

Holographic duality beyond AdS/CFT

Wei Song^a

^aYau Mathematical Sciences Center, Tsinghua University, Beijing 100084, China

ABSTRACT: Lecture notes for the Southeast Summer School on Strings and Stuff, 2021.

Contents

0	Introduction	1
1	Brief review of $\text{AdS}_3/\text{CFT}_2$	2
2	Bottom-up approach: from asymptotic symmetry	2
3	Top-down approach: from string theory	2
3.1	The D1-D5- P system and its IR limit	2
3.1.1	Closed string picture: gravity on AdS_3 background	2
3.1.2	Open string picture: worldvolume CFT_2 and the duality	4
3.2	Type IIB string theory with NS-NS flux	5
3.2.1	The worldsheet CFT_2	6
4	Holographic Dualities Beyond $\text{AdS}_3/\text{CFT}_2$	6
4.1	Top-Down Approach: Deformations of $\text{AdS}_3/\text{CFT}_2$ in string theory	7
4.2	$T\bar{T}$ deformation	8

0 Introduction

Since the discovery of AdS/CFT duality [1], we have greatly furthered our understanding of quantum gravity in asymptotically AdS backgrounds. However, there are plenty of non-AdS geometries in our real worlds, including:

- The Kerr metric of a rotation black hole, **which ...** ;
- The asymptotically flat / Minkowski spacetime, which well approximates our current universe at a smaller length (and time) scale;
- The asymptotically de Sitter spacetime, which well approximates our current (and future) universe at a larger length scale, and also during the period of inflation;
- The FRW metric (or more precisely, the Friedmann-Lemaître-Robertson-Walker [2–5] metric), which describes the evolution of our homogeneous and isotropic universe from the big bang to **its ...** ;
- and more ...

Much less is known about quantum gravity in these backgrounds.

1 Brief review of $\text{AdS}_3/\text{CFT}_2$

2 Bottom-up approach: from asymptotic symmetry

3 Top-down approach: from string theory

String theory is a self-consistent theory of quantum gravity.

The first example of a microscopic counting of black hole entropy, discovered by Strominger-Vafa [6], comes from the D1-D5- P system in string theory.

The first incarnation of holographic principle was realized by Maldacena [1], by a stack of D3 branes in type IIB string theory.

3.1 The D1-D5- P system and its IR limit

Let us look at the D1-D5- P brane configuration in type IIB string theory. This is well-reviewed in [7]. This configuration allows for an open string description and a closed string description.

Geometry	$\mathbb{R}^{4,1}$			S^1	$\mathcal{M}_4 = T^4, \text{K3}$			
Direction	0	1, 2, 3, 4		5	6	7	8	9
# D5 = Q_5	×			×	×	×	×	×
# D1 = Q_1	×			×				
P	×			×				

Table 1. Brane configuration of the D1-D5- P system. Here we are considering type IIB string theory on flat 6D spacetime, with a compactified $x^5 \in S^1$ direction, along with an internal \mathcal{M}_4 manifold. We use “ \times ” to mark the directions x^μ that an object occupies. Here $\mu = 0, 1, \dots, 9$.

The D5 branes wrap the compact \mathcal{M}_4 , while the D1 branes are localized on \mathcal{M}_4 . Both the D1 and D5 branes extend along the fifth direction x^5 , which is compactified to a circle S^1 with a large radius.

Open string excitations on the branes carry momentum and winding. Due to the large radius of S^1 , we can focus on the momentum modes P along $x^5 \in S^1$ and neglect the winding modes. On the other hand, we will neglect momentum modes along the \mathcal{M}_4 directions, since \mathcal{M}_4 is assumed to be compact and small.

3.1.1 Closed string picture: gravity on AdS_3 background

In the IR limit, type IIB string theory is described by the low energy effective action of type IIB supergravity. The field content and the action of type IIB supergravity are well reviewed in the literature; see e.g. Appendix H of [8]. In particular, there is a pair of

2-form gauge potentials in type IIB supergravity. One of them is the NS-NS field B_2 , and the other is the R-R field C_2 . The D1 branes are electrically charged under C_2 , while the D5 branes are magnetically charged under C_2 .

