

- All you need to remember : $\left\{ \begin{array}{l} \text{Snell's Law} \\ \text{Mirror Formula} \\ \text{Lens Formula} \end{array} \right.$
 - Extra : Matrix method for paraxial optics
 - Extra - Extra : Aberration (intro only)
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The formulae

- Reflection on spherical surface $\frac{1}{s} + \frac{1}{s'} = \frac{-2}{R}$
- Refraction on spherical surface $\frac{n_1}{s} + \frac{n_2}{s'} = \frac{n_2 - n_1}{R}$

Sign convention

Object / Image distance (s/s')

+ve = anything real
-ve = anything virtual

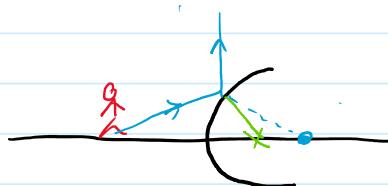
Radius of curvature (R) :

+ve = convex towards real object
-ve = concave towards real object
opposite sign for virtual object

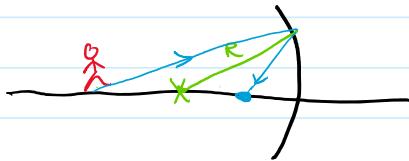
* This convention is subjectively better because

{convex/concave}, {real/virtual} is very commonly asked

Eg. 1 Mirrors



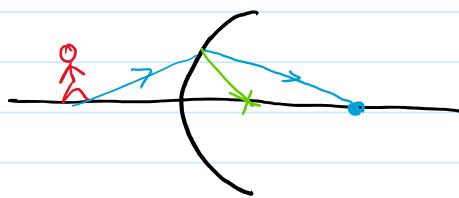
- ~~see~~ convex surface
 $\Rightarrow R = +ve$
- Image form behind mirror
 \Rightarrow Cannot be projected \equiv virtual
 $\Rightarrow s' = -ve$



- see concave surface
 $\Rightarrow R = -ve$

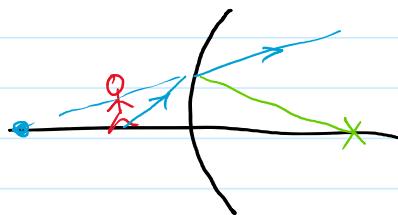
- Image form in front of mirror
 \Rightarrow Can be projected \equiv Real
 $\Rightarrow s' = +ve$

Eg. 2 Lens



- see convex surface
 $\Rightarrow R = +ve$

- Image form behind lens
 \Rightarrow Can be projected \equiv Real
 $\Rightarrow s' = +ve$



- see convex surface
 $\Rightarrow R = +ve$

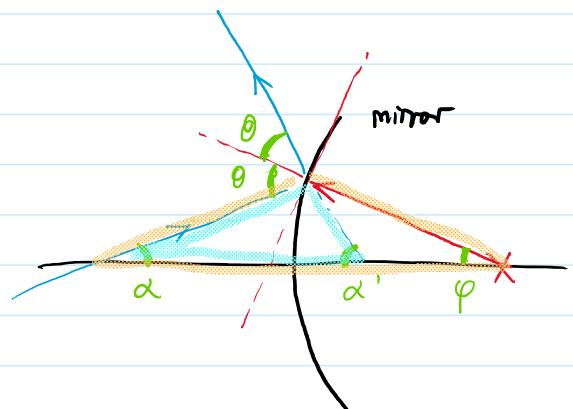
- Image form in front of lens
 \Rightarrow Cannot be projected \equiv Virtual
 $\Rightarrow s' = -ve$

