Assignment 1, Q2, Extragalactic Astronomy

Nasser Mohammed

Feburary 3rd, 2025

The purpose of this question is to familiarize you with spectral synthesis modelling. Spectral synthesis is an essential tool needed to understand how galaxies evolve

```
In [92]: import numpy as np
    import matplotlib.pyplot as plt
    import astropy.units as u
    plt.style.use('dark_background')

In [93]: import os
    os.environ["SPS_HOME"] = "/Users/nasserm/Documents/vscode/extragalactic/fsps
    from fsps import StellarPopulation
In [94]: from astropy.cosmology import WMAP9 as cosmo
```

Assume that a local galaxy with a stellar mass of $10^{10.4}$ Msun was born at z=5.1, and has been forming stars with an exponentially declining star formation history with an e-folding timescale of 6 Gyr.

Key Values:

- Stellar Mass of Galaxy : $10^{10.4}$
- Born at z = 5.1
- e-folding timescale of 6 Gyr

2.1

What is the current star-formation rate of the galaxy? What was its star-formation rate at z=3?

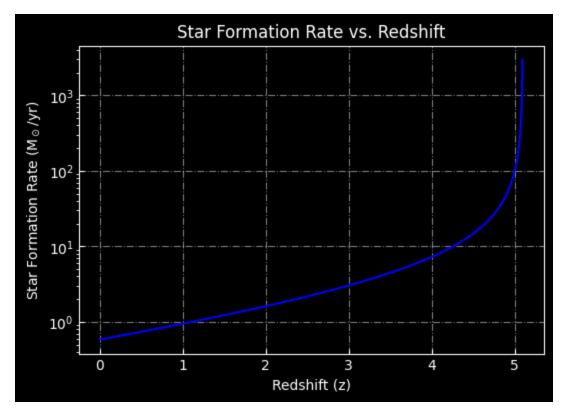
```
dust2=0.2)#,
#tage=cosmo.age(0).value-cosmo.age(5.1).value)
```

To calculate the total stellar mass at any given time, we need to integrate the SFR over time, accounting for mass loss.

$$M(t - \Delta t) = M(t) - SFR(t) \cdot M(t) \cdot \Delta t$$

```
In [96]: mass arr=[10**10.4]
         sf arrate = []
         i=0
         dt = -0.01
         time_range = np.arange(cosmo.age(0).value-cosmo.age(5.1).value, 0, dt)
         for i, t in enumerate(time_range):
             sp.params['tage'] = t
             mass_temp=mass_arr[i]-(sp.sfr*mass_arr[i])*(np.abs(dt)*10**7)
             #print(sp.formed_mass)
             mass arr.append(mass temp)
             sf_arrate.append(sp.sfr*mass_arr[i])
In [97]: # Calculate star formation rate at each redshift in z_array
         z_{array} = np.arange(0, 5.1, 0.004044409199048374)
         # Plot star formation rate as a function of redshift
         plt.figure(figsize=(6, 4))
         plt.plot(z array, sf arrate, 'b-')
         plt.xlabel('Redshift (z)')
         plt.ylabel('Star Formation Rate (M$_\\odot$/yr)')
         plt.grid(ls='-.', alpha=0.5)
         #y axis log
         plt.yscale('log')
         plt.tick params(direction='in')
         plt.title('Star Formation Rate vs. Redshift')
```

Out[97]: Text(0.5, 1.0, 'Star Formation Rate vs. Redshift')

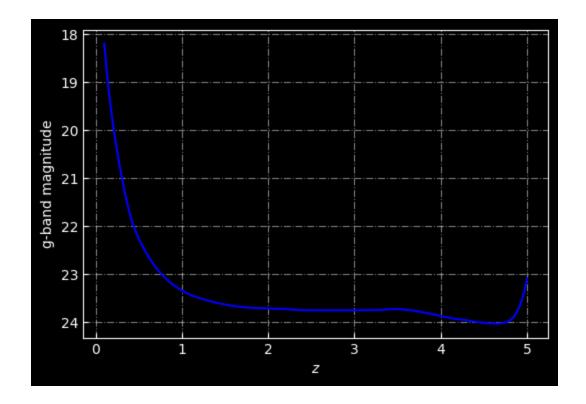


```
In [101... #find the argument where z_array = 3
z3 = np.argmin(np.abs(z_array-3))
print(f'The star forming rate at z=3 is {sf_arrate[z3]:.3f} sollar Masses per
print(f'The star forming rate at z=0 is {sf_arrate[0]:.3f} sollar Masses per
```