The bosonic part of the string frame action is then given by [citations needed]:

$$\frac{1}{16\pi G} \int d^{10}x \sqrt{-g} \left(e^{-2\phi} \left(R + 4(\nabla\phi)^2 - \frac{1}{12} H^2 \right) - \frac{1}{12} F^2 \right), \quad (3.1)$$

$$H = dB_2, \quad F = dC_2 \quad (3.2)$$

where H and F are the 3-form field strengths, $H^2 = H_{\mu\nu\rho} H^{\mu\nu\rho} \propto H \wedge \star H$, and similar for F^2 . After dimension reduction of the compact \mathcal{M}_4 , the equations of motion admit a *black string* solution in 6D, where the metric is given by [citations needed]:

WL: convention for the coefficients?

$$ds^2 = (f_1 f_5)^{-1/2} \left(-dt^2 + d\phi^2 + \frac{r_0^2}{r^2} (\cosh \sigma dt + \sinh \sigma d\phi)^2 \right) + (f_1 f_5)^{+1/2} \left(\frac{dr^2}{1 - r_0^2/r^2} + r^2 d\Omega_3^2 \right), \quad \phi \cong \phi + 2\pi R, \quad (3.3)$$

$$\text{where } f_1 = 1 + \frac{r_1^2}{r^2}, \quad f_5 = 1 + \frac{r_5^2}{r^2} \quad (3.4)$$

The parameters in this supergravity solution can be related to the brane construction as follows:

- $\phi \equiv x^5$ is the compactified S^1 direction along the D1 brane, normalized such that $\phi \cong \phi + 2\pi R$, where R is the large radius of the S^1 circle.

Upon dimension reduction of the ϕ direction, this 6D black string solution will become a 5D black hole solution. In fact the resulting 5D black hole solution is precisely the Strominger-Vafa black hole [6], which serves as the first example of a microscopic counting of the black hole entropy.

- r_0 marks the horizon of the black string, and it is related to the open string momentum P attached to the branes: $P \propto r_0^2 \sinh 2\sigma$.
- r_1^2 and r_5^2 are related to the charges Q_1 and Q_5 .

We further note that the above black string solution is asymptotically flat, consistent with our brane construction in string theory. On the other hand, if we zoom in to the near horizon region of this black string solution, we discover an $\text{AdS}_3 \times S^3$ geometry. This can be achieved by setting:

$$\ell^2 = r_1 r_5, \quad r \mapsto \lambda \ell r, \quad r_0 \mapsto \lambda \ell r_0, \quad t \mapsto t \ell / \lambda, \quad \phi \mapsto \phi \ell / \lambda, \quad (3.5)$$

where ℓ is the AdS radius, and now the ϕ coordinate is normalized such that $\phi \cong \phi + 2\pi$. More specially,

- For extremal black string with $r_0 = 0$ and thus $P = 0$, the near horizon limit leads to the zero mass BTZ geometry, with an additional S^3 factor:

$$ds^2 = \ell^2 \left(r^2 (-dt^2 + d\phi^2) + \frac{dr^2}{r^2} + d\Omega_3^2 \right) \quad (3.6)$$

- For the near-extremal case with generic r_0, σ , the near horizon limit leads to the rotating BTZ geometry, again with an additional S^3 factor:

$$ds^2 = \ell^2 \left(r^2 (-dt^2 + d\phi^2) + \frac{dr^2}{r^2 - r_0^2} + r_0^2 (\cosh \sigma dt + \sinh \sigma d\phi)^2 + d\Omega_3^2 \right) \quad (3.7)$$

It is convenient to define the left and right-moving temperature:

$$T_L = \frac{1}{2\pi} \frac{r_0 e^\sigma}{\ell^2}, \quad T_R = \frac{1}{2\pi} \frac{r_0 e^{-\sigma}}{\ell^2} \quad (3.8)$$

On the other hand, the Hawking temperature T_H of this solution can be computed, and is given in terms of T_L, T_R as follows:

$$\frac{2}{T_H} = \frac{1}{T_L} + \frac{1}{T_R} \quad (3.9)$$

Within the framework of string theory, one can understand the IR black string geometry as the result of “integrating out” the dynamics of the branes, which includes the open string excitations. This process deforms the background geometry, and we end up with a closed string theory on the black string background. The near horizon limit brings us further to the IR fixed point, where the far region dynamics decouple and we are left with the AdS_3 geometry.

WL: accurate?

