

parallel_potential

January 21, 2020

1 Curtain Parallel Potential Model

Here I explore the minimum parallel electric potential necessary to lower the mirror point of an electron from 100 km in the SAA (just trapped), to AC6's altitude in the bounce loss cone.

Further, we assume this potential is at the equator, so it will modify the electron's equatorial pitch angle.

We will first pick a curtain observation and use IRBEM to find the minimum altitude and pitch angle change those electrons experienced such that they were observed by AC6 for a prolonged period.

```
In [1]: import pandas as pd
import numpy as np
import matplotlib.pyplot as plt
from datetime import datetime
import scipy.interpolate #.interp1d

import IRBEM
Re_km = 6_371
```

```
In [2]: t_0 = datetime(2015, 8, 27, 23, 44, 500000)
```

1.0.1 Load catalog file to look up AC6 location

```
In [3]: cat_path = '/home/mike/research/ac6_curtains/data/catalogs/AC6_curtains_sorted_v8.txt'
cat = pd.read_csv(cat_path, index_col=0)
cat.index = pd.to_datetime(cat.index)
cat.head()
```

```
Out [3]:
```

	dos1rate	peak_std	Lm_OPQ	MLT_OPQ	lat	\
dateTime						
2014-12-19 11:45:30.500000	290.002	2.840199	6.66789	10.71360	62.7607	
2014-12-19 18:17:11.099999	340.002	14.832431	7.24842	7.98180	63.5059	
2014-12-19 18:17:12.300000	340.002	11.704743	7.27972	7.97519	63.5759	
2014-12-19 21:33:52.899999	640.003	14.352735	6.97210	7.45659	70.3503	
2015-03-26 07:31:44.500000	170.001	1.455209	7.08831	10.39770	69.2889	
	lon	alt	Dist_In_Track	Lag_In_Track	\	

dateTime						
2014-12-19 11:45:30.500000	-35.6930	662.260	461.176	61.2438		
2014-12-19 18:17:11.099999	-134.1340	663.739	463.721	61.5946		
2014-12-19 18:17:12.300000	-134.1890	663.824	463.718	61.5946		
2014-12-19 21:33:52.899999	170.0450	670.690	464.995	61.7903		
2015-03-26 07:31:44.500000	21.2179	632.797	243.682	32.2168		

	Dist_Total	Loss_Cone_Type	flag	AE	time_cc	\
dateTime						
2014-12-19 11:45:30.500000	461.422	0.0	4.0	51.0	0.094868	
2014-12-19 18:17:11.099999	463.972	1.0	0.0	405.0	0.288474	
2014-12-19 18:17:12.300000	463.968	1.0	0.0	405.0	0.090012	
2014-12-19 21:33:52.899999	465.247	1.0	20.0	99.0	0.536131	
2015-03-26 07:31:44.500000	243.716	1.0	0.0	201.0	0.261325	

	space_cc	time_spatial_A	\
dateTime			
2014-12-19 11:45:30.500000	0.870388	2014-12-19 11:44:29.256200	
2014-12-19 18:17:11.099999	0.892241	2014-12-19 18:16:09.505399	
2014-12-19 18:17:12.300000	0.895632	2014-12-19 18:16:10.705400	
2014-12-19 21:33:52.899999	0.841011	2014-12-19 21:33:52.899999	
2015-03-26 07:31:44.500000	0.888005	2015-03-26 07:31:12.283200	

	time_spatial_B	peak_width_A	\
dateTime			
2014-12-19 11:45:30.500000	2014-12-19 11:45:30.500000	NaN	
2014-12-19 18:17:11.099999	2014-12-19 18:17:11.099999	NaN	
2014-12-19 18:17:12.300000	2014-12-19 18:17:12.300000	NaN	
2014-12-19 21:33:52.899999	2014-12-19 21:34:54.690299	0.565442	
2015-03-26 07:31:44.500000	2015-03-26 07:31:44.500000	0.057192	

	peak_width_B
dateTime	
2014-12-19 11:45:30.500000	0.618749
2014-12-19 18:17:11.099999	0.553333
2014-12-19 18:17:12.300000	0.486890
2014-12-19 21:33:52.899999	0.957143
2015-03-26 07:31:44.500000	0.620818

```
In [4]: curtain_obs_params = cat.loc[t_0]
curtain_obs_params
```

