

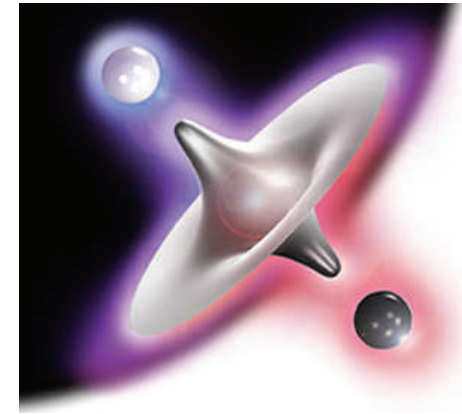
Antimatter

Jan Meier

Seminar: Experimental Methods in Atomic Physics

May, 8th 2007

Overview



- Antimatter and CPT theorie
 - what is antimatter?
 - what physics does it follow to?
- First observations of antimatter
- Natural sources of antimatter
- Artificial sources of antimatter and experiments with antihydrogen
 - PS210, E862 (first detections of $\bar{\text{H}}$)
 - ATHENA, ATRAP (spatial and velocity distribution and temperature measurements)
 - ALPHA (trapping of $\bar{\text{H}}$)
 - AEGIS (gravity measurement)

Antimatter and CPT theorie

Prediction of antimatter

1928 - Paul Dirac
(Nobel prize 1933)



Dirac equation of a free electron

$$\left(i\hbar \frac{\partial}{\partial t} + i\hbar c \vec{\alpha} \cdot \vec{\nabla} - \hat{\beta} m_e c^2 \right) \vec{\Psi}(\vec{r}, t) = \vec{0}$$

solution delivers two energy eigenvalues:

$$E_{\pm} = \pm \sqrt{c^2 |\vec{p}|^2 + m_e^2 c^4}$$

Does negative energy eigenvalue have physical meaning?

(P.A.M. Dirac, Proc. R. Soc. A 117 (1928) 610)

$$\hat{\beta} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix}$$

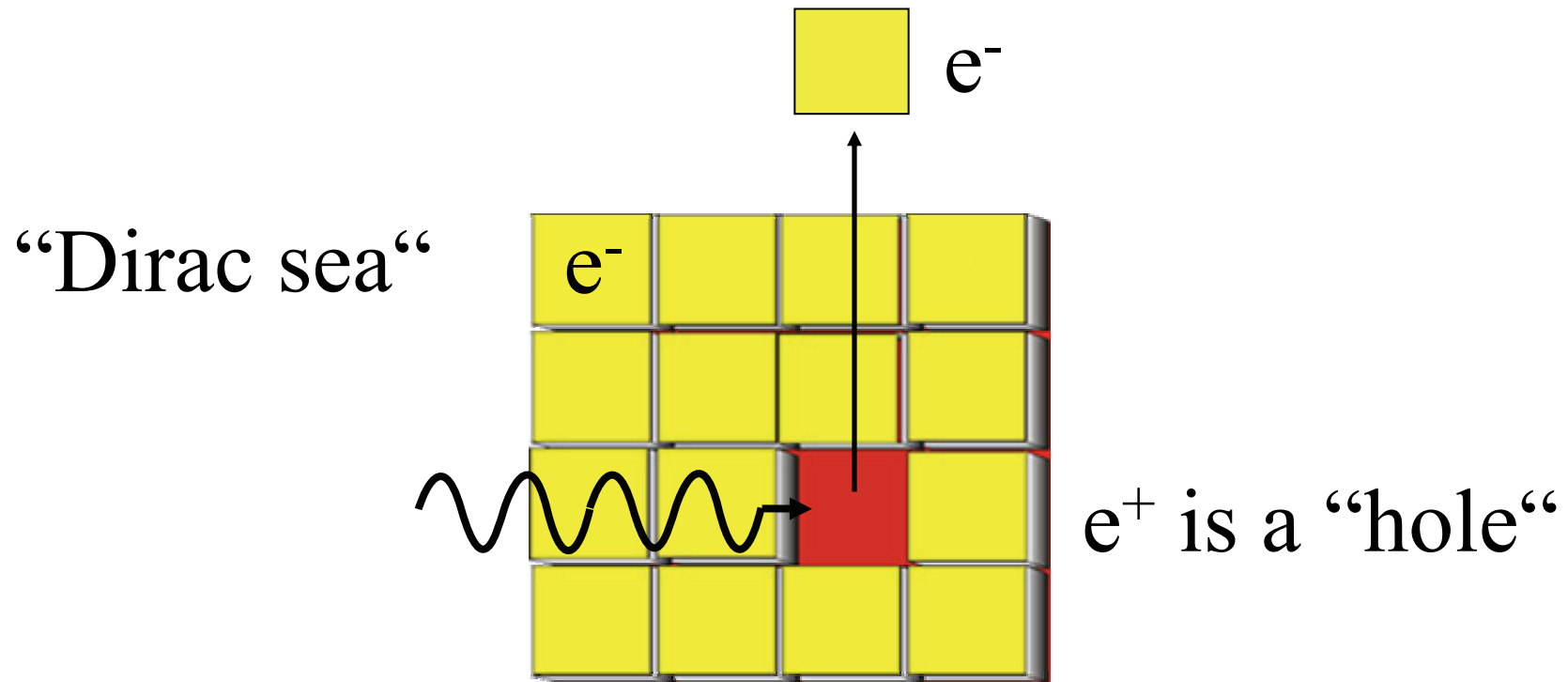
$$\vec{\hat{\alpha}} = \begin{pmatrix} \hat{\alpha}_x \\ \hat{\alpha}_y \\ \hat{\alpha}_z \end{pmatrix}$$

$$\hat{\alpha}_x = \begin{pmatrix} 0 & 0 & 0 & 1 \\ 0 & 0 & 1 & 0 \\ 0 & 1 & 0 & 0 \\ 1 & 0 & 0 & 0 \end{pmatrix}$$

$$\hat{\alpha}_y = \begin{pmatrix} 0 & 0 & 0 & -i \\ 0 & 0 & i & 0 \\ 0 & -i & 0 & 0 \\ i & 0 & 0 & 0 \end{pmatrix}$$

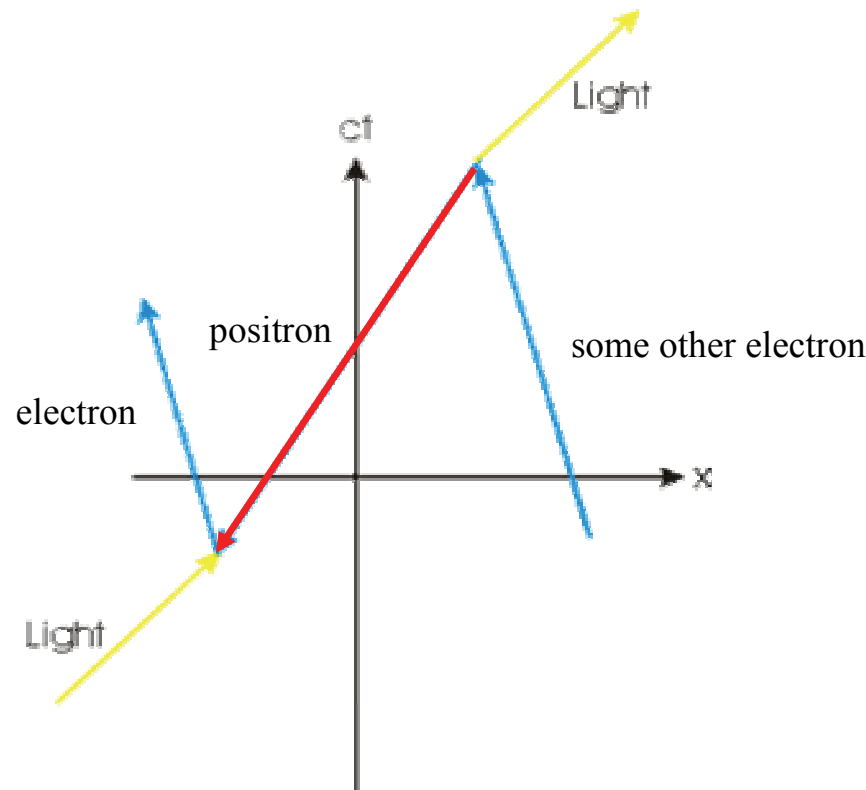
$$\hat{\alpha}_z = \begin{pmatrix} 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & -1 \\ 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \end{pmatrix}$$

Dirac interpretation



1949 Feynman-Stückelberg interpretation

- Positron is a particle (not a “hole”)
- Positron is a (positively charged) electron, travelling backwards in time



CPT-Theory

Transformations C, P and T:

$$\text{C (charge conjugation)} \quad q \rightarrow -q \ ; \ B \rightarrow -B \quad \dots$$

$$\text{P (parity inversion)} \quad \vec{x} \rightarrow -\vec{x}$$

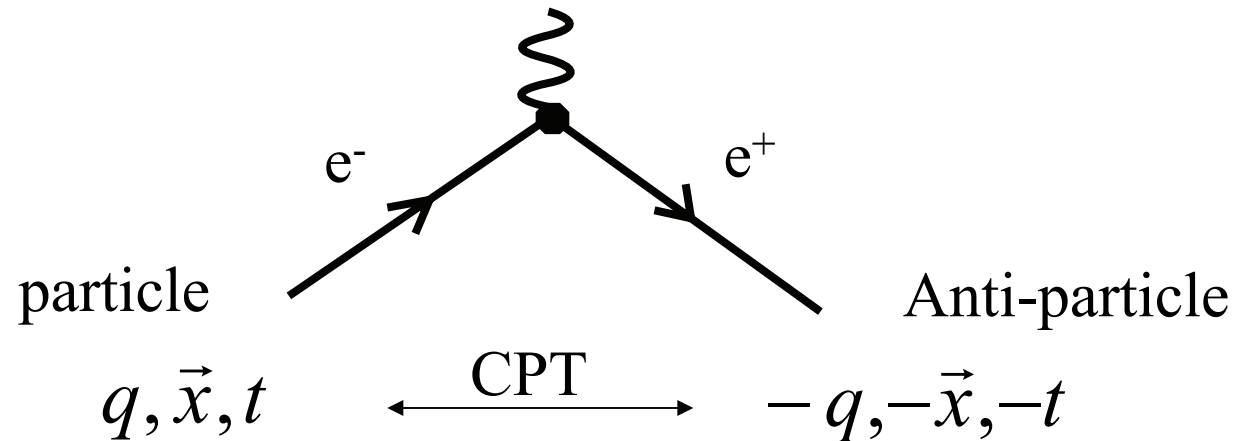
$$\text{T (time inversion)} \quad t \rightarrow -t$$

