

PHY 115A  
Lecture Notes 2B:  
Tunneling and Scattering  
(Griffith's 2.5-2.6,9.1-9.2)

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# Chapter 2

## Tunneling and Scattering

### 2.26 Continuity of the Wave Function

Let's start with the case  $V(x)$  is finite everywhere, then we start from the TISE:

$$-\frac{\hbar^2}{2m} \frac{d^2\psi}{dx^2} + V(x)\psi(x) = E\psi(x)$$

Without loss of generality, we'll investigate continuity at  $x = 0$ , by integrating the TISE from  $-\epsilon$  to  $+\epsilon$ :

$$\int_{-\epsilon}^{+\epsilon} \frac{d^2\psi}{dx^2} dx = \frac{2m}{\hbar^2} \int_{-\epsilon}^{+\epsilon} (V(x) - E) \psi(x) dx$$

We'll assume that we keep  $\epsilon > 0$  here and everywhere below. By the fundamental theorem of calculus the LHS is:

$$\left. \frac{d\psi}{dx} \right|_{+\epsilon} - \left. \frac{d\psi}{dx} \right|_{-\epsilon} = \frac{2m}{\hbar^2} \int_{-\epsilon}^{+\epsilon} (V(x) - E) \psi(x) dx \quad (2.71)$$

In the limit  $\epsilon \rightarrow 0$ , the RHS vanishes since  $V(x)$  is finite, so:

$$\lim_{\epsilon \rightarrow 0} \left( \left. \frac{d\psi}{dx} \right|_{+\epsilon} - \left. \frac{d\psi}{dx} \right|_{-\epsilon} \right) = 0$$

which is to say the derivative of the wave function is continuous, and so the wave function is continuous as well.

But what about infinite (or undefined)  $V(x)$ ? Here we still insist that the wave function be continuous, as otherwise the state of a particle would be undefined at some point. But the derivative need not be continuous, as the  $V(x)$  term in LHS in Equation 2.71 no longer vanishes in the limit  $\epsilon \rightarrow 0$ :

$$\lim_{\epsilon \rightarrow 0} \left( \left. \frac{d\psi}{dx} \right|_{+\epsilon} - \left. \frac{d\psi}{dx} \right|_{-\epsilon} \right) = \lim_{\epsilon \rightarrow 0} \frac{2m}{\hbar^2} \int_{-\epsilon}^{+\epsilon} V(x) \psi(x) dx \quad (2.72)$$

### 2.27 The Dirac Delta Function

The so-called “Dirac Delta Function”  $\delta(x)$  is defined by it's behavior in an integral:

$$\int_{-\infty}^{+\infty} f(x) \delta(x) dx = f(0) \quad (2.73)$$

where it “picks out” the value of  $f(x)$  at  $x = 0$ . It immediately follows (put  $f(x) = 1$ ) that:

$$\int_{-\infty}^{+\infty} \delta(x) dx = 1 \quad (2.74)$$

Also, changing variables to make the substitutions clearer:

$$\int_{-\infty}^{+\infty} g(y) \delta(y) dy = g(0)$$

and putting  $y = x - a$ , we get:

$$\int_{-\infty}^{+\infty} g(x - a) \delta(x - a) dy = g(0)$$

and defining  $f(x) \equiv g(x - a)$  we have:

$$\int_{-\infty}^{+\infty} f(x) \delta(x - a) dy = f(a) \quad (2.75)$$

The Dirac Delta Function isn't real a function at all, but it is often described as one:

$$\delta(x) = \begin{cases} 0, & x \neq 0 \\ \infty, & x = 0 \end{cases}$$

but such a definition shouldn't be taken too seriously. A better way to consider it is as a limit of perfectly reasonable functions with integral one, that get narrower and narrower around 0. Just as the limit of a series of rational numbers can be an irrational number, the  $\delta$ -function is the limit of a sequence of integrable functions, but isn't itself square integrable. We could try:

$$\int_{-\infty}^{+\infty} \delta^2(x) dy = \delta(0) \quad (2.76)$$

but what are we to make of  $\delta(0)$ ? At best, we could say it is in infinity. Mathematicians call the  $\delta$ -function a generalized function or distribution. It only makes sense in the context of its defining integral equation above, and doesn't exist as a function on its own. If you think of what we actually do with wave functions (calculate integrals) this isn't really any limitation at all.

