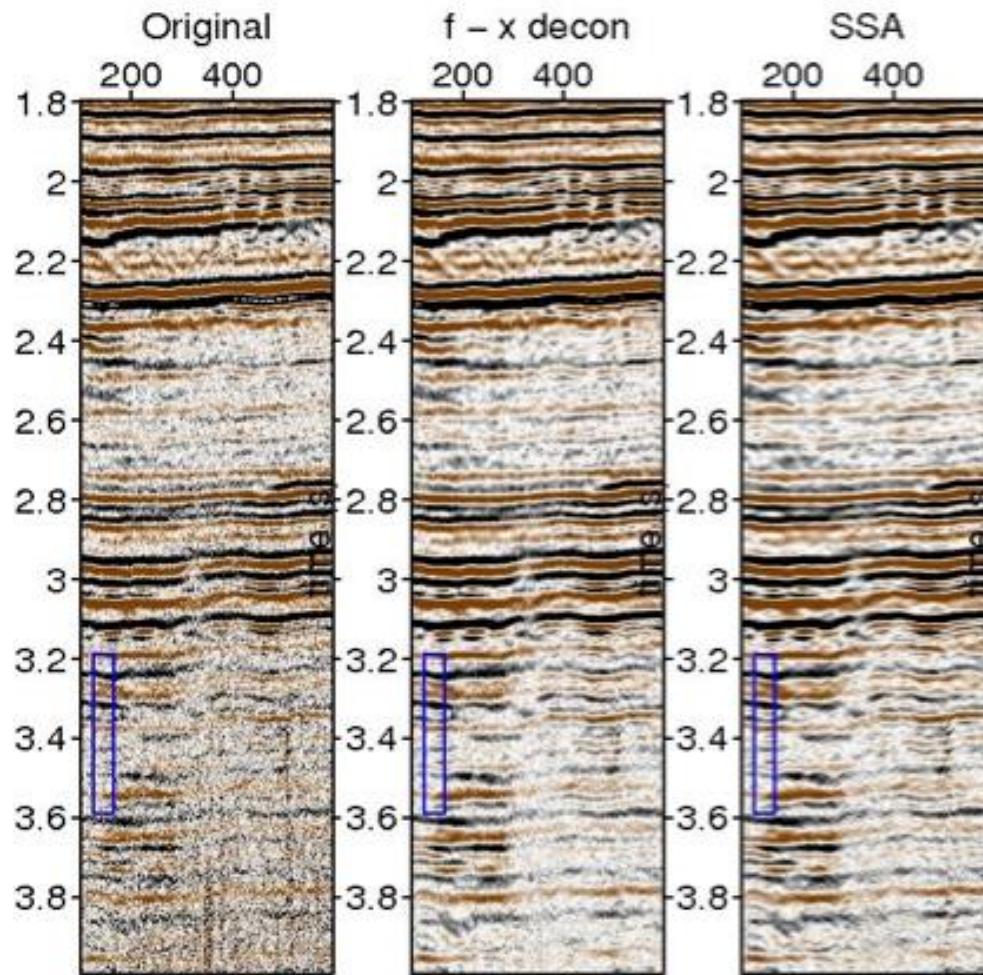


Spectra cut-off is at 4 singular values

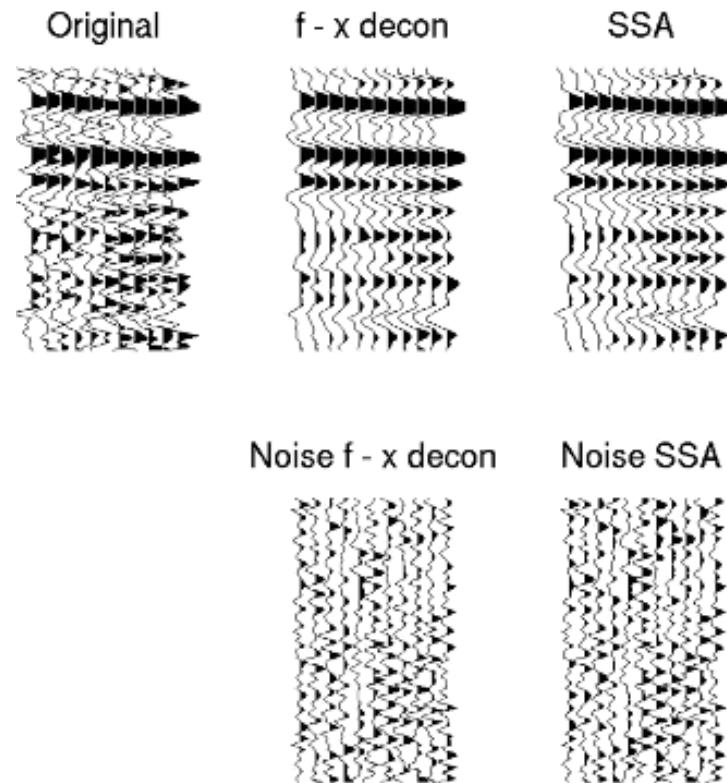
Marmousi gathers.



