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Radiative corrections of heavy scalar decays to gauge bosons in the singlet extension of the Standard Model

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Abstract Assuming the existence of a new real scalar singlet s^0 coupled to the Standard Model via a scalar quartic portal interaction, we compute the radiative corrections to the decay rates of the heavy scalar mass eigenstate to a couple of gauge bosons ZZ and $W^+W^-(\gamma)$, showing that they can give a contribution as large as $\mathcal{O}(5\%)$ and $\mathcal{O}(7\%)$, respectively. We also explicitly analyze in detail their dependence on the heavy mass m_S and on the scalar mixing angle α , finding that, especially in the large-mass region, these depend on the sign and the assumed value of $\sin \alpha$.

1 Introduction

In June of 2012, the LHC experiment [1,2] has finally completed the spectrum of the Standard Model with the discovery of the Higgs boson, predicted in the 60's by Higgs [3,4], Englert, Brout [5], Guralnik, Hagen and Kibble [6]. However, the structure and the physics behind the Higgs sector are not completely clear and this represents a possible gateway to the manifold conceivable extensions of the Standard Model (SM). One of the simplest renormalizable enlargement of the Higgs sector is constructed by adding to the SM Lagrangian one additional spinless real singlet under all groups of the SM, which develops its own vacuum expectation value [7–16].

Beside being easy to implement, the physics of a scalar singlet has received a lot of attention in the recent years for several reasons; among them, it can help in solving the issues related to the metastability of the electroweak vacuum [17–20] if the Higgs potential receives a correction due to new physics which modifies it at large field values [21] and it could provide a door to hidden sectors [22] to which it is coupled. The singlet model has the advantage of depending on relatively few parameters and this implies a feasible

experimental study at the LHC for the analysis of the new physic effects in the Higgs boson couplings, searches for heavy SM-like Higgs bosons [23,24] and direct searches for resonant di-Higgs production [25–27]; in the absence of linear and triple self-interactions, this model possesses a \mathbb{Z}_2 -symmetry and the singlet can be a viable candidate for dark matter, although for masses somehow larger than 500 GeV [28–31] the couplings of the dark matter to the known particles occur only through the mixing of the singlet field with the SM Higgs boson. Without a \mathbb{Z}_2 -symmetry a strong first order electroweak phase transition is allowed and additional sources of \mathcal{CP} violation occur in the scalar potential. In this article we limit ourselves to a situation where the new singlet s^0 communicates with the $SU(2)_L$ doublet ϕ only via a quartic interaction of the form,

$$\kappa (\phi^{\dagger} \phi) (s^0)^2$$
.

This implies that the would-be Higgs boson of the SM mixes with the new singlet leading to the existence of two mass eigenstates, the lighter of which (H) is the experimentally observed Higgs boson whereas the heaviest one (S) is a new state not seen so far in any collider experiments. We call this model the Singlet Extension of the SM (SSM). Since only ϕ is coupled to ordinary matter, the main production mechanisms and decay channels of H and S are essentially the same as those of the usual SM Higgs particle, with couplings rescaled by quantities which depend on the scalar mixing angle, called α , whose bounds have been discussed in details in [12,13,32,33]. We will focus our attention on the mass range: $200 \le m_S \le 1000$ GeV. For masses larger than $\gtrsim 200$ GeV, the most important decay channels of the heavy state S are those to a pair of vector bosons $S \rightarrow VV$ and, when kinematically allowed $(m_S > 2 m_H)$, to a pair of lighter scalars and top quarks, $S \rightarrow HH, \bar{t}t$. With the run II at LHC, the exploration of the scalar sector is expected to reveal more details. Thus the comparison between theory and data requires precise predictions obtained through

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higher-order calculations. To this aim, we evaluated the radiative corrections to the decay rates $\Gamma(S \to VV)$ and studied in details their dependence on the singlet mass m_S as well as on the mixing angle α . For the non-diagonal scalar sector, we use the gauge-independent renormalization scheme called "improved on-shell" (iOS) [34] while for the remaining quantities we work within the standard "on-shell" renormalization scheme (OS).

The main result of this paper is that below $m_S \lesssim 700$ GeV, the above mentioned corrections are positive and reach a maximum of $\mathcal{O}(5\%)$ in the ZZ channel and $\mathcal{O}(7\%)$ in the $W^+W^-(\gamma)$ channel (both for $m_S \sim 200$ GeV), almost independently on the mixing angle α , whereas for larger masses the one-loop contributions drive the decay rates to smaller values with a more pronounced dependence on α .

The structure of the paper is as follows: in Sect. 2 we remind the reader of the relevant features of the Singlet Standard Model; in Sect. 3 we illustrate the details of our renormalization procedure that we apply in Sect. 4 to discuss the structure of the SVV renormalized vertices. The radiative corrections to the $S \rightarrow VV$ decay rates are illustrated in detail in Sect. 5; Sect. 6 is devoted to our conclusions.

2 Description of the singlet Standard Model

The scalar potential of the model analyzed here is given by the usual SM potential $V_{\rm sm}(\phi)$, with ϕ representing the SM scalar field, augmented with the new contributions due to quadratic and quartic terms of the new singlet s^0 , and a portal interaction among s^0 and ϕ , contained in $V_{\rm np}(\phi,s^0)$, as specified below:

$$V_{\rm sm}(\phi) = \mu^2(\phi^{\dagger}\phi) + \lambda(\phi^{\dagger}\phi)^2,$$

$$V_{\rm np}(\phi, s^0) = \mu_s^2 (s^0)^2 + \rho (s^0)^4 + \kappa (\phi^{\dagger}\phi)(s^0)^2.$$

The potential $V_{\rm np}(\phi,s^0)$ exhibits a \mathbb{Z}_2 -symmetry under which $s^0 \to -s^0$, so that linear and trilinear terms in s^0 are not allowed s^0 . Since s^0 is a true isospin singlet, the masses of the W and Z gauge bosons are only due to the coupling with the ϕ fields that we parametrize as $\phi = [\eta^+, \phi^0 + i\eta_3/\sqrt{2}]^T$, where $\phi^0 = (v+h)/\sqrt{2}$ (v=246.22 GeV) and $\eta^\pm = (\eta_1 \pm i\eta_2)/\sqrt{2}$ and η_3 are the Goldstone bosons. Beside the vev of the ϕ field, we also consider the possibility that s^0 acquires a non-zero vacuum expectation value (vev) s^0 , and thus the expansion of the field around its classical minimum is set as $s^0 = (w+s)/\sqrt{2}$. The most immediate consequence of the potential s^0 is that a non-diagonal mass matrix is generated for the two neutral states s^0 and s^0

