Covariances of evaluated nuclear cross-section data for 232 Th. 180,182,183,184,186 W and 55 Mn

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The EMPIRE code system is a versatile package for nuclear model calculations that is often used for nuclear data evaluation. Its capabilities include random sampling of model parameters, which can be utilised to generate a full covariance matrix of all scattering cross sections, including crossreaction correlations. The EMPIRE system was used to prepare the prior covariance matrices of reaction cross sections of ²³²Th, ^{180,182,183,184,186}W and ⁵⁵Mn nuclei for incident neutron energies up to 60 MeV. The obtained modelling prior was fed to the GANDR system, which is a package for a global assessment of nuclear data, based on the Generalised Least-Squares method. By introducing experimental data from the EXFOR database into GANDR, the constrained covariance matrices and cross-section adjustment functions were obtained. Applying the correction functions on the cross sections and formatting the covariance matrices, the final evaluations in ENDF-6 format including covariances were derived. In the resonance energy range, separate analyses were performed to determine the resonance parameters with their respective covariances. The data files thus obtained were then subjected to detailed testing and validation. Described evaluations with covariances of $^{232}\mathrm{Th}.~^{180,182,183,184,186}\mathrm{W}$ and $^{55}\mathrm{Mn}$ nuclei are included into the ENDF/B-VII.1 library release.

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y the designers test new conaterial compog versatile exer licensing repower of modxperiments has been taken over by computational models. Even the best and most accurate models can only be as good as the input data, so there is an increasing demand from nuclear research, industry, safety and regulatory bodies for best estimate predictions of system performance in terms of the design and operational parameters of nuclear re-

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actors, which are to be provided with their confidence bounds. Uncertainties propagated from evaluated microscopic neutron cross-section data are an important component needed to estimate the accuracy of predictions of such integral quantities. The concern about data uncertainties is also related to the need to ensure that nuclear power will continue to be safe, reliable and economically competitive with other alternative energy options. Modern nuclear reactor code systems are now able to accommodate nuclear data uncertainties (e.g. the TSUNAMI module of the SCALE system [1] or the SUSD3D code [2], interfaced to several commonly used deterministic transport codes) thereby providing further stimulus for their production. Along with the growing demand for evaluated nuclear data that include uncertainties and correlations, considerable progress was made in evaluation methodologies. Several subgroups of the NEA Working Party for Evaluation Co-operation (WPEC) have been working on the subject and technical reports have been published recently [3–6].

Addressing the above-mentioned data needs the Nuclear Data Section of the International Atomic Energy Agency (IAEA) actively supported several activities to develop consistent methodologies for evaluating crosssection data including covariances. There is no uniquely correct way of doing this, but the resulting methodology discussed in the present contribution is based on a hybrid approach [4, 7]: modelling uncertainties are incorporated into the model prior, which is calculated stochastically within the EMPIRE code system [8]. Experimental data are incorporated into the evaluation by the Generalized Least-Squares (GLSQ) method as implemented within the GANDR system [9]. The hybrid evaluation method that combines properly-weighted experimental and model covariance matrices was the basic tool in the evaluations of the nuclear reaction data of ²³²Th, W-isotopes and ⁵⁵Mn that includes full covariance information as will be shown below. The choice of the nuclides was driven by specific needs and outstanding deficiencies in the existing data as discussed in several IAEA projects. More detailed information for each studied case will be given in the sections that follow.

Overview of the evaluations

The EMPIRE system [8] is quite flexible in the ways that prior covariance information can be generated. The approach adopted by an informal international consortium under the umbrella of the IAEA is described herein. The first material to be evaluated by this methodology was ²³²Th within the scope of the IAEA Co-ordinated Research Project (CRP) on "Evaluated nuclear data for Th-U fuel cycle" [10]. The evaluation included valuable contributions from different members of the CRP and was truly an international effort. The key point was the inclusion of integral benchmarks into the evaluation loop. In this way the sensitivities to various model pa-

rameters could be assessed early in the evaluation stage. Fine tuning of the input model parameters using integral data (within parameters' uncertainties estimated by expert judgement) was in some cases possible, where differential measurements were insufficient to determine the cross sections more precisely. It should be noted that parameters' variations based on integral data were introduced manually and no integral data were considered in the least-squares fit by GANDR. Within the same IAEA CRP, evaluated data files for ^{231,233}Pa were also produced, albeit with less rigour because of a poorer experimental database available for these two isotopes. All three evaluations were adopted for the ENDF/B-VII.0 library [11]. Since then, minor changes to the cross sections of 232Th were made due to a slight refinement of the assigned correlations to the experimental data. ^{231,233}Pa covariance evaluations were removed during the review process of the ENDF/B-VII.1 beta evaluations as they were considered of lower quality due to the scarcity of available experimental data, and they are not important for current applications.

Subsequently, the evaluations for the five naturallyoccuring isotopes of tungsten were started. The evaluation for ¹⁸⁰W was the first of its kind. Preliminary results were published at the ND2007 conference in Nice [12], and the files performed quite well on the FNG-W benchmark from the SINBAD compilation [13]. Extending the suite of benchmarks to the critical assemblies from the ICSBEP compilation [14] revealed that the benchmarks are sensitive to capture in the 100 keV range, similarly to the FNG-W benchmark, but with opposite indication: the ICSBEP criticality benchmarks over-predicted the reactivity and required an increase in capture while the FNG-W underpredicted the activation of gold foils in a tungsten block and required a decrease in capture. The evaluation procedure was therefore an iterative process of fine-tuning the capture and other cross sections in different energy regions to retain good performance for the FNG-W benchmark, and at the same time improve the performance in the ICSBEP benchmarks. The results of this iterative process were presented at Port Jefferson in 2008 [15]. Since then, only minor changes to the evaluated nuclear data files were made. For example, an error was removed in the assigned gamma-width of the ¹⁸⁴W resonance at 1132 eV, which was discovered through benchmarking, and a few small changes were made to the covariance data for consistency with integral observables.

Very recently the evaluation of the reaction crosssection data of ⁵⁵Mn was performed using the same methodology. Detailed results are presented in the sections that follow.

II. COVARIANCE EVALUATION METHODOLOGY

A. Resonance region

The EMPIRE system has an in-built capability to produce resolved resonance files with covariances from the information available in the Atlas of Neutron Resonances by Mughabghab [16]. This option is commonly applied in the evaluations produced at the National Nuclear Data Center at the Brookhaven National Laboratory. In our case we preferred to use the resonance data with covariance information obtained directly from the analysis of the transmission and capture measurements, when possible. In all discussed cases the resonance evaluations from the Oak Ridge National Laboratory (ORNL) were adopted. The ²³²Th resonance data represent a new evaluation that includes the covariance information [17]. Resonance parameters of the tungsten isotopes were taken from existing evaluations with minor adjustments and the covariance information was generated by the retroactive method [18] because insufficient new measurements were available to justify a re-evaluation. Since then, new measurements have been completed at the IRMM laboratory in Geel [19] and analysis is in progress at ORNL, but is not expected to be finalised in a short time and is not included in the present work. The ⁵⁵Mn resonance data including uncertainties are the result of a new evaluation performed at ORNL [20], including the covariance information.

B. Fast neutron region: Uncertainty of the reaction models and model-parameter uncertainties

In the fast energy region (i.e. above the resonance region) the covariance information is prepared in two steps. In the first step the covariance prior is generated by a Monte Carlo technique based on a random sampling of model parameters and subsequent nuclear model calculations using the EMPIRE code [8]. In the second step the calculated cross sections and the corresponding covariance matrix generated from parameters' sampling in model calculations are fed as a prior information to the GANDR system that combines it with selected experimental data as described in the next Section.

A brief description of the reaction modelling relevant for covariance calculations is given next. Readers interested in a detailed description of theoretical modelling are referred to a companion paper by Chadwick et al. [21]. It is well established that a set of optical model potential (OMP) parameters that can reproduce the nucleon scattering data over a wide energy range is essential for reliable nuclear data predictions. All IAEA evaluations in the fast neutron range described in this contribution have been based on newly derived coupled-channel optical model potentials using dispersion relations (see Refs.[22, 23] and references therein). Start-

ing values for nuclear model parameters and their uncertainties were taken from the RIPL recommendations [23, 24]. Pre-equilibrium emission was considered using a one-component exciton model (PCROSS), which includes nucleon, gamma and cluster emission. Hauser-Feshbach [25] and Hofmann-Richert-Tepel-Weidenmüller [26] versions of the statistical model were used for the compound nucleus cross-section calculations. Both approaches account for the multiple-particle emission and the full gamma-cascade. Level densities were described by the "EMPIRE specific" formalism, which uses the superfluid model below the critical excitation energy and the Fermi gas model above [8, 23]. A modified Lorenzian (MLO) radiative-strength function was taken as recommended by Plujko [23].

The use of model calculations to the determination of covariances for the nuclear reaction observables is very transparent [27]. First, physical model parameters that play a significant role in defining reaction observables of interest are identified. A typical list of model parameters with assigned uncertainties used to calculate the model prior uncertainties for the ²³²Th evaluation is given in Table I of Section III.A. The number of sampled parameters is case-dependent. Usually we include level density parameters for the target and compound nuclei as well as for selected residual nuclei of major competing channels. Other parameters of the reaction model such as the exciton model mean-free path (pre-equilibrium parameter with assumed relative uncertainty of 30%) and optical model parameters (e.g. potential depth, geometric parameters, static and dynamic deformation) are also considered. Adopted model parameter uncertainties are usually those recommended in the RIPL database [23]. Additionally, scaling (or normalisation) parameters (e.g. of the total and the reaction cross sections) with empirically estimated uncertainties take into account systematic deficiencies of nuclear-reaction models. Scaling (nonphysical) parameters may be energy dependent, in which case they are assumed to be fully correlated between different energies. The assumption of fully-correlated scaling parameters is required to avoid fluctuations of calculated cross sections at different energies, i.e. the relevant cross section is scaled but the model shape is preserved. A typical uncertainty of scaling parameters is small: 3%. Once all random variables (parameters) and their corresponding uncertainties are defined, then the EMPIRE code is run for a number of times. Previously selected input parameters are being sampled randomly within the assumed uncertainty around the central (nominal) values of the parameters [28]. A normal distribution of the parameters is assumed; the uncertainty of the parameters is limited to 33 % (if needed) to reduce the number of sampled negative values, which are often unphysical and are discarded. Physical model parameters are assumed to be independent as parameters' correlation does not have a significant impact on already very strong correlation arising from the nuclear reaction model. Sampling of correlated physical parameters is avoided (either the

potential depth or the potential radius is sampled, never both of them).