The star forming rate at z=3 is 3.034 sollar Masses per year The star forming rate at z=0 is 0.583 sollar Masses per year

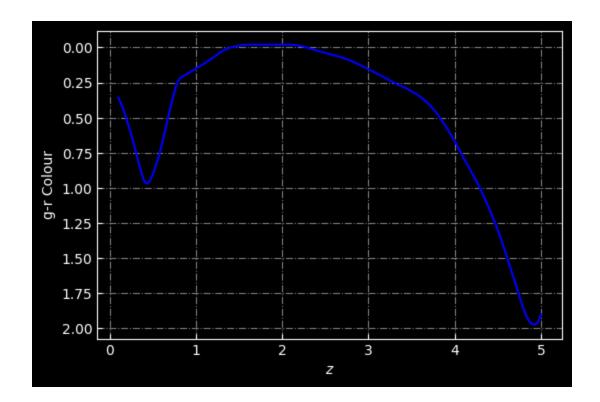
2.2

Plot the rest frame g-band magnitude of the galaxy as a function of redshift, from z=0.1 to z=5.



2.3

Plot the rest frame (g-r) color of the galaxy as a function of redshift over range z=0.1 to z=5.



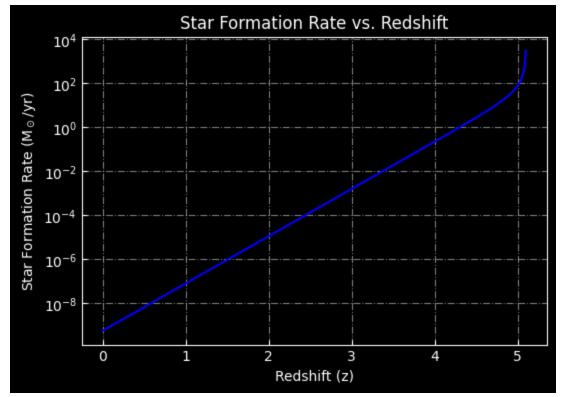
2.4

Repeat 1-3 above assuming the galaxy's star-formation history had an e-folding timescale of 0.5 Gyr.

```
In [109... sp = StellarPopulation(compute_vega_mags=False,
                                      zcontinuous=0, #continuous star formation histor
                                      sfh=1,
                                      tau=0.5,
                                      sf_start=0,
                                      logzsol=0.0,
                                      dust_type=2,
                                      dust2=0.2,
                                      add_stellar_remnants=True,
                                      tage=cosmo.age(0).value-cosmo.age(5.1).value)
In [110... mass_arr=[10**10.4]
         sf arrate = []
         i=0
         dt = -0.01
         time_range = np.arange(cosmo.age(0).value-cosmo.age(5.1).value, 0, dt)
          for i, t in enumerate(time_range):
             sp.params['tage'] = t
              mass_temp=mass_arr[i]-(sp.sfr*mass_arr[i])*(np.abs(dt)*10**7)
              mass arr.append(mass temp)
              sf_arrate.append(sp.sfr*mass_arr[i])
In [111... | # Calculate star formation rate at each redshift in z_array
         z_{array} = np.arange(0, 5.1, 0.004044409199048374)
```

```
# Plot star formation rate as a function of redshift
plt.figure(figsize=(6, 4))
plt.plot(z_array, sf_arrate, 'b-')
plt.xlabel('Redshift (z)')
plt.ylabel('Star Formation Rate (M$_\\odot$/yr)')
plt.grid(ls='-.', alpha=0.5)
#y axis log
plt.yscale('log')
plt.tick_params(direction='in')
plt.title('Star Formation Rate vs. Redshift')
```

Out[111... Text(0.5, 1.0, 'Star Formation Rate vs. Redshift')



```
In [112... #find the argument where z_array = 3
    z3 = np.argmin(np.abs(z_array-3))
    print(f'The star forming rate at z=3 is {sf_arrate[z3]:.3f} sollar Masses pe
    print(f'The star forming rate at z=0 is {sf_arrate[0]:.3f} sollar Masses per
    The star forming rate at z=3 is 0.002 sollar Masses per year
    The star forming rate at z=0 is 0.000 sollar Masses per year
In [113... z_array = np.linspace(0.1, 4.95, 500)
```

```
In [114... plt.figure(figsize=(6, 4))
    plt.plot(z_array, g_band_mag_z, c='b')
    plt.xlabel(r'$z$')
    plt.ylabel(r'g-band magnitude')
    plt.tick_params(direction='in')
```

```
plt.grid(ls = '-.', alpha=0.5)
plt.gca().invert_yaxis()
```