Real data comparison (Processing by V. Oropeza)

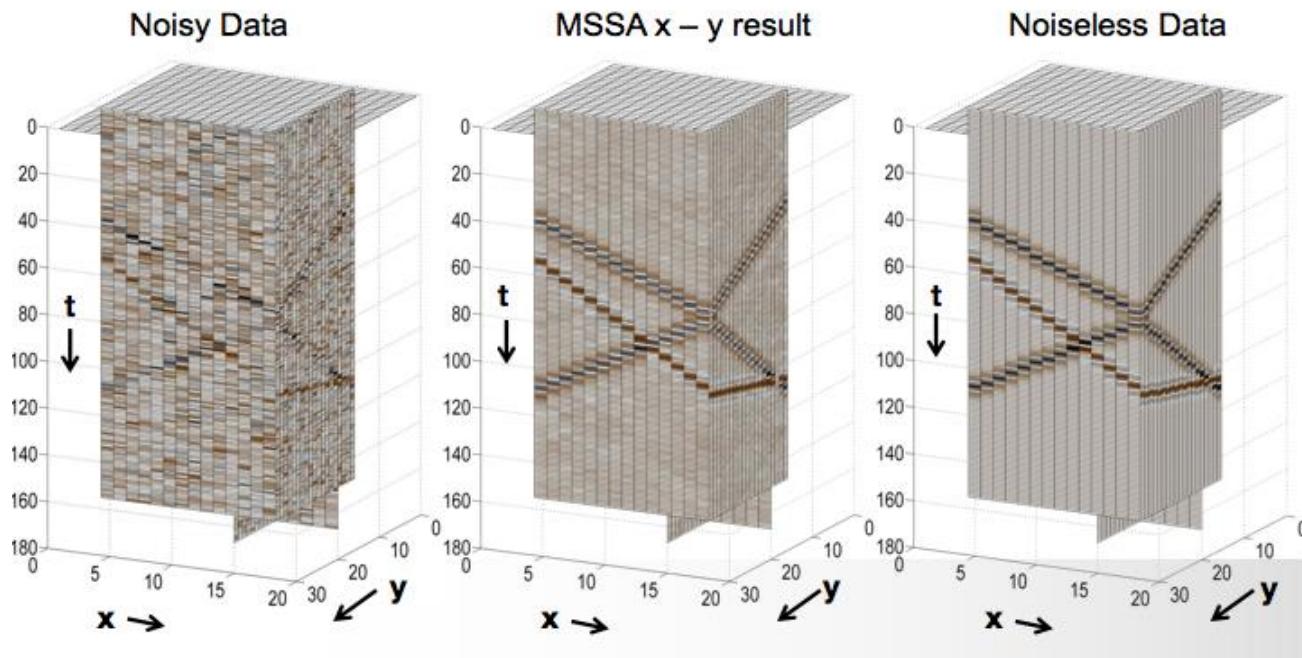


Comparison



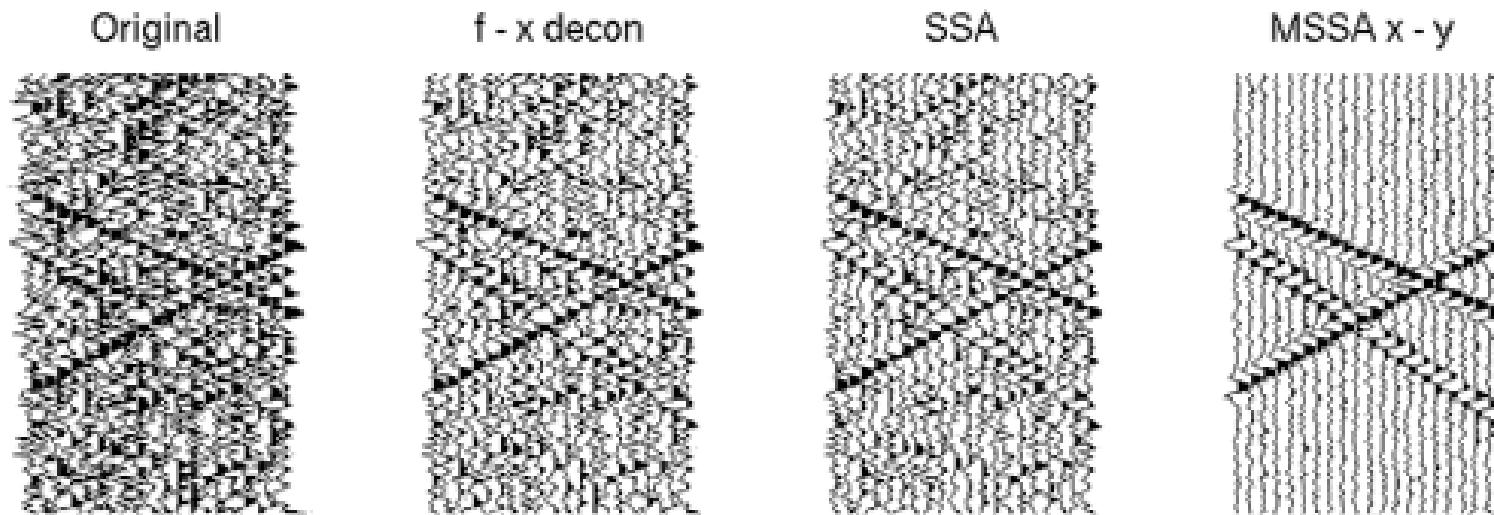
SSA for 3D cubes

MSSA X – Y in 3D data



SSA works better in $x-y-f$ than $x-f$ (more data)

MSSA – XY in 3D data using 6 Singular Values



Notes:

- RSVD (randomized SVD) can be used to speed up the SVD
- For data reconstruction we used fast Lanczos solvers that exploits the structure of the Hankel matrix for fast rank reduction
- These tricks are also used for data reconstruction

Gao, J., Sacchi, M., and Chen, X. (2013). "A fast reduced-rank interpolation method for prestack seismic volumes that depend on four spatial dimensions." *GEOPHYSICS*, 78(1), V21–V30.