Each EMPIRE calculation with a different set of sampled random parameters covers the full incident energy range and produces a full set of cross sections, spectra, angular distributions and other observables. EMPIRE also produces a full ENDF-6 formatted file for each of the sampled sets. Variances and covariances of any cross section x (including the double differential ones) or other parameters can be obtained from the randomly sampled values x_k of the cross section x by

$$Cov(x_i, x_j) = \frac{1}{N} \sum_{i,j=1,N} (x_i - \overline{x})(x_j - \overline{x})$$
(1)

where N is the number of samples and \overline{x} is the average over N sampled values of x_k . The amount of information that needs to be stored should remain reasonable and there are some restrictions in the ENDF-6 format itself, so we choose the energy grid points at which the cross sections are sampled and reaction definitions as commonly used in the GANDR system [9]. This implies that all reactions are given on the same energy grid and that some reactions are lumped together to reduce the size of storage of information. Cross correlations between different reactions are automatically taken into account. The GANDR system works in the cross-section space and does not keep track of the input model parameters.

Energy-range correlations of the model covariance for each reaction are typically very strong and generally positive (at least on the short range), reflecting the stiffness of nuclear reaction models. Such property of the modelling covariance matrix is very general, and is related to the small number of model parameters used to describe a very big number of experimental data points. A typical model correlation matrix for the $^{186}W(n,\gamma)$ produced by random sampling is shown in Figure 1. Strong positive correlations are observed even between distant energy points. Correlations are weakened when experimental data are considered by GANDR as seen in Figure 2 (see the next section for further discussion of the GANDR fit).

The Monte Carlo calculated uncertainty as a function of energy for the $^{186}W(\rm n,\gamma)$ reaction is shown as a dashed line in Figure 3. The prior uncertainty goes from 20%–30% below 1 MeV up to 60% in the energy region of 2–10 MeV. Higher uncertainty above 3-4 MeV results from more competing channels open, with the capture contribution becoming negligible compared to competing channels at higher energies. Uncertainty of 20%–30% below 1 MeV is mainly related to the assumed uncertainty of the level density parameter of the compound nucleus, determining the normalization of the capture cross section at low energies.

The Monte Carlo (MC) calculations defining the prior covariance matrix are conceptually straightforward and free of many simplifying assumptions used in deterministic methods, in particular of the assumption of a linear response of the observables to the variation of model pa-

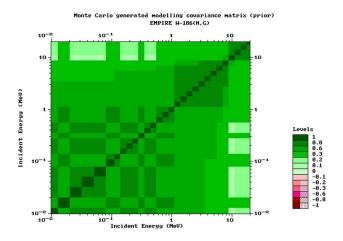


FIG. 1: Correlation matrix of the $^{186}{\rm W}({\rm n},\gamma)$ Monte Carlo prior.

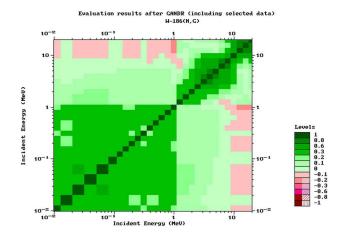


FIG. 2: Correlation matrix of the $^{186}{\rm W(n,}\gamma)$ after the GANDR fit.

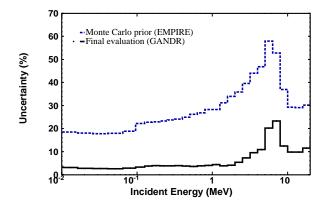


FIG. 3: Uncertainty of the $^{186}\mathrm{W}(\mathrm{n},\gamma)$ reaction in the fast neutron region: Monte Carlo calculated prior with EMPIRE (dashed line) is compared with the evaluation derived by GANDR result (solid line).

rameters [4], which is inherent in the least-squares methods (e.g. KALMAN [29]). There is also no need in MC methods to calculate the sensitivity coefficients, and the computing time is independent of the employed number of model parameters. These advantages come at a price – the number of required MC calculations is in the range of hundreds and the convergence of the results has to be demonstrated.

MC and deterministic methods for generation of the prior information are independent and complementary: the model-based covariances generated by the MC method can be used as a reference benchmark case for validating the faster but linear-model calculations with least-squares codes such as KALMAN.

C. Combining modelling uncertainty and experimental data: the GANDR system

In the second step the calculated cross sections and the corresponding covariance matrix generated from parameters' sampling in model calculations are fed as a prior information to the GANDR system. An alternative to the two-step approach with GANDR would be to use the Unified Monte Carlo method (UMC) [30] in which both the model parameters and the experimental data are sampled together. The UMC method was demonstrated to work efficiently [31, 32] and is free from the effects of nonlinearities, but it is computationally very intensive and needs some additional development for a broad scale application. It should be also remarked that the results of the UMC method converge to the results of the GLSQ method if no experimental ratio data are involved in the evaluation [31].

A careful selection of available experimental data is made (from EXFOR [33]); the experimental data are renormalised to the most recent standards (if needed) or corrected for known deficiencies. Great care is devoted to the assembly of the experimental database that is statistically consistent as least-squares methods have inherent problems in dealing with discrepant data. Unknown experimental uncertainty components are added using expert judgement, if applicable. An estimate of the systematic relative uncertainties is made, separately for the data points from the same measurement and for the data sets from different measurements of the same reaction. Cross correlations between different measurements for the same reaction channel are defined by grouping experiments according to the employed measurement technique (e.g. activation, time-of-flight, etc). Experimental cross correlations between measurements for different reaction channels are not considered, which is probably a reasonable assumption, since the measurement techniques are usually quite different (an exception may be those activation measurements of different reactions undertaken at the same experiment). The selected experimental data are fed into the GANDR system and the adjustment of the cross sections and the covariance matrix is made by

the GLSQ method for each reaction in turn. This can be done without loss of generality since the cross-reaction correlations from model calculations are already included in the prior, and a subsequent addition of experimental data is allowed within the bayesian approach as long as added experimental data are uncorrelated.

The average (nominal) model parameters in the reference EMPIRE calculation are chosen such that they reproduce relatively well the bulk of the experimental data. Therefore, the correction factors for the cross sections after the GANDR least-squares adjustment are generally small (i.e. close to one). The adjustments factors and the covariance matrix are obtained from the GANDR output. Adjusted cross sections and the corresponding covariance matrix are written to the final ENDF-6 file.

Strong energy-range correlations of the prior (model) covariance are significantly softened when experimental data are considered in a Generalized Least-Squares fit by GANDR system. Some anticorrelation of fitted data at distant energies are usually observed as seen in Figure 2, showing the correlation matrix of the evaluation produced by GANDR for the $^{186}W(n,\gamma)$ reaction. Experimental data considered in the fit are described in a relevant section below. A different behavior of the correlation pattern is observed below and above 1 MeV. Below 1 MeV long-range correlations are still quite strong as they were in the prior covariance (see Figure 2). Above 1 MeV correlations became short-range (in fact a quasi-diagonal matrix is obtained). Additionally, anticorrelations are observed between higher and lower energy regions.

The uncertainty of the evaluated data as a function of energy for the $^{186}W(n,\gamma)$ reaction is shown as a solid line in Figure 3. It can be seen that evaluation uncertainties are reduced to 3-4% below 1 MeV from the 20%-30% given for the model prior. This is a direct result of including accurate experimental data in the GANDR fit for this energy region. Uncertainty of the capture cross section is also reduced from 60% (prior) to 20% (final evaluation) for the energy region from 5-10 MeV. There is practically no experimental data available in this energy region. However, strong long-range model correlations from the prior combined with accurate experimental data below 1 MeV lead to the observed uncertainty reduction. There is some additional experimental data around 14 MeV allowing for a further reduction of the capture uncertainty at those higher energies.

Summarizing, we can say that model correlations from the calculated prior define the approximate shape of the evaluated cross sections (including the case when no experimental data are available). On the other side, the existence of accurate and statistically consistent (non-discrepant) experimental data is necessary to obtain cross-section uncertainties of the final evaluation below 10%.

D. Cross-reaction and cross-material covariances

Since all cross sections at the specified energy gridpoints are packed into a single parameter array, the crossreaction covariances are produced naturally by GANDR. The parameter array can also contain Legendre moments of the angle-differential scattering cross section, as well as particle emission spectra. The first angular moment of the elastic cross section and the first angular moment of the first discrete level inelastic cross section are taken into account, which can be stored in the ENDF-6 format as covariances of the first-order Legendre coefficient of the same reaction in the laboratory system, but without correlations with other reactions. Covariances of the emission spectra were not taken into account in the present work (even if they are present in the prior covariance). However, it should be noted that neutron emission spectra are naturally correlated to corresponding cross sections; a fact which may be important for criticality calculations. Further studies of these correlations in future evaluations are warranted.

The most recent version of GANDR can treat experimentally measured elemental cross sections in combination with isotopic cross sections. For tungsten isotopes there are measured total cross-section data available both for the natural element and major isotopes. The inclusion of measurements on natural element in the GLSQ fit leads to cross-material correlations between the evaluated cross sections for different isotopes. Preliminary calculations were made exploiting this option in the case of the tungsten isotopes, but the cross-material covariances were not included in the final evaluation. However, it should be noted that the availability of experiments for all major isotopes of multi-isotopic elements is quite unique; usually only the elemental data exist and derived correlation among different isotopes seem unavoidable in future evaluations of important multi-isotopic elements (e.g. zirconium, iron, lead).

E. Covariance formats and data processing

In general, the ENDF-6 convention is that the total cross section is a redundant reaction. Some evaluators argue that this is the best-known experimental cross section and should be given explicitly, but this is not how the total cross section is usually used in the calculations. The GANDR convention is consistent with the ENDF-6 convention that the total cross section is redundant and that the uncertainty in the total cross section can be reconstructed from the partials. In this process it is critical to account properly for cross-reaction correlations, otherwise the uncertainty in the total cross section could be overestimated (obviously some of the contributing reactions are always anticorrelated). For the same reason it is also important that all reactions that constitute the total cross section are accounted for. To keep the covariance information within reasonable limits in size, reactions are often lumped together. A typical case is the lumping of a number of discrete level reactions. Another example are the two-neutron and three-neutron emission reactions, which are lumped conveniently for transport calculations. This structure could confuse some users, who would be seeking the uncertainty in the (n,2n) reaction, for example. Such a user should be aware that a distinction between the lumped reaction and the (n,2n) covariances occurs only above the threshold for the (n,2n+p) reaction, and similarly for the three-neutron emission case.