CPT transformation:

$$f(q, \vec{x}, t) \rightarrow f(-q, -\vec{x}, -t)$$

CPT Symmetry

A CPT transformation transforms a particle into its corresponding anti-particle



\Rightarrow Standard Model:

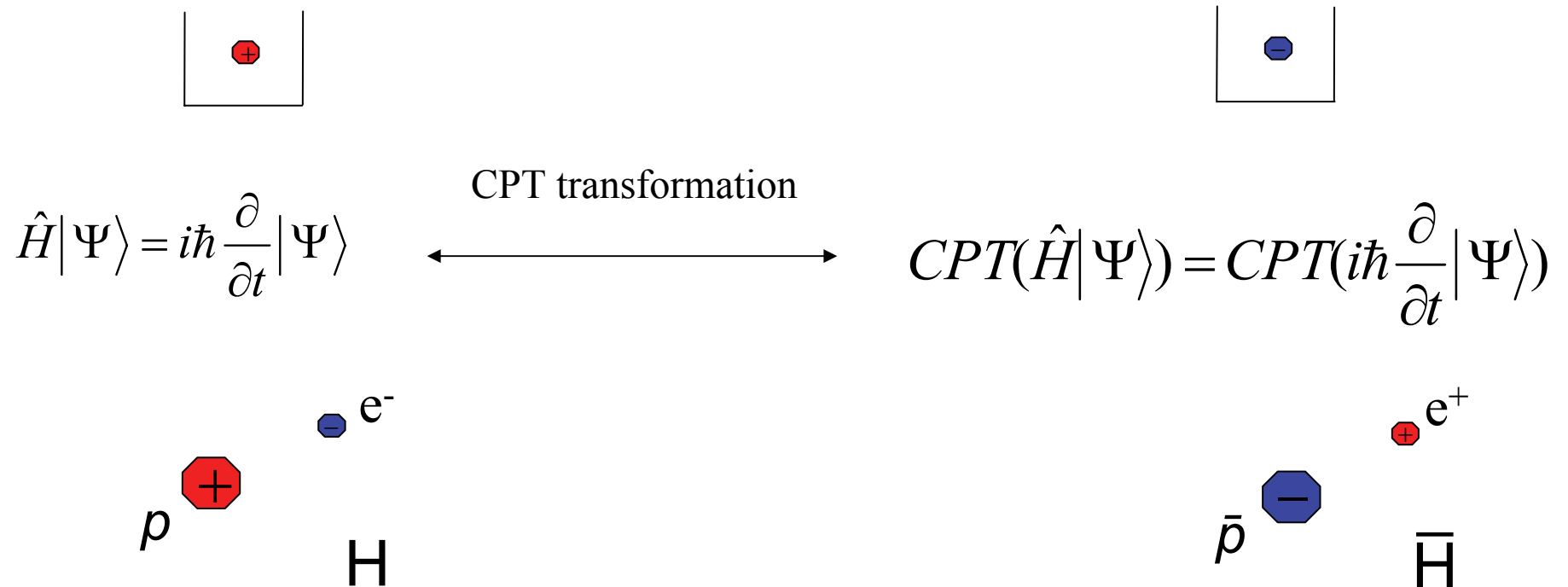
For every particle type there is a corresponding antiparticle type

(some electrically neutral bosons, like Z^0 , γ and η_c are their own antiparticles)

CPT invariance

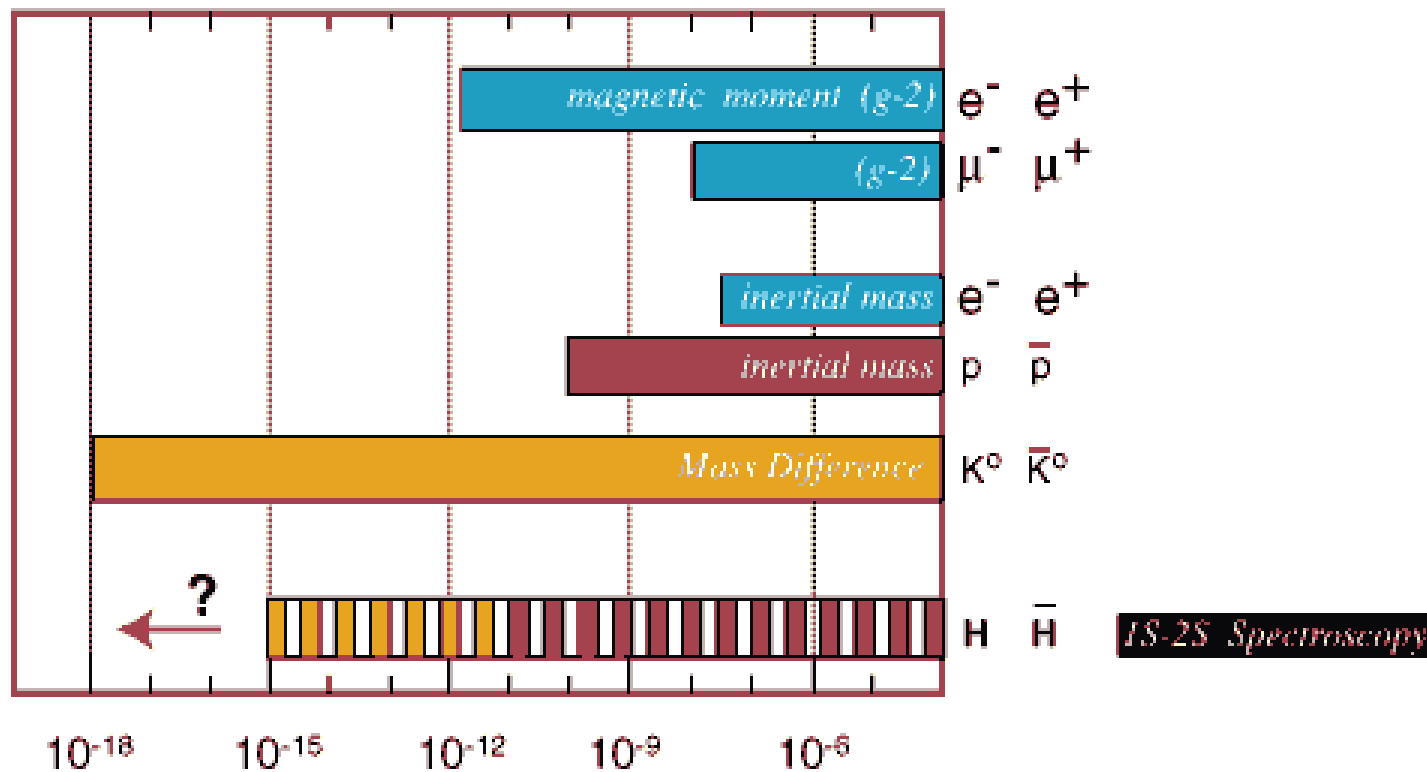
under certain conditions, relativistic quantum field theories say:

Physics (i.g. all physical laws, equations, processes) is invariant under CPT transformations



Prospect of \bar{H} spectroscopy

The most precise CPT Tests



Violation of CPT invariance

Historical: P (theory:Lee,Yang; Exp.:Wu), CP violations (J.H. Christenson)

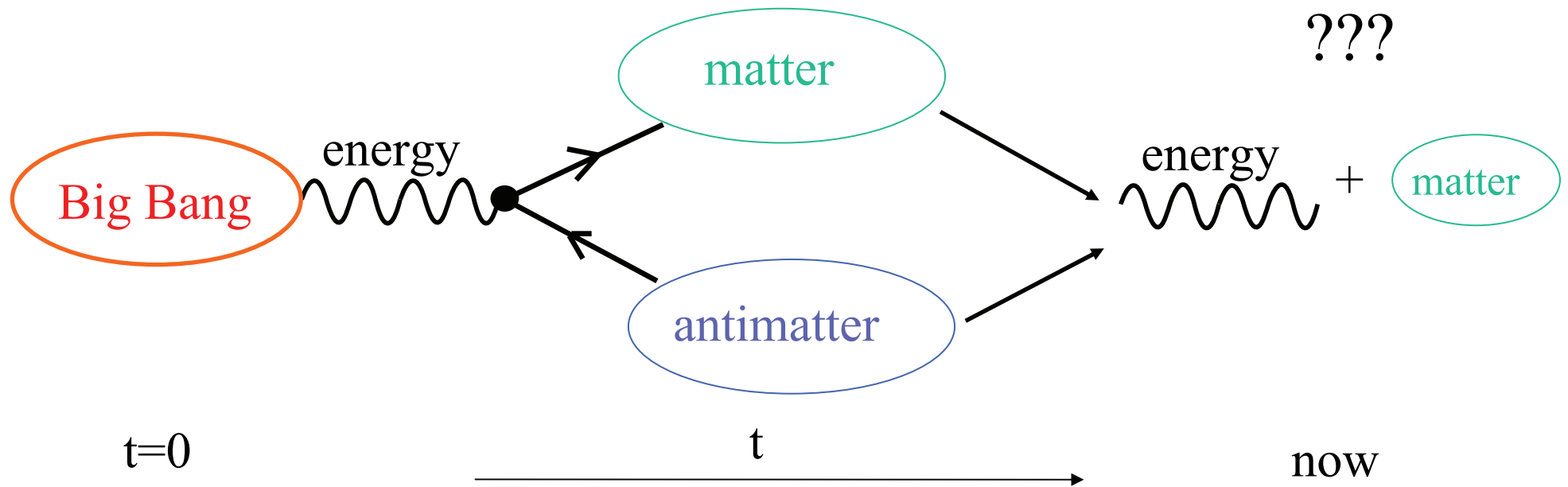
(C.S. Wu et al., Phys. Rev. 105 (1957) 1413)

(J.H. Christenson et al., Phys. Rev. Lett. 13 (1964) 138)

- String theory
- Kostelecky (standard model extension)
- ...