 $^{^{1}\,}$ A similar treatment to the one discussed in this work, including the linear and trilinear terms, can be found in [35–37].



$$\mathcal{M}^2 = \begin{pmatrix} 2\lambda v^2 & \kappa v w \\ \kappa v w & 2\rho w^2 \end{pmatrix}. \tag{1}$$

Considering all couplings as real parameters, the positivity of the mass matrix is ensured requiring that [21]

$$\lambda > \frac{\kappa^2}{4\rho}, \quad \rho > 0. \tag{2}$$

The symmetric mass matrix in Eq. (1) is diagonalized by an orthogonal transformation which, in turn, realizes a mapping between the Lagrangian states and the physical fields H and S [21,34]:

$$m_{H,S}^2 = \lambda v^2 + \rho w^2 \mp \frac{\rho w^2 - \lambda v^2}{\cos 2\alpha},$$
 (3)

$$\phi^0 = \frac{1}{\sqrt{2}}(v + H\cos\alpha + S\sin\alpha),\tag{4}$$

$$s^{0} = \frac{1}{\sqrt{2}}(w - H\sin\alpha + S\cos\alpha),\tag{5}$$

where the mixing angle $\alpha \in [-\pi/2, \pi/2]$. In the rest of this paper we consider the H field as the lightest mass eigenstate and we identify it with the SM Higgs boson, so we always consider $sign\left(\rho w^2 - \lambda v^2\right) \times sign\left(\cos 2\alpha\right) > 0$. Expanding $\mathcal{V}_{np}(\phi,s^0)$ in terms of the physical fields, we get the scalar trilinear and quartic couplings which are reported in Appendix A. The mixing angle α can be expressed in terms of the model parameters and vevs so that

$$\tan 2\alpha = \frac{\kappa v w}{\alpha w^2 - \lambda v^2}.$$
 (6)

Here we limit ourselves to the mass range $200 \le m_S \le 1000$ GeV, which corresponds to the bound $0.018 \le |\sin\alpha| \le 0.36$ [34]. Notice that in the decoupling limit $(v/w) \ll 1$, the expressions for the masses and mixing are well approximated by:

$$m_H^2 \equiv 2v^2 \lambda_{\text{SM}} \simeq 2v^2 \left(\lambda - \frac{\kappa^2}{4\rho}\right)$$

$$m_S^2 \simeq 2\rho w^2 + \frac{\kappa^2 v^2}{2\rho}, \quad \sin \alpha \simeq \frac{\kappa v}{2\rho w},$$
(7)

which clearly shows that the quartic coupling λ_{SM} receives a correction proportional to the ratio among the portal coupling κ and the quartic of the s^0 field [20]. The couplings of the H and S fields with gauge bosons and fermions are similar to the ones of the SM Higgs, rescaled by the appropriate mixing [21]:

$$\mathcal{L}_{\text{SSM}} \supset \frac{Hc_{\alpha} + Ss_{\alpha}}{v} \times \left[m_Z^2 Z_{\mu} Z^{\mu} + 2m_W^2 W_{\mu}^+ W^{\mu-} - \sum_f m_f \bar{f} f \right], \tag{8}$$

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with $s_{\alpha} = \sin \alpha$ and $c_{\alpha} = \cos \alpha$. Then the tree-level amplitude for the $S \to VV$ decays is given by:

$$S(k)V(p,a)V(q,b) \Longrightarrow \rho_V \times [g^{\mu\nu}\epsilon^a_\mu(p)\epsilon^b_\nu(q)],$$
 (9)

where k, p and q are the four momenta and a, b the polarizations. Here ρ_V is the SSM bare coupling defined as

$$\rho_V = e \frac{m_V^2}{s_W m_W} s_\alpha,\tag{10}$$

where $V = W^{\pm}$, Z and $s_W = \sin \theta_W$, θ_W being the Weinberg's angle.

3 Generalities on the renormalization procedure

We start introducing the main renormalized quantities and counterterms of our interest [38]:

$$\mathcal{V}_0 = \mathcal{V}(1 + \delta \mathcal{V}), \tag{11}$$

$$(m_V^2)_0 = m_V^2 + \delta m_V^2, (12)$$

$$(\theta_W)_0 = \theta_W + \delta\theta_W, \tag{13}$$

$$\alpha_0 = \alpha + \delta \alpha, \tag{14}$$

$$e_0 = (1 + \delta Z_e)e, \tag{15}$$

where V_0 is a short-hand notation for a generic coupling and e is the electric charge. The relevant difference with respect to the SM renormalization procedure is the presence of the mixing in the scalar sector. Splitting the bare mixing angle as $\alpha_0 \rightarrow \alpha + \delta \alpha$, the two physical scalar fields S and H are related to the bare ones via mixing specified as

$$\begin{pmatrix} H_0 \\ S_0 \end{pmatrix} \to \begin{pmatrix} 1 + \frac{\delta Z_H}{2} & \frac{\delta Z_{HS}}{2} - \delta \alpha \\ \frac{\delta Z_{SH}}{2} + \delta \alpha & 1 + \frac{\delta Z_S}{2} \end{pmatrix} \begin{pmatrix} H \\ S \end{pmatrix}. \tag{16}$$

We will also need the field renormalization constants for W^{\pm} , Z^0 and γ defined as

$$W_0^{\pm} = \left(1 + \frac{1}{2}\delta Z_W\right)W^{\pm},\tag{17}$$

$$\begin{pmatrix} Z_0 \\ \gamma_0 \end{pmatrix} = \begin{pmatrix} 1 + \frac{\delta Z_Z}{2} & \frac{\delta Z_{Z\gamma}}{2} - \delta \theta_W \\ \frac{\delta Z_{\gamma Z}}{2} + \delta \theta_W & 1 + \frac{\delta Z_{\gamma}}{2} \end{pmatrix} \begin{pmatrix} Z \\ \gamma \end{pmatrix},$$
 (18)

where in the last line we explicitly show the counterterms entering in the mixing matrix of the neutral gauge bosons. Notice that we can rewrite $\delta\theta_W$ as $\delta s_W^2/(2s_Wc_W)$ using: $\delta\theta_W = \delta s_W/c_W$ and $\delta s_W = \delta s_W^2/(2s_W)$. As usual, the counterterms are fixed by renormalization conditions [38] (the "hat" will be used to indicate renormalized quantities). The tadpole of the scalar fields² must also be shifted:

$$\hat{T}_H \to T_H + \delta t_H, \quad \hat{T}_S \to T_S + \delta t_S,$$
 (19)

where T_H and T_S are related with the tadpoles in the gauge basis T_h and T_S by the mixing,

$$\begin{pmatrix} T_h \\ T_s \end{pmatrix} \to \begin{pmatrix} c_{\alpha} & s_{\alpha} \\ -s_{\alpha} & c_{\alpha} \end{pmatrix} \begin{pmatrix} T_H \\ T_S \end{pmatrix}. \tag{20}$$

