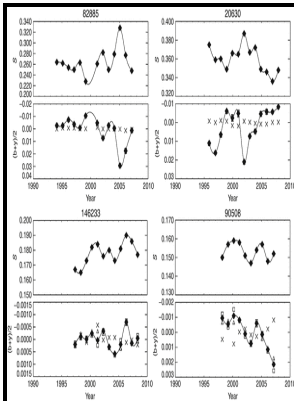


Consequences of a chromospheric temperature gradient on the width of H [alpha] in late-type giants

Big Bear Solar Observatory, California Institute of Technology - Past and recent observations of the solar upper atmosphere at vacuum



Description: -

-
Animals -- Fiction
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Temperature gradients.
Optical properties.
Chromosphere.
Giant stars.
Stellar chromospheres. Consequences of a chromospheric temperature gradient on the width of H [alpha] in late-type giants
-
[NASA contractor report] -- NASA CR-173359. Consequences of a chromospheric temperature gradient on the width of H [alpha] in late-type giants
Notes: Microfiche. [Washington, D.C. : National Aeronautics and Space Administration], 1984. 1 microfiche.
This edition was published in 1984



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Tags: #Magnetism, #dynamo #action #and #the #solar

Polarimetry of cool giants and supergiants

This will occur in roughly in late A and cooler dwarfs, and in more massive stars as they leave the main-sequence and develop convective zones. Thomas in Jordan and Avrett, , p.

The Sun's Atmosphere

Other ideas include that of Lopes and Silk , who revisit the helioseismology problem in the light of revised photospheric abundance for the Sun and solar-like stars. The simulations suggest that neutral-line shear and sigmoidal field arise as a natural by-product of flux emergence.

Model Atmospheres for Intermediate and Late

This activity takes two principal forms.

Polarimetry of cool giants and supergiants

The first of these is discussed in this section: a set of critical theoretical advances that provided a solid basis for the synoptic ground-based programs that soon followed.

h

Asterisks 2003—11 are from XMM-Newton α Cen B only: A was difficult to record because of technical issues.

The Astrophysical Journal, Volume 865, Number 1, 2018 September 20

This is not a big surprise because several dense dust disks have already been observed around progenitor stars.

Yohkoh10

The strength of non-adiabatic effects in a convection zone depends on the distribution of density and temperature, which in turn depends on the convection model.

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