Derivation

Reflection on spherical mirror

II By exterior L of Δ

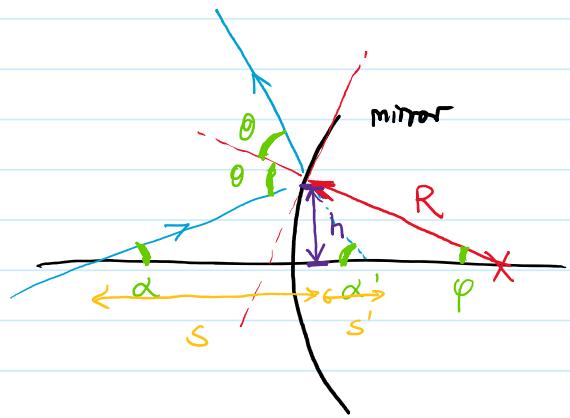
$$\left\{ \begin{array}{l} \alpha + \varphi = \theta \\ \alpha + \alpha' = 2\theta \end{array} \right.$$



$$\Rightarrow \alpha - \alpha' = -2\varphi = \text{Change in angle of incident ray}$$

2] Apply small angle approximation - Assume α very small.

$$\left\{ \begin{array}{l} \tan \alpha \approx \alpha \approx \frac{h}{s} \\ \tan \alpha' \approx \alpha' \approx \frac{h}{s'} \\ \sin \varphi \approx \varphi \approx \frac{h}{R} \\ \Rightarrow \frac{h}{s} - \frac{h}{s'} = \frac{-2h}{R} \\ \boxed{\frac{1}{s} + \frac{1}{s'} = \frac{2}{R}} \end{array} \right.$$



Adopt convention : $s' = -ve$ if virtual (behind mirror)

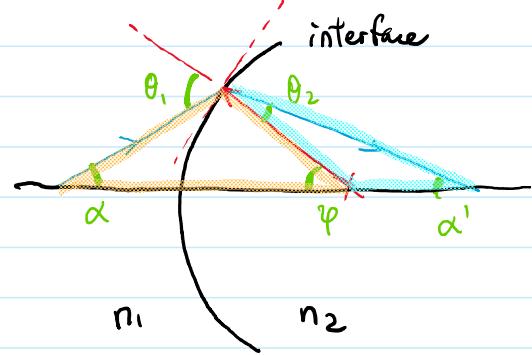
Refraction on spherical lens

1] By Snell's Law

$$n_1 \sin \theta_1 = n_2 \sin \theta_2$$

2] By exterior L of Δ

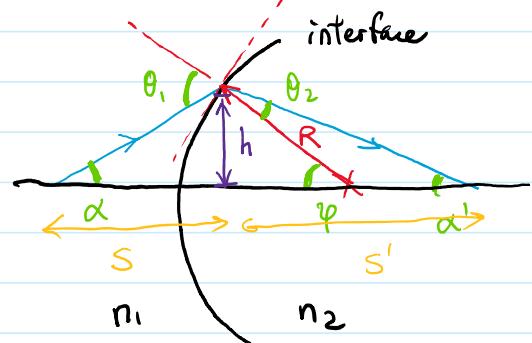
$$\left\{ \begin{array}{l} \theta_1 = \alpha + \varphi \\ \varphi = \theta_2 + \alpha' \end{array} \right.$$



3] Apply small angle approximation , Assume α very small

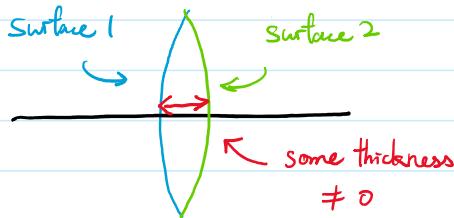
$$\left\{ \begin{array}{l} \tan \alpha \approx \alpha \approx \frac{h}{s} \\ \tan \alpha' \approx \alpha' \approx \frac{h}{s'} \\ \sin \varphi \approx \varphi \approx \frac{h}{R} \\ n_1 \sin \theta_1 \approx n_1 \alpha \approx n_2 \theta_2 \approx n_2 \sin \theta_2 \end{array} \right.$$

$$\Rightarrow \boxed{\frac{n_1}{s} + \frac{n_2}{s'} = \frac{n_2 - n_1}{R}}$$

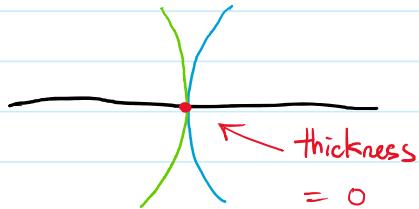


The lens maker's formula (Thin lens formula)