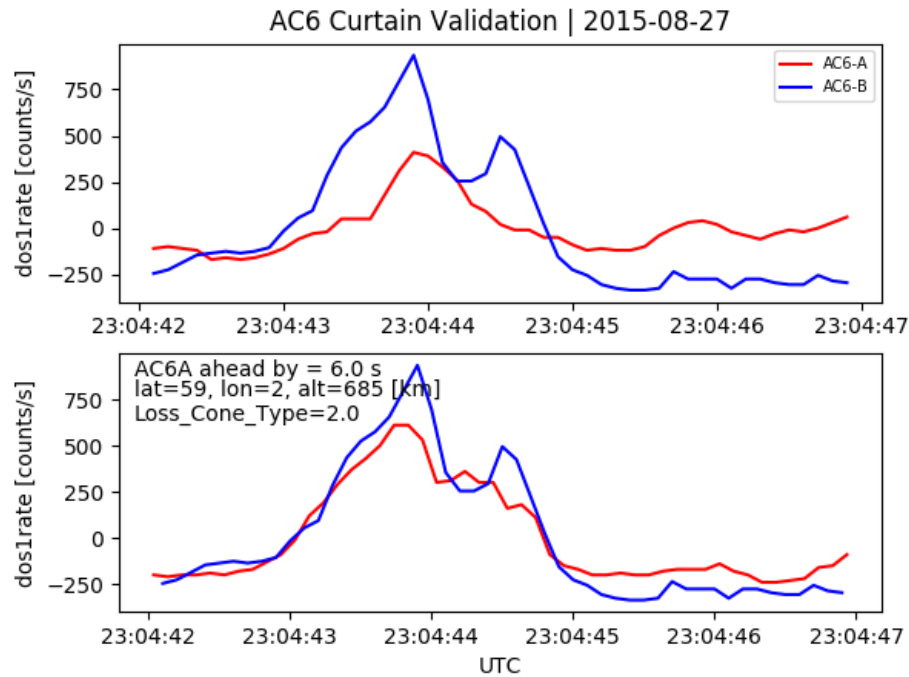
```
Out[4]: dosirate      890.005
peak_std      8.37401
Lm_OPQ      3.83808
MLT_OPQ      0.411383
lat      59.3801
lon      2.14165
```

```

alt                                685.152
Dist_In_Track                      45.3006
Lag_In_Track                       6.03549
Dist_Total                         45.322
Loss_Cone_Type                     2
flag                               0
AE                                 745
time_cc                            0.547989
space_cc                           0.852933
time_spatial_A    2015-08-27 23:04:38.464510
time_spatial_B    2015-08-27 23:04:44.500000
peak_width_A                               NaN
peak_width_B                               0.243307
Name: 2015-08-27 23:04:44.500000, dtype: object

```

From a single spacecraft this looks like two superposed microbursts



1.0.2 BLC Sanity Check

Southern Hemisphere Mirror Point For Locally Mirroring Electrons at AC6

```

In [5]: model = IRBEM.MagFields(kext='OPQ77')
        X = {'dateTime':t_0, 'x1':cat.loc[t_0, 'alt'], 'x2':cat.loc[t_0, 'lat'], 'x3':cat.loc[t_0, 'lon']}

        try:
            model.mirror_point_altitude(X, None)
            print(model.mirrorAlt)
        except ValueError as err:

```

```

if str(err) == 'Mirror point below the ground!':
    print(err)
else:
    raise

```

Mirror point below the ground!

Looks like the mirror point in the southern hemisphere is below the ground - I am very confident that these electrons must have been lost within one bounce.

1.0.3 Mirror points for trapped particles

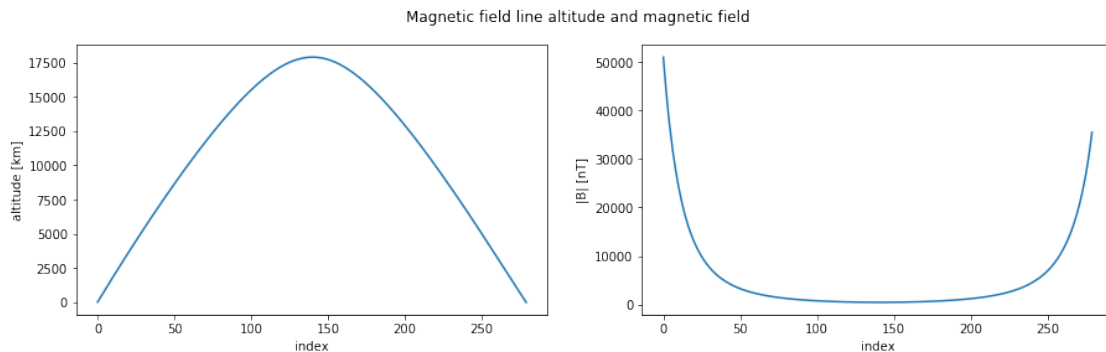
Trace the field line connected to AC6 and find the mirror point altitude above AC6 of barely trapped electrons - defined as electrons with a 100 km mirror point altitude in the SAA.

```

In [6]: output_dictionary = model.trace_field_line(X, None) # Trace the magnetic field line.
        field_line_dict = {'x':output_dictionary['POSIT'][:,0], 'y':output_dictionary['POSIT']
                           'z':output_dictionary['POSIT'][:,2], 'blocal':output_dictionary['blocal'],
                           'alt':Re_km*(np.linalg.norm(output_dictionary['POSIT'], axis=1)-1)}
        field_line_df = pd.DataFrame(data=field_line_dict)
        interp_index = np.linspace(0, len(output_dictionary['POSIT'][:,0])-1, num=10000)

In [7]: _ , ax = plt.subplots(1, 2, figsize=(15, 4))
        ax[0].plot(field_line_df.alt); ax[0].set(xlabel='index', ylabel='altitude [km]')
        ax[1].plot(field_line_df.blocal); ax[1].set(xlabel='index', ylabel='|B| [nT]')
        plt.suptitle('Magnetic field line altitude and magnetic field');