In other words, strings on $\text{AdS}_3 \times S^3 \times \mathcal{M}_4$ is understood as the *closed string description* of the IR fixed point of the D1-D5- P system.

3.1.2 Open string picture: worldvolume CFT_2 and the duality

On the other hand, we can consider the worldvolume theory of the brane construction. This provides the *open string description* of the D1-D5- P system.

After dimension reduction of the compact \mathcal{M}_4 , we have a $(1+1)$ dimensional QFT living on the D1-D5 branes. This is a supersymmetric gauge theory with $\mathcal{N} = (4, 4)$ supersymmetry. Similar to our previous discussions, we can consider the IR limit of this system. It should flow to an IR fixed point, which is a $(1+1)$ dimensional superconformal field theory (SCFT_2).

The central charge of this SCFT_2 can be read off from the field contents of the world-volume theory: the number of bosonic fields is given by [citations needed]:

$$4Q_1Q_5 \quad (3.10)$$

and same for the fermions. The factor Q_1Q_5 comes from the open string excitations between the D1 and D5 branes, while the factor of 4 comes from the fact that the D1 branes can

move inside the D5 branes **along the compact \mathcal{M}_4 directions**. In the end we have the central charge:

$$c = 1 \times 4Q_1Q_5 + \frac{1}{2} \times 4Q_1Q_5 = 6Q_1Q_5 \quad (3.11)$$

WL: is this language accurate?

We have thus discovered two equivalent descriptions of the D1-D5- P system, summarized as follows:

$$\begin{array}{ccc} \left(\begin{array}{c} \text{Far-region dynamics} \\ + \\ \text{Strings on } \text{AdS}_3 \times S^3 \times \mathcal{M}_4 \end{array} \right) & = & \left(\begin{array}{c} \text{Far-region dynamics} \\ + \\ \text{SCFT}_2 \text{ with target } \mathcal{M}_4 \end{array} \right) \\ \text{Closed string picture} & & \text{Open string picture} \end{array}$$

The far-region dynamics decouple on both sides, and we have the following proposal of an $\text{AdS}_3/\text{CFT}_2$ duality **[citations needed]**:

$$\begin{aligned} & \text{Type IIB string theory on } \text{AdS}_3 \times S^3 \times \mathcal{M}_4 \\ & = \text{Marginal deformations of some } \mathcal{N} = (4, 4) \text{ SCFT}_2 \text{ with target } \mathcal{M}_4 \end{aligned} \quad (3.12)$$

Note that both sides of the duality includes a non-trivial moduli **[citations needed]**. In particular, the SCFT_2 on the right-hand side is proposed to be a *symmetric orbifold theory* **[citations needed]**, together with its exactly marginal deformations along the conformal manifold.

WL: elaborate?

3.2 Type IIB string theory with NS-NS flux

Due to its non-trivial coupling to the R-R field C_2 , the D1-D5 system is difficult to deal with on the worldsheet, within the framework of the RNS (or NSR) formalism. This difficulty can be avoided by considering its S-dual, the NS1-NS5 system.

First let us recall the $\text{SL}(2, \mathbb{Z})$ duality of type IIB string theory, reviewed in §12 and §14 of [9]. In particular, the strong-weak S-duality is one of the generators of the $\text{SL}(2, \mathbb{Z})$ action. When acted on the field strengths $H = dB_2$ and $F = dC_2$, it's given by:

$$S: \begin{pmatrix} H \\ F \end{pmatrix} \mapsto \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix} \begin{pmatrix} H \\ F \end{pmatrix} = \begin{pmatrix} F \\ -H \end{pmatrix}, \quad (3.13)$$

$$H = dB_2, \quad F = dC_2$$

We note that S-duality exchanges the field strengths H and F . On the other hand, the D1-D5 system couples to C_2 (electrically for D1, and magnetically for D5), while the NS1-NS5 system couples to B_2 . Therefore S-duality corresponds to a map:

$$S: \begin{array}{ll} Q_5 \text{ D5 branes} & \mapsto k = Q_5 \text{ NS5 branes} \\ Q_1 \text{ D1 branes} & \mapsto p = Q_1 \text{ NS1 branes} \end{array} \quad (3.14)$$