The covariances for all reactions in the fast energy range are stored as the upper-triangular relative covariance matrix (Option LB=5 of File 33 in ENDF terminology). In the resonance range the resonance parameter covariances are stored in ENDF-6 File 32. A background uncertainty in the resonance range is also given in some cases. This option is used to account for sources of uncertainty not included in the resonance parameters' part. Most commonly this includes the scattering radius uncertainty (which could not be accommodated into the resonance covariances because the relevant ENDF-6 format extensions were introduced recently), but also some other systematic uncertainties. It should be noted that the use of File 32 allows for the storage of the full correlation information, and facilitates an accurate calculation of the uncertainty of the integral quantities (e.g. the resonance integral with a high level of self-shielding) as demonstrated in a recent paper by some of the authors [34].

III. EVALUATION OF THORIUM

A. EMPIRE model calculations

The details of the selection of optical model and other parameters for calculating the cross sections in the fast energy range are described in a separate paper [21]. The choice of model parameters may include tuning (scaling) factors, which are used to compensate for the deficiencies of the models to represent the cross sections. The list of model parameters and their uncertainties are listed in Table I; OMP means optical model potential. It is worth to remark that OMP depths are sampled, but not the OMP radii to avoid a well known model correlation. The average value of the parameters is chosen empirically following RIPL guidelines [23].

The full covariance matrix corresponds to a vector of parameters, which consists of cross sections, packed into the vector one after the other. Some of the cross sections in the vector represent lumped reactions. The definitions of each cross-section set in the vector is given in Table II. Note that the energy grid is fixed and is the same for all reactions, but some reactions require fewer points due to their respective thresholds.

To generate the covariance matrix prior, all model parameters in Table I were sampled randomly (and independently) using a normal distribution. From each set of

TABLE I: Model parameters with assigned uncertainties for $^{232}{\rm Th}$ evaluation. Target: $^{232}{\rm Th}$; compound: $^{233}{\rm Th}$; proton emission residual: $^{232}{\rm Ac}$; alpha emission residual: $^{229}{\rm Ra}$. Column "Symbol" designates the EMPIRE keyword for the respective parameter.

EMPIRE input parameter	Crmbol	Volue	Una	Nucleus
EMFIRE input parameter	Symbol	varue	[%]	Nucleus
Preeq. mean-free path	PCROSS	1.6	12	
Level density normalization	ATILNO	-	15	$^{233}\mathrm{Th}$
Level density normalization	ATILNO	1.091	10	$^{232}\mathrm{Th}$
· ·	ATILNO	0.982	5	231 Th
Level density normalization	ATILNO	1.042	5 5	$^{230}\mathrm{Th}$
Level density normalization		0.962	5 5	$^{229}{ m Th}$
Level density normalization	ATILNO		_	$^{227}\mathrm{Th}$
Level density normalization	ATILNO	1.091	10	226 Th
Level density normalization	ATILNO	1.022	10	225 Th
Level density normalization	ATILNO	1.061	10	
Level density normalization	ATILNO	1.041	10	²³² Ac
Level density normalization	ATILNO	1.041	10	²³¹ Ac
Level density normalization	ATILNO	1.041	10	²³⁰ Ac
Preeq. level density normal.	GTILNO	1.05	10	²³² Ac
Preeq. level density normal.	GTILNO	1.05	10	²²⁹ Ra
Gamma-width normalization	TUNE	3.3	5	²³³ Th
Dynamic-deformation normal.	DEFDYN	1.0	15	$^{232}\mathrm{Th}$
Fission barrier normalization		1.0	5	all
Fission level density normal.	ATILFI	1.0		all
Real vol. OMP depth normal.	UOMPVV	1.0	3	$^{232}\mathrm{Th}$
Imag.vol. OMP depth normal.	UOMPWV	1.0	5	$^{232}\mathrm{Th}$
Imag.surf. OMP depth normal.	UOMPWS	1.0	5	$^{232}\mathrm{Th}$
Vol. OMP diffuseness normal.	UOMPAV	1.0	2	$^{232}\mathrm{Th}$
Surf. OMP diffuseness normal.	UOMPAS	1.0	3	$^{232}\mathrm{Th}$
Real vol. OMP depth normal.	UOMPVV	1.0	3	$^{231}\mathrm{Th}$
Imag.vol. OMP depth normal.	UOMPWV	1.0	5	$^{231}\mathrm{Th}$
Imag.surf. OMP depth normal.	UOMPWS	1.0	5	$^{231}\mathrm{Th}$
Vol. OMP diffuseness normal.	UOMPAV	1.0	5	$^{231}\mathrm{Th}$
Surf. OMP diffuseness normal.	UOMPAS	1.0	5	$^{231}\mathrm{Th}$

sampled parameters a full ENDF-6 file was generated. In total 283 files were produced. Relevant cross sections were extracted from the files and lumped reaction cross sections were constructed. The cross sections were packed into the cross-section parameter vector. Using Equation (1) the elements of the covariance matrix were generated.

To account for additional unidentified uncertainties in the thorium evaluation, the variance of the prior was scaled by a factor of two. The scaling is arbitrary and may reduce excessively the correlations from the model parameters in some cases. This practice was changed in newer evaluations. However, the final variances are determined through the adjustment procedure by the method of generalised least squares, so the influence of this practice on the final results was considered less important.

B. Selection of experimental data

Experimental data were retrieved from the EXFOR database [33] for the fission, (n,2n), radiative capture and total cross sections. Each data set was carefully reviewed. Whenever possible, measured cross sections were renormalised to the latest standards, or corrected otherwise, if justified. Measurements that were deemed unreliable were rejected.

Fission cross-section measurements included in the analysis are listed in Table III. Since practically all measurements were measured as ratios to a chosen standard, a 2 % fully correlated uncertainty was added to all measurements.

Measured (n,2n) and radiative capture cross sections included in the analysis are listed in Tables IV and V.

The adopted total cross-section measurements are listed in Table VI. Although the total cross section does not appear explicitly in the list of reactions, it is the sum of all the partial cross sections. The measurements act as a constraint on the sum. Although the measured uncertainties are quoted as total, a 1 % systematic uncertainty was acknowledged by Abfalterer et al. in recent publications [35, 36]. Similar analysis is missing in other experimental reports, therefore 1 % systematic uncertainty is adopted by analogy for all measurements, considering that all measurements were done using the time-of-flight method. In addition, a 0.2 % fully correlated uncertainty was assumed between all measurements.

The covariance matrix prior and the above-listed experimental data were entered into the GANDR system and GANDR was run to calculate the adjustment factors

TABLE II: Cross-section definition in the parameter covariance vector in terms of the ENDF-6 reaction designation MT numbers. The Npnt column designates the number of energy points.

Npnt		Contributing MTs (description)
74	2	(elastic)
74	$P_1(2)$	(first moment of the elastic)
74	$P_{2}(2)$	(second moment of the elastic)
74	$P_{3}(2)$	(third moment of the elastic)
13	5	
16	16	24 41 (lumped)
14	17	
36	51	(first discrete inelastic)
36	$P_1(51)$	(first moment of 51)
32	52	53 54 55 56 57 58 59 60
25	61	62 63 64 65 66 67 68 69
	70	(lumped)
28	22	28 71 72 73 74 75 76 77
	78	79 80 81 82 83 84 85 86
	87	88 89 91 (lumped)
74	102	
74	18	
22	600	649 800 849 (lumped)

TABLE III: Selection of ²³²Th(n,f) cross-section measurements from the EXFOR database.

EXFOR No.	Author	Year	Comment
41455006	Shcherbakov et al.	(2001)	Ratio
	Eismont et al.	(1999)	Priv. comm.
31424002	Sastry et al.	(1992)	
31459025	Garlea et al.	(1992)	
41111002	Fursov et al.	(1991)	Ratio
14176006	Lisowski et al.	(1988)	
40888002	Goverdovskij et al.	(1986)	Ratio
21963002	Kanda et al.	(1985)	Ratio
30813007	Garlea et al.	(1984)	
10658002	Behrens et al.	(1982)	Ratio
20796002	Blons et al.	(1975)	
Systematic uncertainties:			
2 %	common to all mea	sureme	nts

TABLE IV: Selection of the ²³²Th(n,2n) reaction cross-section measurements from the EXFOR database.

EXFOR No.	Author	Year	Comment
22845002	Karamanis et al.	(2003)	
41240118	Filatenkov $et\ al.$	(1999)	renormalized
22235002	Chatani et al.	(1993)	
22637089	Konno et al.	(1993)	
22203002	Chatani et al.	(1991)	
30816002	Raics et al.	(1985)	
20499003	Karius et al.	(1979)	
12305002	Cochran et al.	(1958)	

for the cross sections by the GLSQ method. The covariance matrix of the adjustment factors is by definition the relative covariance matrix of the cross sections.

The adjustment factors were applied to the cross sections from the nominal EMPIRE calculation and the full covariance matrix was processed with a local code ENDCOV to create the final evaluated nuclear data file for ²³²Th.

In the resolved resonance range the resonance parameters were obtained from a sequential Bayes analy-

TABLE V: Selection of 232 Th capture cross-section measurements from the EXFOR database.

EXFOR entry	Author	Year	Comment
22875002	Schillebeeckx et al.	(2004)	
22654003	Wisshak et al.	(2001)	
22663002	Karamanis et al.	(2001)	
41183003	Davletshin et al.	(1993)	
20871002	Kobayashi et al.	(1979)	
10735002	Poenitz et al.	(1978)	
10554003	Macklin et al.	(1977)	

TABLE VI: Selection of the 232 Th total cross-section measurements from the EXFOR database.

EXFOR No.	Author	Year	Comment
13753029	Abfalterer et al.	(2001)	
12853052	Poenitz et al.	(1983)	
10935004	Poenitz et al.	(1981)	
21766002	Iwasaki et al.	(1981)	
10047094	Foster Jr et al.	(1971)	
21088003	Uttley et al.	(1966)	
Systematic u	incertainties:		
1.0 %	for each EXFOR	R set	
0.2 %	correlation betw	een all	experiments
-			

sis [17, 37, 38] with the computer code SAMMY [39]. The experimental data base included Olsen [40] neutron transmission data (ORELA), Schillebeeckx [41] capture data (GELINA), and Gunsing [42] capture data (n_ToF) in the energy range 1 eV to 4 keV. In the thermal energy range the capture data of Chrien [43] and of Lundgren [44] were normalized to a value of 7.35 b recommended by Trkov [45] and fitted by SAMMY along with the total cross section of Olsen [40] in the energy range up to 1 eV. The covariances of the resonance parameters were obtained directly from the resonance analysis [17, 38]. They are stored in ENDF-6 File 32.

