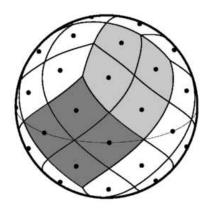
# **HEALPix** Fortran Facility User Guidelines



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Prepared by: Eric Hivon, Frode K. Hansen, Benjamin D. Wandelt,

Krzysztof M. Górski, Anthony J. Banday, Martin

Reinecke

Abstract: This document describes the **HEALPix** Fortran

stand-alone facilities

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## Using the HEALPix Fortran 90 facilities

**Default File Names and Directories**: for some applications, the **HEALPix** facilities require some precalculated input files describing the pixel window function and quadrature ring weight (shipped as Healpix/data/pixel\_window\_n????.fits and Healpix/data/weight\_ring\_n?????.fits respectively).

By default, files with the very same name will be looked for into: the current directory (.), the parent directory (.), ./data, ../data, \$HEALPIX and \$HEALPIX/data where \$HEALPIX is a system variable defined as the full path to the **HEALPix** package (see the installation documentation). The user has the possibility to change both the name of those files and their location.

**Double/Single Precision mode**: several facilities offer the option of switching at run time the precision of the internal variables and arrays and of the I/O data from single to double precision floating point reals. The following points should be noted:

- Facilities running in double precision mode can read indifferently single and double precision data files (and the same is true for single precision facilities). On the other hand, a double (resp. single) precision facility will only output double (single) precision files.
- Since the internal calculations sensitive to numerical round-off error (like the spherical harmonics recurrence) are *always* performed in double precision, switching to double precision mode
  - will have a limited impact on the output accuracy if the input file contains only single precision data,
  - is recommended if the input file contains double precision data, and the precision of the output is critical
  - will not alter the execution speed, but it will almost double the memory consumption of the facility,
  - will obviously double the size of the output file(s).

## Changes between release 2.0 and 2.1

- The anafast facility can now compute the cross-correlations of two different maps.
- The sky\_ng\_sim facility (Rocha et al, 2005), to produce non-Gaussian CMB temperature maps, has been added.

## Changes between release 1.2 and 2.0

- faster implementation of  $a_{lm}$  related facilities, generalization of OpenMP parallelization, and availability of MPI parallelized routines (see mpi\_\* routines in Fortran90 Subroutines Overview document).
- introduction of alteralm facility to modify and/or rotate the spherical harmonics coefficients  $a_{lm}$  and greater flexibility for constraining  $a_{lm}$  in synfast
- single and double precision implementation of most facilities (see **Input and Output Precision** page 3)

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## alteralm

#### Location in HEALPix directory tree: src/f90/alteralm/alteralm.f90

This program can be used to modify a set of  $a_{lm}$  spherical harmonics coefficients, as those extracted by anafast or simulated by synfast, before they are used as constraints on a synfast run. Currently the alterations possible are

- rotation (using Wigner matrices) of the  $a_{lm}$  from the input coordinate system to any other standard astrophysical coordinate system. The resulting  $a_{lm}$  can be used with e.g. synfast to generate a map in the new coordinate system.
- removal of the pixel and beam window functions of the input  $a_{lm}$  (corresponding to the pixel size and beam shape of the map from which they were extracted) and implementation of an arbitrary pixel and beam window function.

$$a_{\ell m}^{\rm OUT} = a_{\ell m}^{\rm IN} \frac{B^{\rm OUT}(\ell) P^{\rm OUT}(\ell)}{B^{\rm IN}(\ell) P^{\rm IN}(\ell)},\tag{1}$$

where  $P(\ell)$  is the pixel window function, and  $B(\ell)$  is the beam window function (assuming a circular beam) or any other  $\ell$  space filter (eg, Wiener filter). For an infinitely small pixel (or beam) one would have  $P(\ell) = 1$  (resp.  $B(\ell) = 1$ ) for any  $\ell$ .

## **FORMAT**

% alteralm [options] [parameter\_file]

#### COMMAND LINE OPTIONS

-d

--double double precision mode (see Notes on double/single precision modes on page 3)

-s

--single single precision mode (default)

## **QUALIFIERS**

infile\_alms = Defines the FITS file from which to read the input  $a_{\ell m}$ .

Defines the FITS file in which to write the altered  $outfile_alms =$  $a_{\ell m}$ .  $fwhm_arcmin_in =$ Defines the FWHM size in arcminutes of the Gaussian beam present in the input  $a_{lm}$ . The output  $a_{lm}$  will be corrected from it, see Eq. (1). (default = value of FWHM keyword in infile\_alms).  $beam_file_in =$ Defines the FITS file describing the Legendre window function of the circular beam present in the input  $a_{lm}$ . The output  $a_{lm}$  will be corrected from it, see Eq. (1). If set to an existing file name, it will override the fhwm\_arcmin\_in given above. (default= value of the BEAM\_LEG keyword in infile\_alms)  $nlmax_out =$ Defines the maximum  $\ell$  value to be used for the output  $a_{lm}$ s. (default= maximum  $\ell$  of input  $a_{\ell m}$ = value of MAX-LPOL keyword in infile\_alms).  $nsmax_in =$ If it can not be determined from the input file infile\_alms, asks for the **HEALPix** resolution parameter  $N_{\text{side}}$  whose window function is applied to the input  $a_{\ell m}$ Defines the **HEALPix** resolution parameter  $N_{\text{side}}$  $nsmax_out =$ whose window function will be applied to the output  $a_{\ell m}$ . Could be set to 0 for infinitely small pixels, ie no pixel window function (default= same as input's  $N_{\text{side}}$ ). fwhm\_arcmin\_out = Defines the FWHM size in arcminutes of the Gaussian beam to be applied to  $a_{lm}$ , see Eq. (1).  $(default = fwhm_arcmin_in).$  $beam_file_out =$ Defines the FITS file describing the Legendre window function of the circular beam to be applied  $a_{lm}$ . If set to an existing file name, it will override the fhwm\_arcmin\_out given above. (default= "" )  $coord_in =$ Defines astrophysical coordinates used to compute the input  $a_{\ell m}$ . Valid choices are 'G' = Galactic, 'E' = Ecliptic, 'C'/'Q' = Celestial = eQuatorial. (default = value of COORDSYS keyword read from input FITS file)  $epoch_in =$ Defines astronomical epoch of input coordinate system (default=2000)

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> $coord_out =$ Defines astrophysical coordinates into which to ro-

> > tate the  $a_{\ell m}$  (default = coord\_in)

 $epoch_out =$ Defines astronomical epoch of output coordinate

system (default=epoch\_in)

 $\overline{\mathbf{DESCRIPTION}}$  Alteralm can modify temperature as well as polarisation  $a_{lm}$ . It will also modify the error on the  $a_{lm}$  if those are provided. It works best if the input FITS file contains the relevant information on the beam size and shape, maximum multipoles,

## **DATASETS**

The following datasets are involved in the **alteralm** processing.

Dataset	Description

/data/pixel\_window\_nxxxx.fitFiles containing pixel windows for various nsmax.

#### **SUPPORT**

This section lists those routines and facilities (including those external to the HEALPix distribution) which can assist in the utilisation of alteralm.

anafast This **HEALPix** Fortran facility can analyse a

**HEALPix** map to extract the  $a_{lm}$  that can be

altered by alteralm.

synfast **HEALPix** facility can generate a

**HEALPix** map from a power spectrum  $C_{\ell}$ , with the possibility of including constraining  $a_{lm}$ 

as those obtained with alteralm.

## EXAMPLES: #1

alteralm

Alteralm runs in interactive mode, self-explanatory.