For  $x \neq 0$ ,  $\delta(x) = 0$  is well defined. But otherwise, just stick to its well defined properties (the numbered equations here) within integrals, and we will see the  $\delta$ -function is extremely useful.

## 2.28 Bound State of the Delta Function Potential

We turn to the very useful example a delta function potential.

$$V(x) = -\alpha\delta(x) \quad (2.77)$$

Since we've agreed to never discuss the delta function at  $x = 0$  outside of an integral, we will just say that  $V(x)$  does not have a defined minimum, and so we are free to see if there are normalizable solution with  $E < 0$ .

Away from  $x = 0$ , where  $\delta(x)$  is well defined,  $V(x) = 0$  and the TISE is:

$$\frac{d^2\psi}{dx^2} = -\frac{2mE}{\hbar^2}\psi(x) = \kappa^2\psi(x), \quad \kappa \equiv \frac{\sqrt{-2mE}}{\hbar}.$$

which has general solutions:

$$\psi(x) = Ae^{-\kappa x} + Be^{\kappa x}$$

But for the wave function to be well defined only:

$$\psi(x) = Ae^{-\kappa x}, \quad x > 0$$

and

$$\psi(x) = Be^{\kappa x}, \quad x < 0$$

are acceptable. From continuity of the wave function at  $x = 0$ , we conclude:

$$A = B$$

and write  $\psi(x)$  as:

$$\psi(x) = \begin{cases} Be^{\kappa x} & x \leq 0 \\ Be^{-\kappa x} & x \geq 0 \end{cases}$$

We saw above that the presence of the  $\delta$ -function means the wave function need not be continuous at  $x = 0$ , and in fact:

$$\lim_{\epsilon \rightarrow 0} \left( \frac{d\psi}{dx} \Big|_{+\epsilon} - \frac{d\psi}{dx} \Big|_{-\epsilon} \right) = \lim_{\epsilon \rightarrow 0} \frac{2m}{\hbar^2} \int_{-\epsilon}^{+\epsilon} V(x) (-\alpha \delta(x)) dx$$

The  $\delta$ -function is well defined in the context of this integral, which can be evaluated as:

$$\lim_{\epsilon \rightarrow 0} \left( \frac{d\psi}{dx} \Big|_{+\epsilon} - \frac{d\psi}{dx} \Big|_{-\epsilon} \right) = \lim_{\epsilon \rightarrow 0} \left( -\frac{2m\alpha}{\hbar^2} \psi(0) \right) = -\frac{2m\alpha}{\hbar^2} \psi(0)$$

In our case:

$$\psi(0) = B$$

and:

$$\frac{d\psi}{dx} = \begin{cases} \kappa Be^{\kappa x} & x \leq 0 \\ -\kappa Be^{-\kappa x} & x \geq 0 \end{cases}$$

so:

$$\lim_{\epsilon \rightarrow 0} \left( \frac{d\psi}{dx} \Big|_{+\epsilon} - \frac{d\psi}{dx} \Big|_{-\epsilon} \right) = -\kappa B - \kappa B = -\frac{2m\alpha}{\hbar^2} B$$

or

$$\kappa = \frac{m\alpha}{\hbar^2}$$

or

$$E = -\frac{\hbar^2 \kappa^2}{2m} = -\frac{m\alpha^2}{2\hbar^2}$$

Normalizing the wave function is left as an exercise, it yields:

$$|B|^2 = \kappa$$

## 2.29 Scattering States of the Delta Function Well

For the case that  $E > 0$ , we have the free particle TISE for  $x < 0$ :

