

Serial EnKF

(51)

Consider implementing a sequential algorithm for EnKF to avoid large matrix inversion in $K = P^b H^T (H P^b H^T + R)^{-1}$.

In EnKF, we use $P^b = \bar{X}^b \bar{X}^{bT}$

Note that $H P^b H^T = (H \bar{X}^b)(H \bar{X}^b)^T$, and $P^b H^T = \bar{X}^b (H \bar{X}^b)^T$

It is $H \bar{X}^b$ that is actually needed in calculation of K .

$$H \bar{X}^b = \begin{pmatrix} -h_1- \\ -h_2- \\ \vdots \\ -h_p- \end{pmatrix} \begin{pmatrix} | & | & \dots & | \\ x'_1 & x'_2 & \dots & x'_N \\ | & | & & | \end{pmatrix}^b$$

$h_j x'_k$ is the perturbation (k^{th} member - mean) of observed value associated with j^{th} observation

Recall that Linearized observation operator h_j is from assumption.

$$h_j \Delta x = \left. \frac{\partial h_j}{\partial x} \right|_x (x + \Delta x - x) \approx h_j(x + \Delta x) - h_j(x)$$

Now we can use the original difference to replace $H \bar{X}^b$ so that Linearized observation operators are no longer needed.

$$(H \bar{X}^b)_{jk} = \underbrace{h_j x'_k}_{\text{linearized observation operator}} = \underbrace{\left. \frac{\partial h_j}{\partial x} \right|_{\bar{x}^b}}_{\text{nonlinear observation function}} (x_k^b - \bar{x}^b) \approx h_j(x_k^b) - h_j(\bar{x}^b) \quad (1)$$

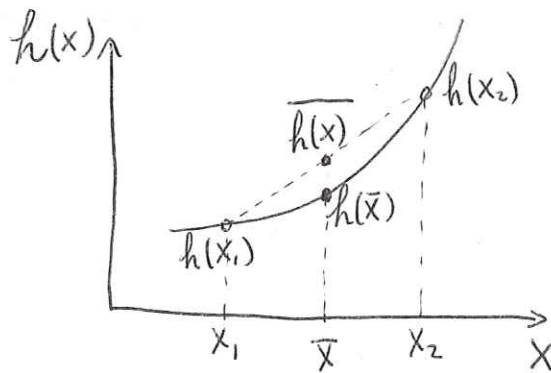
Alternatively, one can use

$$\overline{h_j(x^b)} \equiv \frac{1}{N} \sum_{k=1}^N h_j(x_k^b)$$

to replace $h_j(\bar{x}^b)$.

$$h_j(x_k^b) = h_j(x_k^b) - \overline{h_j(x^b)} \quad (2)$$

If $h()$ is a nonlinear function, $h_j(\bar{x}^b) \neq \overline{h_j(x^b)}$:



$$\bar{x} = \frac{1}{2}(x_1 + x_2).$$

$$\overline{h(x)} = \frac{1}{2}[h(x_1) + h(x_2)]$$

→ In practice, use (1) or (2)?

\bar{x}^b ensemble mean is not necessarily a natural solution of the dynamic model, therefore sometimes $\overline{h_j(x^b)}$ is a better solution than $h_j(\bar{x}^b)$

- for example, let x_k^b be ensemble members simulating a convective cloud, the location of these clouds are different in each member. \bar{x}^b will have a smoothed cloud field with wide-spread thin cloud instead of convective cloud. Let $h(x)$ be the radiative transfer model (RTM) that gives radiance from cloud top, which is strongly nonlinear (takes only a very thin cloud to reduce the brightness temperature a lot). → $h(\bar{x}^b)$ will be cold biased.

If R is diagonal (observation errors are uncorrelated) (53)
 a sequential algorithm can be used to update \bar{x} and x'
 \Rightarrow Serial EnKF.

Initial input: $x_{(1),k} = x_k^b$ for $k=1, 2, \dots, N$

Assimilate observations one at a time:

for $j=1, 2, \dots, p$

$$\bar{x}_{(j)} = \frac{1}{N} \sum_{k=1}^N x_{(j),k} \quad \text{prior mean}$$

for $k=1, 2, \dots, N$

$$x'_{(j),k} = x_{(j),k} - \bar{x}_{(j)} \quad \text{prior perturbations}$$

$$y_{(j),k} = h_j(x_{(j),k}) \quad \text{observation priors}$$

$$\bar{y}_{(j)} = \frac{1}{N} \sum_{k=1}^N y_{(j),k} \quad \text{observation prior mean}$$

$$\text{for } k=1, 2, \dots, N : y'_{(j),k} = y_{(j),k} - \bar{y}_{(j)} \quad \dots \text{perturbations}$$

$$\text{var}(y_j^o) = R_{jj} \quad \text{observation error variance} \quad \text{scalars}$$

$$\text{var}(y_{(j)}) = \frac{1}{N-1} \sum_{k=1}^N (y'_{(j),k})^2 \quad \text{observation prior variance}$$

$$\text{cov}(x_{(j)}, y_{(j)}) = \frac{1}{N-1} \sum_{k=1}^N \underbrace{(x'_{(j),k})}_{\text{vector}} \underbrace{y'_{(j),k}}_{\text{scalar}}$$

$$K_j = \frac{\text{cov}(x_{(j)}, y_{(j)})}{\text{var}(y_j^o) + \text{var}(y_{(j)})}$$

for $k=1, 2, \dots, N$

$$x_{(j+1),k} = x_{(j),k} + K_j \underbrace{(y_{j,k}^o - y_{(j),k})}_{\text{perturbed observation}}$$

$$\sim N(h_j(x^{\text{tr}}), R_{jj})$$

output: $x_k^a = x_{(p+1),k}$

Note on generating random perturbations that fit a Gaussian distribution:

$$\vec{\epsilon} \sim N(0, R)$$

1. Draw random ensemble $\vec{\eta}_k$, $k=1, 2, \dots, N$

2. Recenter ensemble mean (zero in this case)

$$\vec{\delta}_k = \vec{\eta}_k - \frac{1}{N} \sum_{k=1}^N \vec{\eta}_k \quad \text{for } k=1, 2, \dots, N$$

3. Rescale ensemble covariance:

$$R' = \frac{1}{N-1} \sum_{k=1}^N \vec{\delta}_k \vec{\delta}_k^T$$

Take eigenvalue decomposition of $R' = U S U^T$

$$\vec{\epsilon}_k = R^{\frac{1}{2}} U S^{-\frac{1}{2}} U^T \vec{\delta}_k$$

Note on update Equation $\delta X = K(y^o - h(x^b))$

Anderson (2003) derived the impact of a single observation y on a single state variable $x \rightarrow$ this is sufficient to describe all commonly used ensemble filter algorithms, without loss of generality:

The update happens in two steps:

(1) Analysis in observation space: $\delta y = \frac{\text{var}(y^b)}{\text{var}(y^b) + \text{var}(y^o)} (y^o - y^b)$

(2) regression of increment to state space: $\delta x = \frac{\text{cov}(x^b, y^b)}{\text{var}(y^b)} \delta y$