The covariance matrix of the average resonance parameters in the unresolved resonance region (URR) was deduced from the covariance matrix of the experimental data and the sensitivity matrix or the partial derivatives with respect to the adjusted parameters [46]. The covariance matrix of the experimental data in the URR was constructed by assuming the following uncertainties:

Total cross section:

- correlated : 1.0 % - uncorrelated : 0.5 %

Capture cross section - correlated : 1.5 % - uncorrelated : 1.5 %

The sensitivity matrix was deduced with the RECENT code of the PrePro series [47] around the final average resonance parameters. The covariance and sensivitivy matrix were constructed at the energy points of the tabulated cross-section data.

Since in the fitting procedure only the neutron strength functions S_l for l=0,1 and 2 and the radiation widths $\Gamma_{g,l}$ for l=0,1 were adjusted, only the partial derivatives with respect to these parameters were considered. To construct the covariance matrix in File 32 a very small relative uncertainty on the level distance was introduced and it was assumed that the relative covariance elements for the reduced neutron widths can be approximated by the relative covariance elements of the neutron strength functions.

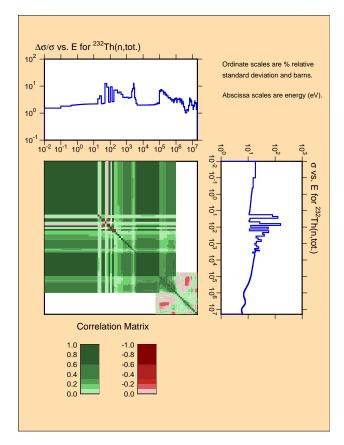


FIG. 4: Uncertainty and correlation matrix of the $^{232}\mathrm{Th}$ total cross section.

A 2 % uncertainty was added from 10^{-5} eV to 897 keV to the elastic cross section in File 33 to account for the scattering radius uncertainty.

The data on the average number of neutrons per fission (including covariances) were adopted from the BROND-2 evaluation, which is based on a least-squares fit of selected experimental data sets.

C. Analysis of obtained covariances

The full evaluated files with covariances were verified by the ENDF Utility Codes [48] to check for the formal correctness. In the next step the covariance plots were generated on a fine energy grid with the NJOY processing system [49] to check processability and the general features of the covariance matrices, including cross-reaction correlations. The covariance matrix of the total cross section is the sum of the covariance matrices of the partial reaction cross sections. The variance in the total cross section is smaller than the sum of the variances of the partials due to cross-reaction anti-correlations. The correlation matrix of the total cross section is shown in Figure 4 as an example of a derived result.

The next step in testing the covariance matrices was to generate with the NJOY code [49] the average cap-

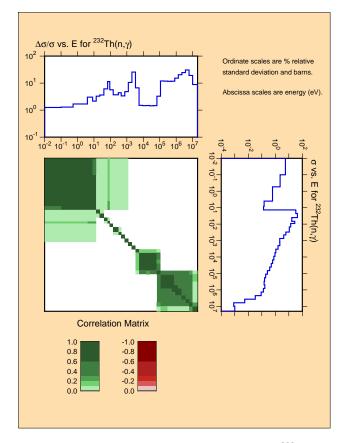


FIG. 5: Uncertainty and correlation matrix of the $^{232}\mathrm{Th}(\mathrm{n},\gamma)$ reaction.

ture cross sections in a thermal Maxwellian spectrum and the resonance integral in a pure 1/E spectrum from 0.55 eV to 2 MeV, including uncertainties, which can be compared to the values derived from the database of constants for neutron activation analysis recommended by the International Union of Pure and Applied Chemistry (IUPAC) [50]. The comparison is described in Table VII. The thermal capture cross-section uncertainty derived from the k_0 constants in the IUPAC database is lower because it does not include the systematic uncertainty in the gamma-emission probabilities. The resonance integral calculated from the evaluated cross sections is lower, but since the evaluation is based on two very recent data sets measured on GELINA at IRMM in Geel and on n_ToF at CERN in Geneva, which are in excellent agreement, we believe that the uncertainty estimate is reliable. Particularly, the accuracy quoted in Table VII is observed also in the unresolved resonance region. From Figure 5 it is evident that the uncertainties at the upper end of the resolved resonance range (10 eV to 4 keV) are mainly statistical (correlations are negligible) and contribute little to the total uncertainty of the resonance integral.

TABLE VII: Comparison of thermal capture cross section and resonance integral uncertainties of ²³²Th with values derived from constants for neutron activation analysis [50].

	IUPAC	C/Nudat	ENDF/	B-VII.1	
		${\bf Uncert.}$		${\bf Uncert.}$	Difference
	$[{\rm barns}]$	[%]	$[{\rm barns}]$	[%]	[%]
σ_0	7.34	0.5	7.338	1.2	0.0
RI	84.64	3.6	84.23	1.9	-0.5

IV. EVALUATION OF TUNGSTEN ISOTOPES

A. EMPIRE model calculations

As in the case of the ²³²Th nucleus, the details of the selection of optical model and other parameters are described in a separate paper [21]. The choice of model parameters includes tuning (scaling) factors to compensate for the deficiencies of the reaction models used to represent the cross sections. The absolute value of correction is always less than 5 %. These "non-physical" parameters could be energy dependent. It is important to realise that the values of tuning parameters at different energies are correlated. Full correlation was assumed when energy-dependent parameters were employed. The list of parameters and their uncertainties is similar to those listed for thorium, except fission parameters, as tungsten was assumed to be a non-fissile element. The threshold for fission is at about 60 MeV of neutron incident energy and was not included in the models. It was added subsequently as a dosimetry reaction in ENDF-6 File 10, since the cross section is so small that it contributes negligibly to the summation of the total cross section.

The definitions of cross sections in the cross-section parameter vector for all tungsten isotopes are given in Tables VIII-XII. Note that the energy grid is fixed and is the same for all reactions, but some reactions require fewer points due to their respective thresholds.

The covariance matrix prior was generated in the same way as for ²³²Th. In total, 449 files were produced for ¹⁸⁰W, 498 for ¹⁸²W, 499 for ¹⁸³W, ¹⁸⁴W and ¹⁸⁶W. Relevant cross sections were extracted from the files, lumped reaction cross sections were constructed. The cross sections were packed into the cross-section parameter vector and the elements of the covariance matrix were generated as described for ²³²Th,

To account for additional unidentified uncertainties, the relative variance of all cross-section parameters was increased by 0.0009, which corresponds to 3 % standard deviation in the prior.

TABLE VIII: Cross-section definition in the parameter covariance vector of $^{180}\mathrm{W}$ in terms of the ENDF-6 reaction designation MT numbers. The Npnt column designates the number of energy points.

$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$			
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	Npnt		Contributing MTs
74 P ₂ (2) (second moment of the elastic) 74 P ₃ (2) (third moment of the elastic) 17 5 37 (lumped) 14 16 41 (lumped) 12 17 33 51 (first discrete inelastic) 33 P ₁ (51) (first moment of 51) 28 52 53 54 55 56 57 58 26 28 91 (lumped)	74	2	(elastic)
74 P ₃ (2) (third moment of the elastic) 17 5 37 (lumped) 14 16 41 (lumped) 12 17 33 51 (first discrete inelastic) 33 P ₁ (51) (first moment of 51) 28 52 53 54 55 56 57 58 26 28 91 (lumped)	74	$P_1(2)$	(first moment of the elastic)
17 5 37 (lumped) 14 16 41 (lumped) 12 17 33 51 (first discrete inelastic) 33 P ₁ (51) (first moment of 51) 28 52 53 54 55 56 57 58 26 28 91 (lumped)	74	$P_{2}(2)$	(second moment of the elastic)
14 16 41 (lumped) 12 17 33 51 (first discrete inelastic) 33 P ₁ (51) (first moment of 51) 28 52 53 54 55 56 57 58 26 28 91 (lumped)	74	$P_{3}(2)$	(third moment of the elastic)
12 17 33 51 (first discrete inelastic) 33 P ₁ (51) (first moment of 51) 28 52 53 54 55 56 57 58 26 28 91 (lumped)	17	5	37 (lumped)
33 51 (first discrete inelastic) 33 P ₁ (51) (first moment of 51) 28 52 53 54 55 56 57 58 26 28 91 (lumped)	14	16	41 (lumped)
33 P ₁ (51) (first moment of 51) 28 52 53 54 55 56 57 58 26 28 91 (lumped)	12	17	
28 52 53 54 55 56 57 58 26 28 91 (lumped)	33	51	(first discrete inelastic)
26 28 91 (lumped)	33	$P_1(51)$	(first moment of 51)
(1 /	28	52	53 54 55 56 57 58
74 102	26	28	91 (lumped)
	74	102	
20 600 601 602 603 604 605 606 607 608	20	600	601 602 603 604 605 606 607 608
609 610 611 612 613 614 649 (lumpe		609	610 611 612 613 614 649 (lumped)
52 800 801 802 803 804 805 806 807 808	52	800	801 802 803 804 805 806 807 808
809 810 811 812 813 814 849 (lumpe		809	$810\ 811\ 812\ 813\ 814\ 849\ (lumped)$

TABLE IX: Cross-section definition in the parameter covariance vector of $^{182}\mathrm{W}$ in terms of the ENDF-6 reaction designation MT numbers. The Npnt column designates the number of energy points.

Npnt		Contributing MTs
74	2	
74	$P_1(2)$	(first moment of the elastic)
74	$P_{2}(2)$	(second moment of the elastic)
74	$P_{3}(2)$	(third moment of the elastic)
16	5	37 (lumped)
14	16	41 (lumped)
12	17	
33	51	(first discrete inelastic)
33	$P_1(51)$	(first moment of 51)
28	52	53 54 55 56 57 58 59 (lumped)
25	28	91 (lumped)
74	102	
19	600	$601\ 602\ 603\ 604\ 605\ 606\ 607\ 608$
	609	$610\ 611\ 612\ 613\ 614\ 615\ 616\ 617$
	618	$619\ 620\ 621\ 622\ 623\ 624\ 625\ 626$
	627	628 629 649 (lumped)
44	800	801 802 803 804 805
		806 849 (lumped)

B. Selection of experimental data

1. W-180

No experimental data on $^{180}\mathrm{W}$ were available that could be used to improve the evaluation. The final file is identical to the prior.