If we impose the renormalization conditions $\hat{T}_H = T_H + \delta t_H = 0$ and $\hat{T}_S = T_S + \delta t_S = 0$, no scalar one-point insertions need to be explicitly computed since it is equivalent to require that v and w are the physical vacuum expectation values of the doublet and the singlet fields, respectively. The next conditions involve the renormalized one-particle irreducible two-point functions of the scalar and vector fields; in the 't Hooft-Feynman gauge (that will be used throughout the rest of this paper) we have:

$$\hat{\Gamma}^{ii'}(k) = i(k^2 - m_i^2)\delta^{ii'} + i\hat{\Sigma}^{ii'}(k^2), \tag{21}$$

$$\hat{\Gamma}_{\mu\nu}^{jj'}(k) = -ig^{\mu\nu}(k^2 - m_i^2)\delta^{jj'}$$

shell³:

$$-i\left(g^{\mu\nu} - \frac{k^{\mu}k^{\nu}}{k^2}\right)\hat{\Sigma}_{T}^{jj'}(k^2) - i\frac{k^{\mu}k^{\nu}}{k^2}\hat{\Sigma}_{L}^{jj'}(k^2),\tag{22}$$

where $m_{i/j}$ is the mass of the incoming particle and (ii', jj') can be one of the combinations $\{HH, SS, HS, SH\}$ and $\{WW, ZZ, \gamma\gamma, \gamma Z, Z\gamma\}$, respectively. $\hat{\Sigma}_{T}^{jj'}$ and $\hat{\Sigma}_{L}^{jj'}$ are the transverse and longitudinal contributions to the self-energies. Following Ref. [38], we can impose the following conditions on the self-energy functions in the OS renormalization scheme, in which all renormalization conditions are formulated for the external fields on their mass

$$\begin{split} Re\,\hat{\Sigma}^{HH}(m_H^2) &= 0, \quad Re\,\hat{\Sigma}^{'HH}(k^2)|_{k^2=m_H^2} = 0 \;, \\ Re\,\hat{\Sigma}^{SS}(m_H^2) &= 0, \quad Re\,\hat{\Sigma}^{'SS}(k^2)|_{k^2=m_H^2} = 0 \;, \\ \widetilde{Re}\,\hat{\Sigma}^{WW}_T(m_W^2) &= 0, \quad \widetilde{Re}\,\hat{\Sigma}^{'WW}_T(k^2)|_{k^2=m_W^2} = 0 \;, \\ Re\,\hat{\Sigma}^{ZZ}_T(m_Z^2) &= 0, \quad Re\,\hat{\Sigma}^{'ZZ}_T(k^2)|_{k^2=m_Z^2} = 0 \;, \\ Re\,\hat{\Sigma}^{\gamma\gamma}_T(0) &= 0, \quad Re\,\hat{\Sigma}^{'\gamma\gamma}_T(k^2)|_{k^2=0} = 0 \;, \\ Re\,\hat{\Sigma}^{Z\gamma}_T(m_Z^2) &= 0, \quad Re\,\hat{\Sigma}^{\gamma\gamma}_T(0) &= 0 \;. \end{split}$$

The function \widetilde{Re} takes the real part of the loop integrals only and it does not remove the imaginary parts arising from the various couplings of the theory (e.g. from complex CKM matrix elements); $\hat{\Sigma}'$ instead is a short-hand nota-

³ The renormalization conditions for the mixing in the scalar sector are discussed below Eq. (37).



² The tadpoles are given by the following relations: $T_h = \mu v^2 + v^3 \lambda + v w^2 \kappa / 2$, $T_s = \mu_s w^2 + w^3 \rho + v^2 w \kappa / 2$.

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tion for $\hat{\Sigma}'(k^2) = \partial \hat{\Sigma}(k^2)/\partial k^2$. The definitions of $\hat{\Sigma}(k^2)$ are [34,38]⁴:

$$\widetilde{Re}\,\hat{\Sigma}_T^{WW}(p^2) = \widetilde{Re}\,\Sigma_T^{WW}(p^2) + \delta Z_W(p^2 - m_W^2) - \delta m_W^2,$$
(23)

$$Re\hat{\Sigma}_{T}^{ZZ}(p^{2}) = Re\Sigma_{T}^{ZZ}(p^{2}) + \delta Z_{Z}(p^{2} - m_{Z}^{2}) - \delta m_{Z}^{2},$$
 (24)

$$Re\hat{\Sigma}_{T}^{\gamma\gamma}(p^2) = Re\Sigma_{T}^{\gamma\gamma}(p^2) + p^2\delta Z_{\gamma},$$
 (25)

$$Re\hat{\Sigma}_{T}^{\gamma Z}(p^2) = Re\Sigma_{T}^{\gamma Z}(p^2)$$

$$+\frac{1}{2}\delta Z_{\gamma Z}(2p^2 - m_Z^2) + m_Z^2 \delta \theta_W, \tag{26}$$

$$Re\hat{\Sigma}^{HH}(p^2) = Re\Sigma^{HH}(p^2) + \delta Z_H(p^2 - m_H^2) - \delta m_H^2,$$
(27)

$$Re\hat{\Sigma}^{SS}(p^2) = Re\Sigma^{SS}(p^2) + \delta Z_S(p^2 - m_S^2) - \delta m_S^2, (28)$$

$$Re\hat{\Sigma}^{HS}(p^2) = Re\hat{\Sigma}^{SH}(p^2) = Re\Sigma^{HS}(p^2)$$

$$+ (m_H^2 - m_S^2)\delta\alpha + + \left[\frac{\delta Z_{HS}}{2} (p^2 - m_H^2) + \frac{\delta Z_{SH}}{2} (p^2 - m_S^2) \right], \tag{29}$$

where we can put (for example, see Eq. (38) of [35]):

$$\delta Z_W = \delta Z_{\gamma} + \frac{c_W}{s_W} \delta Z_{\gamma Z}. \tag{30}$$