(R. Bluhm et al., Phys. Rev. Lett. 82 (1999) 2254)

Matter-antimatter asymmetry



possible explanations:

- CPT violation, breaking of Baryon number conservation
- CP violation, breaking of Baryon number conservation, out of equilibrium situation

First observations of antimatter

First detection of antimatter

1932 - Carl Anderson
(Nobel prize 1936)



secondary cosmic rays

cloud chamber

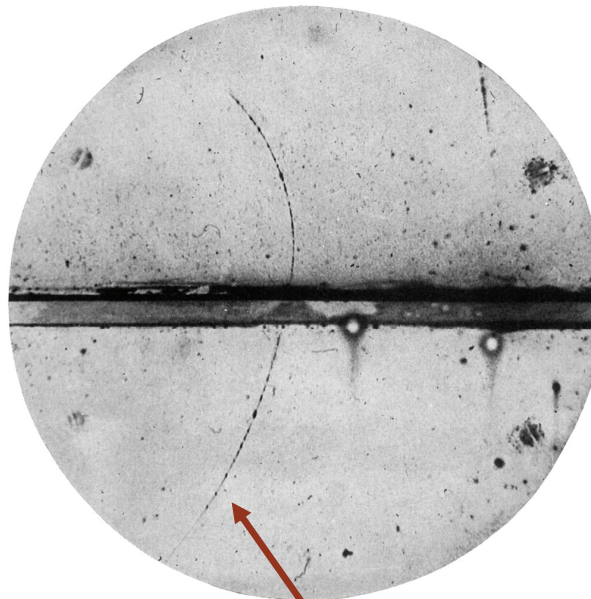
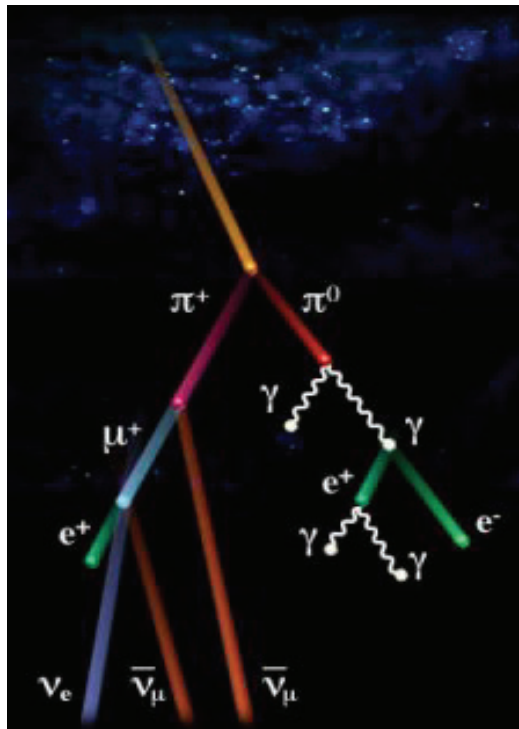


Fig. 1. A 63 million volt positron ($H\rho=2.1\times 10^8$ gauss-cm) passing through a 6 mm lead plate and emerging as a 23 million volt positron ($H\rho=7.5\times 10^4$ gauss-cm). The length of this latter path is at least ten times greater than the possible length of a proton path of this curvature.

“Positron“ (e^+)

From particle track:

$$q_{Positron} < +2e$$

$$m_{Positron} < 20m_e$$

(C.D. Anderson, Phys. Rev. 43 (1933) 491)

Further detections of antiparticles

- **1955** - antiproton at Lawrence Berkeley National Laboratory (Chamberlain, Sergé, Wiegand, Ypsilantis)
- **1956** - antineutron (B. Cork)

(S.L. Chamberlain, Phys. Rev. 100, 947 (1955))

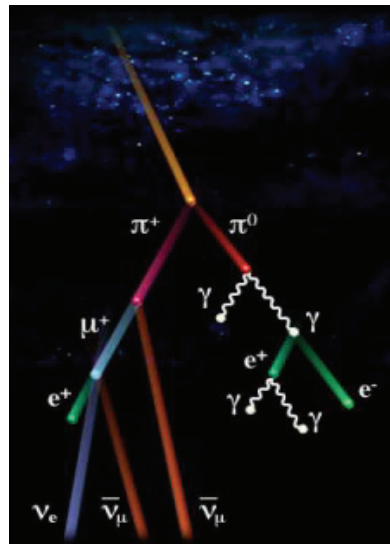
(B. Cork, Phys. Rev. 104, 1193 (1956))

Natural sources of antimatter

- Beta(plus)-decay



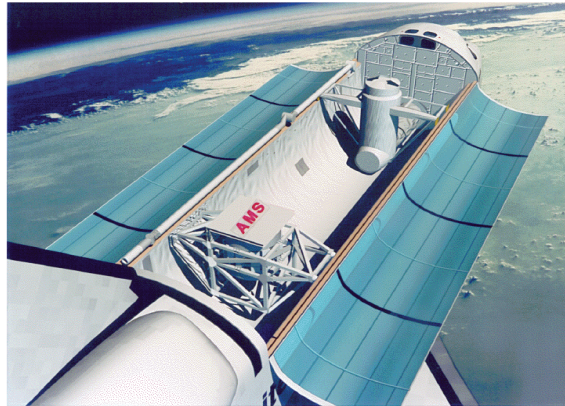
- secondary cosmic rays



$$e^{+}, \mu^{+}, \pi$$

Is there antimatter in (primary) cosmic rays?

1998 - AMS-01

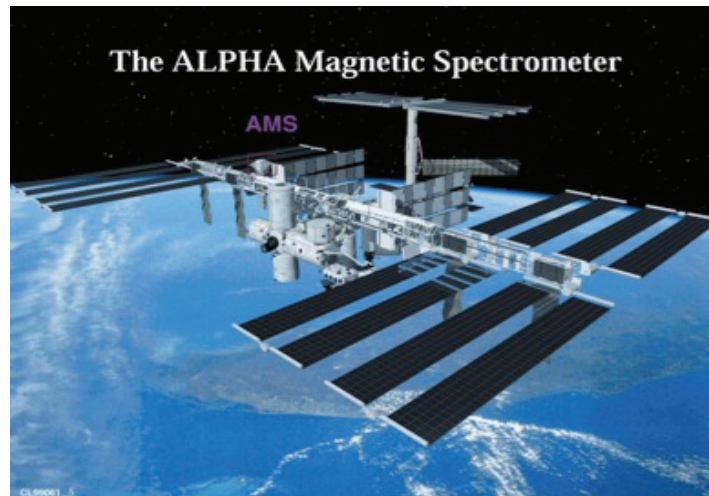


ten day flight on Discovery

“prototype“ $m \sim 3$ tons

No evidence for primary antimatter!

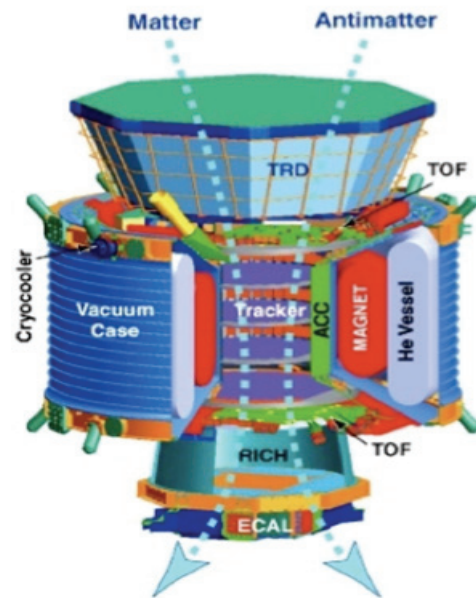
2009 – 2012
AMS-02



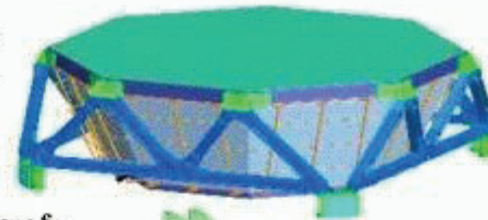
three years on ISS

$m \sim 7$ tons

Goal:
detection of He, \overline{He}
and heavier nuclea



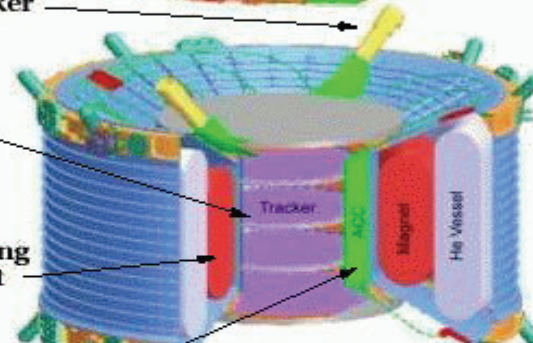
Transition
Radiation
Detector



Upper Time-of-
Flight



Star tracker



Silicon
tracker

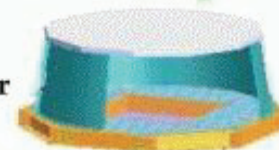
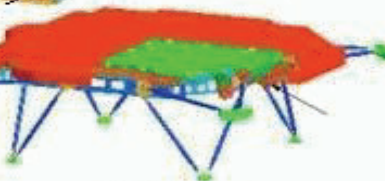
Super-
conducting
Magnet

Anti-coincidence
Counter

Lower Time-
Of-Flight

Ring-imaging
Cerenkov detector

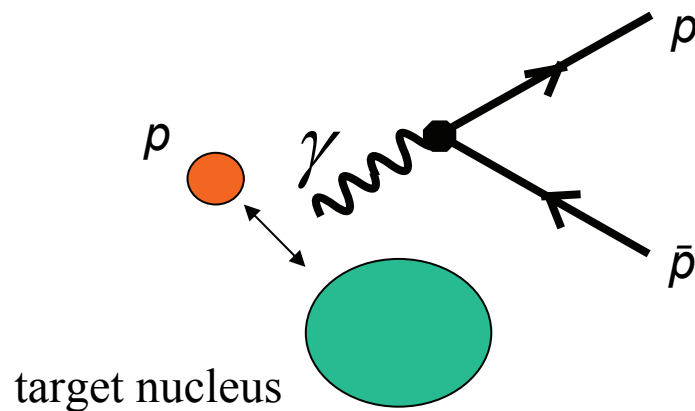
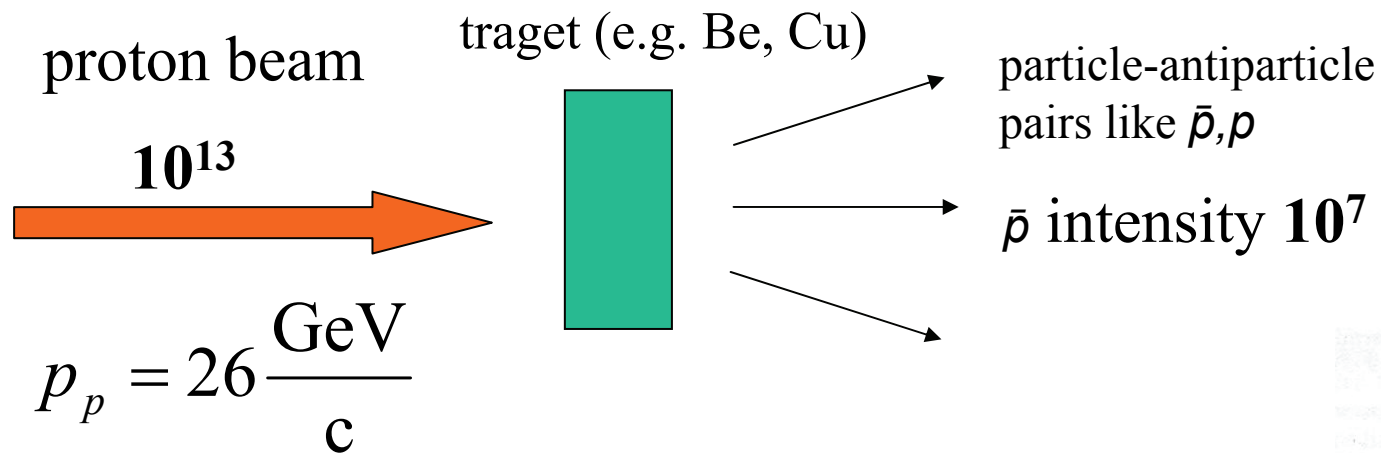
Electromagnetic
calorimeter