The resonance parameters were formatted from the nu-

TABLE X: Cross-section definition in the parameter covariance vector of 183 W in terms of the ENDF-6 reaction designation MT numbers. The Npnt column designates the number of energy points.

Npnt		Contributing MTs
74	2	
74	$P_1(2)$	(first moment of the elastic)
74	$P_{2}(2)$	(second moment of the elastic)
74	$P_{3}(2)$	(third moment of the elastic)
17	5	37 (lumped)
16	16	41 (lumped)
12	17	
35	51	(first discrete inelastic)
35	$P_1(51)$	(first moment of 51)
34	52	53 54 55 56 57 58 59 60 (lumped)
	61	
27	28	91 (lumped)
74	102	
20	600	601 602 603 604 605 606 607 608
	609	610 611 612 613 614 649 (lumped)
44	800	801 802 803 804 805 806 807 808
	809	810 811 812 813 814 815 816 817
	818	819 820 821 822 823 824 849 (lumped)

TABLE XI: Cross-section definition in the parameter covariance vector of $^{184}\mathrm{W}$ in terms of the ENDF-6 reaction designation MT numbers. The Npnt column designates the number of energy points.

Npnt		Contributing MTs
74	2	
74	$P_1(2)$	(first moment of the elastic)
74	$P_{2}(2)$	(second moment of the elastic)
74	$P_{3}(2)$	(third moment of the elastic)
16	5	37 (lumped)
15	16	41 (lumped)
12	17	
33	51	(first discrete inelastic
33	$P_1(51)$	(first moment of 51)
28	52	53 54 55 56 (lumped)
25	28	91 (lumped)
74	102	
17	600	649 (lumped)
29	800	801 802 803 804 805 806 807 849 (lumped)

merical information in the Atlas of Resonance Parameters by S.F. Mughabghab [16]. No resonance covariances are given.

2. W-182

The covariance matrix prior and the experimental data listed in Tables XIII-XIX were processed by the GANDR system. In the tables, the list includes all data found in the EXFOR database. Only the data sets with a "plus"

TABLE XII: Cross-section definition in the parameter covariance vector of $^{186}\mathrm{W}$ in terms of the ENDF-6 reaction designation MT numbers. The Npnt column designates the number of energy points.

Npnt		Contributing MTs
74	2	
74	$P_1(2)$	(first moment of the elastic)
74	$P_{2}(2)$	(second moment of the elastic)
74	$P_{3}(2)$	(third moment of the elastic)
20	5	37 (lumped)
15	16	41 (lumped)
12	17	
33	51	(first discrete inelastic)
33	$P_1(51)$	(first moment of 51)
27	52	53 54 55 56 57 58 59 60
	61	(lumped)
25	28	91 (lumped)
74	102	
17	600	649 (lumped)
22	800	849 (lumped)

sign in column-1 were included; those beginning with a "dash" were rejected.

Fission cross-section measurements exist above 60 MeV, but the cross section is small. The nuclide was treated as non-fissile; the measured fission cross section was fitted by a polynomial and added as an activation reaction in File 10 with no assigned uncertainties.

Adjustment factors were applied to the cross sections and the data were formatted to create the final evaluated nuclear data file for $^{182}{\rm W}.$

Resonance parameters were taken from JENDL-3.3 with small adjustments. The resonance covariances were prepared by the retro-active method [18] with the SAMMY code as follow:

- Retrieve best set of available resonance parameters.
- Convert the resonance parameters into the Reich-Moore form.
- Generate "experimental" cross sections for total, scattering, and capture cross sections with the NJOY code [49], for use as "experimental data" in SAMMY.
- Assign a global "experimental" uncertainty to the "experimental data." The uncertainties used were based on uncertainty in the thermal capture cross section and uncertainty in the capture resonance integrals.
- Run SAMMY code with the option to generate resonance-covariances retroactively.
- Convert the resonance-covariance results from SAMMY into the ENDF format for File 32 and MT=151.

TABLE XIII: Selection of total cross-section measurements of $^{182}\mathrm{W}$ from the EXFOR database.

	EXFOR No.	Author	Year	Comment			
+	13887002	F.S.Dietrich et al.	(03)	(42 points)			
-	22045008	K.Knopf et al.	(87)	$E=145~\mathrm{keV}$ only			
+	11540009	J.F.Whalen et al.	(82)	(272 points)			
				Added 0.1 b			
				statist. uncert.			
-	10803017	P.T.Guenther $et\ al.$	(82)				
-	40585034	A.N.Djumin et al.	(77)				
+	10047105	D.G.Foster Jr et al.	(71)	(252 points)			
-	12159003	R.C.Martin et al.	(67)	Old, scattered			
-	12168003	W.Selove	(51)	Resonance reg.			
Sys	Systematic uncertainties:						
3%	3% for each EXFOR set						
1%	common to a	all measurements					

TABLE XIV: Selection of radiative capture cross-section measurements of $^{182}{
m W}$ from the EXFOR database.

Author	Year	EXFOR No.	Comment
- J.Voignier et al.	(92)	22006018	repeated
- K.Knopf et al.	(87)	22045013	Thermal
+ M.V.Bokhovko et al.	(86)	40925003	(32 points)
			Remove
			resonance data
			and data above
			inel. threshold
+ R.L.Macklin et al.	(83)	10999006	(32 points)
			Remove
			resonance data
- S.Joly et al.	(81)	21619012	
- Z.M.Bartolome et al.	(69)	10002035	Old, no uncert.
- S.J.Friesenhahn	(66)	12167011	Low energy
- S.J.Friesenhahn	(66)	12167003	Thermal
+ R.C.Block et al.	(66)	11458006	(4 points)
			Remove
			resonance data
- S.V.Kapchigashev	(66)	40755002	Old, uncertain
- V.N.Kononov et al.	(66)	40076005	No uncert.
- H.Pomerance		11507091	Thermal
+ Artificial point at 1 M			
for normalisation, bas	sed on	integral bene	chmark
considerations			
Systematic uncertainties:			
3% for each EXFOR set			

The resulting resonance parameters and covariances were used in the calculations of average cross sections and uncertainties. The calculations were done with SAMMY [39], NJOY [49], and PUFF-IV [51] codes. The results of these three codes agreed very well.

3% common to all measurements

TABLE XV: Selection of elastic cross-section measurements of $^{182}{
m W}$ from the EXFOR database.

Author	Year	EXFOR No.	Comment
+ P.T.Guenther et al.	(82)	10803008	
- D.Lister et al.	(67)	12174003	No uncert.
Systematic uncertainties:			
10% for each EXFOR set	t		

TABLE XVI: Selection of first discrete level inelastic cross-section measurements of $^{182}{
m W}$ from the EXFOR database.

	Author	Year	EXFOR No. Comment				
+	P.T.Guenther et al.	(82)	10803014				
+	D.Lister et al.	(67)	12174004				
Syst	Systematic uncertainties:						
10%	for each EXFOR set	t					
10%	common to all meas	surem	ents (altogether 14.1 %)				

TABLE XVII: Selection of (n,2n+x) cross-section measurements of $^{182}{\rm W}$ from the EXFOR database.

Author	Year	EXFOR No.	Comment			
+ Kong Xiangzhong	(97)	31484005				
+ J.Frehaut et al.	(80)	20416050	Corrected [52]; added uncert. $0.12\%^{252} Cf \overline{\nu} + 0.9\%^{238} U(n,f)$			
+ S.M.Qaim et al.	(75)	20668005				
+ G.N.Maslov et al.	(72)	40136019				
+ W.Dilg et al.	(68)	20802015				
- A.A.Druzhinin et al.	(66)	40174002	Outlier, no monitor x.s.			
- M.Lindner et al.	(59)	12185004	Outlier, old			
Systematic uncertainties:						
1% for each EXFOR set						
4% common to all measu	ireme	nts (altogethe	r 4.1 %)			

TABLE XVIII: Selection of (n,p) cross-section measurements of $^{182}{\rm W}$ from the EXFOR database.

	Author	Year	EXFOR No.	Comment
+	V.Semkova et al.	(07)	22877	(4 points)
-	V.Avrigeanu et al.	(06)	41469003	=41428045
+	A.A.Filatenkov	(03)	41428044	(1 point)
+	A.A.Filatenkov	(03)	41428045	(6 points)
+	Kong Xiangzhong	(97)	31484002	(5 points)
+	A.Grallert et al.	(93)	31496056	
	M.Lindner et al.	(59)	12185003	(1 point)

Systematic uncertainties:

2% for each EXFOR set

10% common to all measurements

TABLE XIX: Selection of (n,α) cross-section measurements from the EXFOR database.

Only data for isomer production are available and are not included in the fitting procedure.

TABLE XX: Selection of total cross-section measurements of $^{183}\mathrm{W}$ from the EXFOR database.

Author	Year	EXFOR No.	Comment
+ K.Knopf et al.	(87)	22045009	E=143 keV
			point only
- M.Ohkubo et al.	(78)	20825002	Resonance
+ M.Ohkubo[a] et al.	(78)	20825	Use broad
			bin averages at higher energies
- W.Selove	(51)	12168004	Resonance
Systematic uncertainties	s:		
1% for each EXFOR se	t		
1% common to all measure	surem	ents (altogeth	ner 2.8 %)

W-183

The covariance matrix prior and the experimental data listed in Tables XX-XXVI were processed by the GANDR system. As before, the data sets with a dash in column-1 were rejected.

Fission cross-section measurements exist above 60 MeV, but the cross section is small. The nuclide was treated as non-fissile; the measured fission cross section was fitted by a polynomial and added as an activation reaction in File 10 with no assigned uncertainties.

Adjustment factors were applied to the cross sections and the data were formatted to create the final evaluated nuclear data file for ¹⁸³W.

Resonance parameters were taken from JENDL-3.3 with small adjustments. The resonance covariances were prepared by the retro-active method [18] in the same way as for 182 W.

TABLE XXI: Selection of (n,2n+x) cross-section measurements of $^{183}\mathrm{W}$ from the EXFOR database.

Author	Year	EXFOR No.	Comment
+ J.Frehaut et al.	(80)	20416051	Corrected [52]; added uncert. $0.12\%^{252}$ Cf $\overline{\nu}$ $+ 0.9\%^{238}$ U(n,f)
Systematic uncertain	nties:		
1% for each EXFOR	l set		

TABLE XXII: Selection of fission cross-section measurements of $^{183}\mathrm{W}$ from the EXFOR database.