Finally, from the renormalization conditions above and Eqs. (23–29) we can extract the counterterms:

$$\delta m_H^2 = Re \Sigma^{HH}(m_H^2) ,$$

 $\delta Z_H = -Re \Sigma^{'HH}(k^2)|_{k^2 = m_H^2} ,$ (31)

$$\delta m_S^2 = Re \Sigma^{SS}(m_S^2),$$

$$\delta Z_S = -Re \Sigma^{'SS}(k^2)|_{k^2 = m_o^2}, \qquad (32)$$

$$\delta m_W^2 = \widetilde{Re} \Sigma_T^{WW}(m_W^2),$$

$$\delta Z_W = -\widetilde{R}e\Sigma_T^{'WW}(k^2)|_{k^2 = m_W^2}, \qquad (33)$$

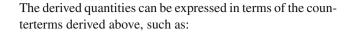
$$\delta m_Z^2 = Re \Sigma_T^{ZZ}(m_Z^2) ,$$

$$\delta Z_Z = -Re \Sigma_T^{'ZZ}(k^2)|_{k^2 = m_Z^2} , \qquad (34)$$

$$\delta Z_{\gamma Z} = 2Re \frac{\Sigma_T^{\gamma Z}(0)}{m_Z^2} + \frac{\delta s_W^2}{s_W c_W},$$

$$\delta Z_{Z\gamma} = -2Re \frac{\Sigma_T^{\gamma Z}(m_Z^2)}{m_Z^2} - \frac{\delta s_W^2}{s_W c_W},$$
 (35)

$$\delta Z_{\gamma} = -Re \Sigma_T^{'\gamma\gamma}(k^2)|_{k^2=0}. \tag{36}$$



$$\delta s_W^2 = -\delta c_W^2 = c_W^2 \left(\frac{\delta m_Z^2}{m_Z^2} - \frac{\delta m_W^2}{m_W^2} \right),$$

$$\delta Z_e = -\frac{1}{2} \delta Z_\gamma + \frac{s_W}{c_W} \frac{Re \Sigma_T^{\gamma Z}(0)}{m_Z^2}.$$
(37)

Now, to fix the non-diagonal scalar field renormalization δZ_{HS} , we consider the improved on-shell (iOS) renormalization scheme [34] which requires that loop-induced S-H or H-S transitions vanish for on-shell external scalar states:

$$Re\hat{\Sigma}^{HS}(p^2)\big|_{p^2=m_H^2} = 0$$
 , $Re\hat{\Sigma}^{HS}(p^2)\big|_{p^2=m_S^2} = 0$. (38)

The iOS definitions of $\delta Z_{HS,SH}^{\rm ios}$ (we indicate with the superscript "ios" the counterterm of the mixing scalar sector arising from the iOS renormalization conditions) are determined using Eqs. (29) and (38). These equations lead to [34],

$$\frac{\delta Z_{HS}^{\text{ios}}}{2} = \frac{Re \Sigma^{HS}(m_S^2)}{m_H^2 - m_S^2} + \delta \alpha^{\text{ios}},$$

$$\frac{\delta Z_{SH}^{\text{ios}}}{2} = \frac{Re \Sigma^{HS}(m_H^2)}{m_S^2 - m_H^2} - \delta \alpha^{\text{ios}}.$$
(39)

On the other hand, the mixed mass (mixing angle) counterterm is defined in the following way [34]:

$$\delta m_{HS}^{2 \text{ ios}} = (m_S^2 - m_H^2) \, \delta \alpha^{\text{ ios}} = \left. Re \Sigma^{HS} (p^{*2}) \right|_{p^{*2} = \frac{m_H^2 + m_S^2}{2}},$$
(40)

where p^{*2} is fixed to the average of the squared masses. The reason for such a choice of p^{*2} lies on the fact that the mixed scalar self-energy at p^{*2} is independent on the gauge-fixing scheme. The gauge independence of the iOS scheme is also discussed in [39] where it is shown how the mixed scalar self energy at p^{*2} coincides with the gauge invariant part of the same quantity obtained through the so-called *pinch technique*, which generally allows the construction of off-shell Green's functions in non-Abelian gauge [40] or extended scalar [41] theories that are independent of the gauge-fixing parameter.

The analytic results presented in this paper have been obtained using FEYNRULES [42] to generate the Feynman rules for the SSM model. All amplitudes are computed with FEYNARTS [43] while their analytical processing was done with FORMCALC [43]. The outputs, written in terms of standard loop integrals, have been evaluated with the help of PACKAGE- X [44].



⁴ Differently from the approach of Ref. [34], where $Re\hat{\Sigma}^{HS}(k^2)$ is expressed in terms of the mixed mass counterterm δm_{HS}^2 , we used the mixing angle counterterm $\delta \alpha$. The two approaches are related as follows: $\delta m_{HS}^2 = (m_S^2 - m_H^2)\delta \alpha$.

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4 Renormalization of the SVV vertex in the SSM

In this section we apply the renormalization procedure to the vertex of the scalar field S with two gauge bosons. The related Feynman diagrams are reported in Appendix B, where we only show the contributions due to the insertion of the S field in the loops because loops with the SM fields are as those quoted in [45] where the external Higgs leg can be replaced with the new scalar singlet. The bare V_0 and one-loop corrections to the SVV vertex can be put in the following form [45]:

$$\mathcal{V}^{\mu\nu} = \mathcal{V}_0^{\mu\nu} + \rho_V T^{\mu\nu} \,, \tag{41}$$

where ρ_V has been defined in Eq. (10) and $\mathcal{V}_0^{\mu\nu} = \rho_V g^{\mu\nu}$. The generic expansion of $T^{\mu\nu}$ in terms of 2-index tensors is given by [45]:

$$T_V^{\mu\nu} = \mathcal{A}_V p^\mu p^\nu + \mathcal{B}_V q^\mu q^\nu + \mathcal{C}_V p^\mu q^\nu + \mathcal{D}_V q^\mu p^\nu + \mathcal{E}_V g^{\mu\nu} + i \mathcal{F}_V \epsilon^{\mu\nu\rho\sigma} p_\rho q_\sigma, \tag{42}$$

where p and q are the four-vectors of the external gauge bosons. The coefficients $\mathcal{A}_V,...,\mathcal{F}_V$ have to be ultra-violet (UV) finite whereas the term proportional to the antisymmetric tensor $\epsilon^{\mu\nu\rho\sigma}$ vanishes due to the charge conjugation invariance for external Z bosons and also if the gauge bosons are on the mass-shell. We decide to set the external squared momenta $[k^2, p^2, q^2]$ in the S(k)V(p, a)V(q, b) vertex as $[m_S^2, m_V^2, m_V^2]$, respectively. We take physical gauge bosons, so that only the coefficients \mathcal{D}_V and \mathcal{E}_V become relevant. Since the counterterms arising from the quantities in Eq. (10) are included in the coefficient \mathcal{E}_V , we put it in the form $\mathcal{E}_V = \delta \rho_V^{CT} + \delta \mathcal{V}_V^{\mathcal{E}}$, where the symbol $\delta \mathcal{V}$ is used to indicate the three point function contributions at the one loop level (whose cumbersome expressions are not reported here) and $\delta \rho_V^{CT}$ are given by:

$$\delta\rho_{W} = \frac{\delta m_{W}^{2}}{2m_{W}^{2}} - \frac{\delta s_{W}^{2}}{2s_{W}^{2}} + \delta Z_{W} + \delta Z_{e}$$

$$+ \frac{\delta Z_{S}}{2} + \frac{c_{\alpha}}{s_{\alpha}} \left(\frac{\delta Z_{HS}^{\text{ios}}}{2} - \delta \alpha^{\text{ios}} \right) + \frac{\delta s_{\alpha}^{\text{ios}}}{s_{\alpha}}, \qquad (43)$$

$$\delta\rho_{Z} = \frac{\delta m_{Z}^{2}}{m_{Z}^{2}} - \frac{\delta m_{W}^{2}}{2m_{W}^{2}} - \frac{\delta s_{W}^{2}}{2s_{W}^{2}} + \delta Z_{Z} + \delta Z_{e}$$

$$+ \frac{\delta Z_{S}}{2} + \frac{c_{\alpha}}{s_{\alpha}} \left(\frac{\delta Z_{HS}^{\text{ios}}}{2} - \delta \alpha^{\text{ios}} \right) + \frac{\delta s_{\alpha}^{\text{ios}}}{s_{\alpha}}. \qquad (44)$$

Notice that we obtain $\delta \rho_V$ independent from the mixing angle counterterm using the following substitution: $\delta s_{\alpha}^{\rm ios} = c_{\alpha} \delta \alpha^{\rm ios}$. The other counterterms entering the previous expressions have been listed in Eqs. (32–37) and Eq. (39). In the following we will work in the *modified on-shell mass scheme*

(MOMS), in which the electric charge dependence in the coupling is replaced by the Fermi constant G_F via

$$\frac{G_F}{\sqrt{2}} = \frac{e^2}{8s_W^2 m_W^2} \frac{1}{1 - \Delta r}. (45)$$

This prevents the appearance of ambiguities associated with the definition of the light-quark masses and mass singularities due to light fermions in $\ln(m_Z^2/m_f^2)$ terms in δZ_e (more precisely, these logarithms appear in δZ_γ which is part of δZ_e). The factor Δr represents finite corrections to G_F ; these are well known and up to $\mathcal{O}(\alpha_{em}^2)$ are given by $[45,46]^5$:

$$\Delta r = \frac{\widetilde{Re} \hat{\Sigma}_{T}^{WW}(0)}{m_{W}^{2}} + \frac{\alpha_{em}}{4\pi s_{W}^{2}} \left[\left(\frac{7}{2s_{W}^{2}} - 2 \right) \ln[c_{W}^{2}] + 6 \right], \tag{46}$$

where $\widetilde{Re}\widehat{\Sigma}_T^{WW}(p^2)$ is the renormalized transverse self-energy of the W boson at momentum transfer p defined in Eq. (23) and the last term is due to the vertex-box loop corrections in the muon decay process. The use of G_F instead of the electric charge amounts to shift $\delta \rho_V^{CT} \to \delta \rho_V^{CT'} = \delta \rho_V^{CT} - \Delta r/2$; if in Δr we use the form of δZ_W as given in Eq. (30), the cancellation of the δZ_γ in the final counterterm contributions is guaranteed and no problem arises from the light fermion loop contributions. So finally we get:

$$\delta \rho_{V}' = \delta Z_{V} + \frac{\delta Z_{S}}{2} + \frac{c_{\alpha}}{s_{\alpha}} \frac{\delta Z_{HS}^{\text{ios}}}{2} + \frac{\delta m_{V}^{2}}{m_{V}^{2}} - \frac{\widetilde{Re} \Sigma_{T}^{WW}(0)}{2m_{W}^{2}} + \frac{Re \Sigma_{T}^{\gamma Z}(0)}{s_{W} c_{W} m_{Z}^{2}} - \frac{\alpha_{em}}{8\pi s_{W}^{2}} \left[\left(\frac{7}{2s_{W}^{2}} - 2 \right) \ln[c_{W}^{2}] + 6 \right].$$
(47)

The coefficient \mathcal{E}_V is UV-finite both for Z and W external boson pairs, as it can be explicitly verified from the expressions of the bosonic and fermionic divergent parts quoted in Table 1 for all the counterterms (which are divided by a common factor $g^2/(16\pi^2 \epsilon)$).

Regarding the finite parts, we know that the S field gives negligible contributions to the corrections of the muon decay and since there is no S field dependence in $\Sigma_T^{\gamma Z}(0)$ [32], the new scalar contributions only affect the bosonic parts of $\widetilde{Re}\Sigma_T^{WW}(0)$, δm_Z^2 , δZ_Z , δZ_S , δZ_{HS}^{ios} and $\delta \mathcal{V}_V^{\mathcal{E}}$. The fermionic contributions of $\widetilde{Re}\Sigma_T^{WW}(0)$, δm_Z^2 , δZ_Z and $\delta \mathcal{V}_V^{\mathcal{E}}$ are identical to those associated to the HVV vertex in the SM; in addition, their contributions to δZ_S and δZ_{HS}^{ios} can be determined from the fermion loop terms in the SM δZ_H expression but now multiplied by s_α^2 and $c_\alpha s_\alpha$ and with external

⁵ In the SSM, the new contributions to Δr generate a maximum variation of $\mathcal{O}(0.1)\%$ for $|s_{\alpha}| \sim 0.2$ and $m_S \geq 200$ GeV [32].



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Table 1 Coefficients of the bosonic and fermionic UV divergent parts of the relevant counterterms (which are divided by the common factor $g^2/(16\pi^2\,\epsilon)$). The symbol \checkmark (×) indicates that the corresponding counterterm is present (absent) in $\mathcal{E}_{W,Z}$