We can then repeat the analysis in §3.1 and obtain a similar black string solution in 6D. By taking the near horizon limit, we further obtain an asymptotically AdS_3 geometry,

namely the BTZ solutions as described in §3.1, but now with NS-NS flux instead of R-R flux. In particular, for $P = 0$ we have the massless BTZ solution [citations needed]:

$$ds^2 = k (d\rho^2 + e^{2\rho} d\gamma d\bar{\gamma}), \quad (3.15)$$

$$B_2 = \frac{1}{2} B_{\mu\nu} dx^\mu \wedge dx^\nu = -\frac{k}{2} e^{2\rho} d\gamma \wedge d\bar{\gamma}, \quad e^{2\Phi} = \frac{k}{p}, \quad (3.16)$$

where we note that there is a non-trivial B -field and a dilaton profile Φ .

Why are we interested in this S-dual system? As we've mentioned before, unlike the original D1-D5 system, the NS1-NS5 system admits a worldsheet description as an RNS superstring theory. Therefore we have a well understood *worldsheet* CFT_2 for the bulk string theory. Furthermore, based on this worldsheet CFT_2 , we are able to construct explicitly the *dual* CFT_2 on the asymptotic boundary of the spacetime. This is sometimes referred to as the *spacetime* CFT_2 . Therefore the NS1-NS5 system serves as a realization of $\text{AdS}_3/\text{CFT}_2$ where both sides of the holographic duality are, in some sense, under control.

3.2.1 The worldsheet CFT_2

The bosonic action of the worldsheet theory is given by the familiar non-linear sigma model [citations needed]:

$$\begin{aligned} S &= -\frac{1}{4\pi\alpha'} \int d^2z \sqrt{-\eta} (\eta^{ab} G_{\mu\nu}(X) + \epsilon^{ab} B_{\mu\nu}(X)) \partial_a X^\mu \partial_b X^\nu \\ &= \frac{k}{2\pi} \int d^2z (\partial\rho \bar{\partial}\rho + e^{2\rho} \partial\bar{\gamma} \bar{\partial}\gamma) \end{aligned} \quad (3.17)$$

WL:
conventions?

where in the second line we've restricted to the AdS_3 part and plugged in the BTZ solution given in (3.15) and (3.16). We see that ρ is a Liouville field. We can further introduce an auxiliary field β , which leads us to the Wakimoto representation [citations needed]:

$$S[\rho, \gamma, \bar{\gamma}, \beta, \bar{\beta}] = \frac{k}{2\pi} \int d^2z (\partial\rho \bar{\partial}\rho + \beta \bar{\partial}\gamma + \bar{\beta} \partial\bar{\gamma} - \beta \bar{\beta} e^{-2\rho} \bar{\partial}\bar{\gamma} \partial\gamma) \quad (3.18)$$

WL:
conventions?

It returns to the original action after we integrate out β . From this action it is clear that at large radius (large ρ), the last term in the action can be dropped, and the theory is well approximated by free fields, including a free scalar ρ , a pair of left-moving free fermions (β, γ) , and a pair of right-moving ones $(\bar{\beta}, \bar{\gamma})$.

An alternative representation is given in terms of the $\widehat{\text{SL}}(2, \mathbb{R})_k$ WZW model.

4 Holographic Dualities Beyond $\text{AdS}_3/\text{CFT}_2$

In the previous sections, we have laid down some basic aspects of $\text{AdS}_3/\text{CFT}_2$ dualities in string theory. Starting from this section, we are going to introduce recent salient developments on holographic dualities beyond $\text{AdS}_3/\text{CFT}_2$. In particular, we will focus on $T\bar{T}$ deformation. This section is organized as follows:

4.1 Top-Down Approach: Deformations of $\text{AdS}_3/\text{CFT}_2$ in string theory

In order to introduce deformations, we here remind the basic triangle among $\text{AdS}_3/\text{CFT}_2$.