## EXAMPLES: #2

alteralm filename

When 'filename' is present, alteralm enters the non-interactive mode and parses its inputs from the file 'filename'. This has the following structure: the first entry is a qualifier which announces to the parser which input immediately follows. If this input is omitted in the input file, the parser assumes the default value. If the equality sign is omitted, then the parser ignores the entry. In this way comments may also be included in the file. In this example, the file contains the following qualifiers:

```
infile_alms = alm.fits
nlmax_out = 512
fwhm_arcmin_out = 20.0
coord_out = G
outfile_alms = newalm.fits
```

Alteralm reads the  $a_{lm}$  from 'alm.fits'. Since  $nsmax_in$   $nsmax_out$   $fwhm_arcmin_in$   $beam_file_in$   $coord_in$   $epoch_in$ 

have their default values, the pixel size will remain the same, the  $a_{lm}$  will be corrected from its input beam (whatever it was, assuming the relevant information can be found), and a gaussian beam of 20.0 arcmin will be applied instead, the  $a_{lm}$  will also be rotated from their original coordinate system (whatever it was, assuming the relevant information can be found) into Galactic coordinates, assuming a year 2000 epoch for both, and only the multipoles up to 512 will be written in 'newalm.fits'.

## RELEASE NOTES

Revision 1: Initial release (HEALPix 2.00)

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## Messages

This section describes error messages generated by **alteralm**.

Message	Severity	Text
can not allocate memory for array xxx	Fatal	You do not have sufficient system resources to run this facility at the map resolution you required. Try a lower map resolution.
this is not a binary ta- ble		the fitsfile you have specified is not of the proper format
there are undefined values in the table!		the fitsfile you have specified is not of the proper format
the header in xxx is too long		the fitsfile you have specified is not of the proper format
XXX-keyword not found		the fitsfile you have specified is not of the proper format
found xxx in the file, expected:yyyy		the specified fitsfile does not contain the proper amount of data.
alteralm> no informa- tion found on input alms beam	Fatal	no information on the input beam was found, neither from parsing the FITS file header, nor from what the user pro- vided.

## anafast

#### Location in HEALPix directory tree: src/f90/anafast/anafast.f90

This program performs harmonic analysis of the **HEALPix** maps up to a user specified maximum spherical harmonic order  $\ell_{max}$ . The integrals are computed on the whole sphere, unless the user chooses a provided option to excise from the input map(s) a simple, constant latitude, symmetric cut, and/or apply an arbitray cut read from an external file. Scalar, or scalar and tensor, spherical harmonic coefficients are evaluated from the map(s) if the input provides, respectively, only the temperature, or temperature and polarisation maps. The total operation count scales as  $\mathcal{O}(N_{\rm pix}^{3/2})$  with a prefactor depending on  $\ell_{max}$ .

Anafast reads one (two) file(s) containing the map(s) and produces a file containing the temperature auto- (or cross-) power spectrum  $C_l^T$  and, if requested, also the polarisation power spectra  $C_l^E$ ,  $C_l^B$ ,  $C_l^{T \times E}$ ,  $C_l^{T \times B}$  and  $C_l^{E \times B}$ . The  $a_{lm}$  coefficients computed during the execution also can be written to a (two) file(s) if requested.

Anafast executes an approximate, discrete point-set quadrature on a sphere sampled at the **HEALPix** pixel centers. Spherical harmonic transforms are computed using recurrence relations for Legendre polynomials on co-latitude,  $\theta$ , and Fast Fourier Transforms on longitude,  $\phi$ .

Anafast is provided with an option to use precomputed Legendre Polynomials to decrease the total execution time.

Anafast permits two execution options which allow a significant improvement of accuracy of the approximate quadrature performed by this facility:

- An improved analysis using the provided ring weights, which correct the quadrature on latitude, and/or
- An iterative scheme using in succession several backward and forward harmonic transforms of the maps.

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## RECOMMENDATIONS FOR USERS

Execution of anafast requires a user to specify the maximum spherical harmonic order  $\ell_{max}$  up to which the harmonic decomposition of the input maps will be performed. Since there are no formal limits on parameter  $\ell_{max}$  enforced by anafast, the user should make his/her choices judiciously. Hereafter it is convenient to specify  $\ell_{max}$  in terms of the **HEALPix** map resolution parameter nsmax.

If the function to be analysed is strictly band-width limited, or nearly band-width limited (as in the case of a Gaussian beam smoothed signal discretized at a rate of a few pixels per beam area), it is sufficient to run anafast with  $\ell_{max} \approx 2 \cdot nsmax$ , with a very good  $C_{\ell}$  error performance already in the raw (i.e. uncorrected quadrature) harmonic transform mode. If quadrature corrections are still desired in this case, it should be sufficient to use, at no extra cost in execution time, the ring-weighted quadrature scheme. This is the recommended mode of operation of anafast for essentially error and worry free typical applications, e.g. CPU-intensive Monte Carlo studies.

If more aggressive attempts are undertaken to extract from a map the spectral coefficients at  $\ell > 2 \cdot nsmax$  (for example, as in a possible case of an attempt to analyse an existing map, which was irreversibly binned at a suboptimal resolution) the following should be kept in mind:

- Spherical harmonics discretized using **HEALPix** (either sampled at pixel centers, or avaraged over pixel areas) form a linearly independent system up to  $\ell_{max} = 3 \cdot nsmax 1$ . Hence, the functions which are strictly band-width limited to  $\ell_{max} = 3 \cdot nsmax 1$  can be fully spectrally resolved with anafast, albeit with integration errors in the uncorrected quadrature mode, which grow up to  $\delta C_{\ell} \propto \epsilon \cdot C_{\ell}$ , with  $\epsilon < 0.1$ , at the highest values of  $\ell$ . These integration errors can be efficiently reduced using anafast in the iterative mode. Although this  $\ell_{max}$  range  $-2 \cdot nsmax < \ell_{max} < 3 \cdot nsmax 1$  is easily manageable with anafast used on strictly band-width limited functions, it should be used with caution in basic and automated applications, e.g. Monte Carlo simulations.
- As with any discrete Fourier transform, anafast application to functions which are not band-width limited results with aliasing of power, which can not be remedied. If the particular case of interest may result in such a band-width violation (i.e. there is significant power in the function at  $\ell > 3 \cdot nsmax 1$ ), the function should be smoothed before the application of anafast, or discretized and then analysed, on a refined **HEALPix** grid (with larger nsmax).
- REMEMBER: A peculiar property of the sphere, which usually surprizes those whose intuition is built on experience with FFTs on a segment, or on a Euclidean multidimensional domains, is the lack of a regular and uniform point-set at arbitrary resolution, and the resulting non-commutativity of the forward and backward discrete Fourier transforms on nearly-uniform point-sets, e.g. **HEALPix**. Hence, as in any case of attempting an extreme application of an off-the-shelf software, use caution and understand your problem well before executing anafast under such cicumstances!

## **FORMAT**

## % anafast [options] [parameter\_file]

#### **COMMAND LINE OPTIONS**

-d

--double double precision mode (see Notes on double/single precision modes on page 3)

-s

--single single precision mode (default)

## **QUALIFIERS**

infile = Defines the input map file. (default= map.fits)

infile2 = Defines the 2nd input map file, to be cross-

correlated with the first one. The 2 maps should match in resolution (nsmax) and coordinate. (default='', only the auto-correlation of the first map

will be computed)

outfile = Defines the output file with the power spectrum.