Approximation : Separation between the spherical surfaces = 0



In reality



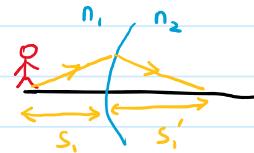
The ideal case in the formulae

Let surrounding refractive index = n_1

Lens material's refractive index = n_2

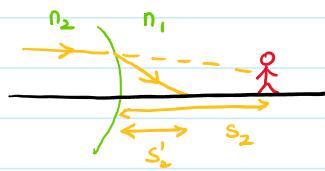
① Through the 1st interface (Radius R_1)

$$\frac{n_1}{s_1} + \frac{n_2}{s'_1} = \frac{n_2 - n_1}{R_1}$$



② Through the 2nd interface (Radius R_2)

$$\frac{n_2}{s_2} + \frac{n_1}{s'_2} = \frac{n_1 - n_2}{R_2}$$



Note that object of 2nd interface = image of 1st interface

- This image is behind 2nd interface \equiv virtual object

\Rightarrow Object distance get -ve sign

$$\text{i.e., } s_2 = -s'_1$$

- Virtual object sees 2nd interface as convex

\Rightarrow Radius of curvature get -ve sign

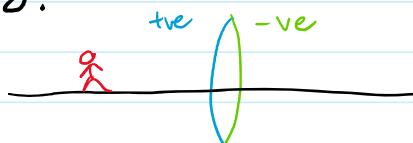
i.e. Change to $-R_2$

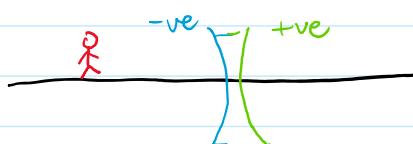
③ Substitute to remove s'_1 , get

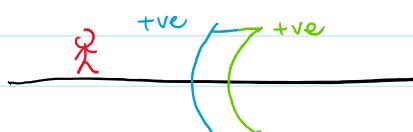
$$\frac{1}{s_1} + \frac{1}{s_2} = \frac{n_2 - n_1}{n_1} \left(\frac{1}{R_1} - \frac{1}{R_2} \right) = \frac{1}{f}$$

Easy remembering : Sign of R = what is seen by original object

Eg.


$$\Rightarrow \frac{1}{+R_1} - \frac{1}{-R_2} = \text{very tve}$$


$$\Rightarrow \frac{1}{-R_1} - \frac{1}{+R_2} = \text{very -ve}$$


$$\Rightarrow \frac{1}{+R_1} - \frac{1}{+R_2} \approx 0$$

Extra : Matrix method for paraxial optics

Paraxial = Light rays have very small angle of elevation ($< 5^\circ$)
 s.t. almost parallel to principle axis

(Also called Gaussian optics because proposed by Gauss)

★ Consequence = Can use small angle approximation

- All the previous formulae belong to this domain
- Optical system can be approximated by linear transform.

The framework

Represent distance from principle axis & elevation angle of light ray

as a column vector

$$\begin{bmatrix} y \\ \alpha \end{bmatrix} \quad \begin{array}{l} \leftarrow \text{Vertical distance} \\ \leftarrow \text{Elevation angle} \end{array}$$