	\mathcal{E}_W	\mathcal{E}_Z	$UV_{bosonic}$	$UV_{fermionic}$
δZ_W	\checkmark	×	19/6	-4
δZ_Z	×	\checkmark	$\frac{-1+2c_W^2+18c_W^4}{6c_W^2}$	$\frac{-20+40c_W^2-32c_W^4}{3c_W^2}$
$\delta Z_S/2$	\checkmark	\checkmark	$\frac{s_{\alpha}^{2}(2c_{W}^{2}+1)}{4c_{W}^{2}}$	$-\frac{s_{\alpha}^2 \sum_f N_c m_f^2}{4m_W^2}$
$c_{\alpha}\delta Z_{HS}^{\mathrm{ios}}/2s_{\alpha}$	\checkmark	\checkmark	$\frac{c_{\alpha}^{2}(2c_{W}^{2}+1)}{4c_{W}^{2}}$	$-rac{c_{lpha}^2\sum_f^N N_c m_f^2}{4m_W^2}$
$\delta m_W^2/m_W^2$	\checkmark	×	$\frac{6-31c_W^2}{6c_W^2}$	$4 - \frac{\sum_{f} N_{c} m_{f}^{2}}{2m_{W}^{2}}$
$\delta m_Z^2/m_Z^2$	×	\checkmark	$\frac{7 + 10c_W^2 - 42c_W^4}{6c_W^2}$	$\frac{20 - 40c_W^2 + 32c_W^4}{3c_W^2} - \frac{\sum_f N_c m_f^2}{2m_W^2}$
$-\widetilde{Re}\Sigma_T^{WW}(0)/2m_W^2$	\checkmark	\checkmark	$\frac{2c_W^2-1}{2c_W^2}$	$\frac{\sum_f N_c m_f^2}{4m_W^2}$
$Re\Sigma_T^{\gamma Z}(0)/s_W c_W m_Z^2$	\checkmark	\checkmark	-2	0
$\delta {\cal V}^{\cal E}_W$	\checkmark	×	$\frac{-3+10c_W^2}{4c_W^2}$	$\frac{\sum_f N_c m_f^2}{2m_W^2}$
$\delta \mathcal{V}_Z^{\mathcal{E}}$	×	\checkmark	$\frac{-3 - 6c_W^2 + 16c_W^4}{4c_W^2}$	$\frac{\sum_f N_c m_f^2}{2m_W^2}$

momenta fixed to m_S^2 (the definition of $\delta Z_{HS}^{\rm ios}$ also contains the mixed two-point function with external momenta fixed to $(m_H^2 + m_S^2)/2$).

Notice that the renormalization of the vertex SW^+W^- is more complicated than the SZZ vertex since contributions due to the photons in the loop integrals, which are plagued by infrared (IR) singularities when the W bosons are on-shell, must be taken into account. The IR-cancellation is obtained considering soft-photon bremsstrahlung contributions [47] which, for the model under discussion, are shown in Fig. 1.

5 The decay $S \rightarrow VV$ in the SSM

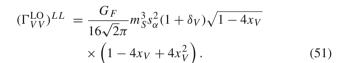
The decay rate of the scalar S into two physical gauge bosons gets contributions from longitudinally (L) and transversally $(\pm\pm)$ polarized gauge bosons. Using the LO amplitude of Eq. (9), a straightforward computation of the decay width gives,

$$\Gamma_{VV}^{\text{LO}} = \frac{G_F}{16\sqrt{2}\pi} m_S^3 s_\alpha^2 (1 + \delta_V)
\times \sqrt{1 - 4x_V} \left(1 - 4x_V + 12x_V^2 \right),$$
(48)

where $x_V = m_V^2/m_S^2$, $\Gamma_{ii} = \Gamma(S \to ii)$ and $\delta_V = 0, 1$ for V = Z, W^{\pm} respectively. The longitudinally and transversally polarized gauge boson contributions to Γ_{VV}^{LO} are given by [45,47]:

$$(\Gamma_{VV}^{\text{LO}})^{\pm L} = (\Gamma_{VV}^{\text{LO}})^{\pm \mp} = 0, \tag{49}$$

$$(\Gamma_{VV}^{\text{LO}})^{\pm \pm} = \frac{G_F}{16\sqrt{2}\pi} m_S^3 s_\alpha^2 (1 + \delta_V) \sqrt{1 - 4x_V} \times (4x_V^2), \tag{50}$$



As a consequence of Eq. (42) and of the polarization conditions ($\epsilon_{\mu}(p) \cdot p^{\mu} = \epsilon_{\nu}(q) \cdot q^{\nu} = 0$), the physical amplitude is reduced to:

$$\mathcal{V}^{\mu\nu}\epsilon_{\mu}(p)\epsilon_{\nu}(q) = \rho_{V}[g^{\mu\nu}(1+\mathcal{E}_{V}) + q^{\mu}p^{\nu}\mathcal{D}_{V}]\epsilon_{\mu}(p)\epsilon_{\nu}(q). \tag{52}$$

As mentioned above, the radiative corrections to $S \to W^+W^-$ are guaranteed to be IR-finite if we include softphoton bremsstrahlung contributions which, for the model under discussion, are shown in Fig. 1. We call the photon momenta as q_γ whose maximum value is $q_\gamma^{\max} = m_S(1-4x_W)/2$. To set an ideal boundary between the soft and hard region, we introduce a cutoff Λ_γ in such a way that the soft region corresponds to $m_\gamma \leq q_\gamma \leq \Lambda_\gamma$ while the hard region to $\Lambda_\gamma \leq q_\gamma \leq q_\gamma^{\max}$, where m_γ is a fake mass assigned to the photon. The total photon-bremsstrahlung decay rate is then given by the sum of the soft and hard contributions:

$$\Gamma_{WW}^{\text{brem}} = \Gamma_{WW}^{\text{soft}} + \Gamma_{WW}^{\text{hard}} = \Gamma_{WW}^{\text{LO}}(\delta_W^{\text{soft}} + \delta_W^{\text{hard}}),$$
 (53)

where the correction factors $\delta_W^{\rm soft}$ and $\delta_W^{\rm hard}$ are reported in Appendix C and are extracted from [47]. The m_γ and Λ_γ dependences show up in δZ_W , $\delta \mathcal{V}_W^{\mathcal{E}}$, $\delta_W^{\rm soft}$ and $\delta_W^{\rm hard}$, as detailed in Table 2. Here $r=m_S^2/4m_W^2$ and the function $\mathcal{G}(r)$ is defined in Appendix C.



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Fig. 1 Feynman diagrams of the photon bremsstrahlung associated to the first-order radiative corrected $SW^+W^-(\gamma)$ vertex

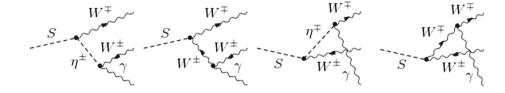


Table 2 IR-dependence on m_{γ} and Λ_{γ} in δZ_W , $\delta \mathcal{V}_W^{\mathcal{E}}$, δ_W^{soft} and δ_W^{hard}

	m_{γ} (IR regulator)	$Λ_γ$ (IR cutoff)
δZ_W	$\frac{\alpha_{em}}{2\pi} \ln \left(\frac{m_W^2}{m_Y^2} \right)$	-
$\delta \mathcal{V}_W^{\mathcal{E}}$	$rac{lpha_{em}}{2\pi}\left[\mathcal{G}(r)+1 ight]\ln\left(rac{m_{\gamma}^{2}}{m_{W}^{2}} ight)$	-
$\delta_W^{\rm soft}$	$rac{lpha_{em}}{\pi} \mathcal{G}(r) \ln \left(rac{m_W^2}{m_\gamma^2} ight)$	$\frac{\alpha_{em}}{\pi} \mathcal{G}(r) \ln \left(\frac{4\Lambda_{\gamma}^2}{m_W^2} \right)$
$\delta_W^{ m hard}$	-	$\frac{\alpha_{em}}{\pi} \mathcal{G}(r) \ln \left(\frac{m_S^2}{4\Lambda_\gamma^2} \right)$