If only one input map is provided, outfile will contains its auto-spectra, if 2 maps are provided, outfile will contain their cross-spectra. Note in particular that in the latter case, the  $C_l^{T\times E}$  power spectrum will be build from the T field of the 1st (possibly polarized) map, and the E field of the

second polarized map. (default= cl\_out.fits)

 $simul\_type = Defines which map(s) to analyse, 1=tempera-$ 

ture only, 2=temperature AND polarisation. (de-

fault = 1

nlmax = Defines the maximum  $\ell$  value to be used. See the

Recommendations for Users. (default= 64)

maskfile = Defines the FITS file containing the pixel mask(s)

or weighting scheme(s) by which the map(s) read from infile will be multiplied before analysis. If the file contains several fields, the first one in which at least one pixel is non-zero will be used.

This option can be used to, for instance, apply

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 $theta\_cut\_deg =$ 

 $iter\_order =$ 

 $outfile\_alms =$ 

 $outfile\_alms2 =$ 

plmfile =

w8file =

w8filedir =

 ${\rm regression} =$ 

won =

the WMAP Kp intensity mask to the data (see http://lambda.gsfc.nasa.gov), but it will not handle the WMAP composite mask correctly. Can be used in conjonction with theta_cut_deg. Masked or weighted pixels will be correctly accounted when performing the monopole and dipole regression. (default= ": no mask, all valid pixels are used)
Defines the latitude (in degrees) of an optional, straight symmetric cut around the equator. Pixels located within that cut ( $ b $ <theta_cut_deg (default="0°:" are="" cut)<="" ignored.="" no="" td=""></theta_cut_deg>
Defines the maximum order of quadrature iteration to be used. (default=0, no iteration)
Defines the name of the file containing the $a_{lm}$ coefficients of the first map which can be written optionally. (default= no entry — $a_{lm}$ s are not written to a file)
Defines the name of the file containing the $a_{lm}$ coefficients of the second map, if any, which can be written optionally. (default= no entry — $a_{lm}$ s are not written to a file)
Defines the name for an input file containing pre- computed Legendre polynomials $P_{lm}$ . (default= no entry — anafast executes the recursive evalua- tion of $P_{lm}$ s)
Defines name for an input file containing ring weights in the improved quadrature mode (default= no entry — the name is assumed to be 'weight_ring_n0xxxx.fits' where xxxx is nsmax)
Gives the directory where the ring weight files are to be found (default= no entry — anafast searches the default directories, see introduction)
Set this to 1 if ring weight files are to be used, otherwise set it to 0 (or 2). (default= 0)
Sets the degree of the regression made on the input map before doing the power spectrum analysis. The regression is a minimal variance fit (assuming a uniform noise) made on valid (unflagged and unmasked) pixels, out of the symmetric cut (if any). In case of cut sky analysis, such a regres-

the WMAP Kp intensity mask to the data (see

sion reduces the monopole and dipole leakage to higher  $\ell$ 's.

0: no regression, does the alm analysis on the

1: removes the best fit monopole first

2: removes the best fit monopole and dipole first default = 0.

**DESCRIPTION** Anafast reads one of two binary FITS-files containing a **HEALPix** map. These files can each contain a temperature map or both temperature and polarisation (Q,U) maps. Anafast analyses the map(s) and makes an output ascii-FITS file containing the angular auto or cross power spectra  $C_l^T$ s (and  $C_l^E$ ,  $C_l^B$ ,  $C_l^{T \times E}$ ,  $C_l^{T \times B}$  and  $C_l^{E \times B}$  if specified). Anafast produces  $C_l$ s up to a specified maximum  $\ell$ -value (see Recommendations for Users). Note in particular that if 2 maps are provided, the  $C_{I}^{T \times E}$  power spectrum will be build from the T field of the 1st (polarized) map, and the E field of the second polarized map. If requested, the computed  $a_{lm}$  coefficients can be written to a FITS file. This file can be used in the constrained realisation mode of synfast.

> Anafast permits two execution modes that allow to improve the quadrature accuracy: (1) the ring weight corrected quadrature, and (2) the iterative scheme. Using the ring weights does not increase the execution time. The precomputed ring weights to be used for each **HEALPix** resolution nsmax are provided in the \$HEALPIX/data directory. The more sophisticated iterative scheme increases the accuracy more effectively than the weighted ring scheme, but its disadvantage is that the time for the analysis increases, 1 iteration takes 3 times as long, 2 iterations 5 times as long on so forth, since each order of iteration requires one more forward and backward transform.

> The spherical harmonics evaluation uses a recurrence on associated Legendre polynomials  $P_{lm}(\theta)$ . This recurrence consumes most of the CPU time used by anafast. We have therefore included an option to load precomputed values for the  $P_{lm}(\theta)$  from a file generated by the **HEALPix** facility plmgen. This represents a tradeoff between memory usage and execution speed of anafast.

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#### **DATASETS**

The following datasets are involved in the anafast processing.

Dataset	Description
, 0	Files containing ring weights for the anafast improved quadra- ture mode.

## **SUPPORT**

This section lists those routines and facilities (including those *external* to the HEALPix distribution) which can assist in the utilisation of **anafast**.

synfast	This <b>HEALPix</b> facility can generate a map for analysis by anafast.
alteralm	This $\mathbf{HEALPix}$ Fortran facility can be used to
	modify the $a_{lm}$ extracted by anafast before they
	are used as constraints on a synfast run.
plmgen	This <b>HEALPix</b> facility can be used to generate
	precomputed Legendre polynomials.

## EXAMPLES: #1

anafast

Anafast runs in interactive mode — self-explanatory.

## EXAMPLES: #2

#### anafast filename

When 'filename' is present, anafast enters the non-interactive mode and parses its inputs from the file 'filename'. This has the following structure: the first entry is a qualifier which announces to the parser which input immediately follows. If this input is omitted in the input file, the parser assumes the default value. If the equality sign is omitted, then the parser ignores the entry. In this way comments may also be included in the file. In this example, the file contains the following qualifiers:

```
simul_type = 1
nlmax = 64
theta_cut_deg = 0
iter_order = 0
infile = map.fits
outfile = cl_out.fits
regression = 0
```

Anafast reads the map from map.fits, makes an analysis and produces  $C_l^T$ s up to l=64. This powerspectrum is saved in the file  $cl_out.fits$ . No galactic cut is excised and no iterations are performed. As regression is set to 0 (its default value) the map is analyzed as is, without prior best fit removal of the monopole nor the dipole.

```
Since
infile2
outfile_alms
outfile_alms2
w8file
w8filedir
plmfile
maskfile
```

were omitted, they take their default values (empty strings). This means that no file for precomputed Legendre polynomials is read, no second map is read, no mask is applied, and anafast does not save the  $a_{lm}$  values from the analysis.

Also since

#### won

is not given, it takes it default value 2, which means that ring weights are not used.

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## RELEASE NOTES

Revision 1: Initial release (HEALPix 0.90)

**Revision 2:** Optional non-interactive operation. Proper FITS file support. Improved recurrence algorithm for  $P_{lm}(\theta)$  which can compute to higher  $\ell$  values. New functionality: arbitrary order of iterations, precomputed  $P_{lm}$ , dumping of  $a_{lm}$ . (**HEALPix** 1.00)

Revision 3: New functionality: possibility of removing the best fit monopole and dipole. New Parser. Can be linked to FFTW (**HEALPix** 1.20)

Revision 4: New functionality: addition of maskfile (HEALPix 2.0)

**Revision 5:** Bug correction: correct interaction of iterative scheme with masked pixels (**HEALPix** 2.01)

Revision 6: New functionality: cross-correlation of 2 maps; Correction of this documentation: the code expects maskfile and not mask\_file (HEALPix 2.1)

## Messages

This section describes error messages generated by anafast.

Message	Severity	Text
can not allocate memory for array xxx	Fatal	You do not have sufficient system resources to run this facility at the map resolution you required. Try a lower map resolution.