$$\frac{d^2\psi}{dx^2} = -\frac{2mE}{\hbar^2}\psi(x) = -k^2\psi(x), \quad k \equiv \frac{\sqrt{2mE}}{\hbar}.$$

with general solution:

$$\psi(x) = Ae^{ikx} + Be^{-ikx}$$

Similarly, for  $x > 0$  the general solution is:

$$\psi(x) = Fe^{ikx} + Ge^{-ikx}$$

and so:

$$\psi(x) = \begin{cases} Ae^{ikx} + Be^{-ikx} & x \leq 0 \\ Fe^{ikx} + Ge^{-ikx} & x \geq 0 \end{cases}$$

and:

$$\frac{d\psi}{dx} = \begin{cases} (iAk)e^{ikx} + (-iBk)e^{-ikx} & x \leq 0 \\ (iFk)e^{ikx} + (-iGk)e^{-ikx} & x \geq 0 \end{cases}$$

Continuity of  $\psi(x)$  at  $x = 0$  requires:

$$F + G = A + B$$

and from:

$$\lim_{\epsilon \rightarrow 0} \left( \frac{d\psi}{dx} \Big|_{+\epsilon} - \frac{d\psi}{dx} \Big|_{-\epsilon} \right) = -\frac{2m\alpha}{\hbar^2}\psi(0)$$

so:

$$\begin{aligned} ik(F - G - A + B) &= -\frac{2m\alpha}{\hbar^2}(A + B) \\ F - G &= (A - B) + i\frac{2m\alpha}{k\hbar^2}(A + B) \end{aligned}$$

Finally:

$$F - G = A(1 + 2i\beta) - B(1 - 2i\beta)$$

where:

$$\beta = \frac{m\alpha}{\hbar^2 k}$$

To measure scattering, let  $A$  represent the (known) incident wave and set

$$G = 0$$

so now we have two equations and two unknowns:

$$F = A + B$$

and:

$$F = A(1 + 2i\beta) - B(1 - 2i\beta)$$

Solving for  $F$  in terms of  $A$ :

$$\begin{aligned} F &= A(1 + 2i\beta) - (F - A)(1 - 2i\beta) \\ 2F(1 - i\beta) &= 2A \end{aligned}$$

and so:

$$F = \frac{A}{1 - i\beta}$$

similarly:

$$B = \frac{i\beta}{1 - i\beta}$$

The reflection coefficient is:

$$R = \frac{|B|^2}{|A|^2} = \frac{\beta^2}{1 + \beta^2}$$

and:

$$T = \frac{|F|^2}{|A|^2} = \frac{1}{1 + \beta^2}$$

Notice that:

$$R + T = 1$$

Now let's look at what happens for:

$$V(x) = +\alpha\delta(x)$$

Nothing changes until we reach the boundary condition on the derivative, which becomes:

$$\lim_{\epsilon \rightarrow 0} \left( \frac{d\psi}{dx} \Big|_{+\epsilon} - \frac{d\psi}{dx} \Big|_{-\epsilon} \right) = +\frac{2m\alpha}{\hbar^2} \psi(0)$$

So we can read off the solutions for this case by putting  $\beta \rightarrow -\beta$  in the solutions:  
so the boundary conditions (keeping  $G = 0$ ) become:

$$\begin{aligned} F &= A + B \\ F &= A(1 - 2i\beta) - B(1 + 2i\beta) \end{aligned}$$

## 2.30 Bound States of the Finite Square Well

Next we consider the finite square well potential:

$$V(x) = \begin{cases} -V_0, & -a \leq x \leq a \\ 0 & \text{otherwise} \end{cases}$$

