

# Gravitational waves from topological defects

Any short subtitle

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## Contents

# Notation

**Constants and units.** We use ‘natural units’ where  $\hbar = c = 1$ , where  $\hbar$  is the reduced Planck constant and  $c$  is the speed of light in vacuum. The Newtonian constant of gravitation  $G_N$  is referenced explicitly, and we use Planck units such as the Planck mass  $M_{\text{Pl}} = (\hbar c/G)^{1/2} = G^{-1/2} \sim 10^{-8} \text{ kg}$ .

**Tensors.** The metric signature  $(-, +, +, +)$  is considered. A four-vector  $p^\mu =$

$$\lambda \equiv \lambda/(2\pi); (\lambda = 1/f) \iff k = 2\pi/\lambda = 1/\lambda \quad (1)$$

## Frequently used abbreviations

CDM	<u>c</u> old <u>d</u> ark <u>m</u> atter
CMB	<u>c</u> osmic <u>m</u> icrowave <u>b</u> ackground (radiation)
DW	<u>d</u> omain <u>w</u> all
GR	<u>g</u> eneral <u>r</u> elativity
GW	<u>g</u> ravitational <u>w</u> ave
$\Lambda$ CDM	‘ <u>L</u> ambda’ (cosmological constant) <u>C</u> DM (i.e. standard model of cosmology)

## Nomenclature

In the table below is listed the most frequently used symbols in this paper, for reference.

**Table 1:** helo

Symbol	Referent	SI-value or definition
<i>Natural constants</i>		
$G_N$	Newtonian constant of gravitation	1.2 kg
$k_B$	Boltzmann's constant	1.2 K
<i>Fiducial quantities</i>		
$h_0$	Reduced Hubble constant	0.67
<i>Subscripts</i>		
$Q_{\text{gw}}$	Quantity $Q$ related to gravitational wave	
<i>Functions and operators</i>		
$\Theta(\xi)$	Heaviside step function	$\begin{cases} 1 & \xi > 0 \\ 0 & \xi < 0 \end{cases}$
$\delta(\xi)$	Dirac-Delta function of $\xi \in \mathbb{R}^n$ , $n \in \mathbb{N}$ .	
$\delta^{\mu\nu}$	Kronecker delta.	



# **Part I**

## **Introduction**



# Chapter 1

## Introduction

We use the convention that  $(-, +, +, +)$  is the metric signature.

$\text{SO}(3)$  is a group

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The order  $O(1)$  is large

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The Planck unit  $M_{\text{Pl}}$

We have a GW with  $\rho_{\text{GW}}$  or  $\rho_{\text{gw}} \dots \rho_{\text{gw}}$

$$\begin{aligned} f(x) &= \int \frac{d^4 k}{(2\pi)^4} e^{-ik \cdot x} \tilde{f}(k) \\ \tilde{f}(k) &= \int d^4 x e^{ik \cdot x} f(x) \end{aligned} \tag{1.1}$$



## Chapter 2

# Theoretical Background

In the context of ... blah ... blah

$$d^3, d^3x, dx, d^2x \tag{2.1}$$

### 2.1 High-Performance Computing



# **Part II**

## **Analytical work**





## Chapter 3

# Dummy chapter

### 3.1 Dummy section

Hello this is a nice thesis

$$h_{ij} = \boldsymbol{a}^s \square \hat{=} \tag{3.1}$$

**Nanna** is cool

**Alma** is lame



**Figure 3.1:** .

helo  
[1, 2]  
See Eq. (3.1). fdf [3]  
Boltzmann equation

$$\hat{L}[f] = \hat{C}[f]; \quad \hat{L}[f] \equiv \frac{\mathrm{d}f}{\mathrm{d}\lambda} \tag{3.2}$$

$$\mathrm{e}^{\mathrm{i}\pi} \tag{3.3}$$

**Hei** jeg heter Nanna



# **Part III**

## **Numerical work**



# **Part IV**

## **Finishing**



# Bibliography

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