## hotspot

#### Location in HEALPix directory tree: src/f90/hotspot/HotSpots.f90

This Fortran facility provides a means to find local extrema of a map in **HEALPix** format. It also serves to illustrate the use of the following parts of the **HEALPix** toolkit: fast neighbour and extrema finding in the nested scheme, in-place conversion between RING and NESTED pixel schemes

## **FORMAT**

 $\sqrt{\%}$  hotspot

## **QUALIFIERS**

infile =Defines the input map file.

 $extrema_outfile =$ Defines the output map file.

 $maxima_outfile =$ Defines the output ascii list of maxima.

minima\_outfile= Defines the output ascii list of minima.

**DESCRIPTION** hotspot reads a healpix map in FITS format and generates the following outputs: 1) a **HEALPix** map in FITS format which is zero everywhere, except at pixels which contain local extrema. These pixels have the same values as in the input map. 2) an ASCII file which lists the pixel numbers and values of maxima, and 3) an ASCII file which lists the pixel numbers and values of minima.

> The facility can be used in both an interactive mode and a command mode, where command qualifiers are fed to the facility using an input file.

> Note the following limitations: hotspot (and the toolkit neighbour finder which it uses) will only work on maps with  $N_{side} \geq 2$ .

## DATASETS

The following datasets are involved in the **hotspot** processing.

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Description

## **SUPPORT**

This section lists those routines and facilities (including those *external* to the HEALPix distribution) which can assist in the utilisation of **hotspot**.

synfast This **HEALPix** facility can generate a FITS for-

mat sky map to be input to hotspot.

map2gif This **HEALPix** Fortran facility can be used to

visualise the output map.

mollview This **HEALPix** IDL facility can be used to visu-

alise the output map.

## EXAMPLES: #1

hotspot

hotspot runs in interactive mode.

## $\overline{\text{EXAMPLES}}$ : #2

#### hotspot filename

When 'filename' is present, hotspot enters the non-interactive mode and parses its inputs from the file 'filename'. This has the following structure: the first entry is a qualifier which announces to the parser which input immediately follows. If this input is omitted in the input file, the parser assumes the default value shown below. If the equality sign is omitted, then the parser ignores the entry. In this way comments may also be included in the file. In this example, the file contains the following qualifiers:

infile = map.fits
extrema\_outfile = pixlminmax.fits
maxima\_outfile = maxima.dat
minima\_outfile = minima.dat

hotspot reads in the map 'map.fits' and generates an output map with name 'pixlminmax.fits', and two ASCII files, 'maxima.dat' and 'minima.dat'.

#### RELEASE NOTES

Revision 1: Initial release (HEALPix 0.90)

Revision 2: Optional non-interactive operation. Proper FITS file support for input and output maps. (HEALPix 1.00)

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# Messages

This section describes error messages generated by  ${\bf hotspot}.$ 

Message	Severity	Text
can not allocate memory for array xxx	Fatal	You do not have sufficient system resources to run this facility at the map resolution you required. Try a lower map resolution.

# map2gif

#### Location in HEALPix directory tree: src/f90/map2gif/map2gif.f90

This Fortran facility provides a means to generate a gif image from an input **HEALPix** sky map. It is intended to allow some primitive visualisation for those with limited or no access to IDL. It is also useful for image generation in a pipeline environment.

## **FORMAT**

## % map2gif

## **QUALIFIERS**

3	
- bar	Logical which determines whether a color bar is displayed (default: .false.).
- col	The number of the color table to utilise ( <b>default:</b> 5).
- hlp	Print on-line help for the facility.
- inp	The file name of the input FITS sky map (default: none).
- log	Logical to use the log of the signal when plotting (default: .false.)
- ash	Logical to use the hyperbolic arc sine of the signal when plotting. Cannot be true when -log is true. (default: .false.)
- add	Real value to add to the signal before performing any other operation to it (like taking the logarithm etc.) ( <b>default:</b> 0.0)
- mul	Real value to multiply the signal with directly after adding the offset (see above). ( <b>default:</b> 1.0)
- min	Set the minimum value for the plotted signal (default: is to use the actual signal minimum).
- max	Set the maximum value for the plotted signal (default: is to use the actual signal maximum).
- pro	Select the projection scheme ( <b>default:</b> Mollweide).
- out	The file name of the output gif image ( <b>default:</b> none).

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> - sig The identifier of the signal to plot: for a polarisation map, then the mapping is 1 = I; 2 = Q; 3 =U (default: 1). - ttl A string specifying the title for the plot (**default:** The x-dimension of the image in pixels (**default:** - XSZ 800).

**DESCRIPTION** map2gif reads in a **HEALPix** sky map in FITS format and generates an image in GIF format. map2gif allows the selection of the projection scheme (Mollweide or Gnomonic for small patches of the sky), color table, color bar inclusion, linear or log scaling, maximum and minimum range for the plot and plottitle. The facility utilises a command-line interface.

## **DATASETS**

The following datasets are involved in the **map2gif** processing.

Dataset	Description	
None required		

#### **SUPPORT**

This section lists those routines and facilities (including those external to the HEALPix distribution) which can assist in the utilisation of **map2gif**.

xv or a similar facility is required to view the gif XVimage generated by map2gif (a browser can also be used). synfast This **HEALPix** facility will generate the FITS format sky map to be input to map2gif.

EXAMPLES: #1

map2gif reads in the map 'planck100GHZ-LFI.fits' and generates an output gif image with name 'planck100GHZ-LFI.gif' in which the temperature scale has been set to lie between  $\pm$  100 ( $\mu$ K), a color bar has been drawn and the title 'Simulated Planck LFI Sky Map at 100GHz' appended to the image.

#### RELEASE NOTES

Revision 1: Initial release (HEALPix 1.00)

## Messages

This section describes error messages generated by map2gif.

Message	Severity Text	
None at present		

median\_filter 25

## median\_filter

Location in HEALPix directory tree: src/f90/median\_filter/median\_filter.f90

This program produces the median filtered map of an input **HEALPix** map (polarised or unpolarised). The neighborhood on the which the median is computed is defined as a disk of user-defined radius

## **FORMAT**

% median\_filter [options] [parameter\_file]

#### COMMAND LINE OPTIONS

-d

--double double precision mode (see Notes on double/single precision modes on page 3)

-s

--single single precision mode (default)

## **QUALIFIERS**

simul\_type = Either 1 or 2. If set to 1, only the temperature component of the input map will be filtered. If set to 2, all the Stokes components available in the input file will be filtered (default = 1)

infile = Name of the FITS file containing the map to be filtered (default = ", no default input file).

mf\_radius\_arcmin = Radius in arcmin of the disk over which the median is computed (default =  $3\theta_{pix}$  where  $\theta_{pix}$  is the input map pixel size).

fill\_holes = If set to true, flagged pixels take for value the median of the valid pixels surrounding them (if any). Otherwise they are left unchanged. (default = .false.). Note that y, yes,t, true, .true. and 1 are interpreted as true, while n, no, f, false, .false. and 0 stand for false.

mffile = Name of the FITS file containing the median filtered map

 ${f DESCRIPTION}$  Median\_Filter produces a median filtered map in which the value of each pixel is the median of the input map valid pixels found within a disk of given radius centered on that pixel. A pixel flagged as 'non-valid' in the input map can either be left unchanged or 'filled in' with the same scheme, if at least one valid pixel is found among its neighbors.

> If the map is polarized, each of the three Stokes components is filtered separately.

## **SUPPORT**

This section lists those routines and facilities (including those external to the HEALPix distribution) which can assist in the utilisation of median\_filter.

anafast This **HEALPix** Fortran facility can analyse a

**HEALPix** map.

synfast **HEALPix** facility can generate

**HEALPix** map from a power spectrum  $C_{\ell}$ .

## EXAMPLES: #1

median\_filter [option]

Median\_Filter runs in interactive mode, self-explanatory.