Author	Year	EXFOR No.	Comment
- V.P.Eismont et al.	(06)	22937	Not fitted.

TABLE XXIII: Selection of (n,n+p) cross-section measurements of ¹⁸³W from the EXFOR database.

Author	Year	EXFOR No.	Comment
- V.Semkova et al.	(07)	22877	Not fitted

TABLE XXIV: Selection of the first discrete level inelastic cross-section measurements of $^{183}\mathrm{W}$ from the EXFOR database.

Author	Year	EXFOR No.	Comment
- B.Anders,et.al.	(82)	21818007	Metastable isomer product. (not fitted)

TABLE XXV: Selection of radiative capture cross-section measurements of ¹⁸³W from the EXFOR database.

	Author	Year	EXFOR No.	Comment
-	J.Voignier et al.	(92)	22006	Equal to 10
-	K.Knopf et al.	(87)	22045	Thermal
+	M.V.Bokhovko et al.	(86)	40925004	Removed data
				above inelastic
				threshold
-	R.L.Macklin et al.	(83)	10999	(34 points)
-	S.Joly et al.	(81)	21619	(3 points)
-	Grenier et al.	(81)	-	(3 points)
-	S.J.Friesenhahn	(66)	12167	Low energy
				(32 points)
-	S.J.Friesenhahn	(66)	12167	Thermal
+	R.C.Block et al.	(66)	11458009	Removed
				resonance data
-	S.V.Kapchigashev	(66)	40755	Low energies
				with no uncert.
				removed
+	V.N.Kononov et al.	(66)	40076	(5 points)
-	Z.M.Bartolome et al.	(69)	10002	raw data
-	H.Pomerance	(52)	11507	Thermal
Syst	tematic uncertainties:			

Systematic uncertainties:

1 % for each EXFOR set

3 % common to all measurements

TABLE XXVI: Selection of (n,α) cross-section measurements of $^{183}\mathrm{W}$ from the EXFOR database.

Only data for isomer production given; no data fitted.

TABLE XXVII: Selection of (n,p) cross-section measurements of $^{183}\mathrm{W}$ from the EXFOR database.

	Author	Year	EXFOR No.	Comment
+	V.Semkova et al.	(07)	22877	
-	V.Avrigeanu et al.	(06)	41469006	=41428009
+	A.A.Filatenkov	(03)	41428009	
+	A.Grallert et al.	(93)	31496057	
+	S.M.Qaim et al.	(75)	20668009	
+	M.Lindner et al.	(59)	12185005	

Added covariance of each measurement equals minimum variance of the same.

10 % common to all measurements

TABLE XXVIII: Selection of total cross-section measurements of $^{184}\mathrm{W}$ from the EXFOR database.

	Author	Year	EXFOR No.	Comment
+	F.S.Dietrich et al.	(03)	13887003	
+	Bessheyoko et al.	(92)	-	
+	K.Knopf et al.	(87)	22045010	Only $143~\mathrm{keV}$
-	V.A.Anufriev et al.	(83)	40850004	Resonance
-	J.A.Harvey et al.	(82)	12897005	Resonance
+	J.F.Whalen et al.	(82)	11540019	Resonance
+	P.T.Guenther et al.	(82)	10803018	
+	A.N.Djumin et al.	(77)	40585035	
+	D.G.Foster Jr et al.	(71)	10047106	
-	R.C.Martin et al.	(67)	12159004	Old

Systematic uncertainties:

3% for each EXFOR set

1% common to all measurements (altogether 2.8 %)

4. W-184

The covariance matrix prior and the experimental data listed in Tables XXVIII-XXXIX were processed by the GANDR system. As before, the data sets with a dash in column-1 were rejected.

Fission cross section measurements exist above 60 MeV, but the cross section is small. The nuclide was treated as non-fissile; the measured fission cross section was fitted by a polynomial and added as an activation reaction in File 10 with no assigned uncertainties.

Adjustment factors were applied to the cross sections and the data were formatted to create the final evaluated nuclear data file for 184 W.

Resonance parameters were taken from JENDL-3.3 with small adjustments. The resonance covariances were prepared by the retro-active method [18] in the same way as for 182 W. In the unresolved resonance range a 5 % systematic uncertainty was added to capture.

TABLE XXIX: Selection of elastic cross-section measurements of $^{184}\mathrm{W}$ from the EXFOR database.

	Author	Year	EXFOR No.	Comment
-	V.A.Anufriev et al.	(83)	40850007	Thermal
				(1point)
+	P.T.Guenther et al.	(82)	10803009	(13 points)
-	D.Lister et al.	(67)	12174006	No uncert.
Syst	ematic uncertainties:			
10%	for each EXFOR set	t		

TABLE XXX: Selection of (n,2n+x) cross-section measurements of $^{184}\mathrm{W}$ from the EXFOR database.

Author	Year E	EXFOR No.	Comment
+ Frehaut	et al. (80) 2	20416052	Corrected [52];
			added uncert.
			$0.12\%^{-252} {\rm Cf} \overline{\nu}$
			+ 0.9% ²³⁸ U(n,f)
Systematic u	ncertainties:		
1% for each l	EXFOR set		

5. W-186

The covariance matrix prior and the experimental data listed in Tables XL-XLVI were processed by the GANDR system. As before, the data sets with a dash in column-1 were rejected.

Fission cross section measurements exist above 60 MeV, but the cross section is small. The nuclide was treated as non-fissile; the measured fission cross section was fitted by a polynomial and added as an activation reaction in File 10 with no assigned uncertainties.

Adjustment factors were applied to the cross sections and the data were formatted to create the final evaluated nuclear data file for $^{186}{
m W}.$

Resonance parameters were taken from IRDF-2002 with small adjustments. The resonance covariances were prepared by the retro-active method [18] in the same way as for 182 W.

TABLE XXXI: Selection of 184 W(n,2n) 183m W cross-section measurements from the EXFOR database.

Author	Year	EXFOR No.	Comment
- B.Anders et al.	(82)	21818008	Cannot be
- R.Prasad et al.	(66)	30015008	processed Cannot be processed

TABLE XXXII: Selection of (n,3n+x) cross-section measurements of ^{184}W from the EXFOR database (not fitted).

Author		EXFOR No.	Comment
- J.Frehaut et al.	(80)	20416060	Corrected [52]; added uncert. $0.12\%^{252} Cf \overline{\nu} + 0.9\%^{238} U(n,f)$

TABLE XXXIII: Selection of fission cross-section measurements of $^{184}\mathrm{W}$ from the EXFOR database.

Author	Year	EXFOR No.	Comment
- V.P.Eismont et al.	(06)	22937004	Not fitted

TABLE XXXV: Selection of first discrete level inelastic cross-section measurements of $^{184}{
m W}$ from the EXFOR database.

	Author	Year	EXFOR No. Comment		
+	P.T.Guenther et al.	(82)	10803015		
+	D.Lister et al.	(67)	12174007		
Syst	ematic uncertainties:				
10%	for each EXFOR set	5			
10%	10% common to all measurements (altogether 14.1 %)				

TABLE XXXIV: Selection of (n,n+p) cross-section measurements of ^{184}W from the EXFOR database.

Author	Year	EXFOR No. Comment
- V.Semkova et al.	(07)	22877
- H.Sakane et al.	(03)	22827018
- S.M.Qaim et al.	(75)	20668012
- A.K.Ganguly et al.	(63)	31329005

TABLE XXXVI: Selection of radiative capture cross-section measurements of $^{184}{
m W}$ from the EXFOR database.

	Author	Year	EXFOR No.	Comment
-	V.Bondarenko et al.	(05)	31572004	Thermal
-	J. Voignier et al.	(92)	22006020	
-	K.Knopf et al.	(87)	22045015	Thermal
+	M.V.Bokhovko et al.	(86)	40925005	
-	V.A.Anufriev et al.	(83)	40850008	Thermal
+	R.L.Macklin et al.	(83)	10999008	
-	H.Beer et al.	(81)	21768002	
-	H.Beer et al.	(81)	21768003	
-	S.Joly et al.	(81)	21619014	
-	S.J.Friesenhahn	(66)	12167006	Low energy
+	R.C.Block et al.	(66)	11458007	
-	S.V.Kapchigashev	(66)	40755004	
-	S.J.Friesenhahn	(66)	12167007	Thermal
-	Z.M.Bartolome et al.	(69)	10002039	
-	V.N.Kononov et al.	(66)	40076006	No uncert.
-	R.Booth et al.	(58)	11429027	
-	H.Pomerance	(52)	11507093	Thermal
CI /	, , , , , , , , , , , , , , , , , , , ,			

1% for each EXFOR set

10% common to all measurements

TABLE XXXVII: Selection of $^{184} W(n,\gamma)^{185m} W$ cross-section measurements from the EXFOR database.

Author	Year	EXFOR No.	Comment
- S.K.Mangal et al.	(62)	31248015	Cannot be
			processed

TABLE XXXVIII: Selection of (n,p) cross-section measurements of $^{184}\mathrm{W}$ from the EXFOR database.

	Author	Year	EXFOR No.	Comment
+	V.Semkova et al.	(07)	22877	
-	V.Avrigeanu et al.	(06)	41469008	=41428011
+	R.Eichin et al.	(05)		
+	A.A.Filatenkov	(03)	41428011	
+	A.A.Filatenkov	(99)	41298091	
+	A.A.Filatenkov	(99)	41240105	
+	Kong Xiangzhong	(97)	31484003	
+	Y.Kasugai et al.	(94)	22351007	
+	A.Grallert et al.	(93)	31496058	
+	Y.Ikeda et al.	(88)	22089107	
+	S.M.Qaim et al.	(75)	20668012	Corrected
+	S.K.Mukherjee et al.	(63)	31330010	
-	A.Poularikas et al.	(60)	12181002	Old
-	R.F.Coleman et al.	(59)	21440024	Old
Syst	ematic uncertainties:	` /		
٠	for each EXFOR set.			

2% for each EXFOR set

10% common to all measurements

TABLE XXXIX: Selection of (n,α) cross-section measurements of $^{184}\mathrm{W}$ from the EXFOR database.

Author	Year	EXFOR No.	Comment		
+ V.Semkova et al.	(07)	22877			
- V.Avrigeanu et al.	(06)	41469009	=41428010		
+ R.Eichin et al.	(05)				
+ A.A.Filatenkov	(03)	41428010			
+ Y.Kasugai et al.	(94)	23011004	=22351009		
+ A.Grallert et al.	(93)	31496028			
- Y.Kasugai et al.	(92)	22351009			
+ S.M.Qaim et al. [r]	(75)	20668014	Corrected		
+ M.Lindner et al.	(59)	12185006			
Systematic uncertainties:					
2% for each EXFOR set					
3% common to all measures	surem	ents			

TABLE XL: Selection of total cross-section measurements of ¹⁸⁶W from the EXFOR database.