The NLO total decay width, which we call $\Gamma_{VV}^{\rm NLO}$, is given by the sum of $\Gamma_{VV}^{\pm\pm}$ and Γ_{VV}^{L} at the one-loop order and $\Gamma_{WW}^{\rm brem}$ of Eq. (53). So we have,

$$\Gamma_{VV}^{\text{NLO}} = \frac{G_F}{16\sqrt{2}\pi} m_S^3 s_\alpha^2 (1+\delta_V) \sqrt{1-4x_V}
\times (1-4x_V+12x_V^2)
\times \left\{ 1+2 \left[\delta \rho_V' + \delta \mathcal{V}_V^{\mathcal{E}} + \frac{m_S^2}{2} \left(\frac{1-6x_V+8x_V^2}{1-4x_V+12x_V^2} \right) \delta \mathcal{V}_V^{\mathcal{D}} \right] \right\}
+ \delta_V \Gamma_{WW}^{\text{brem}},$$
(54)

where $\delta \mathcal{V}_V^{\mathcal{D}}$ are the corrections from the coefficient \mathcal{D}_V . Now, two comments are in order:

1. using the m_{γ} -dependent contributions, reported in Table 2, we can verify the cancellation of the IR-divergences:

$$\{\Gamma_{WW}^{\text{NLO}}\}_{\text{IR}} \propto \left[1 + \delta_{W}^{\text{soft}} + 2(\delta Z_{W} + \delta \mathcal{V}_{W}^{\mathcal{E}})\right]$$

$$\propto \left[1 + \frac{\alpha_{em}}{\pi} \mathcal{G}(r) \ln\left(\frac{4\Lambda_{\gamma}^{2}}{m_{W}^{2}}\right)\right]; \tag{55}$$

2. the combination of all terms in Table 2 is Λ_{γ} -independent at $\mathcal{O}(\alpha_{em})$.

5.1 Numerical results

In the evaluation of the corrections to the total decay rate we make use of the following quantity:

$$\mathcal{R}_{VV}^{\text{SSM}} = \left[\left(\Gamma_{VV}^{\text{NLO}} / \Gamma_{VV}^{\text{LO}} \right) - 1 \right]. \tag{56}$$

Table 3 Values of m_S considered in our numerical analysis, the ranges of $|s_{\alpha}|$ and the corresponding w_{min} . Table extracted from Table I of [34]

m_S [GeV]	$ s_{\alpha} $	w_{min} [GeV]	
200	[0.09,0.36]	0.85 v	
300	[0.067,0.31]	1.25 v	
400	[0.055,0.27]	1.69 v	
500	[0.046,0.24]	2.13 v	
600	[0.038,0.23]	2.56 v	
700	[0.031,0.21]	3.03 v	
800	[0.027, 0.21]	3.45 v	
900	[0.022,0.19]	3.85 v	
1000	[0.018, 0.17]	4.34 v	

As a set of independent input parameters we choose w, m_S and α and express at tree-level λ , ρ and κ according to [34]:

$$\lambda = \frac{m_H^2 c_\alpha^2 + m_S^2 s_\alpha^2}{2v^2}, \quad \rho = \frac{m_S^2 c_\alpha^2 + m_H^2 s_\alpha^2}{2w^2},$$

$$\kappa = \frac{(m_S^2 - m_H^2) s_{2\alpha}}{2vw}.$$
(57)

The mass of the light scalar field is kept fixed to $m_H = 125.09$ GeV. We then evaluate $\mathcal{R}_{VV}^{\text{SSM}}$ as a function of s_{α} for different values of m_S and w. It has to be considered that the maximally allowed ranges for $|s_{\alpha}|$ depend on the assumed singlet mass [12] and have been derived considering W boson mass measurement, electroweak precision observables tested via the oblique parameters S, T and U, perturbativity of the RG-evolved coupling λ evaluated for the exemplary choice w/v = 10, perturbative unitarity, direct LHC searches and Higgs signal strength measurement [34]. On the other hand, perturbative unitarity poses a lower limit on the ratio w/vwhich, again, depends on the singlet mass and s_{α} . We summarize such informations on Table 3, extracted from Table I of [34], where we report the values of m_S considered in our numerical analysis as well as the ranges of $|s_{\alpha}|$ and the corresponding w_{min} .

The numerical results for $\mathcal{R}_{VV}^{\text{SSM}}$ as a function of s_{α} are reported in Fig. 2; in the upper panels we show the case V=W (where for simplicity we fixed $q_{\gamma}=q_{\gamma}^{\text{max}}$) whereas in the lower ones V=Z. For both cases we considered four possible values of m_S : a low mass region with $m_S=200,300 \text{ GeV}$ (plots on the left) and a high mass region with $m_S=900,1000 \text{ GeV}$ (plots on the right). In order to roughly



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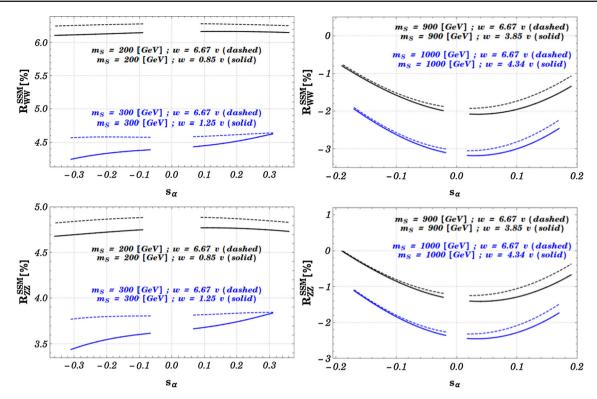


Fig. 2 $\mathcal{R}_{VV}^{\rm SSM}$ as a function of s_{α} , for different values of m_S (and the corresponding vev w). The range of s_{α} is the one deduced from Table 3. $\mathcal{R}_{WW}^{\rm SSM}$ is computed with $q_{\gamma} = q_{\gamma}^{\rm max}$

analyze the dependence on w, in the same plots we also show $\mathcal{R}_{VV}^{\rm SSM}$ computed for two different values of the singlet vev w: the smallest one (solid lines) is chosen according to the minimum reported in Table 3 while the largest is kept fixed at w=6.67v (dashed lines), which is the value used in [34] to determine the allowed intervals of s_{α} and, according to Table 3, valid for every m_S . First of all, we clearly see that the dependence on w is not dramatic for every value of m_S . In fact, the absolute differences between the solid and the dashed lines amount to a maximum of $\sim 0.3\%$ for both $\mathcal{R}_{ZZ}^{\rm SSM}$ and $\mathcal{R}_{WW}^{\rm SSM}$ when $m_S=300$ GeV, $s_{\alpha}\sim -0.3$ and $m_S=900$ GeV, $s_{\alpha}\sim 0.2$. The reason for such a dependence is simply due to the fact that κ (defined in Eq. (57) and entering in $\mathcal{R}_{VV}^{\rm SSM}$) are inversely proportional to w but grow with the mixing angle.