## $\overline{\text{EXAMPLES: } \#}2$

median\_filter [option] filename

median\_filter 27

When 'filename' is present, median\_filter enters the non-interactive mode and parses its inputs from the file 'filename'. This has the following structure: the first entry is a qualifier which announces to the parser which input immediately follows. If this input is omitted in the input file, the parser assumes the default value. If the equality sign is omitted, then the parser ignores the entry. In this way comments may also be included in the file. In this example, the file contains the following qualifiers:

```
simul_type = 1
infile = map.fits
mf_radius_arcmin = 20.0
mffile = med.fits
```

Median\_Filter reads the sky map from 'map.fits'. Since fill\_holes

has its default value, ..., The median will be computed on a disk of 20 arcmin in radius, and the result will be written in 'med.fits'.

#### RELEASE NOTES

Revision 1: Initial release (HEALPix 2.00)

## Messages

This section describes error messages generated by **median\_filter**.

Message	Severity	Text
can not allocate memory for array xxx	Fatal	You do not have sufficient system resources to run this facility at the map resolution you required. Try a lower map resolution.
this is not a binary ta- ble		the fitsfile you have specified is not of the proper format
there are undefined values in the table!		the fitsfile you have specified is not of the proper format
the header in xxx is too long		the fitsfile you have specified is not of the proper format
XXX-keyword not found		the fitsfile you have specified is not of the proper format
found xxx in the file, expected:yyyy		the specified fitsfile does not contain the proper amount of data.

plmgen 29

# plmgen

#### Location in HEALPix directory tree: src/f90/plmgen/plmgen.f90

This program can be used to create a file containing the precomputed values of the associated Legendre polynomials  $P_{lm}(\theta)$ (and, if requested, of the tensor spherical harmonics) for faster execution of the **HEALPix** map analysis/synthesis. The map resolution parameter, nsmax, and the maximum value of the spherical harmonic order  $\ell_{max}$  must be specified.

#### **FORMAT**

% plmgen

## **QUALIFIERS**

2100	
nsmax =	Defines the resolution parameter for the map to be analysed/synthesized with the precomputed harmonics. (default= 32)
nlmax =	Defines the $\ell_{max}$ value for the execution. (default= 64)
$simul_type =$	Defines whether only scalar, or scalar and tensor harmonics are to be precomputed, 1=scalar only, 2=scalar AND tensor. (default=1)
outfile =	Defines the name for the file that will contain the precomputed harmonics. (default='plm.fits')

 $\overline{
m DESCRIPTION}$  The recursion of Legendre polynomials and tensor harmonics during the analysis and synthesis of **HEALPix** maps can be time consuming. Especially when repetitive applications are desired there is no need to compute the recursions every time. For such applications the values of  $P_{\ell m}(\theta)$  can be precomputed with plmgen and stored in a file. When using synfast or anafast this file can be read in to shorten the analysis/synthesis execution time.

> The memory (and disc) consumption of plmgen is  $8N_{\lambda}N_{p}$ bytes, with  $N_{\lambda} = \operatorname{nsmax}(\operatorname{nlmax} + 1)(\operatorname{nlmax} + 2)$  and  $N_p$  is either 1 or 3, depending whether tensor harmonics are computed.

> Currently an extra limitation  $N_{\lambda} < 2^{31} = 2147483648$  also applies, corresponding to, eg,  $lmax \le 1446$  for nsmax = 1024.

## **DATASETS**

The following datasets are involved in the **plmgen** processing.

Dataset	Description	
None required		

#### **SUPPORT**

This section lists those routines and facilities (including those *external* to the HEALPix distribution) which can assist in the utilisation of **plmgen**.

synfast This **HEALPix** facility can generate a map using precomputed harmonics made from plmgen.

anafast This **HEALPix** facility can analyse a map using precomputed harmonics.

plm\_gen Fortran subroutine used to generate the harmon-

## EXAMPLES: #1

plmgen

plmgen runs in interactive mode, self-explanatory.

ics

plmgen 31

## EXAMPLES: #2

plmgen filename

When 'filename' is present, plmgen enters the non-interactive mode and parses its inputs from the file 'filename'. This has the following structure: the first entry is a qualifier which announces to the parser which input immediately follows. If this input is omitted in the input file, the parser assumes the default value. If the equality sign is omitted, then the parser ignores the entry. In this way comments may also be included in the file. In this example, the file contains the following qualifiers:

```
simul_type = 1
nsmax = 32
nlmax = 86
outfile = plm.fits
```

Creates a binary FITS file called 'plm.fits' containing Legendre polynomials up to  $\ell$  and m values of 86 for a nsmax = 32 map. Legendre polynomials for all  $\ell$  and m values for each angle  $\theta$  corresponding to all of the **HEALPix** pixel center rings will be created.

#### RELEASE NOTES

Revision 1: Initial release HEALPix 1.00

## Messages

This section describes error messages generated by **plmgen**.

Message	Severity	Text
can not allocate memory for array xxx	Fatal	You do not have sufficient system resources to run this facility at the map resolution you required. Try a lower map resolution.
Error: these values of Nside and l_max are too large	Fatal	You are exceeding the limitation on Nside and l_max. Try a lower l_max.

smoothing 33

# smoothing

#### Location in HEALPix directory tree: src/f90/smoothing/smoothing.f90

This program can be used to convolve a map with a gaussian beam. The input map can be given in RING or NESTED scheme and the smoothed map is written to a FITS file in the RING scheme.

NOTE: This automated facility is susceptible to problems with non-commutativity of discrete spherical harmonics transforms, described in the Recommendations for Users of the anafast facility. If very high accuracy of the results is required in the spectral regime of  $\ell > 2 \cdot nsmax$ , it is recommended to choose an iterative computation of the  $a_{\ell m}$  coefficients.

## $\overline{FORMAT}$

% smoothing [options] [parameter\_file]

#### COMMAND LINE OPTIONS

-d

--double double precision mode (see Notes on double/single precision modes on page 3)

-s

--single single precision mode (default)

## **QUALIFIERS**

 $simul_type =$ Defines which map(s) to analyse, 1=temperature only, 2=temperature AND polarisation. (default = 1infile =Defines the filename for the FITS file containing the map to be smoothed. (default='map.fits') nlmax =Defines the  $\ell_{max}$  value for the application. (default = 64 $iter\_order =$ Defines the maximum order of quadrature iteration to be used. (default=0, no iteration)  $fwhm_arcmin =$ Defines the FWHM in arcminutes of the gaussian beam for the convolution. (default=10)

beam\_file = Defines the FITS file describing the Legendre window function of the circular beam to be used for

	the simulation. If set to an existing file name, it will override the fhwm_arcmin given above. default="'
outfile =	Defines the filename for the file that will contain the smoothed map. (default='map_smoothed.fits')
plmfile =	Defines the name for an input file containing pre- computed Legendre polynomials $P_{lm}$ . (default= no entry — anafast executes the recursive evalua- tion of $P_{lm}$ s)
w8file =	Defines name for an input file containing ring weights in the improved quadrature mode (default= no entry — the name is assumed to be 'weight_ring_n0xxxx.fits' where xxxx is nsmax)
w8filedir =	Gives the directory where the weight files are to be found (default= no entry — smoothing searches in the default directories, see introduction)
won =	Set this to 1 if weight files are to be used, otherwise set it to 0 (or 2). (default= $0$ )

# **DESCRIPTION** A FITS file containing a **HEALPix** map in RING or NESTED scheme is read in. The map is analysed and smoothed in fourier space with a gaussian beam of a given FHWM. A new map is then synthesized using the smoothed $a_{\ell m}$ coeffecients. For a more accurate application, an iteration of arbitrary order can be applied. The output map is stored in the same scheme as the input map.