Author	Year	EXFOR No.	Comment
+ F.S.Dietrich et al.	(03)	13887	
+ K.Knopf et al.	(87)	22045	Only 143 keV
+ J.F.Whalen et al.	(82)	11540	Add 0.1 b
			uncertainty
+ P.T.Guenther et al.	(82)	10803	
+ D.G.Foster Jr et al.	(71)	10047	
- A.N.Djumin et al.	(77)	40585	Single point
- R.C.Martin et al.	(67)	12159	Old
- W.Selove	(51)	12168	Old

2% for each EXFOR set

2% common to all measurements

Analysis of obtained covariances

The full evaluated files with covariances were verified by the ENDF Utility Codes [48] to check for the formal correctness. In the next step the covariance plots were generated on a fine energy grid with the NJOY processing system [49] to check processability and the general features of the covariance matrices, including cross-reaction correlations. An example of the cross-correlation matrix between the ¹⁸³W elastic and inelastic cross sections is shown in Figure 6. Swings of the correlation pattern from strong anti-correlations around 1 MeV to strong correlations around 4-5 MeV are observed. Similarly, the crosscorrelation matrix between the ¹⁸³W first-discrete-level inelastic and (n,2n) cross sections is shown in Figure 7. In this particular case the (n,2n) reaction is included in the lumped reaction designated MT=852, but since the threshold of the (n,2n+p) reaction is 14.71 MeV, the matrix corresponds purely to the (n,2n) reaction below this energy. At higher energies some approximation is needed to obtain the uncertainty of the pure (n,2n) cross sec-

TABLE XLI: Selection of (n,2n+x) cross-section measurements of $^{186}\mathrm{W}$ from the EXFOR database.

Author	Year	EXFOR No.	Comment
+ A.A.Filatenkov	(03)	41428012	(5 points)
+ Kong Xiangshong	(97)	31484006	(5 points)
+ J.Frehaut et al.	(80)	20416053	Corrected [52];
			added uncert.
			$0.12\%^{-252}{ m Cf} \overline{\nu}$
			+ 0.9% ²³⁸ U(n,f)
- A.A.Druzhinin et al.	(66)	40174061	Old
- M.Lindner	(59)	12185007	Old

1% systematic for each data set.

4% common to all measurements

TABLE XLII: Selection of (n,3n+x) cross-section measurements of ¹⁸⁶W from the EXFOR database.

Author	Year	EXFOR No.	Comment
+ J.Frehaut et al.	(80)	20416061	Corrected [52];
			added uncert.
			$0.12\%^{-252} Cf \overline{\nu}$
			+ 0.9% ²³⁸ U(n,f)
Systematic uncertai	nties:		
1% for each EXFO	R set		
4% common to all r	neasu	rements	

tion, but in this energy range the (n,2n) cross section is decreasing rapidly to a small value and its uncertainty is usually less important. Note the relatively strong and uniform anti-correlation between the inelastic and the (n,2n) reactions, as expected for the dominant channels in the given energy range, arising from the physics constraint that the total cross section equals to the sum of the partials. An example of the correlation matrix for the capture reaction of ¹⁸⁶W is shown in Figure 8. Additional plots of the covariance matrices for the capture reaction of ¹⁸⁶W in the fast neutron region were generated online at the IAEA website (see Figures 1, 2 and 3 in previous

Broad-group testing of the calculated covariances was

TABLE XLIII: Selection of first discrete level inelastic crosssection measurements of $^{186}\mathrm{W}$ from the EXFOR database.

	Author	Year	EXFOR No.	Comment
+	D.Lister et al.	(67)	12174	(59 points)
+	P.T. Guenther $et\ al.$	(82)	10803	(15 points)

Systematic uncertainties:

10% for each EXFOR set

10% common to all measurements (altogether 2.8~%)

TABLE XLIV: Selection of radiative capture cross-section measurements of ¹⁸⁶W from the EXFOR database.

	Author	Year	EXFOR No.	Comment
+	G.D.Kim et al.	(07)	20673035	(3 points)
-	M.Karadag et al.	(04)	22842002	Thermal
-	Vuong Huu Tan et al.	(04)	31542002	
-	Guohui Zhang et al.	(01)	32550002	
-	J. Voignier et al.	(92)	22006021	
-	K.Knopf et al.	(87)	22045016	Thermal
+	Yu.N.Trofimov	(87)	41001037	(1 point)
+	M.V.Bokhovko et al.	(86)	40925006	(19 points)
				> 50 keV
-	M.V.Bokhovko et al.	(86)	40925006	(38 points)
				> 50 keV
+	R.L.Macklin et al.	(83)	10999009	(34 points)
+	S.Joly et al.	(81)	21619015	(3 points)
-	V.A.Anufriev et al.	(81)	40608003	Thermal
+	G.Magnusson et al.	(80)	21004004	(1 point)
-	T.Bradley et al.	(79)	10938006	
-	G.Gleason	(77)	10662014	Thermal
-	M.Lindner et al.	(76)	10221044	
-	O.Schwerer et al.	(76)	20670016	
+	M.Valkonen	(76)	20673035	(1 point)
-	K.Siddappa et al.	(74)	30502018	
-	C.H.Hogg et al.	(70)	12602007	Thermal
-	M.Diksic et al.	(70)	30023012	
-	Z.M.Bartolome et al.	(69)	10002041	
-	G.G.Zaikin et al.	(68)	40248004	
-	P.P.Damle et al.	(67)	20184002	Thermal
-	V.N.Kononov et al.	(66)	40076007	
-	S.J.Friesenhahn	(66)	12167009	Thermal
-	S.J.Friesenhann	(66)	12167008	Resonance
-	S.V.Kapchigashev	(66)	40755005	Resonance
+	R.C.Block et al.	(66)	11458008	(9 points)
-	A.K.Chaubey et al.	(65)	30079035	
-	J.A.Miskel et al.	(62)	12115005	
-	J.A.Miskel et al.	(62)	12115006	
-	W.S.Lyon	(60)	11625032	Thermal
-	Yu.Ya.Stavisskii	(60)	40452005	
-	A.E.Johnsrud et al.	(59)	11675025	
-	W.S.Lyon et al.	(59)	11407030	
-	R.Booth et al.	(58)	11429028	
-	A.I.Leipunskij	(58)	40244036	
-	V.N.Kononov et al.	(58)	40421023	
-	M.V.Pasechnik et al.	(58)	40336093	
-	A.I.Leipunskij	(58)	40244017	
-	A.I.Leipunskij	(58)	40244058	
-	J.L.Perkin <i>et al.</i>	(58)	21438025	/- · · · ·
-	R.L.Macklin et al.	(57)	11399045	(1 point)
-	H.Pomerance	(52)	11507094	Thermal
-	L.E.Beghian et al.	(49)	22580014	mı ı
- C	L.Seren et al.	(47)	11447115	Thermal
Sys	stematic uncertainties:			

TABLE XLV: Selection of (n,p) cross-section measurements of $^{186}\mathrm{W}$ from the EXFOR database.

	Author	Year	EXFOR No.	Comment		
+	V.Semkova et al.	(07)	22877	(3 points)		
+	A.A.Filatenkov et al.	(03)	41428014	(7 points)		
+	A.A.Filatenkov et al.	(99)	41240106	(8 points)		
+	Y.Satoh et al.		22365012	(5 points)		
+	A.A.Filatenkov et al.	(99)	41298012	(1 point)		
+	S.Murahira et al.	(95)	22311014	(6 points)		
+	Y.Kasugai et al.	(92)	22351008	(5 points)		
+	S.M.Qaim et al.	(75)	20668017	(1 point)		
-	S.K.Mukherjee et al.	(63)	31330012	(1 point)		
-	A.Poularikas et al.	(60)	12181004	(1 point)		
-	R.F.Coleman et al.	(59)	21440025	(1 point)		
Syst	ematic uncertainties:					
	for each EXFOR set					
10%	10% common to all measurements					

TABLE XLVI: Selection of (n,α) cross-section measurements of ¹⁸⁶W from the EXFOR database.

	Author	Vear	EXFOR No.	Comment		
+	V.Semkova et al.	(07)	22877	(6 points)		
+	A.A.Filatenkov et al.	(03)	41428013	(7 points)		
+	Kong Xiangzhong	(97)	31484004	(5 points)		
+	A.Grallert et al.	(93)	31496029	(1 point)		
+	Y.Kasugai et al.	(92)	22351010	(6 points)		
+	Y.Ikeda et al.	(88)	22089108	(4 points)		
+	S.M.Qaim et al.	(75)	20668019	(1 point)		
-	S.K.Mukherjee et al.	(63)	31330011	(1 point)		
-	A.Poularikas et al.	(60)	12181003	(1 point)		
Sys	tematic uncertainties:					
1%	1% for each EXFOR set					
5%	common to all measu	remer	nts			

performed by calculating the thermal cross sections and the resonance integrals with their respective uncertainties. The values are given in Table XLVII. The capture reaction of ¹⁸⁶W is the only one that can be measured by the activation technique because the intensities of the gamma rays from the activation of ¹⁸⁴W are low. The thermal cross section and the resonance integral are about 10 % higher than the values derived from the activation analysis constants recommended by IUPAC [50], but the absolute γ -emission probabilities are difficult to measure. The recommended value of γ -emission probabilities in ENSDF database has changed in recent years by as much as 20 %.

^{3%} for each EXFOR set

^{1%} common to all measurements

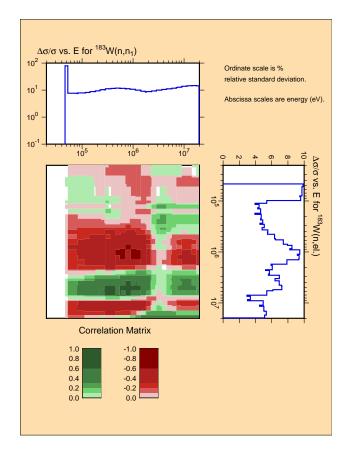


FIG. 6: Uncertainties and cross-correlation matrix of the $^{183}{\rm W}$ elastic and first-discrete-level inelastic cross sections.