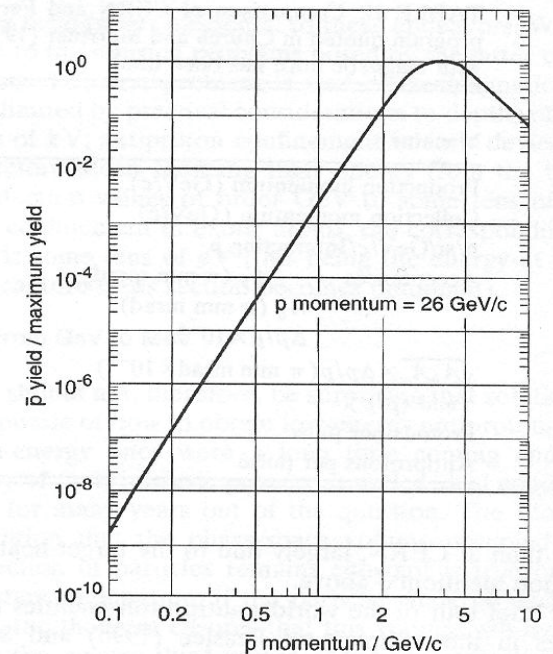
Artificial sources of antimatter and
experiments with \bar{H}

antiproton production

principle (since 1954):

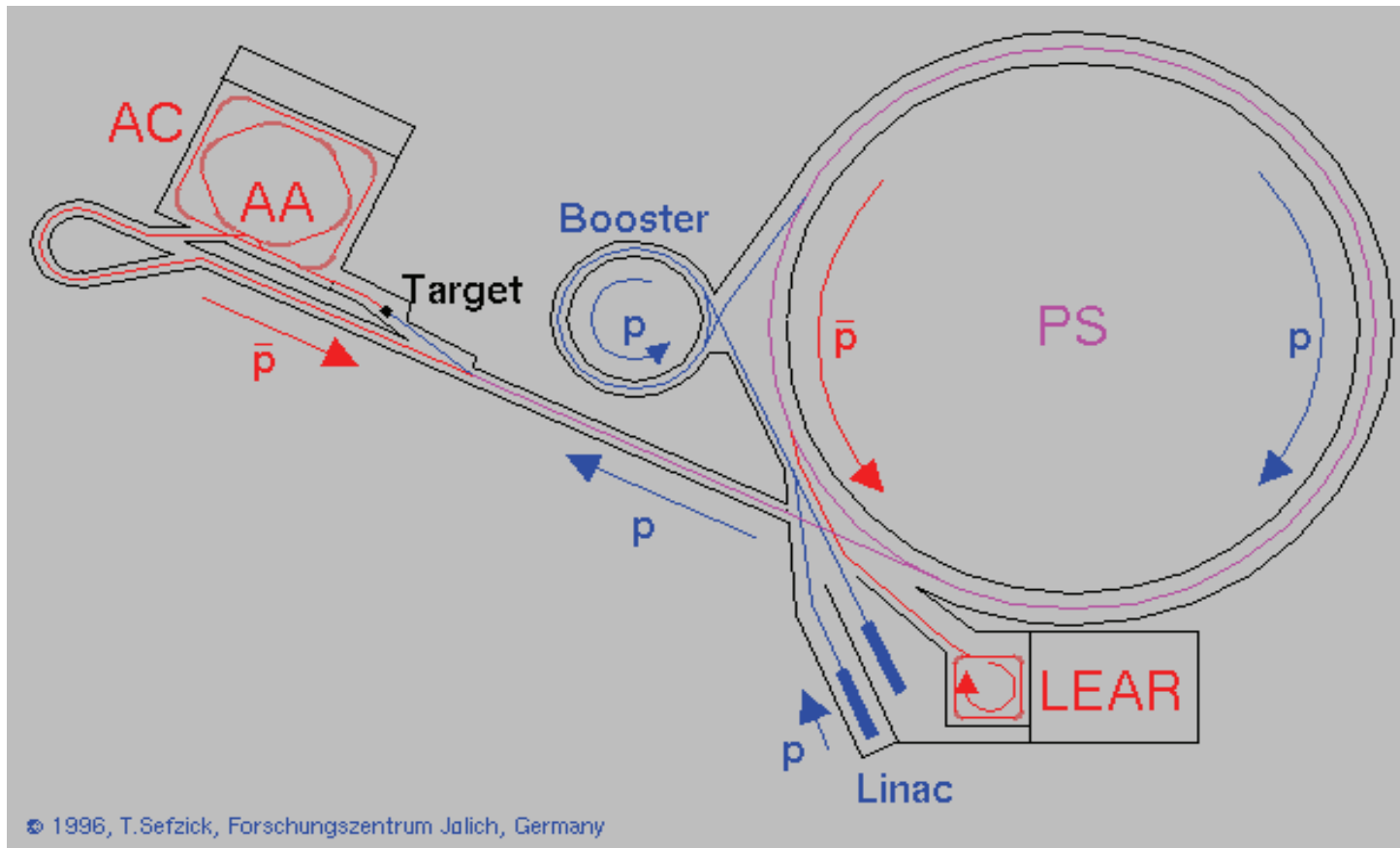


(J. Eades, Rev. of modern phys. Vol.71, No.1 (1999))

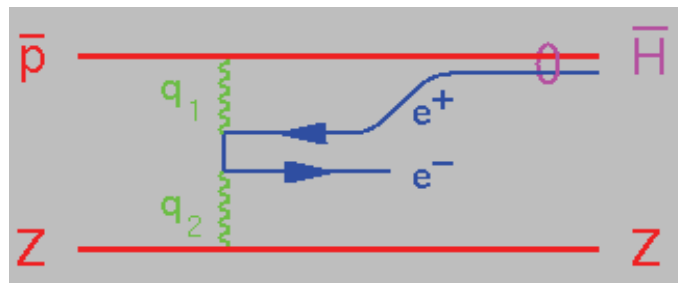
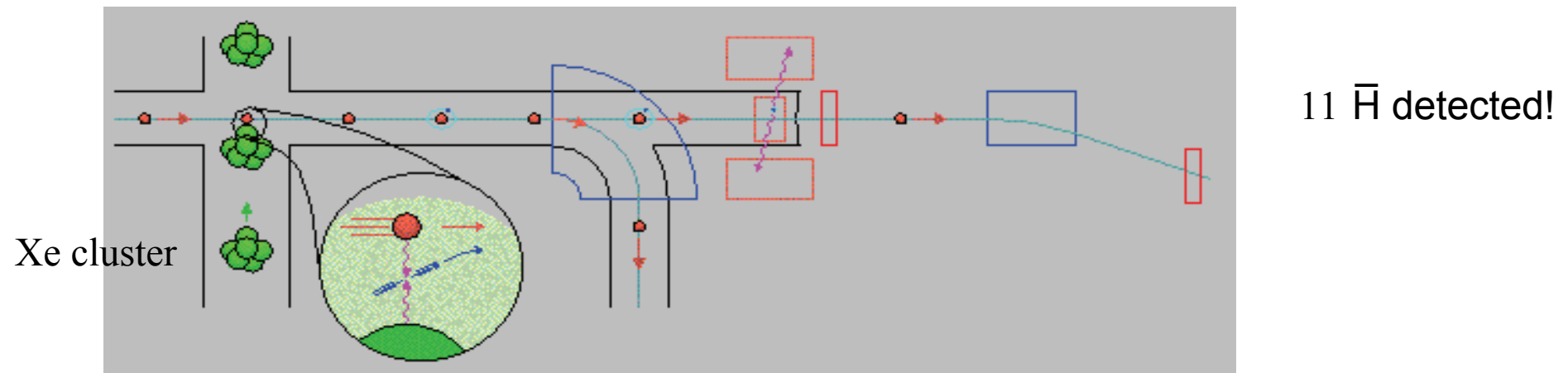


PS210 Experiment (1995 first \bar{H} detection)