## DATASETS

The following datasets are involved in the **smoothing** processing.

Dataset	Description
data/weight_ring_n0xxxx.fit	s Files containing ring weights for the smoothing improved quadrature mode.

smoothing 35

## **SUPPORT**

This section lists those routines and facilities (including those *external* to the HEALPix distribution) which can assist in the utilisation of **smoothing**.

synfast This **HEALPix** facility can generate a map and

also do the smoothing.

anafast This **HEALPix** facility can analyse a smoothed

map.

## EXAMPLES: #1

smoothing

Smoothing runs in interactive mode, self-explanatory.

## EXAMPLES: #2

#### smoothing filename

When 'filename' is present, smoothing enters the non-interactive mode and parses its inputs from the file 'filename'. This has the following structure: the first entry is a qualifier which announces to the parser which input immediately follows. If this input is omitted in the input file, the parser assumes the default value. If the equality sign is omitted, then the parser ignores the entry. In this way comments may also be included in the file. In this example, the file contains the following qualifiers:

```
simul_type = 1
nlmax = 64
infile = map.fits
outfile = map_smoothed.fits
fwhm_arcmin = 10.
iter_order = 1
```

Smoothes the **HEALPix** temperature map contained in 'map.fits' with a 10 arcmin FWHM beam. The resulting map is saved in 'map\_smoothed.fits'. The map analysis/synthesis was carried out using fourier coeffecients up to an  $\ell$  value of 64. A first order iteration of the quadrature was performed.

#### RELEASE NOTES

Revision 1: Initial release (HEALPix 0.90)

**Revision 2:** Extension to polarization and arbitrary *circular* beams (**HEALPix** 1.20)

smoothing 37

# Messages

This section describes error messages generated by **smoothing**.

Message	Severity	Text
can not allocate memory for array xxx	Fatal	You do not have sufficient system resources to run this facility at the map resolution you required. Try a lower map resolution.

#### Location in HEALPix directory tree: src/f90/ngsims\_full\_sky/sky\_ng\_sim.F90

This program can be used to create temperature **HEALPix** maps computed as realisations of random *Non-Gaussian* fields on a sphere (either even power of a Gaussian distribution, or Simple Harmonics Oscillator PDF, see Description section for details).

It is directly adapted from the NGSIMS code described in Rocha et al, MNRAS, 357, 1 (2005)

The operation count is dominated by a term scaling as  $\mathcal{O}(N_{\rm pix}^{1/2}\ell_{\rm max}^2)$ . The map angular power spectrum, resolution, Gaussian beam FWHM or arbitrary beam window and random seed for the simulation can be selected by the user.

#### **FORMAT**

# % sky\_ng\_sim [parameter\_file]

# **QUALIFIERS**

infile = Defines the input power spectrum file, (default= HEALPIX/test/cl.fits).

outfile = Defines the output map file, (default= test.fits).

nsmax = Defines the resolution of the map. (default= 32)

nlmax = Defines the maximum  $\ell$  value to be used in the simulation. WARN-ING:  $\ell_{max}$  can not exceed the value  $4 \cdot nsmax$ , because the coefficients of the average Fourier pixel window functions are precomputed and provided up to this limit. (default= 2\*nsmax)

fwhm\_arcmin = Defines the FWHM size in arcminutes of the simulated Gaussian beam. (default= 0.0)

beam\_file = Defines the FITS file describing the Legendre window function of the circular beam to be used for the simulation. If set to an existing file name, it will override the fhwm\_arcmin given above. (default="')

windowfile = Defines the input filename for the pixel smoothing windows (default= pixel\_window\_n????.fits, see Notes on default files and directories on page 3)

winfiledir = Defines the directory in which windowfile is located (default : see Notes on default files and directories on page 3).

iseed = Defines the seed of the pseudo-random sequence to be used for the generation of the non-gaussian white noise (default= 1)

plot = If sky\_ng\_simwas linked with the PGPLOT library during compilation, and if plot is set to (case unsensitive) .true., t, yes, y or 1, then the histogram of the simulated non-gaussian is produced, overplotted with the theoretical PDF; the histogram of the final map pixel values, overplotted with a PDF of a gaussian of same mean and variance is subsequently produced. (default=.false.)

pdf\_choice=1 Choice of non-Gaussian PDF to use: 1= Simple Harmonics oscillator (see Eq 2 below)

sigma0= RMS of oscillator ground state

na= Integer in  $\{0, 20\}$ . Number of  $\alpha$  coefficients to be given (default=3). Note: analytical calculation of the PDF moments can only be done for  $\mathtt{na} \leq 3$ .

alpha\_1=, alpha\_2=, ... Real values of  $\alpha_i$  coefficients for i in [1, na]

pdf\_choice=2 Choice of non-Gaussian PDF to use: 2=Power of a Gaussian (see Eq 3 below)

npower = Positive integer in {1,4} (default=1). The gaussian will be set to the power 2\*npower.

# DESCRIPTION A random non-Gaussian white noise map is generated, using either

• a simple linear harmonic oscillator, where the PDF of the pixel temperature t is

$$\rho_{\text{SHO}}(t) = |\psi_n|^2 = e^{-t^2/2\sigma_0^2} \left| \sum_{i=0}^n \alpha_i C_i H_i \left( \frac{t}{\sqrt{2}\sigma_0} \right) \right|^2$$
 (2)

where  $H_i$  are the Hermite polynomials,  $C_i$  their normalization constants,  $\sigma_0^2$  the variance of the (Gaussian) ground state  $|\psi_0|^2$ ,  $\alpha_i$  for  $i \geq 1$  are free parameters, while  $\alpha_0 = (1 - \sum_i^n |\alpha_i|^2)^{1/2}$  is constrained;

• or, an even power of a gaussian PDF, where the temperature of pixel q is

$$t_a = g_a^{2P} \tag{3}$$

where g is a zero mean, unit variance Gaussian variable, and P is a user-defined positive integer.

The resulting map is analyzed into its  $a_{lm}$  coefficients, which are then multiplied by the beam, pixel and spectrum window

$$a_{lm} \longrightarrow a_{lm} \left[ C(l) \right]^{1/2} B(l) w_{pix}(l)$$
 (4)

The resulting  $a_{lm}$  coefficients are turned back into a map, which is therefore non-gaussian, with an effective angular power spectrum  $C(l)B(l)^2w_{\text{pix}}(l)^2$  (Rocha et al, 2005).

Notes: the code has been modified from the original NGSIMS package in two respects: the seed parameter is named iseed instead of idum, to be consistent with other **HEALPix** simulation codes; and the SHO generator has been dramatically sped up, without loss of accuracy.

#### **DATASETS**

The following datasets are involved in the **sky\_ng\_sim** processing.

Dataset	Description
/data/pixel_window_r	nxxxx.fitFiles containing pixel windows for various nsmax.

# **SUPPORT**

This section lists those routines and facilities (including those external to the HEALPix distribution) which can assist in the utilisation of **sky\_ng\_sim**.

map2gif This **HEALPix** Fortran facility can be used to visualise the output map.

This **HEALPix** IDL facility can be used to visualise the output map.

This **HEALPix** Fortran facility can analyse a **HEALPix** map and save the  $a_{lm}$  and  $C_{l}$ s to be

read by sky\_ng\_sim.

# EXAMPLES: #1

 $sky_ng_sim$ 

sky\_ng\_sim runs in interactive mode, self-explanatory.