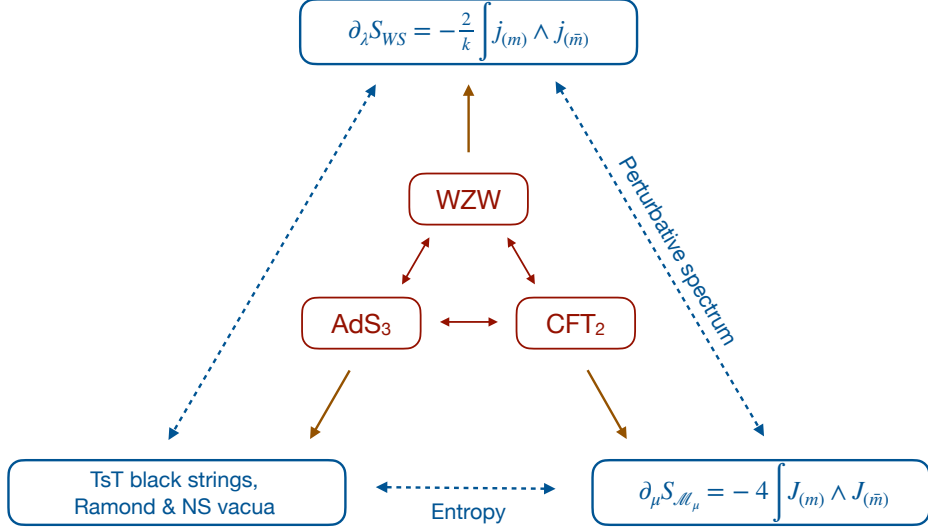


Figure 1. The holographic correspondence between string theory on $\text{AdS}_3 \times M^7$ backgrounds and two-dimensional CFTs (inner triangle), and its relationship to the conjectured duality between TsT transformations of $\text{AdS}_3 \times M^7$ spacetimes, bi-current deformations of the worldsheet action, and single-trace irrelevant deformations of two-dimensional CFTs (outer triangle). Adapted from [10]

$$D(x, \bar{x}) = \int d^2 z (\partial_x J \partial_x + 2\partial_x^2 J)(\partial_{\bar{x}} \bar{J} \partial_{\bar{x}} + 2\partial_{\bar{x}}^2 \bar{J}) \Phi_1 \quad (4.1)$$

$$A^a(x, \bar{x}) = \int d^2 z k^a(z) (\partial_{\bar{x}} \bar{J} \partial_{\bar{x}} \Phi_1 + 2\partial_{\bar{x}}^2 \bar{J} \Phi_1) \quad (4.2)$$

The dual spacetime CFT at large p is argued to be the symmetric product orbifold \mathcal{M}^p/S_p , denoting the seed theory \mathcal{M} , the energy-momentum tensor can be written as

$$T(x) = \sum_{i=1}^p T_i(x) \quad (4.3)$$

which can be also written as

$$T(x) = \frac{1}{2k} (\partial_x J \partial_x \Phi_1 + 2\partial_x^2 J \Phi_1) \bar{J}(\bar{x}; \bar{z}) \quad (4.4)$$

The OPEs are

$$T(x, \bar{x}) D(y, \bar{y}) \sim \frac{c_{\mathcal{M}}}{(x-y)^4} \bar{T}(\bar{y}) + \dots \quad (4.5)$$

4.2 $T\bar{T}$ deformation

In a Lorentzian invariant theory, the stress tensor $T^{\mu\nu}$ is conserved, i.e. $\partial_\mu T^{\mu\nu} = 0$. One can define its determinant $\text{Det}T_{\mu\nu}$ as

$$\begin{aligned}\mathcal{O}_{T\bar{T}} &:= \text{Det}T_{\mu\nu} \\ &= T_{xx}T_{\bar{x}\bar{x}} - T_{x\bar{x}}T_{\bar{x}x} \\ &= \text{Lim}_{x \rightarrow y} (T_{xx}(x)T_{\bar{x}\bar{x}}(y) - T_{x\bar{x}}(x)T_{\bar{x}x}(y)) + \text{derivatives}\end{aligned}\tag{4.6}$$

which is well-defined, and has

$$\langle n | \mathcal{O}_{T\bar{T}} | \rangle = \langle n | T | n \rangle \langle n | \bar{T} | n \rangle - \langle n | \Theta | n \rangle \langle n | \bar{\Theta} | n \rangle\tag{4.7}$$

The $T\bar{T}$ deformations can be described by deforming the CFT by an extra term

$$\frac{\partial S_{\mathcal{M}_\mu}}{\partial \mu} = \int d^2x \mathcal{O}_{T\bar{T}}\tag{4.8}$$

where we have combine the 2d spacetime coordinate (ϕ, t) into $x = \phi + t$ and $\bar{x} = \phi - t$.