PS (Proton Synchrotron) at CERN



PS210 at LEAR (Low Energy Antiproton Ring) $p = 1.94 \text{ GeV}/c$



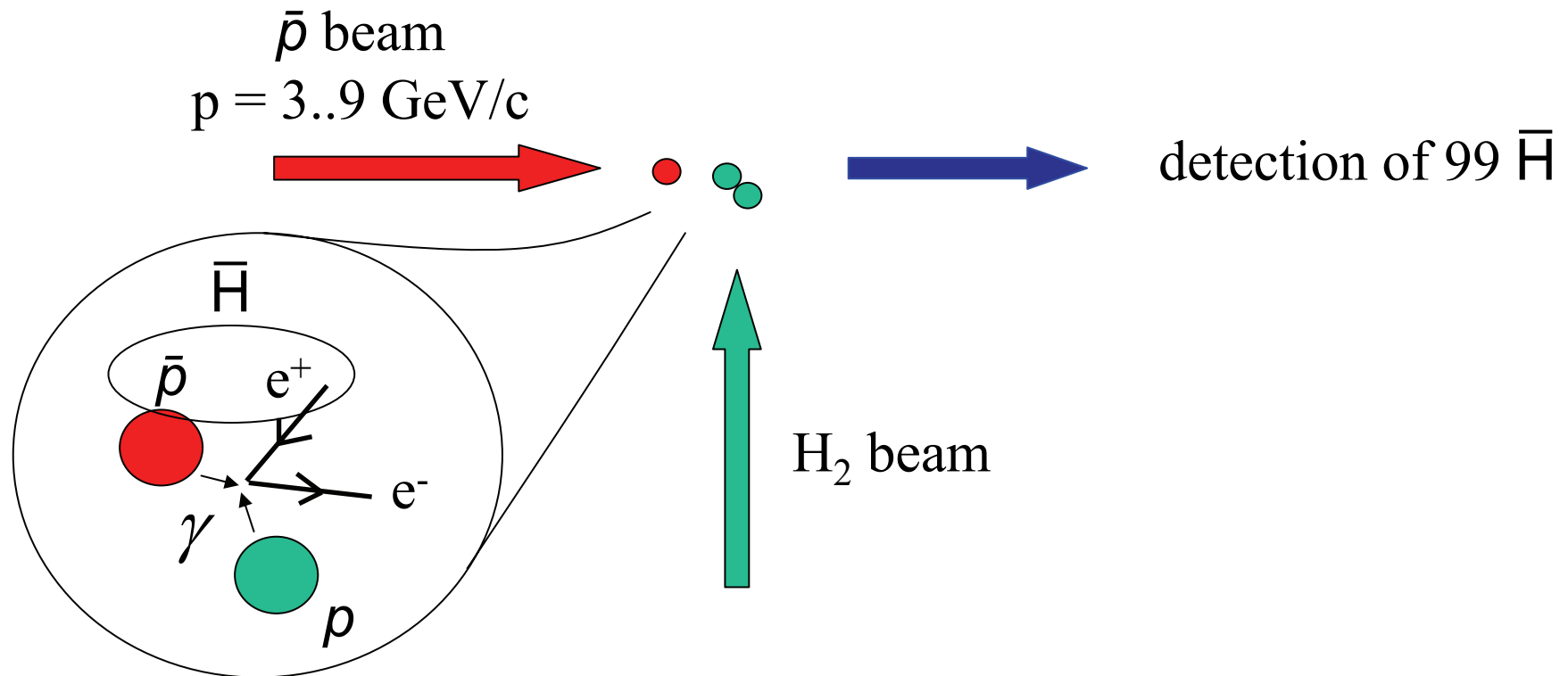
e^-, e^+ pair creation is a rare process

- only if \bar{p} , Z get close
- two photon collision

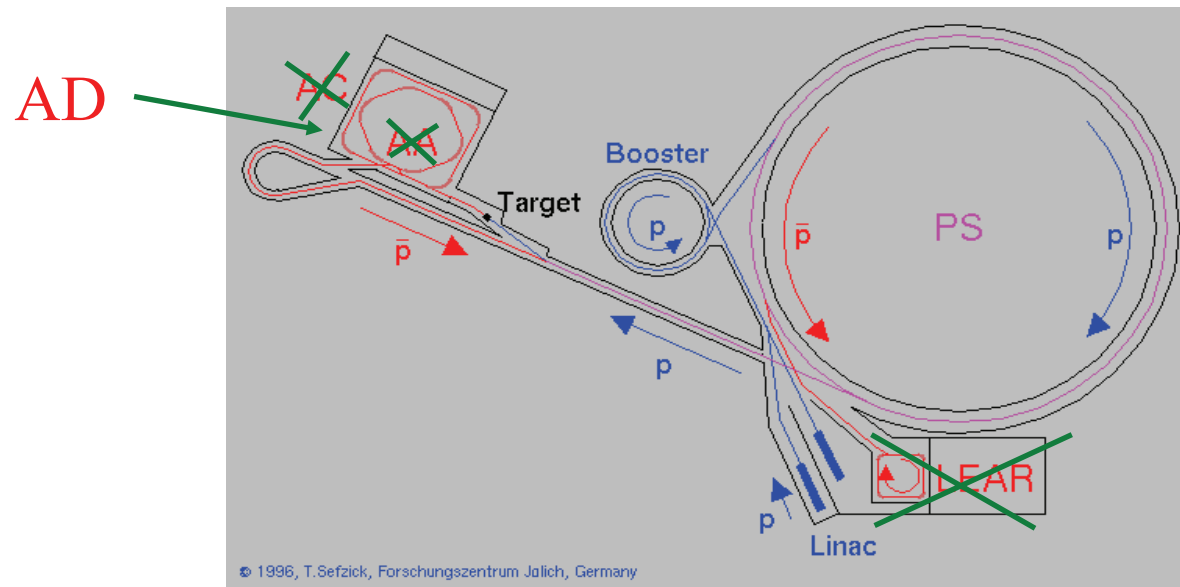
$\bar{\text{H}}$ production only if
rel. energy \bar{p} , $e^+ < 13.7 \text{ eV}$

probability = 0.000 000 000 000 000 01 %

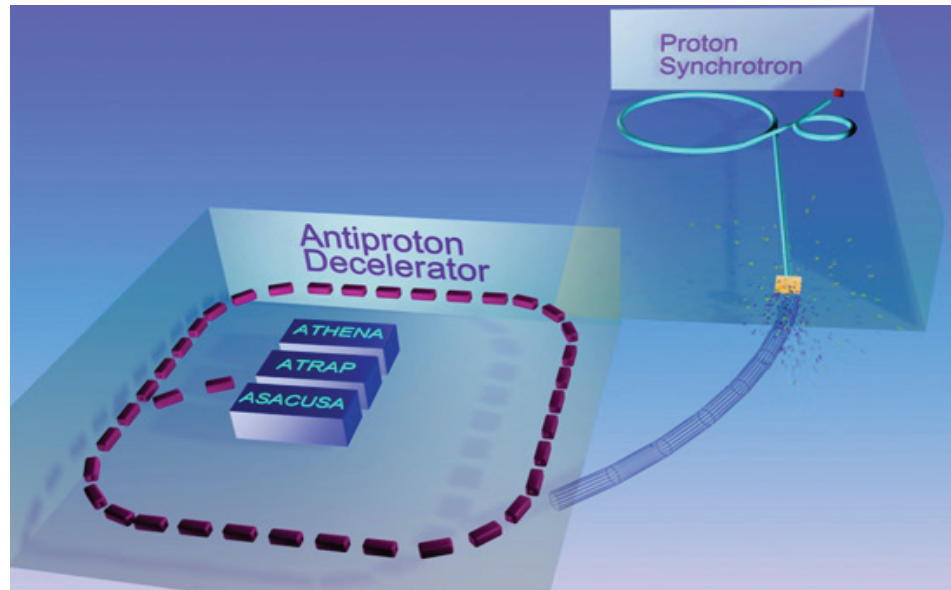
E862 at Fermilab (1996)



AD (Antiproton Decelerator) (since July 2000)

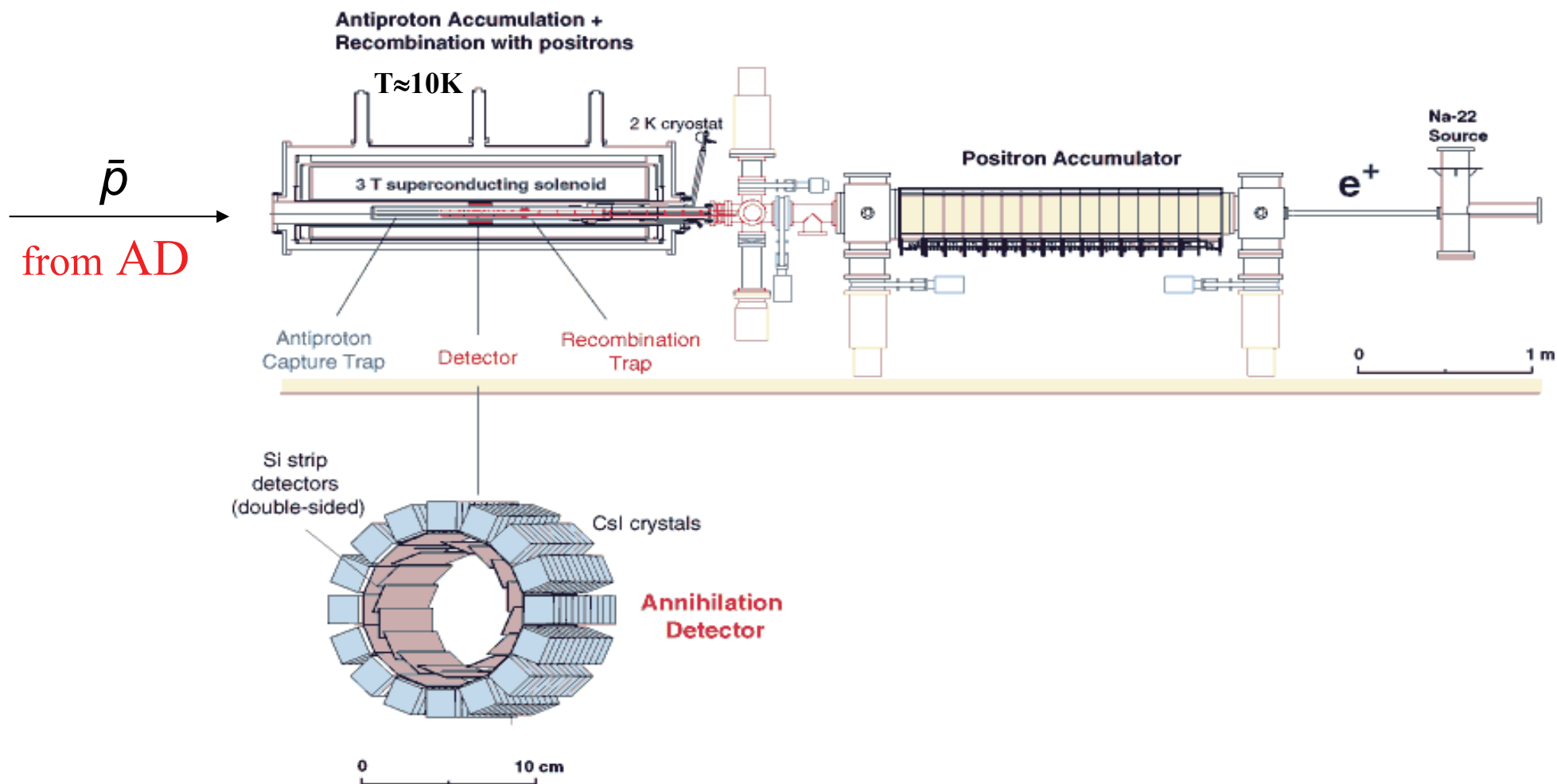


deceleration and cooling
 $p_{\bar{p}}: 3.5 \rightarrow 0.1 \text{ GeV/c}$



- ATHENA (2002)
(\bar{H} detection by detector)
- ATRAP (2002)
(\bar{H} detection by reionization)