# EXAMPLES: #2

sky\_ng\_sim filename

When 'filename' is present, sky\_ng\_sim enters the non-interactive mode and parses its inputs from the file 'filename'. This has the following structure: the first entry is a qualifier which announces to the parser which input immediately follows. If this input is omitted in the input file, the parser assumes the default value. If the equality sign is omitted, then the parser ignores the entry. In this way comments may also be included in the file. In this example, the file contains the following qualifiers:

```
nsmax = 128
nlmax = 256
fwhm_arcmin = 30.0
infile = cl.fits
pdf_choice = 1
iseed = 1
na = 3
sigma0 = 1.0
alpha_1 = 0.0
alpha_2 = 0.0
alpha_3 = 0.2
outfile = !test_ngfs.fits
```

sky\_ng\_sim reads in the  $C_l$  power spectrum in 'cl.fits' up to l256, and produces the map 'map.fits' which has  $N_{\rm side} = 128$ . The non-gaussian white noise was generated assuming a SHO PDF (see Eq 2 above) with  $\sigma_0 = 1$  and  $\alpha_i = (0,0,0.2)$ . The map is convolved with a beam of FWHM 30.0 arcminutes. The iseed = 1 sets the random seed for the realisation. A different iseed would have given a different realisation from the same power spectrum.

Since

beam\_file

windowfile

were omitted, they take their default values (empty strings). This means that no beam were read and sky\_ng\_sim attempts to find the pixel window files in the default directories (see page 3).

# RELEASE NOTES

Revision 1: Initial release (HEALPix 2.10)

# $\overline{\text{Messages}}$

This section describes error messages generated by **sky\_ng\_sim**.

Message	Severity	Text
can not allocate memory for array xxx	Fatal	You do not have sufficient system resources to run this facility at the map resolution you required. Try a lower map resolution.
this is not a binary ta- ble		the fitsfile you have specified is not of the proper format
there are undefined values in the table!		the fitsfile you have specified is not of the proper format
the header in xxx is too long		the fitsfile you have specified is not of the proper format
XXX-keyword not found		the fitsfile you have specified is not of the proper format
found xxx in the file, expected:yyyy		the specified fitsfile does not contain the proper amount of data.

#### Location in HEALPix directory tree: src/f90/synfast/synfast.f90

This program can be used to create **HEALPix** maps (temperature only or temperature and polarisation) computed as realisations of random Gaussian fields on a sphere characterized by the user provided theoretical power spectra, or as constrained realisations of such fields characterised by the user provided  $a_{\ell m}$  coefficients and/or power spectra. Total operation count scales as  $\mathcal{O}(N_{\rm pix}^{3/2})$  with a prefactor dependent on the limiting spherical harmonics order  $\ell_{max}$  of the actual problem. The map resolution, Gaussian beam FWHM, and random seed for the simulation can be selected by the user. Spherical harmonics are either generated using the recurrence relations during the execution of spectral synthesis, or precomputed and read in before the synthesis is executed to shorten the computation in repetitive applications.

# **FORMAT**

% synfast [options] [parameter\_file]

# **COMMAND LINE OPTIONS**

-d

--double double precision mode (see Notes on double/single precision modes on page 3)

-s

--single single precision mode (default)

### **QUALIFIERS**

infile = Defines the input power spectrum file, (default= cl.fits). Note that infile is now optional: synfast can run even if only almsfile is provided.

outfile = Defines the output map file, (default= map.fits). Note that outfile is now optional: if it set to " (empty string), mo map is synthesized but the  $a_{lm}$  generated can be output.

outfile\_alms = Defines the FITS file in which to output  $a_{lm}$  used for the simulation (default= '')

simul_type =	Defines the simulation type, 1=temperature only, 2=polarisation, 3=temperature and its first spatial derivatives, 4=temperature and its first and second spatial derivatives, 5=temperature and polarisation, and their first and second derivatives, 6=same as 5 plus the second derivatives of $(T,Q,U)$ . (default= 1).
nsmax =	Defines the resolution of the map. (default= 32)
nlmax =	Defines the maximum $\ell$ value to be used in the simulation. WARN-ING: $\ell_{max}$ can not exceed the value $4\cdot nsmax$ , because the coefficients of the average Fourier pixel window functions are precomputed and provided up to this limit. (default= 64)
iseed =	Defines the random seed to be used for the generation of $a_{lm}$ s from the power spectrum. (default= -1)
fwhm_arcmin =	Defines the FWHM size in arcminutes of the simulated Gaussian beam. (default= $420.0)$
beam_file =	Defines the FITS file describing the Legendre window function of the circular beam to be used for the simulation. If set to an existing file name, it will override the ${\tt fhwm\_arcmin}$ given above. (default='')
almsfile =	Defines the input filename for a file containing $a_{lm}$ s for constrained realisations. (default= ''). If apply_windows is false those $a_{lm}$ s are used as they are, without being multiplied by the beam or pixel window function (with the assumption that they already have the correct window functions). If apply_windows is true, the beam and pixel window functions chosen above are applied to the constraining $a_{lm}$ (with the assumption that those are free of beam and pixel window function). The code does not check the validity of these asumptions; if none is true, use the alteralm facility to modify or remove the window functions contained in the constraining $a_{lm}$ .
apply_windows=	Determines how the constraining $a_{lm}$ read from almsfile are treated with respect to window functions; see above for details. y, yes, t, true, .true. and 1 are considered as $true$ , while n, no, f, false, .false. and 0 are considered as $false$ , (default = .false.).
plmfile =	Defines the input filename for a file containing precomputed Legendre polynomials $P_{lm}$ . (default= '')
windowfile =	Defines the input filename for the pixel smoothing windows (default= pixel_window_n?????.fits, see Notes on default files and directories on page 3)
${\rm winfiledir} =$	Defines the directory in which windowfile is located (default : see Notes on default files and directories on page $3$ ).

DESCRIPTION Synfast reads the power spectrum from a file in ascii FITS format. This can contain either just the temperature power spectrum  $C_l^T$ s or temperature and polarisation power spectra:  $C_l^T$ ,  $C_l^E$ ,  $C_l^B$  and  $C_l^{T \times E}$ . If simul\_type = 2 synfast generates Q and U maps as well as the temperature map. The output map(s) is (are) saved in a FITS file.  $C_l$ s are used up to the specified  $\ell_{lmax}$ , which can not exceed  $4 \times$  nsmax. If simul\_type = 3 or 4 the first derivatives of the temperature field or the first and second derivatives respectively are output as well as the temperature itself: T(p),  $(\partial T/\partial \theta, \partial T/\partial \phi/\sin \theta)$ ,  $(\partial^2 T/\partial \theta^2, \partial^2 T/\partial \theta \partial \phi/\sin \theta)$ ,  $\partial^2 T/\partial \phi^2/\sin^2 \theta$ ). If simul\_type = 5 or 6 the first derivatives of the (T,Q,U) fields or the first and second derivatives respectively are output as well as the field themself:  $T(p), Q(p), U(p), (\partial T/\partial \theta, \partial Q/\partial \theta, \partial U/\partial \theta; \partial T/\partial \phi/\sin \theta, \ldots),$  $(\partial^2 T/\partial \theta^2, \dots; \partial^2 T/\partial \theta \partial \phi/\sin \theta, \dots; \partial^2 T/\partial \phi^2/\sin^2 \theta \dots)$ 

The random sequence seed for generation of  $a_{lm}$  from the power spectrum should be non-zero integer. If 0 is provided, a seed is generated randomly by the code, based on the current date and time. The map can be convolved with a gaussian beam for which a beamsize can be specified, or for an arbitrary circular beam for which the Legendre transform is provided. The map is automatically convolved with a pixel window function. These are stored in FITS files in the healpix/data directory. If synfast is not run in a directory which has these files, or from a directory which can reach these files by a '.../data/' or './data/' specification, the system variable HEALPIX is used to locate the main **HEALPix** directory and its data subdirectory is scanned. Failing this, the location of these files must be specified (using winfiledir). In the interactive mode this is requested only when necessary (see Notes on default directories on page 3).