Transition = Matrix multiplication

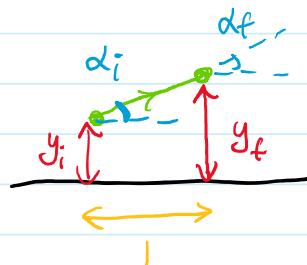
① Translation

After travelling for a horizontal distance L

$$\left\{ \begin{array}{l} y_f = y_i + L \tan \alpha_i \approx y_i + L \alpha_i \\ \alpha_f = \alpha_i \end{array} \right.$$

$$\Rightarrow \begin{bmatrix} y_f \\ \alpha_f \end{bmatrix} = \underbrace{\begin{bmatrix} 1 & L \\ 0 & 1 \end{bmatrix}}_{\text{Matrix for translation}} \begin{bmatrix} y_i \\ \alpha_i \end{bmatrix}$$

Matrix for translation



② Reflection

Consider the light ray right before reflect us, right after reflect

- Change in vertical distance = 0 (No translation yet)

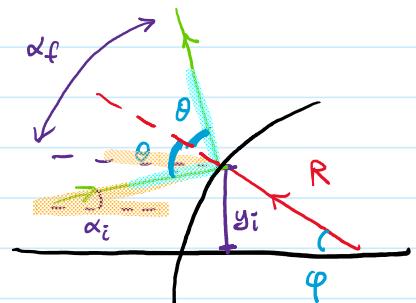
$$\Rightarrow y_f = y_i$$

- Change in elevation angle depends on where it is reflected

By Geometry ,

$$\left\{ \begin{array}{l} \theta = \alpha_i + \varphi \\ 2\theta = \alpha_i + \alpha_f \end{array} \right.$$

And small angle approximation



$$\sin \varphi \approx \varphi \approx \frac{y_i}{R}$$

$$\Rightarrow \alpha_f = \alpha_i + \frac{2y_i}{R}$$

\therefore Express reflection by the matrix

$$\begin{bmatrix} y_f \\ \alpha_f \end{bmatrix} = \underbrace{\begin{bmatrix} 1 & 0 \\ \frac{2}{R} & 1 \end{bmatrix}}_{\text{Matrix of reflection}} \begin{bmatrix} y_i \\ \alpha_i \end{bmatrix}$$

Matrix of reflection

③ Refraction

Similar to reflection

- Change in vertical distance = 0 (No translation yet)

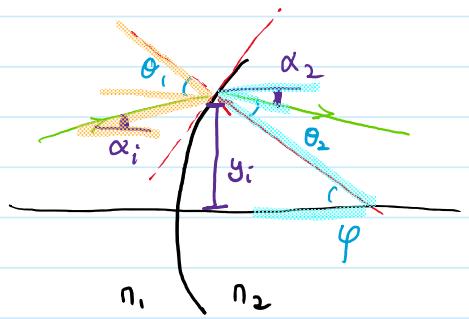
$$\Rightarrow y_f = y_i$$

- Change in elevation angle depends on where it is refracted

- Snell's Law : $n_1 \sin \theta_1 = n_2 \sin \theta_2$

By geometry

$$\begin{cases} \theta_1 = \alpha_i + \varphi \\ \varphi = \alpha_2 + \theta_2 \end{cases}$$



And small angle approximation

$$\sin \varphi \approx \varphi \approx \frac{y_i}{R}$$

After substitution

$$\Rightarrow \alpha_f = \left[\frac{1}{R} \left(\frac{n_1}{n_2} - 1 \right) \right] y_i + \left[\frac{n_1}{n_2} \right] \alpha_i$$

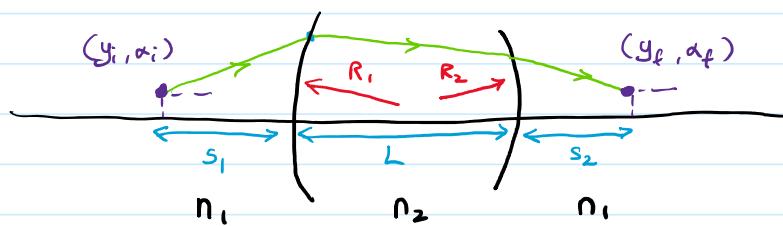
\therefore Express refraction by matrix

$$\begin{bmatrix} y_f \\ \alpha_f \end{bmatrix} = \begin{bmatrix} 1 & 0 \\ \frac{1}{R} \left(\frac{n_1}{n_2} - 1 \right) & \frac{n_1}{n_2} \end{bmatrix} \begin{bmatrix} y_i \\ \alpha_i \end{bmatrix}$$