The ATHENA experiment



positron production

Solid Neon is best known positron moderator

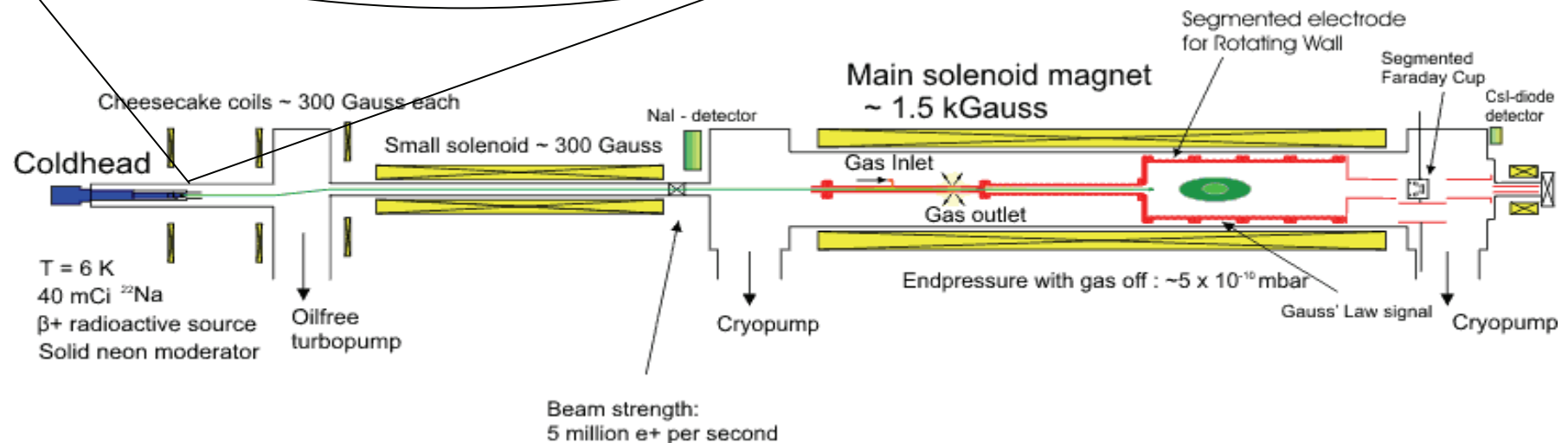
Neon gas

^{22}Na

$T=6\text{K}$

solid Neon layer

“low energy” positrons

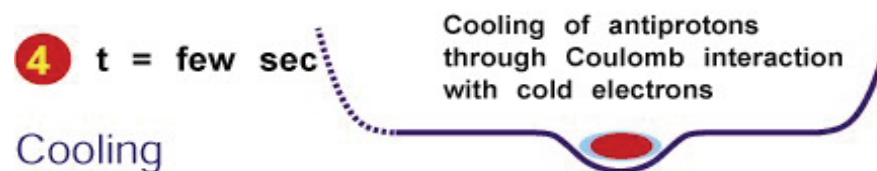
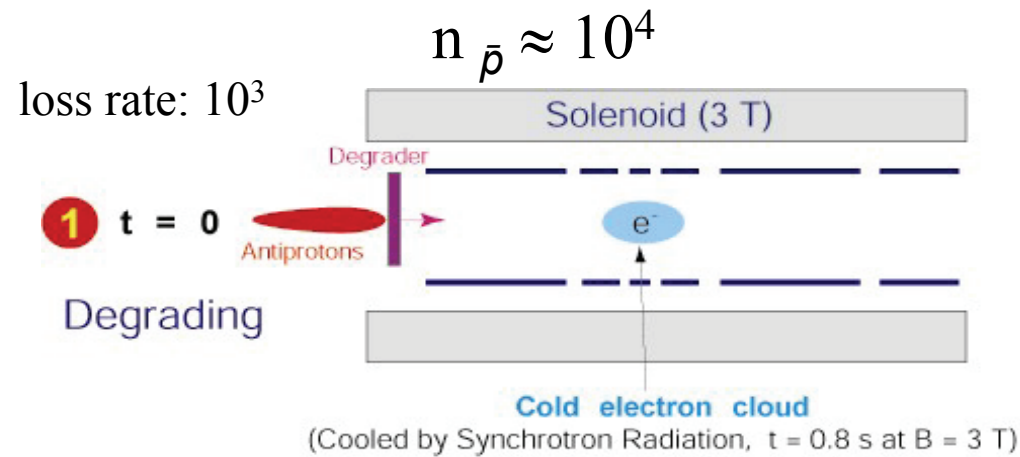


Antiproton capturing

$$E_{\bar{p}} = 5 \text{ MeV}$$

$$t_p = 200 \text{ ns}$$

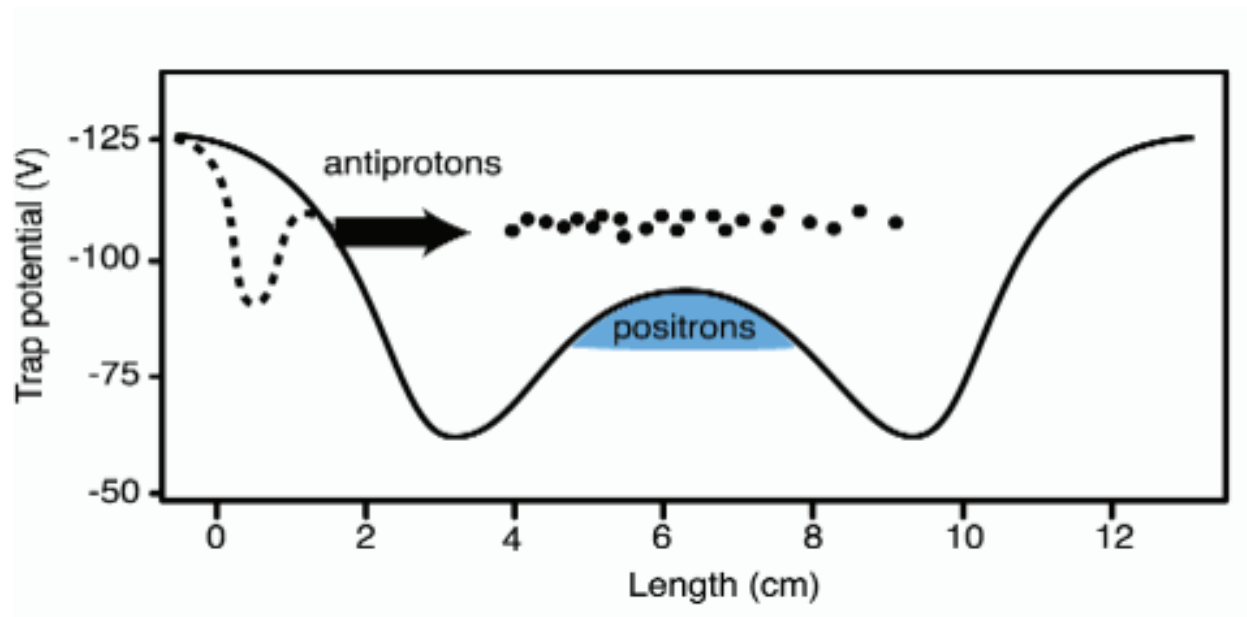
$$n_{\bar{p}} \approx 10^7$$



5kV

Antiproton positron mixing

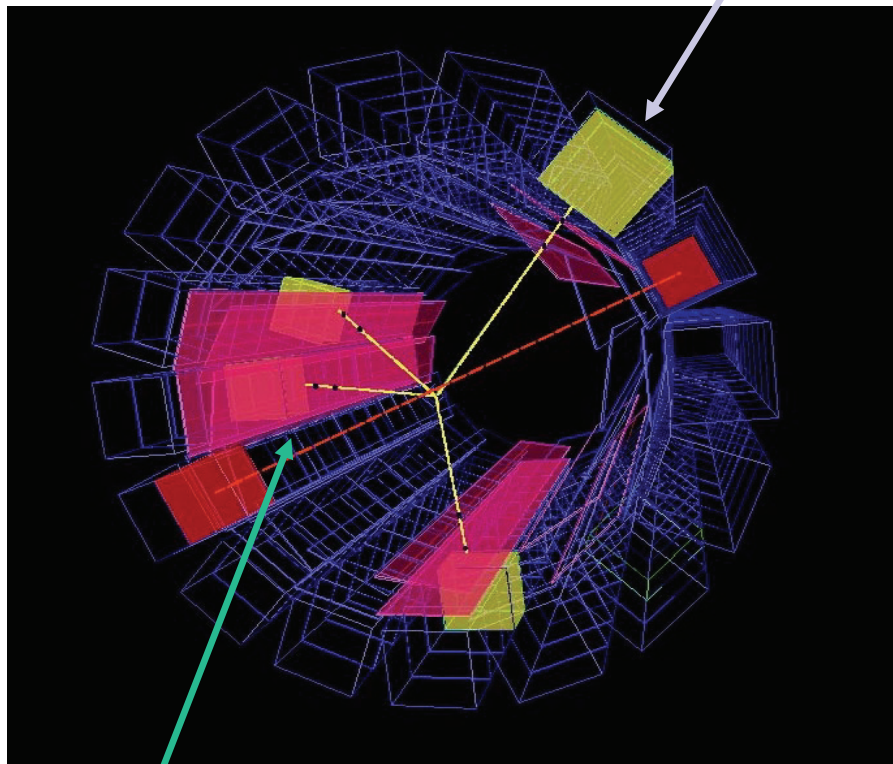
“Mixing trap”: nested potential with both positive and negative ions