If some of the  $a_{lm}$  in the simulations are constrained eg. from observations, a FITS file with these  $a_{lm}$  can be read. This FITS file contains the  $a_{lm}$  for certain  $\ell$  and m values and also the standard deviation for these  $a_{lm}$ . The sky realisation which synfast produces will be statistically consistent with the constraining  $a_{lm}$ .

The code can also be used to generate a set of  $a_{lm}$  matching the input power spectra, beam size and pixel size with or without actually synthesizing the map. Those  $a_{lm}$  can be used as an input (constraining  $a_{lm}$ ) to another synfast run.

Spherical harmonics values in the synthesis are obtained from a recurrence on associated Legendre polynomials  $P_{lm}(\theta)$ . This recurrence consumes most of the CPU time used by synthest. We have therefore included the option to read in the precomputed values of  $P_{lm}(\theta)$  from a file generated by the **HEALPix** facility

Synfast will issue a warning if the input FITS file for the power spectrum does not contain the keyword POLNORM. This keyword indicates that the convention used for polarization is consistent with CMBFAST (and consistent with **HEALPix** 1.2). See the **HEALPix** Primer for details on the polarization convention and the interface with CMBFAST. If the keyword is not found, no attempt will be made to renormalize the power spectrum. If the keyword is present, it will be inherited by the simulated map.

#### **DATASETS**

The following datasets are involved in the **synfast** processing.

Dataset	Description		
	C(D)		

/data/pixel\_window\_nxxxx.fitFiles containing pixel windows for various nsmax.

#### **SUPPORT**

This section lists those routines and facilities (including those *external* to the HEALPix distribution) which can assist in the utilisation of **synfast**.

map2gif	This <b>HEALPix</b> Fortran facility can be used to visualise the output map.
mollview	This <b>HEALPix</b> IDL facility can be used to visualise the output map.
alteralm	This <b>HEALPix</b> Fortran facility can be used to implement the beam and pixel window functions on the constraining $a_l ms$ (almsfile file).
anafast	This <b>HEALPix</b> Fortran facility can analyse a <b>HEALPix</b> map and save the $a_{lm}$ and $C_l$ s to be read by synfast.
plmgen	This <b>HEALPix</b> Fortran facility can be used to generate precomputed Legendre polynomials.

# EXAMPLES: #1

synfast

Synfast runs in interactive mode, self-explanatory.

# EXAMPLES: #2

#### synfast filename

When 'filename' is present, synfast enters the non-interactive mode and parses its inputs from the file 'filename'. This has the following structure: the first entry is a qualifier which announces to the parser which input immediately follows. If this input is omitted in the input file, the parser assumes the default value. If the equality sign is omitted, then the parser ignores the entry. In this way comments may also be included in the file. In this example, the file contains the following qualifiers:

```
simul_type = 1
nsmax = 32
nlmax = 64
iseed = -1
fwhm_arcmin = 420.0
infile = cl.fits
outfile = map.fits
```

Synfast reads in the  $C_l$  power spectrum in 'cl.fits' up to l = 64, and produces the map 'map.fits' which has  $N_{\rm side} = 32$ . The map is convolved with a beam of FWHM 420.0 arcminutes. The iseed = -1 sets the random seed for the realisation. A different iseed would have given a different realisation from the same power spectrum.

Since
outfile\_alms
almsfile
apply\_windows
plmfile
beam\_file
windowfile

were omitted, they take their default values (empty strings). This means that no file for constrained realisation or precomputed Legendre polynomials are read, the  $a_{lm}$  generated in the process are not output, and synfast attempts to find the pixel window files in the default directories (see page 3).

## RELEASE

#### **NOTES**

Revision 1: Initial release (HEALPix 0.90)

Revision 2: Optional non-interactive operation. Proper FITS file support. Improved reccurence algorithm for  $P_{lm}(\theta)$  which can compute to higher  $\ell$  values. Improved pixel windows averaged over actual HEALPix pixels. New functionality: constrained realisations, precomputed  $P_{lm}$ . (HEALPix 1.00)

Revision 3: New functionality: constrained realisations and pixel windows are now available for polarization as well. Arbitrary circular beams can be used. New parser (HEALPix 1.20)

**Revision 4:** New functionnality: the generated  $a_{lm}$  can be output, and the map synthesis itself can be skipped. First and second derivatives of the temperature field can be produced on demand.

# Messages

This section describes error messages generated by **synfast**.

Message	Severity	Text
can not allocate memory for array xxx	Fatal	You do not have sufficient system resources to run this facility at the map resolution you required. Try a lower map resolution.
this is not a binary ta- ble		the fitsfile you have specified is not of the proper format
there are undefined values in the table!		the fitsfile you have specified is not of the proper format
the header in xxx is too long		the fitsfile you have specified is not of the proper format
XXX-keyword not found		the fitsfile you have specified is not of the proper format
found xxx in the file, expected:yyyy		the specified fitsfile does not contain the proper amount of data.

# ud\_grade

Location in HEALPix directory tree: src/f90/ud\_grade/ud\_grade.f90

This program can upgrade or degrade the resolution of a HEALPix map.

#### **FORMAT**

% ud\_grade [options] [parameter\_file]

#### COMMAND LINE OPTIONS

-d

--double double precision mode (see Notes on double/single precision modes on page 3)

single precision mode (default) --single

### **QUALIFIERS**

 $nside\_out =$ Defines the resolution parameter for the output

map. (default = 64)

infile =Defines the name of the file containing the map to

be up/degraded. (default='map.fits')

outfile =Defines the filename for the output up/degraded

map. (default='outmap.fits')

**DESCRIPTION** This facility transforms the resolution of an input **HEALPix** map. At each step of map resolution upgrade the four output map pixels nested in one pixel of the input map are given the values of the input pixel. At each step of map resolution degradation the four input map pixels nested in one output map pixel are averaged to produce the pixel value in the output map. **Caution** Beware that, at this stage, the parallel transport of the polarization (Q and U Stokes vectors) that would be necessary to describe the change in local coordinates is **not** implemented.

# **DATASETS**

The following datasets are involved in the **ud\_grade** processing.

 $ud\_grade$  53

Dataset	Description
None required	

# **SUPPORT**

This section lists those routines and facilities (including those external to the HEALPix distribution) which can assist in the utilisation of  $\mathbf{ud\_grade}$ .

mollview IDL routine to view an up/downgraded map.

This **HEALPix** facility can analyse an up/downgraded map.

# EXAMPLES: #1

ud\_grade

ud\_grade runs in interactive mode, self-explanatory.

# EXAMPLES: #2

ud\_grade filename

When 'filename' is present, ud\_grade enters the non-interactive mode and parses its inputs from the file 'filename'. This has the following structure: the first entry is a qualifier which announces to the parser which input immediately follows. If this input is omitted in the input file, the parser assumes the default value. If the equality sign is omitted, then the parser ignores the entry. In this way comments may also be included in the file. In this example, the file contains the following qualifiers:

nside\_out = 64
infile= map.fits
outfile = outmap.fits

Ud\_grade transforms the **HEALPix** map in 'map.fits' to resolution Nside = 64, no matter what the input map resolution was. The up/downgraded map is stored in 'outmap.fits'.

#### RELEASE NOTES

Revision 1: (Initial release HEALPix 0.90)

Revision 2: Extension to multi-dimensional maps (HEALPix 1.20)

# Messages

This section describes error messages generated by ud\_grade.

Message	Severity	Text
can not allocate memory for array xxx	Fatal	You do not have sufficient system resources to run this facility at the map resolution you required. Try a lower map resolution.