

# Bayesian Variable Selection

Practical Workshop

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# Basics

## Notation:

- ▶  $\mathbf{t} \in \mathbb{R}^n$ : outcome variable for  $n$  observations
- ▶  $\Phi \in \mathbb{R}^{n \times p}$ : design matrix with  $p$  input variables
- ▶  $M_j$ : model  $j \in \{1, \dots, 2^p\}$  out of  $2^p$  possible models. Each model has associated one (and only one) binary vector of predictor inclusion  $\gamma_j \in \{0, 1\}^p$ . All models lie in the model space  $\mathcal{M}_p$ , composed of linear models with at most  $p$  active predictors
- ▶  $\Phi_j$ : design matrix with only the active predictors under  $M_j$
- ▶  $\theta$ : set of parameters of the model, for standard linear models this is equivalent to regression coefficients and residual variance, i.e.  
 $\theta \equiv \{\mathbf{w}, q\} \in \mathbb{R}^{p+1}$

# Bayesian model selection

For any given model  $M_k \in \mathcal{M}_p$ , BMS relies on computing:

$$\begin{aligned}\mathbb{P}(M_k \mid \mathbf{t}) &= \frac{\mathbb{P}(\mathbf{t} \mid M_k)\mathbb{P}(M_k)}{\mathbb{P}(\mathbf{t})} \quad (\text{Bayes' Theorem}) \\ &= \frac{\mathbb{P}(\mathbf{t} \mid M_k)\mathbb{P}(M_k)}{\sum_{j=1}^{2^p} \mathbb{P}(\mathbf{t} \mid M_j)\mathbb{P}(M_j)} \\ &= \frac{\int \mathbb{P}(\mathbf{t} \mid \boldsymbol{\theta}, M_k)\mathbb{P}(\boldsymbol{\theta} \mid M_k)d\boldsymbol{\theta}\mathbb{P}(M_k)}{\sum_{j=1}^{2^p} \mathbb{P}(\mathbf{t} \mid M_j)\mathbb{P}(M_j)}\end{aligned}$$

This calculation poses (at least) **four** delicate issues to solve.

# Computing posterior probabilities

Each part of the expression needs to be carefully specified:

$$\mathbb{P}(M_k \mid \mathbf{t}) = \frac{\overbrace{\int \mathbb{P}(\mathbf{t} \mid \boldsymbol{\theta}, M_k) \overbrace{\mathbb{P}(\boldsymbol{\theta} \mid M_k) \mathrm{d}\boldsymbol{\theta}}^1 \overbrace{\mathbb{P}(M_k)}^2}^3}{\underbrace{\sum_{j=1}^{2^p} \mathbb{P}(M_j) \mathbb{P}(\mathbf{t} \mid M_j)}_4}$$

Concerns:

1. Setting a prior on the parameter space:  $\mathbb{P}(\boldsymbol{\theta} \mid M_k)$
2. Setting a prior on the model space:  $\mathbb{P}(M_k)$
3. Computing the integrated likelihood, i.e.  $p + 1$  ugly integrals
4. Fully exploring  $\mathcal{M}_p$ , which has  $2^p$  elements

# Priors on parameters (I)

Consider the linear model with  $\boldsymbol{\theta} \equiv \{\mathbf{w}, q\}$ . Then:

$$\mathbf{t} \mid \mathbf{w}, q, M_k \sim \mathcal{N}(\Phi_k \mathbf{w}_k, q\mathbf{I})$$

Usually, we seek conjugacy to compute the integrated likelihood easily:

$$\mathbb{P}(\mathbf{w}, q \mid M_k) = \underbrace{\mathbb{P}(\mathbf{w} \mid q, M_k)}_{\mathbf{w} \sim \mathcal{N}(\mathbf{0}, q\mathbf{D})} \underbrace{\mathbb{P}(q \mid M_k)}_{q^{-1} \sim \text{Gam}\left(\frac{a_q}{2}, \frac{b_q}{2}\right)}$$

With *conjugate priors*, the integral is in closed form<sup>1</sup>. And so:

- ▶ This is nice with  $2^P$  models to deal with...
- ▶ ...but it requires to set hyperparameters:  $\{a_q, b_q, \mathbf{D}\}$

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<sup>1</sup>See Bishop 2.3.7

# Priors on parameters (II)

Generally,  $\{a_q, b_q\}$  are not very influential for moderate  $n$ :

- ▶ Set them “small”, but positive (to be proper)
- ▶ Say,  $a_q = b_q = 10^{-3}$

The battle is with **D**.

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<sup>2</sup>See lecture on prior modelling

<sup>3</sup>Main advantage: they favour nested (smaller) models even as  $n$  is not “large”, whenever they fit better

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Two families of priors:

## 1. Local priors

- ▶ Zellner's  $g$ -prior<sup>2</sup>
- ▶ Unit information prior: Zellner's with  $g = n$
- ▶ ...

## 2. Non-local priors<sup>3</sup>

- ▶ MOM, eMOM, iMOM (all same idea, changing shape, tails)

Here we will work with Unit Information and iMOM priors as illustration.