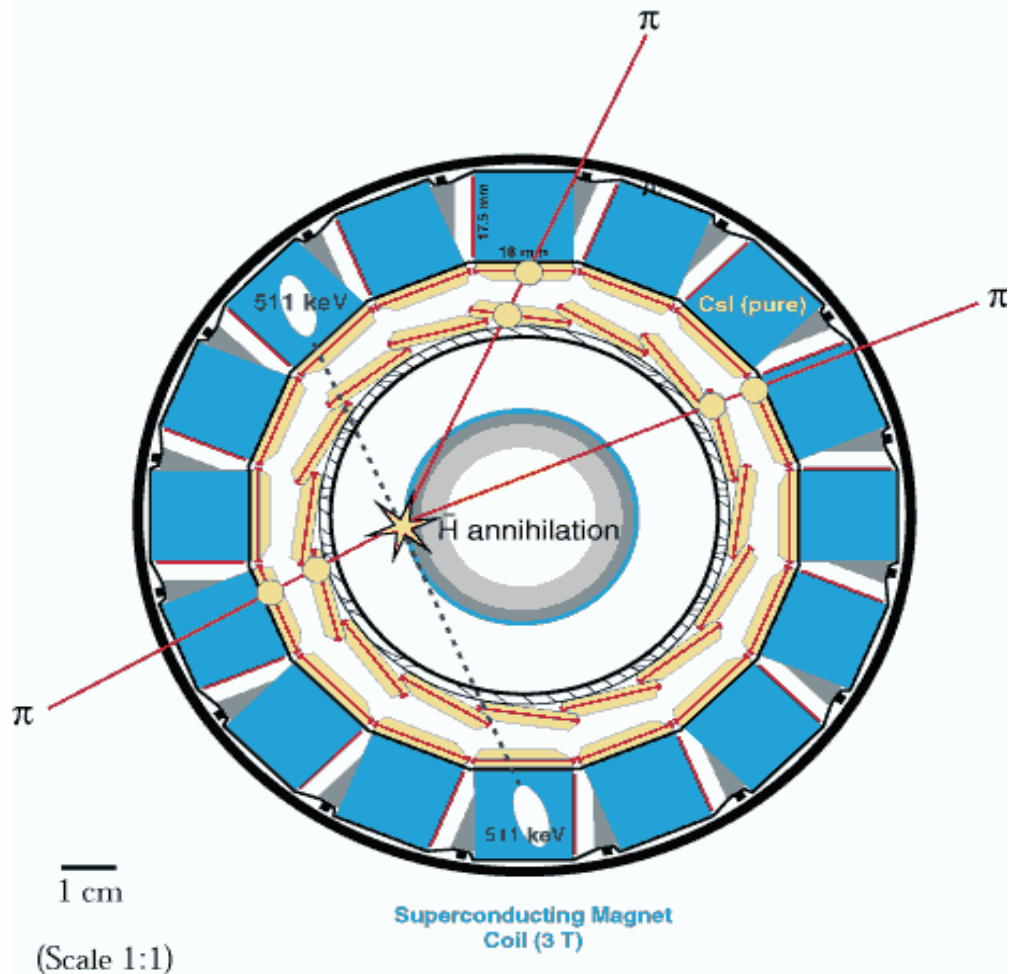
- production of several million \bar{H} between 2002 and 2004

\bar{H} detection

CsI crystal calorimeter



Si strips to “follow the path”



Discriminate Antihydrogen Annihilation from background of **Antiproton annihilation** and Positron annihilation

Good spatial resolution (< 1 cm) of vertex for

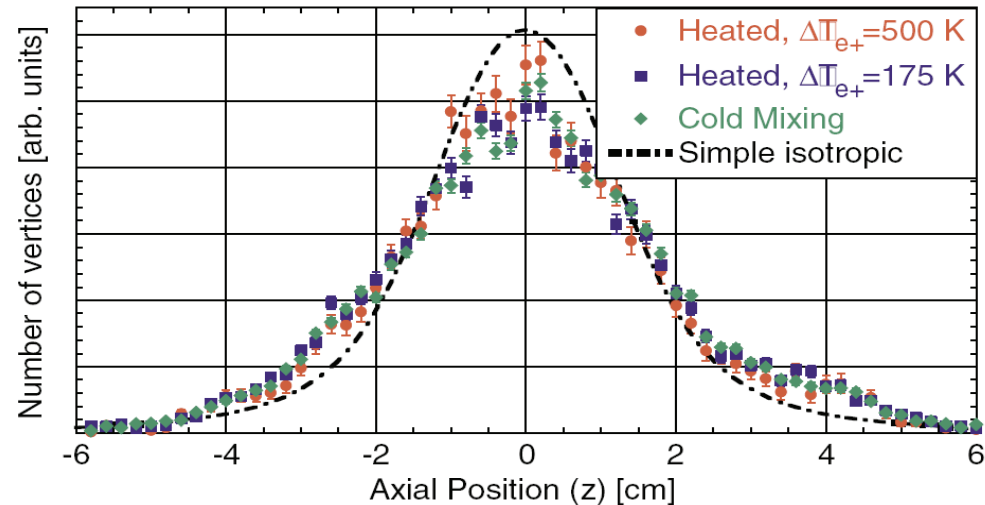
- Antiproton Annihilation (≥ 2 prongs)
- Positron Annihilation (2×511 keV γ)

Time coincidence (~ 1 μ sec)

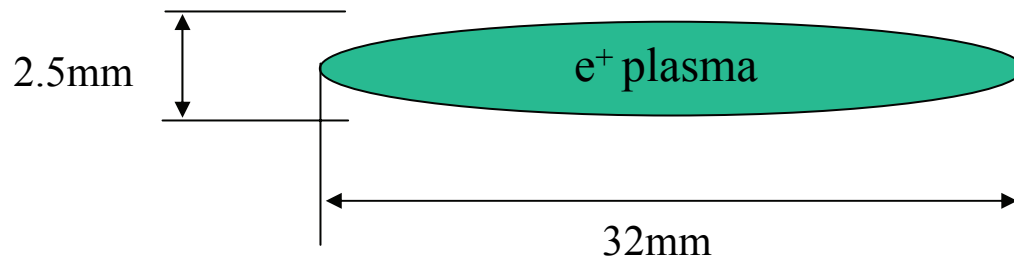
High rate capability

ATHENA measurements

Measurement of the spatial distribution of \bar{H} in dependence of e^+ plasma temperature



model:



and

isotropical
emission of \bar{H}

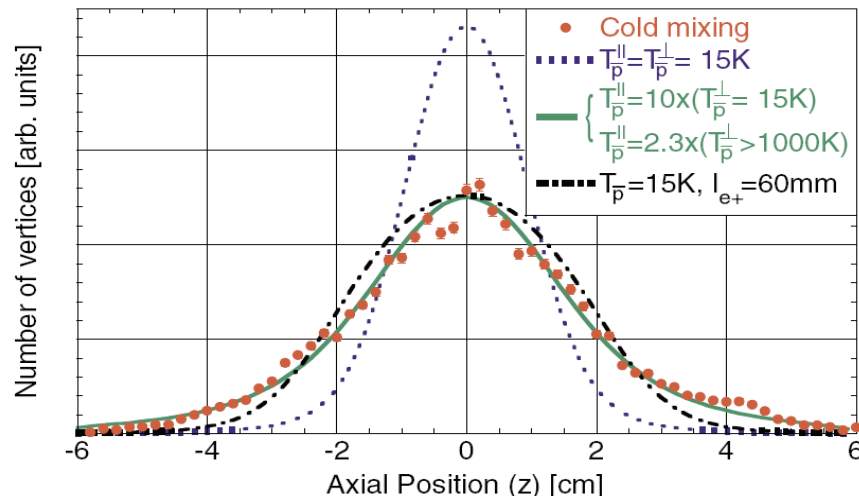
- spatial distribution is independent of e^+ plasma temperature
- \bar{H} is not emitted isotropically

temperature measurement

Model:

- Recombining \bar{p} rotate with e^+ plasma; isotropically produced \bar{H} propagates with momentum of \bar{p}
- using two temperatures to describe nonequilibrium conditions
- spatial distribution measurement provides temperature ratio

from measurement: $T_{\bar{p}}^{para} = (10 \pm 2) T_{\bar{p}}^{perp}$

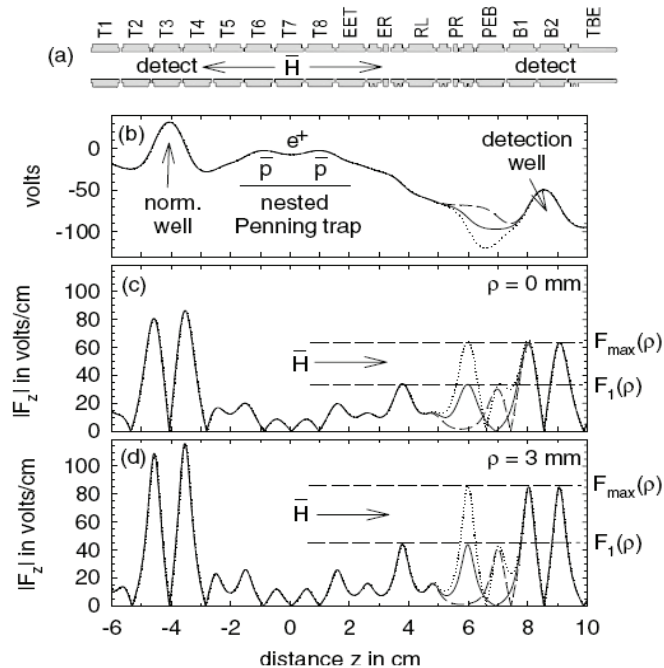


With $T_{\bar{p}}^{perp} \geq 15K$ surrounding temperature

$$\Rightarrow T_{\bar{p}}^{para} \geq 150K$$

\bar{p} and e^+ are not in thermal equilibrium i.g. cooling rate is much lower than recombination rate!