As the potential is finite everywhere, we expect both  $\psi$  and  $d\psi/dx$  to be continuous. The bound state potentials will have  $-V_0 < E < 0$ . For  $x < -a$  and  $x > a$ , we encounter a familiar form of the TISE:

$$\frac{d^2\psi}{dx^2} = -\frac{2mE}{\hbar^2} \psi(x) = \kappa^2 \psi(x), \quad \kappa \equiv \frac{\sqrt{-2mE}}{\hbar}.$$

with the solutions

$$\psi(x) = Ae^{\kappa x}$$

for  $x < -a$  and

$$\psi(x) = Be^{-\kappa x}$$

for  $x > a$ . Within the potential well at  $-a \leq x \leq a$  we have:

$$\frac{d^2\psi}{dx^2} = -\frac{2m(V_0 + E)}{\hbar^2}\psi(x) = -k^2\psi(x), \quad k \equiv \frac{\sqrt{2m(V_0 + E)}}{\hbar}.$$

with general solution:

$$\psi(x) = C \sin(kx) + D \cos(kx)$$

At this point, we will save ourselves a lot of hassle by noting that  $V(x)$  is even, and so the general solutions can be constructed as either even or odd solutions. (See Griffith's P2.1c). One major benefit of this, is that we need only establish continuity for  $x \geq 0$  and we will automatically have it for  $x \leq 0$ .

Starting with the even solutions and  $x \geq 0$ , we write:

$$\psi(x) = \begin{cases} Be^{-\kappa x} & x > a \\ D \cos(kx) & 0 \leq x \leq a \end{cases}$$

and continuity implies:

$$Be^{-\kappa a} = D \cos(ka) \tag{2.78}$$

We see that the dimensionless quantities  $\kappa a$  and  $ka$  are of interest, so it will simplify things to note that:

$$(ka)^2 + (\kappa a)^2 = \frac{-2mEa^2}{\hbar^2} + \frac{2m(E + V_0)a^2}{\hbar^2} = \frac{2mV_0a^2}{\hbar^2} \equiv z_0^2$$

The derivative is:

$$\frac{d\psi}{dx} = \begin{cases} -\kappa B e^{-\kappa x} & x > a \\ -kD \sin(kx) & 0 \leq x \leq a \end{cases}$$

which both must be continuous at  $x = a$ , so:

$$\begin{aligned} -\kappa B e^{-\kappa a} &= -kD \sin(ka) \\ \kappa B e^{-\kappa a} &= kD \sin(ka) \\ (\kappa a) B e^{-\kappa a} &= (ka) D \sin(ka) \end{aligned}$$

where we have multiplied by  $a$  so that the equation is in terms of the dimensionless constants  $ka$  and  $\kappa a$ . Dividing by Equation 2.78 we obtain:

$$\begin{aligned} \kappa a &= (ka) \tan(ka) \\ \sqrt{z_0^2 - (ka)^2} &= (ka) \tan(ka) \\ \sqrt{z_0^2 - z^2} &= z \tan(z) \end{aligned}$$

where we have defined

$$z \equiv ka = \frac{2m(E + V_0)a}{\hbar^2}$$



and so finally:

$$\sqrt{\left(\frac{z_0}{z}\right)^2 - 1} = \tan(z) \quad (2.79)$$

(Interpretation, Limiting cases...)

## 2.31 Scattering States of the Finite Square Well

For  $E > 0$ , we have only free particle solutions, but with two different wave numbers, so we'll take:

$$k \equiv \frac{\sqrt{2mE}}{\hbar^2}, \quad \text{and} \quad k' \equiv \frac{\sqrt{2m(E + V_0)}}{\hbar^2}$$

with:

$$\frac{k'}{k} = \sqrt{\frac{E + V_0}{E}} \equiv \eta$$

and so:

$$k' = \eta k$$

Note that if this was an optics problem we would call  $\eta$  the refractive index. The piecewise general solution to the TISE is:

$$\psi(x) = \begin{cases} Ae^{ikx} + Be^{-ikx} & x < -a \\ Ce^{i\eta kx} + De^{-i\eta kx} & -a \leq x \leq a \\ Fe^{ikx} + Ge^{-ikx} & x > a \end{cases}$$

but we are going to write it this way instead:

$$\psi(x) = \begin{cases} Ae^{ik(x+a)} + Be^{-ik(x+a)} & x < -a \\ Ce^{i\eta kx} + De^{-i\eta kx} & -a \leq x \leq a \\ Fe^{ik(x-a)} + Ge^{-ik(x-a)} & x > a \end{cases}$$

which will make it much easier to evaluate at the boundaries ( $x = \pm a$ ). We are free to do this because:

$$Ae^{ik(x+a)} = (Ae^{ika})e^{ikx}$$

where  $A$  is a constant we have not yet determined. The derivative is:

$$\frac{d\psi}{dx} = \begin{cases} ik(Ae^{ik(x+a)} - Be^{-ik(x+a)}) & x < -a \\ i\eta k(Ce^{i\eta kx} - De^{-i\eta kx}) & -a \leq x \leq a \\ ik(Fe^{ik(x-a)} - Ge^{-ik(x-a)}) & x > a \end{cases}$$

Define:

$$\alpha \equiv e^{i\eta ka}$$

and note that:

$$|\alpha|^2 = 1$$

which will be useful later. Now the four continuity equations are:

$$\begin{aligned} A + B &= C\alpha^* + D\alpha \\ A - B &= \eta(C\alpha^* - D\alpha) \\ C\alpha + D\alpha^* &= F + G \\ \eta(C\alpha - D\alpha^*) &= F - G \end{aligned}$$

The incoming waves are  $A$  and  $G$ . In principle, we can use two equations to eliminate the intermediate waves  $C$  and  $D$ . Then we can use the remaining two to calculate the outgoing waves  $B$  and  $F$  from the incoming waves  $A$  and  $G$ .

Note that the first two equations can be added to eliminate  $B$ :

$$A = \frac{(\eta + 1)\alpha^*}{2} C - \frac{(\eta - 1)\alpha}{2} D$$

But now let's make our lives a little easier and set  $G = 0$ . Then the last two equations become:

$$\begin{aligned} \alpha C + \alpha^* D &= F \\ \eta\alpha C - \eta\alpha^* D &= F \end{aligned}$$

Which we use to solve for  $C$  and  $D$  in terms of  $F$ :

$$C = \frac{\eta + 1}{2\eta} \alpha^* F$$

and:

$$D = \frac{\eta - 1}{2\eta} \alpha F$$

Plugging these back in to our expression for  $A$ :

$$\begin{aligned} A &= \frac{(\eta + 1)^2 (\alpha^2)^* - (\eta - 1)^2 \alpha^2}{4\eta} F \\ &= \frac{\eta^2 + 1}{2\eta} \frac{(\alpha^2)^* - \alpha^2}{2} + \frac{(\alpha^2)^* + \alpha^2}{2} \end{aligned}$$

Recalling our definition for  $\alpha$ :

$$A/F = \cos(2\eta ka) - i \frac{\eta^2 + 1}{2\eta} \sin(2\eta ka)$$

and so the transmission is:

$$\frac{1}{T} = 1 + \frac{\eta^2 - 1}{4\eta^2} \sin^2(2\eta ka)$$

Notice that the transmission is one when either:

$$\eta = 1$$

which corresponds to no change in index of refraction, i.e.  $V_0 = 0$ . Or:

$$\sin^2(2\eta ka)$$

that is when:

$$2a \frac{\sqrt{2m(E + V_0)}}{\hbar} = n\pi$$

or:

$$E + V_0 = \frac{n^2 \pi^2 \hbar^2}{2m(2a)^2}$$

which you may recognize as the allowed energies from the infinite potential well.