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# Priors on models

Let:

- ▶  $d_k := \|\gamma_k\|$  (number of active predictors in  $M_k$ )
- ▶  $\omega := \mathbb{P}(w_i \neq 0)$ , for  $i = 1, \dots, p$  uniformly (though not necessarily)

Common specifications on model priors:

1. Uniform for all models:  $\mathbb{P}(M_k) = 1/2^p$
2. Binomial:  $d_k \sim \text{Binom}(p, \omega)$ 
  - ▶ Models of size  $\approx \omega p$  have higher probability
3. Beta-Binomial(1, 1): as if  $d_k \sim \text{Binom}(p, \omega)$  and  $\omega \sim \text{Unif}(0, 1)$ 
  - ▶ This gives equal probability to models of any size
  - ▶ If Beta-Binomial(1,  $p$ ), then one would favour smaller models
4. Same but with other distributions on  $d_k$ : Beta-binomial, Poisson...



# A large model space $\mathcal{M}_p$

When  $p$  grows, the set of models to explore gets **huge**:

$$\begin{aligned}1,024 &= 2^{10} \\1,073,741,824 &= 2^{30} \\1,152,921,504,606,846,976 &= 2^{60} \\1,267,650,600,228,229,401,496,703,205,376 &= 2^{100} \\&\dots\end{aligned}$$

To gain intuition:  $2^{265} \approx \#$  particles in the observable Universe.

Therefore:

- ▶ For **small**  $p$ : enumerate all models and compute  $\mathbb{P}(M_j \mid \mathbf{t}), \forall j$
- ▶ For **large**  $p$ : stochastic search within  $\mathcal{M}_p$ , i.e. MCMC, Gibbs...

# Choosing a model

Comparing two models  $\{k, l\}$  is easy, regardless (!) of the size of  $\mathcal{M}_p$ :

$$\frac{\mathbb{P}(M_k | \mathbf{t})}{\mathbb{P}(M_l | \mathbf{t})} = \frac{\mathbb{P}(\mathbf{t} | M_k) \mathbb{P}(M_k) / \mathbb{P}(\mathbf{t})}{\mathbb{P}(\mathbf{t} | M_l) \mathbb{P}(M_l) / \mathbb{P}(\mathbf{t})} = \frac{\mathbb{P}(\mathbf{t} | M_k)}{\mathbb{P}(\mathbf{t} | M_l)} \frac{\mathbb{P}(M_k)}{\mathbb{P}(M_l)}$$

And even evaluating a set of them<sup>4</sup>:

$$\left\{ \frac{\mathbb{P}(M_{j_1} | \mathbf{t})}{\mathbb{P}(M_1 | \mathbf{t})}, \frac{\mathbb{P}(M_{j_2} | \mathbf{t})}{\mathbb{P}(M_1 | \mathbf{t})}, \dots, \frac{\mathbb{P}(M_k | \mathbf{t})}{\mathbb{P}(M_1 | \mathbf{t})} \right\}$$

But to actually choose one, we need to either (a) compute them all, or (b) at least the most relevant candidates.

Common selection strategies for **explanatory** models:

1. Highest probability model (HPM)
2. Thresholding: fix  $\alpha \in (0, 1)$  and pick all  $i : \mathbb{P}(\gamma_i = 1 | \mathbf{t}) > \alpha$
3. Bayesian false discovery rate (FDR)

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<sup>4</sup>Say model  $M_1$  is the null model, with no variables

# Bayesian model averaging

For **predictive** models, BMA is the most common choice.

**Idea:** take the weighted average with our posterior model probs.:

$$\begin{aligned}\mathbb{E}(t_{n+1} \mid \mathbf{t}) &= \boldsymbol{\phi}_{n+1}^T \mathbb{E}(\mathbf{w} \mid \mathbf{t}) \\ &= \boldsymbol{\phi}_{n+1}^T \left( \sum_{j=1}^{2^p} \mathbb{E}(\mathbf{w} \mid M_j, \mathbf{t}) \mathbb{P}(M_j \mid \mathbf{t}) \right) \\ &= \boldsymbol{\phi}_{n+1}^T \left( \sum_{j=1}^{2^p} \left( (\boldsymbol{\Phi}_j^T \boldsymbol{\Phi}_j)^{-1} \boldsymbol{\Phi}_j^T \mathbf{t} \right) \mathbb{P}(M_j \mid \mathbf{t}) \right)\end{aligned}$$

This prediction minimises error under  $\ell_2$  loss.

Because this is computationally costly, some approximations exist:

- ▶ Median probability model
- ▶ Model with closest prediction to BMA on average

# Heuristics

A few methods on avoiding (sometimes utopic) computations:

## 1. Stepwise methods

- ▶ Forward
- ▶ Backward
- ▶ Hybrid

## 2. Model space restriction

- ▶ LASSO, SCAD, other penalised methods
- ▶ DECO (Wang, Dunson and Leng, 2016)
- ▶ Block-search (Papaspiliopoulos and Rossell, 2016)

## 3. Pre-screening variables

- ▶ SIS (Fan and Lv, 2008)
- ▶ HOLP (Wang and Leng, 2015)

**(Switch to code now)**

- ▶ Main code: `bvsw.R`
- ▶ Auxiliary functions: `bvsf.R`