Application

Matrix method simplify design process & computer programmable

E.g. Design a thick lens with a desired thickness



(y_f, α_f) is related (y_i, α_i) simply by a series of matrix

$$\begin{bmatrix} y_f \\ \alpha_f \end{bmatrix} = \begin{bmatrix} 1 & s_2 \\ 0 & 1 \end{bmatrix} \begin{bmatrix} 1 & 0 \\ -\frac{1}{R_2} \left(\frac{n_2}{n_1} - 1 \right) & \frac{n_2}{n_1} \end{bmatrix} \begin{bmatrix} 1 & L \\ 0 & 1 \end{bmatrix} \begin{bmatrix} 1 & 0 \\ \frac{1}{R_1} \left(\frac{n_1}{n_2} - 1 \right) & \frac{n_1}{n_2} \end{bmatrix} \begin{bmatrix} 1 & s_1 \\ 0 & 1 \end{bmatrix} \begin{bmatrix} y_i \\ \alpha_i \end{bmatrix}$$

translate by s_2
retract at 2nd interface
translate by L
retract at 1st interface
translate by s_1

- Easy to obtain object-image distance relation, instead of applying lens formula many times
- Easy to tune parameters for desired purposes

E.g. $\left\{ \begin{array}{l} \text{Can vary} = R_1, R_2, L, \frac{n_2}{n_1} \\ \text{Require equivalent focal length} = \frac{1}{s_1 + \frac{L}{2}} + \frac{1}{s_2 + \frac{L}{2}} = \frac{1}{f} \end{array} \right.$

Extra-Extra : Aberration Theory

Problem: Perfect imaging only happen at paraxial condition ($< 5^\circ$)

Aberration = Imaging error between object and the image in reality

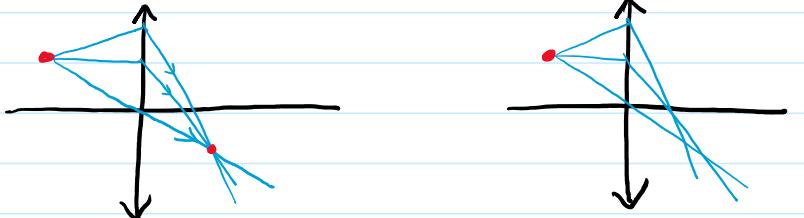
2 Main groups of aberration

① Achromatic aberration

= Error unrelated to colour of light.

= Error from small angle approximation.

E.g.



Perfect focus

All light from same source go to same spot

Reality

Geometrically they cannot arrive at the same spot

How to mathematically describe this error?

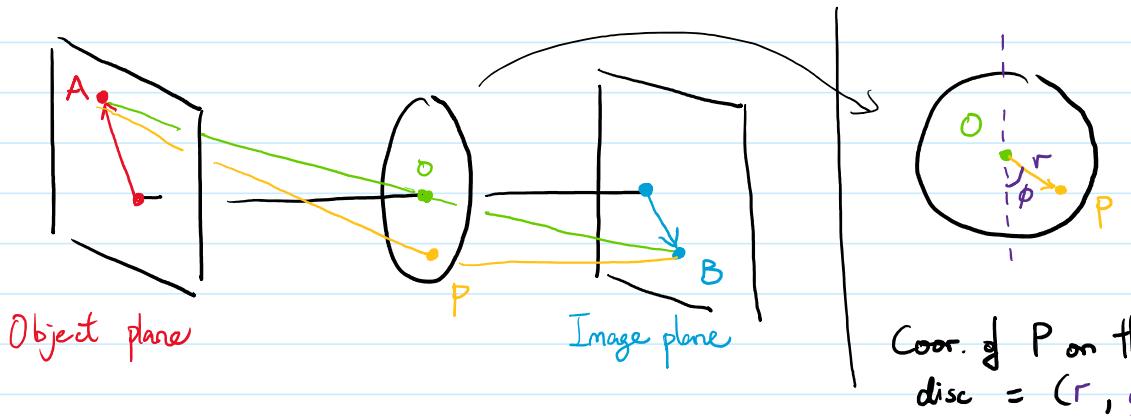
By the light path length difference between a light ray through the image device's center vs a light ray pass through elsewhere.