Temperature measurement at ATRAP

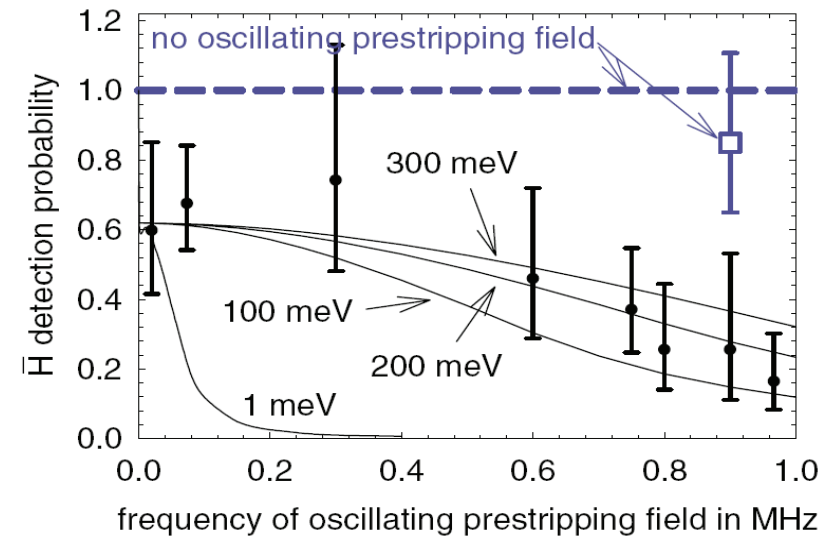


Measurement of velocity distribution:

- Oscillating field (at radiofrequencies)
- Ionization probability in oscillating field is higher for slower \bar{H}
- detection of ionized \bar{H} (antiprotons)

Best fit for : $\bar{E}_{kin} = 200 \text{ meV}$

corresponds to $T_{\bar{H}} = 2400 \text{ K}$



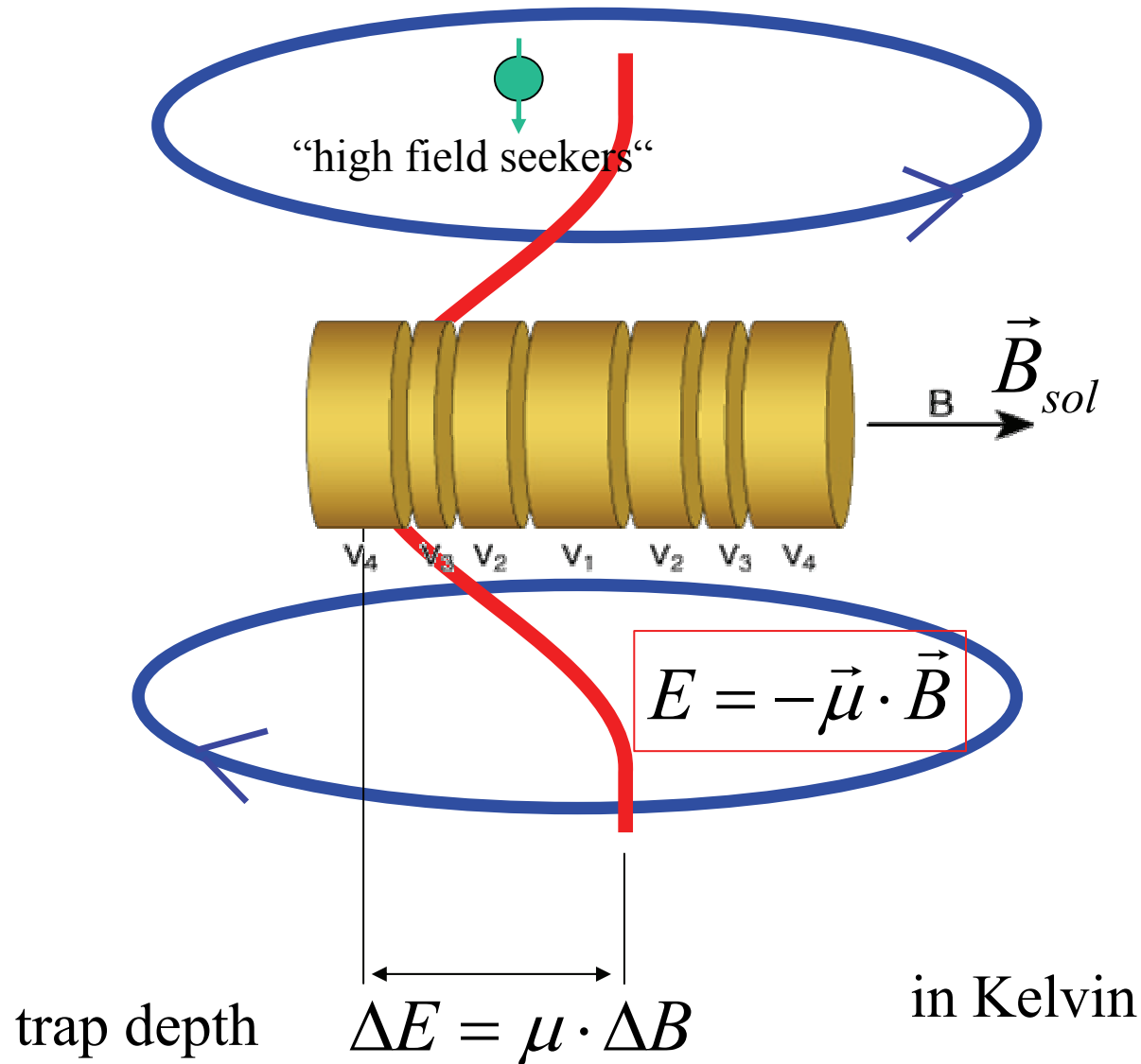
The ALPHA Project (since 2006, successor to ATHENA)

- Goal: Trapping $\bar{\text{H}}$ for spectroscopy!
- Since $\bar{\text{H}}$ is neutral it can not be trapped in a Penning trap!
- Other method: using magn. momentum of $\bar{\text{H}}$
- But: How deep is such a trap? How hot is $\bar{\text{H}}$ allowed to be?

Summary:

- ATRAP:
 - $T_{\bar{\text{H}}} \approx 2400 \text{ K}$
- ATHENA:
 - $T_{\bar{\text{H}}} \geq 150 \text{ K}$

\bar{H} trapping with magn. quadrupole



For $B_r = 1\text{T}$, $B_{sol} = 6\text{T}$

$$\Rightarrow \Delta B \approx 0.1\text{T}$$

$$T = 0.07\text{K}$$

$$T \propto \frac{\Delta E}{k_B}$$

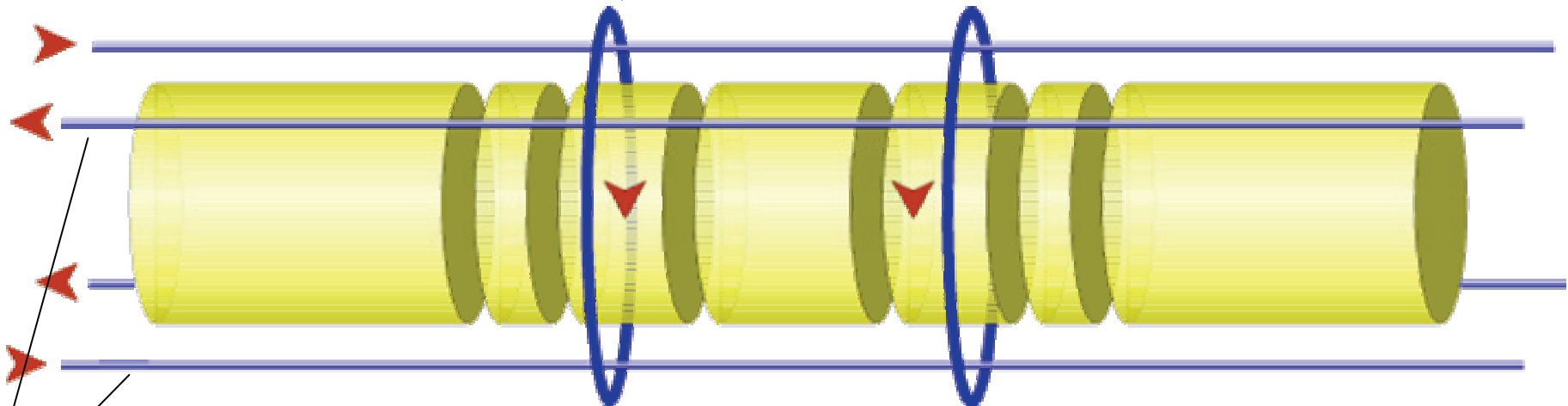
$$T = 0.7 \frac{\text{K}}{\text{T}} \cdot \Delta B$$



$$|B| = \sqrt{B_{sol}^2 + B_r^2}$$

$$\Rightarrow \Delta B = \sqrt{B_{sol}^2 + B_r^2} - B_{sol}$$

Helmholz configuration for axial trapping



Anti-Helmholz configuration for radial trapping

AEGIS project (planned to be in AD)

Goal: direct measurement of g for \bar{H}

Rydberg
positronium
Beam (laser
exited)

e^+e^-

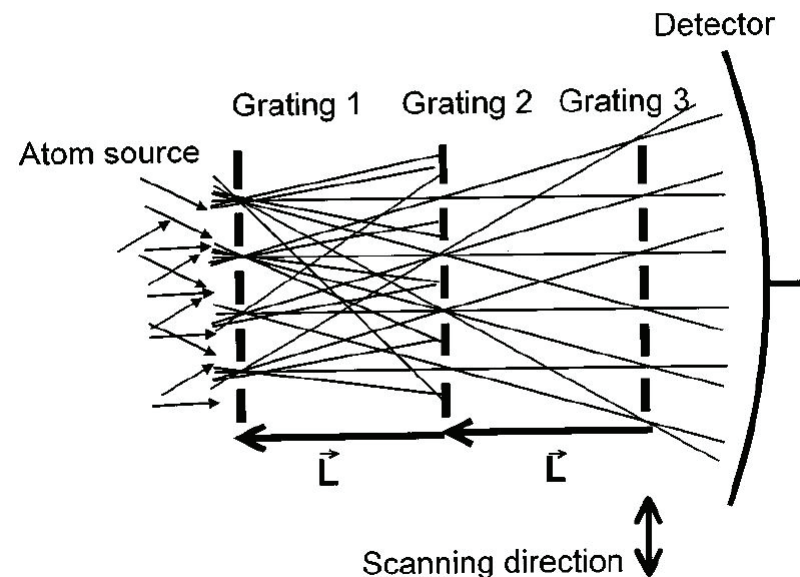
\bar{p}

Cooled in Penning trap
 $T \approx 100 \text{ mK}$

Stark accelerator

\bar{H}^*
 $v \sim 100 \text{ m/s}$

Moiré-deflectometer



In AEGIS: With two gratings and position dependent detector

Precision: $\sim 1\%$

conclusion

- Low temperatures needed for trapping $\bar{\text{H}}$ and to
 - do spectroscopy experiments (CPT test; precision $10^{13}!!!$)
 - test gravity for antimatter
- Temperatures still too high for trapping!
- Challenge: Cooling of negative ions (\bar{p}) to build $\bar{\text{H}}$ at very low temperatures ($\sim\text{mK}$)
- Further cooling of $\bar{\text{H}}$ with lasercooling (if convenient lasersystems are developed)

Thank You for Your attention!