```



Interpolate the field line values

```

In [8]: def interp_df(df, new_index):
        """
        Return a new DataFrame with all columns values interpolated
        to the new_index values.
        """
        df_out = pd.DataFrame(index=new_index)

```

```

df_out.index.name = df.index.name

for colname, col in df.iteritems():
    df_out[colname] = np.interp(new_index, df.index, col, right=np.nan)

return df_out

field_line_interp = interp_df(field_line_df, interp_index)

```

1.0.4 Find the magnetic field strength at 100 km in the southern hemisphere and estimate by how much the electron's mirror point must have lowered

```

In [9]: southern_mp_trapped = field_line_interp[field_line_interp.alt > 100].iloc[-1]
        southern_mp_trapped

```

```

Out [9]: x          0.455764
        y          0.309929
        z         -0.853187
        blocal    33716.817123
        alt       100.354565
        Name: 278.3024302430243, dtype: float64

```

Now use the local b field at 100 km altitude in the SAA to find the altitude in the northern hemisphere that has the same magnetic field strength

```

In [10]: northern_mp_trapped = field_line_interp[field_line_interp.blocal < southern_mp_trapped.blocal]
        northern_mp_trapped

```

```

Out [10]: x          0.603544
         y          0.023170
         z          0.982677
         blocal    33654.932946
         alt       977.715098
         Name: 5.3294329432943295, dtype: float64

```

Now find out the difference in altitude - by how much the electron's mirror point needed to lower in one bounce and be observed by AC6

```

In [11]: ac6_blocal = field_line_interp[field_line_interp.alt > curtain_obs_params.alt].iloc[0]
        print(f"The electron's mirror point must have decreased by at least "
              f"{round(northern_mp_trapped.alt - curtain_obs_params.alt)} km "
              f"into a {round(ac6_blocal - northern_mp_trapped.blocal)} nT stronger field strength")

```

The electron's mirror point must have decreased by at least 293.0 km into a 4283.0 nT stronger field strength

1.0.5 Estimate the change in pitch angle from the trapped to the precipitating particle.

First find the equatorial magnetic field strength

```
In [12]: b_equator = model.find_magequator(X, None)['bmin']
         print(f'Equatorial magnetic field strength = {round(b_equator)} nT')
```

Equatorial magnetic field strength = 478 nT

The equatorial pitch angle is defined as

$$\alpha_{eq} = \sin^{-1} \left(\sqrt{\frac{B_{eq}}{B_m}} \right)$$

```
In [13]: alpha_initial = np.arcsin(np.sqrt(b_equator/southern_mp_trapped['blocal']))
         alpha_final = np.arcsin(np.sqrt(b_equator/ac6_blocal))
         d_alpha = alpha_initial - alpha_final

         print(f"Initial pitch angle = {round(np.rad2deg(alpha_initial), 2)} degrees")
         print(f"Final pitch angle = {round(np.rad2deg(alpha_final), 2)} degrees")
         print(f"Pitch angle change = {round(np.rad2deg(d_alpha), 2)} degrees")
```

Initial pitch angle = 6.84 degrees

Final pitch angle = 6.45 degrees

Pitch angle change = 0.39 degrees

Now find the parallel potential (I worked out the math on paper).

```
In [14]: potential = 35E3 * np.cos(alpha_initial) * ((np.tan(alpha_initial)/np.tan(alpha_final))
         print(f'Minimum electric potential is {round(potential)} Volts')
```

Minimum electric potential is 4413.0 Volts

2 Misc

2.0.1 35 keV electron bounce period in the area near the BLC

Calculate the bounce period for locally mirroring, 35 keV electrons near the BLC region. This is only an approximation to see if the bounce period is anywhere near the 4-8 second AC6 in-track separations when these curtains were observed.

```
In [15]: X = {'dateTime':t_0, 'x1':cat.loc[t_0, 'alt'], 'x2':cat.loc[t_0, 'lat'], 'x3':-30}
         model.bounce_period(X, None, 35)
         print(f'35 keV electron bounce period is {round(model.Tb, 2)} seconds')
```

35 keV electron bounce period is 1.53 seconds