= The distance light travelled in Fermat's light principle.
Not just geometrical distance.

If the image is focused, any rays should have the same light path distance, no matter where it passes through.

For the error due to spherical device, we can write it as

$$\begin{aligned} a(r, \phi) &= (\text{length of light path through device center}) \\ &\quad - (\text{length of light path through anywhere else}) \\ &= \frac{\text{light}}{AOB} - \frac{\text{light}}{APB} \end{aligned}$$



This difference in light path length can be expressed as an infinite sum if we keep all the terms in small angle approximation

$$a(r, \phi) = \sum_m \sum_n C_{mn} \left(\frac{r}{R}\right)^m \cos^n \phi$$

$\curvearrowleft R = \text{radius of the disc so } 0 \leq \frac{r}{R} \leq 1$

where C_{mn} are the constants that determine the size of
the aberration type $(\frac{r}{R})^m \cos^n \phi$

Note that when m, n go bigger, $(\frac{r}{R})^m \cos^n \phi$ becomes smaller
i.e. the higher order aberration types are more negligible.

Siedel Aberration

The first 5 terms in the sum contribute to the major
imaging error to most optical devices.

They are grouped to called Siedel Aberration or

3rd order aberration ($\sin \theta \approx \theta - \frac{\theta^3}{3!} + \dots$ error begin with θ^3)

They are given individual name mostly by their geometric effect.

① Spherical Aberration $\propto (\frac{r}{R})^4$

② Coma $\propto (\frac{r}{R})^3 \cos \phi$

③ Astigmatism $\propto (\frac{r}{R})^2 \cos^2 \phi$

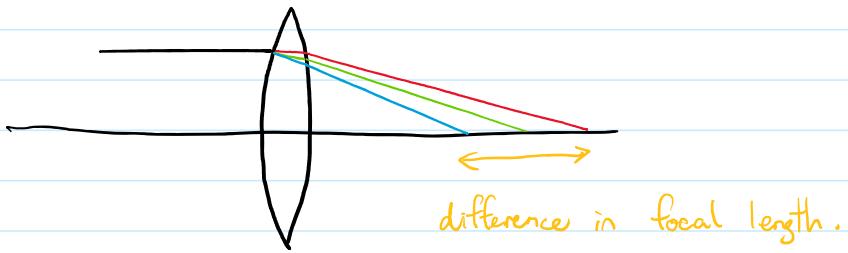
④ Curvature of field $\propto (\frac{r}{R})^2$

⑤ Distortion $\propto (\frac{r}{R}) \cos \phi$

② Chromatic Aberration

= Error by refractive index dependency on wavelength.

Usually quantified by the difference in focal length
between several standard wavelength.



Idea of optical device = Meet design requirement + minimize aberration.

- Most calculation handle by computer software nowadays.

- Most aberrations have standard "solution"

E.g. Chromatic aberration \rightarrow flint glass + crown glass combination

- But different types of aberration are mathematically related

Reducing one type may enlarge another type.

\Rightarrow Need to sacrifice certain reduction depending on needs.

E.g. Telescope targets = paraxial objects.

\hookrightarrow Can safely ignore most achromatic aberration

when choosing the right lens.

- Also need to care other engineering requirement / production cost.

E.g. Weight / Size of device? Adjustable focal length?