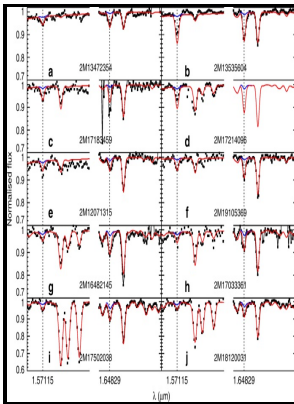


Spectra of peculiar strontium stars.

- - Peculiar star



Description: -

-Spectra of peculiar strontium stars.

-Spectra of peculiar strontium stars.

Notes: Thesis (M.A.) -- University of Toronto, 1940.

This edition was published in 1940



Filesize: 15.109 MB

Tags: #U of #Toronto #Astronomy #Department #Theses #1930

Przybylski's Star

R; Hubrig, S; Bord, D.

Alioth (ε Ursae Majoris)

Nearly all of the measured polarization is believed to be due to interstellar dust, a conclusion that is consistent with previous studies. The excitation of the CO is probably the result of the propagation of a shock wave at the third maximum.

Alioth (ε Ursae Majoris)

Thus, according to the team, HR8844 resembles both a very hot Am star and a very cool HgMn star. Formation Alioth formed around 300 million years ago from a molecular cloud of gas and dust. One of the best months to observe them is during April.

Spectral evolution of V838 Monocerotis in the optical and near

Thus a given radiation field can maintain a higher degree of ionization in stars with more extended atmospheres. It is assumed that the centers of the stars, and the bulk compositions of the entire star, have more normal chemical abundance mixtures which reflect the compositions of the gas clouds from which they formed.

U of Toronto Astronomy Department Theses 1930

In stars later than A0 it appears to be indicative of the presence of Eu II. In both spectra the line is stronger than in any normal luminosity class at F0.

U of Toronto Astronomy Department Theses 1930

The two spectropolarimetric epochs were taken 6 and 27 d after the second outburst on 2002 February 8. The blend at 4171 is strong; this is an indication of abnormal strength of Cr II.

Related Books

- [Selling Today Sales Connection S/G](#)
- [Meaning of relativity.](#)
- [Need for roots - prelude to a declaration of duties towards mankind](#)
- [Genesis of the theory of art for arts sake in Germany and in England](#)
- [White flock](#)