## 2.32 save...

$$\begin{aligned} B - \alpha^* C - \alpha D &= -A, \\ F - \alpha C - \alpha^* D &= -G, \\ B + \eta \alpha^* C - \eta \alpha D &= A, \\ F - \eta \alpha C + \eta \alpha^* D &= G, \end{aligned}$$

Also, since:

$$(k'a)^2 - (ka)^2 = \frac{2mV_0a^2}{\hbar^2} \equiv z_0^2$$

it's related to our previous constant by:

$$n^2 - 1 = (z_0/z)^2$$

## 2.33 The WKB Approximation and the “Classical” Region

## 2.34 Tunneling in the WKB Approximation

## 2.35 The Fourier Transform Revisited

Our inner product now extends between positive and negative infinity:

$$\langle \Psi, \phi \rangle \equiv \int_{-\infty}^{\infty} \Psi^*(x) \phi(x) dx \quad (2.80)$$

Our basis functions, which are now defined for any value of  $k$ ,

$$e_k = \frac{1}{\sqrt{2\pi}} \exp(ikx) \quad (2.81)$$

are still orthonormal, but the condition looks a bit different in the continuum case:

$$\langle e_k, e_{k'} \rangle = \delta(k - k')$$

See the appendix for more details on the Dirac delta function  $\delta(x)$ , which is zero everywhere but at  $x = 0$ , where it is infinite. It is the continuous version of  $\delta_{nm}$ .

Our basis functions are also still complete. In the discrete case we have a complex Fourier coefficient for every integer  $n$ . Now we have a complex Fourier coefficient for any real value of  $k$ .

In place of Fourier coefficients, we have instead a function of  $k$  which we call the Fourier transform:  $\tilde{\Psi}(k)$ . Instead of a sum over discrete terms, we now have to integrate over all values of  $k$ :

$$\Psi(x) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} \tilde{\Psi}(k) \exp(ikx) dk. \quad (2.82)$$

Just as in the discrete case, we determine the Fourier transform from the inner product:

$$\tilde{\Psi}(k) = \langle e_k, \Psi \rangle = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} \Psi(x) \exp(-ikx) dx \quad (2.83)$$

Equation 2.83 is generally referred to as the *Fourier Transform*, while Equation 2.82 is referred to as the *Inverse Fourier Transform*.

## 2.36 The Fourier Transform in Quantum Mechanics

So far we have been considering the Fourier transform with respect to position  $x$  and wave-number  $k$ . A much more useful pair of variables for Quantum Mechanics turns out to be momentum  $p$  and position  $x$ . To relate  $p$  to  $k$  we need only apply the DeBroglie relation to the wavelength in the definition of the wavenumber:

$$k \equiv \frac{2\pi}{\lambda} = \frac{2\pi p}{h} = \frac{p}{\hbar}$$

We could therefore make the substitution  $k \rightarrow p/\hbar$  (and  $dk \rightarrow dp/\hbar$ ) in Equations 2.82 and 2.83. It turns out that a marginally more useful equation results if we make the normalization factors symmetric, by splitting the normalization factor of  $1/\hbar$  across both equations with  $1/\sqrt{\hbar}$  applied to each:

$$\Psi(x) = \frac{1}{\sqrt{2\pi\hbar}} \int_{-\infty}^{\infty} \tilde{\Psi}(p) \exp(ipx/\hbar) dp \quad (2.84)$$

$$\tilde{\Psi}(p) = \frac{1}{\sqrt{2\pi\hbar}} \int_{-\infty}^{\infty} \Psi(x) \exp(-ipx/\hbar) dx \quad (2.85)$$

The major benefit of this symmetric form is that the normalization of  $\Psi(x)$  and  $\tilde{\Psi}(p)$  in this case turns out to be the same:

$$\int_{-\infty}^{\infty} |\Psi(x)|^2 dx = \int_{-\infty}^{\infty} |\tilde{\Psi}(p)|^2 dp = 1$$

Because we can always calculate  $\Psi(x)$  from  $\tilde{\Psi}(p)$  either one completely describes the quantum mechanical state. We call  $\tilde{\Psi}(p)$  the momentum wave function. Whereas  $|\Psi(x)|^2$  gives us the probability density for the quanton to be at position  $x$ ,  $|\tilde{\Psi}(p)|^2$  gives us the probability density for the quanton to have momentum  $p$ .