When putting the deformed CFT on a cylinder, with $(x, \bar{x}) \sim (x + 2\pi R, \bar{x} + 2\pi R)$, one can then evaluate the energy spectra via inviscid Burger's equation:

$$\partial E_n(R, \mu) = E_n(R, \mu) \partial E(R, 0) + \frac{1}{R} J_n^2(R)\tag{4.9}$$

The spectrum

$$E(\mu) = -\frac{R}{2\mu} \left[1 - \sqrt{1 + \frac{4\mu}{R} E(0) + \frac{4\mu^2}{R^4} J(0)^2} \right], \quad J(\mu) = J(0),\tag{4.10}$$

where $E = E_L + E_R$ and $J = R(E_L - E_R)$ denote the energy and angular momentum.

Let us make some comments on the $T\bar{T}$ deformation: • it is irrelevant, since the dimension of μ is $2 > 0$.

- It is also solvable, which is the main reason why people are interested in studying them.

- It can also be reformulated as coupling to random geometry, flat JT gravity, (non-) critical strings.

- The partition function after the deformation is still modular invariant.

- The scattering amplitude can be shown as dressed by a phase factor.

From the historic viewpoint,

It was conjectured in [11] that performing the following TsT transformation on the string theory is holographically equivalent to deforming the seed of the dual symmetric orbifold,

$$\text{TsT}_{(X^m, X^{\bar{m}}; \check{\mu})} \iff \frac{\partial S_{\mathcal{M}_\mu}}{\partial \mu} = -4 \int J_{(m)} \wedge J_{(\bar{m})}, \quad (4.11)$$

where $\text{TsT}_{(X^m, X^{\bar{m}}; \check{\mu})}$ denotes T-duality on X^m , shift $X^{\bar{m}} \rightarrow X^{\bar{m}} - 2\check{\mu}X^m$, and T-duality on $X^{\bar{m}}$, $S_{\mathcal{M}_\mu}$ is the action of the deformed seed, $\mu(\check{\mu})$ is the dimensionful(less) deformation parameter, and $J_{(m)}$, $J_{(\bar{m})}$ are the Noether currents generating translations along X^m , $X^{\bar{m}}$. It is interesting to note that from a purely worldsheet perspective, a TsT transformation is equivalent to a *marginal* bi-current deformation of the worldsheet action, the latter of which can be shown to be equivalent to a gauged $(SL(2, R) \times U(1))/U(1)$ WZW model [11].

References

- [1] J.M. Maldacena, *The Large N limit of superconformal field theories and supergravity*, *Adv. Theor. Math. Phys.* **2** (1998) 231 [[hep-th/9711200](#)].
- [2] A. Friedmann, *On the Possibility of a world with constant negative curvature of space*, *Z. Phys.* **21** (1924) 326.
- [3] G. Lemaitre, *The expanding universe*, *Annales Soc. Sci. Bruxelles A* **53** (1933) 51.
- [4] H.P. Robertson, *Kinematics and World-Structure*, *Astrophys. J.* **82** (1935) 284.
- [5] A.G. Walker, *On Milne's Theory of World-Structure*, *Proceedings of the London Mathematical Society* **42** (1937) 90.
- [6] A. Strominger and C. Vafa, *Microscopic origin of the Bekenstein-Hawking entropy*, *Phys. Lett. B* **379** (1996) 99 [[hep-th/9601029](#)].
- [7] J.R. David, G. Mandal and S.R. Wadia, *Microscopic formulation of black holes in string theory*, *Phys. Rept.* **369** (2002) 549 [[hep-th/0203048](#)].
- [8] E. Kiritsis, *Introduction to superstring theory*, vol. B9 of *Leuven notes in mathematical and theoretical physics*, Leuven U. Press, Leuven (1998), [[hep-th/9709062](#)].
- [9] J. Polchinski, *String theory. Vol. 2: Superstring theory and beyond*, Cambridge Monographs on Mathematical Physics, Cambridge University Press (12, 2007), [10.1017/CBO9780511618123](#).
- [10] L. Apolo and W. Song, *TsT , black holes, and $T\bar{T} + J\bar{T} + T\bar{J}$* , [2111.02243](#).
- [11] L. Apolo, S. Detournay and W. Song, *TsT , $T\bar{T}$ and black strings*, *JHEP* **06** (2020) 109 [[1911.12359](#)].