V. EVALUATION OF MANGANESE

A. EMPIRE model calculations

As in the case of the ²³²Th evaluation, the details of the selection of optical model and other parameters are described in a separate paper [21]. The list of parameters and their uncertainties is similar to those parameters used in thorium (fission parameters are obviously not required for manganese).

The definitions of cross sections in the cross-section parameter vector for manganese are given in Table XLVIII. Note that the energy grid is fixed and is the same for all reactions, but some reactions require fewer points due to their respective thresholds.

B. Selection of experimental data

In the fast energy range the experimental data for radiative capture show significant discrepancies, comparable to the uncertainty of the prior. Preliminary analysis of integral experiments is also to some extent contradictory. For other reaction channels - except perhaps the (n,2n) reaction - the experimental data would not have a significant impact on the covariances, therefore the fi-

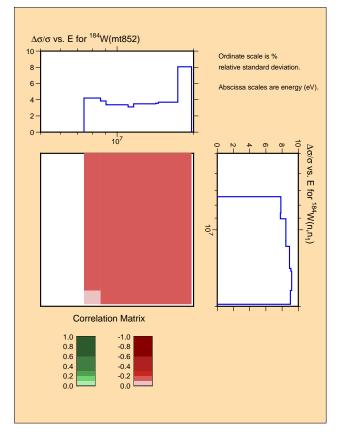


FIG. 7: Uncertainties and cross-correlation matrix of the ¹⁸⁴W first-discrete-level inelastic and (n,2n) cross sections.

nal covariance matrix in the fast energy range is equal to the prior. Note that a new comprehensive evaluation of the ⁵⁵Mn(n,2n) reaction for reactor dosimetry has been undertaken by Zolotarev, including a careful assessment of available experimental data [53]. The new dosimetry evaluation has smaller uncertainties, but the cross sections are in good agreement at least in the plateau region. Slight differences are observed near threshold, but they are within the assigned uncertainties.

The resonance parameters were obtained from a sequential Bayes SAMMY analysis of the most recent experimental neutron transmissions and capture cross sections in the resolved energy range (1.0-5 eV to 125 keV): Harvey et al. transmissions measured at ORELA in 1980 [54], Aerts et al. capture cross sections measured at GELINA in 2006 [55], and Guber et al. capture cross section measured at ORELA in 2007 [56]. For the evaluation in the thermal energy range, Widder et al. (1975) capture cross sections [57], Coté et al. (1964) total cross sections [58], and Rainwater et al. (1947) total cross sections [59] were considered. The thermal capture cross sections were normalized at the 0.0253 eV value of 13.27 b which is an average of the available experimental values.

The uncertainties and covariances of the resonance parameters in File 32 were generated only from the statistical uncertainties of the experimental database.

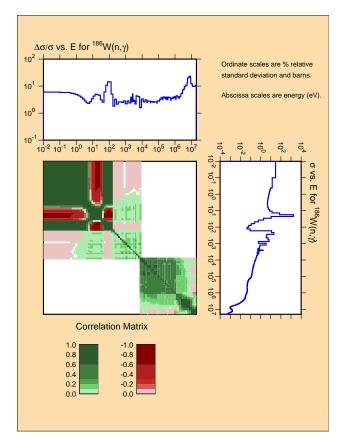


FIG. 8: Uncertainty and correlation matrix of the $^{186} W(n,\gamma)$ reaction.

The systematic uncertainties were estimated from results obtained in the SAMMY analysis of the experimental database, from some information found in the publications, and from discussion with some authors. These uncertainties were included as additional uncertainty in File 33 for the total, elastic and radiative capture cross-section covariances, and are to be combined with the cross-section uncertainties obtained from the resonance parameters stored in File 32 to obtain the corresponding total uncertainties.

C. Analysis of obtained covariances

The full evaluated files with covariances were verified with the ENDF Utility Codes [48] to check for the formal correctness. In the next step the covariance plots were generated on a fine energy grid with the NJOY processing system [49] to check processability and the general features of the covariance matrices, including cross-reaction correlations. No problems were identified. An example of the covariance matrix for the capture reaction up to 60 MeV is shown in Figure 9. The uncertainty (variance) is higher near 10 MeV, where the cross section has a dip (and experimental data are not available), and at higher energies where the cross section becomes very small.

TABLE XLVII: Thermal cross sections and resonance integrals with uncertainties of the tungsten isotopes.

Reaction	σ_0 [b]	RI [b]
180 W(n,tot)	$39.61 \pm 26 \%$	$537.9 \pm 7.2 \%$
180 W(n,el)	$9.96\pm36~\%$	$284.6\pm13.4~\%$
$^{180}\mathrm{W}(\mathrm{n},\gamma)$	$29.65\pm33~\%$	$249.6~\pm~4.8~\%$
$^{182}\mathrm{W}(\mathrm{n,tot})$	$29.55\pm3.0~\%$	$1124.~\pm~2.9~\%$
¹⁸² W(n,el)	$8.84 \pm 6.3 \%$	$492. \pm 5.5 \%$
$^{182}\mathrm{W}(\mathrm{n},\gamma)$	$20.71\pm3.5~\%$	$628. \pm 2.7 \%$
¹⁸³ W(n,tot)	$12.49\pm4.6~\%$	$822.8~\pm~5.0~\%$
$^{183}W(n,el)$	$2.38 \pm 8.1 \%$	$478.5 \pm 7.8 \%$
$^{183}\mathrm{W}(\mathrm{n},\gamma)$	$10.11\pm5.5~\%$	$338.9~\pm~5.0~\%$
184 W(n,tot)	$8.854\pm2.9~\%$	$331.1 \pm \ 2.0 \ \%$
184 W(n,el)	$7.353 \pm 3.2 \%$	$311.2 \pm 2.1 \%$
184 W(n, γ)	$1.501\pm9.4~\%$	$15.6~\pm~5.6~\%$
¹⁸⁶ W(n,tot)	$38.17\pm5.9~\%$	$3645.~\pm~2.2~\%$
¹⁸⁶ W(n.el)	.0852 \pm 100 $\%$	$3158. \pm 2.8 \%$
186 W(n, γ)	$38.08 \pm 5.9 \%$	$483. \pm 4.4 \%$

The thermal capture cross section and the resonance integral are shown in Table XLIX. They are in excellent agreement with the values derived from the activation analysis constants recommended by IUPAC [50]. The uncertainty in the thermal cross section derived from the neutron activation constants seems low. We consider the assigned uncertainty in the evaluated cross sections file to be reasonable.

VI. SUMMARY AND CONCLUSIONS

A self-consistent method of generating covariance information that combines the theoretical and experimental uncertainties was developed and first applied in the evaluation of ²³²Th within an IAEA coordinated research project. It includes medium and long-range correlations (in terms of energy) and correlations between different reaction channels that originate from physics constraints in the nuclear reaction models. These correlations are introduced through the covariance matrix prior calculated by a Monte Carlo method based on the EMPIRE code that implements the nuclear reaction modelling. In this way the physical consistency of the cross sections is preserved even after a subsequent least-squares adjustment. Small adjustments to the cross sections are calculated by the introduction of experimental data using the generalised least-squares method as implemented in the GANDR code. The main function of the fitting is to produce the evaluated mean values consistent with produced covariances, and to constrain the covariances based on accurate experimental data. Using statistically consistent (non-discrepant) experimental data is of primary importance. Failure to correctly account for correlations between experiments or to underestimate the systematic

TABLE XLVIII: Cross-section definition in the parameter covariance vector of $^{55}\mathrm{Mn}$ in terms of the ENDF-6 reaction designation MT numbers. The Npnt column designates the number of energy points.

Npnt		Contributing MTs		
74	2	(elastic)		
74	$P_1(2)$	(first moment of the elastic)		
74		(second moment of the elastic)		
74	$P_3(2)$	(third moment of the elastic)		
17	5	37 (lumped)		
13	16			
11	17			
32	51	(first discrete inelastic)		
32	$P_1(51)$	(first moment of 51)		
23	52	53 54 55 56 57 58 59 60		
20	61	62 63 64 65 66 67 68 69		
	70	(lumped)		
19	71	72 73 74 75 76 77 78 79		
	80	81 82 83 84 85 86 87 88		
		91 (lumped)		
74	102			
19	600	601 602 603 604 605 606 607 608		
	609	610 611 612 613 614 615 616 617		
	618	619 620 621 622 623 624 625 626		
	627	628 629 630 631 632 649 (lumped)		
19	800	801 802 803 804 805 806 807 808		
	809	810 811 812 813 814 815 816 817		
	818	819 820 821 822 823 824 825 826		
	827	828 829 830 831 (lumped)		

TABLE XLIX: Comparison of thermal capture cross section and resonance integral uncertainties of ⁵⁵Mn with values derived from constants for neutron activation analysis [50].

	IUPAC	Nudat	ENDF/	B-VII.1	
		${\bf Uncert.}$		${\bf Uncert.}$	Difference
	$[\mathrm{barns}]$	[%]	$[{\rm barns}]$	[%]	[%]
σ_0	13.19	0.6	13.27	1.1	0.7
RI	13.9	3.1	13.26	3.8	-0.5

uncertainties may lead to unrealistically small uncertainties in resulting calculations.

An extra bonus of the developed methodology is that new (independent) experimental data (both differential and integral) can be introduced to further improve the cross sections and constrain the uncertainties in the future, as customary in Bayesian approaches. Improvements in one reaction channel will be correctly propagated to all other channels through correlations, including the total cross section. The price to pay is that all cross section covariances are given on the same energy grid and that in some cases several reactions are lumped together. This may be inconvenient for users, who are specifically interested in a particular reaction channel,

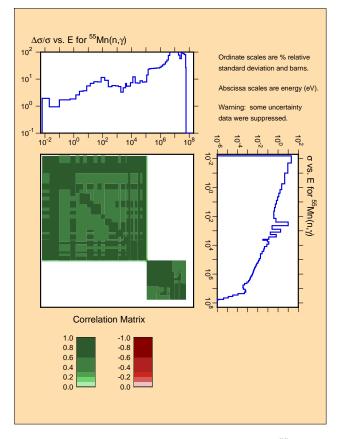


FIG. 9: Uncertainty and correlation matrix of the $^{55}{\rm Mn}({\rm n},\gamma)$ reaction.

but is well suited for applications in deterministic neutron transport.

Developed methodology has been applied to obtain high-quality nuclear data evaluations that include covariances for $^{232}{\rm Th},\,^{180,182,183,184,186}{\rm W}$ and $^{55}{\rm Mn}$ nuclei. These evaluated nuclear data files will be included into the ENDF/B-VII.1 library release.

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