We also observe a different behavior with respect to $sign\ (s_{\alpha});$ in particular, the ratios \mathcal{R}_{VV}^{SSM} are weakly dependent on s_{α} in the low mass range while for the high mass region they increase or decrease proportionally to the mixing angle, especially for $s_{\alpha}<0$. To be more quantitative: the two plots on the left-hand side of Fig. 2 show a maximum variation of each \mathcal{R}_{VV}^{SSM} with respect to negative s_{α} of $\sim 0.2\%$, while for those on the right-hand side such a variation becomes ~ 1.2 and $\sim 1.3\%$ (for V=W and Z, respectively). The high mass region also shows a variation with respect to the mixing angle in the region of $s_{\alpha}>0$ amount-

ing to a maximum of $\sim 0.7\%$ for both ratios. The reason of this different behavior with respect to $sign(s_{\alpha})$ has to be ascribed to those diagrams which contain odd powers of the coupling κ which, according to Eq. (57), grows with m_S^2 and whose sign is only determined by $sign(s_{\alpha})$ (for $m_S > m_H$, as it is the case in this paper). Typical Feynman diagrams with such a structure and that contribute to the mixing angle dependence of \mathcal{R}_{VV}^{SSM} are depicted in Fig. 3. Neglecting the loop integrals for simplicity, the couplings evaluated up to $\mathcal{O}(v^2/w^2)$ (we used the approximate expressions in Eq. (7)) are the following:

(SSH)
$$\sim \kappa v$$
, $(HV\eta^i) \propto \frac{m_V}{v}$, $(SV\eta^i) \propto s_\alpha \frac{m_V}{v} \sim \frac{\kappa m_V}{2\rho w}$,

which in turn imply an overall dependence given by:

$$(SSH) \times (SV\eta^i) \times (HV\eta^i) \propto \frac{\kappa^2 m_V^2}{2\rho w} \sim \rho_V \kappa,$$
 (58)

where $i = (3, \pm)$ for $V = Z, W^{\pm}$, respectively.

In the rest of this section we will scrutinize more in detail the dependence of $\mathcal{R}_{VV}^{\text{SSM}}$ on the singlet mass, separating the cases of positive and negative s_{α} .



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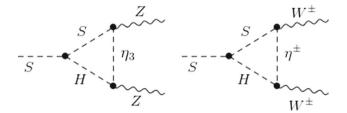
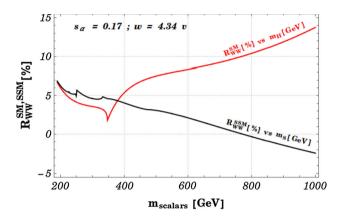


Fig. 3 Examples of the Feynman diagrams contributing to the mixing angle dependence of \mathcal{R}_{VV}^{SSM}



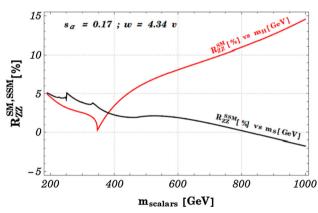


Fig. 4 \mathcal{R}_{VV}^{SM} (red line) and \mathcal{R}_{VV}^{SSM} (black line) as a function of the scalar mass $m_{scalars}$. In the case of the SSM, we fixed ($s_{\alpha}=0.17,\ w=4.34\ v$)

Notice that, increasing m_S , the range of the $|s_{\alpha}|$ taken into account is restricted to the following interval: $|s_{\alpha}| \in [|s_{\alpha}|_{\min}^{m_S=200 \text{ GeV}}, |s_{\alpha}|_{\max}^{m_S=1000 \text{ GeV}}] = [0.09, 0.17]$, which is valid for every choice of m_S (see Table 3).

Case $s_{\alpha} > 0$ In Fig. 4 we show the behavior of $\mathcal{R}_{VV}^{\text{SSM}}$ (V = W on the left, V = Z on the right) as a function of m_S for a fixed $s_{\alpha} = 0.17$ and two different values of w, namely w = (4.34, 6.67) v. For the sake of comparison, we also computed the same ratio $\mathcal{R}_{VV}^{\text{SM}} = [(\Gamma_{VV}^{\text{NLO}}/\Gamma_{VV}^{\text{LO}}) - 1]$ in the SM (red line) leaving the Higgs mass as a free parameter (in practice, the SM with a heavy Higgs [45]). In the plots, on the common x-axis we use the label m_{scalars} to indicate either m_H or m_S .

We observe two main differences.

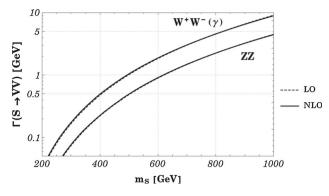


Fig. 5 Leading order (dashed line) and next-to-leading order (solid line) results for $\Gamma(S \to VV)$ ($s_{\alpha} = 0.17$, w = 4.34 v). In the case of V = W the photon momenta is fixed to $q_{\gamma} = q_{\gamma}^{\text{max}}$

The first one is the finite peak at $m_S = 2m_H$ (which corresponds to fix in the loop integrals: $k^2 = m_S^2 = 4m_H^2$) in \mathcal{R}_{VV}^{SSM} due to the new coupling SHH which is obviously absent in the SM. The second is the different behavior in the heavy scalar mass region. This is mainly due to the new scalar contributions arising from the coefficient \mathcal{D}_V (see [45] for an explicit evaluation in the SM). For example, setting the mass of the heavy scalars to $m_{\text{scalars}} = 10^3$ GeV and considering V = Z, we have:

$$\{\delta V_Z^{\mathcal{D}}\}^{\text{SM}} \sim (3.31 - 0.16 \,\lambda) \times 10^{-5} > 0 \,,$$

whose positivity is determined by the fact that $\lambda = m_{\text{scalars}}^2/2v^2$. Instead, in the case of the SSM, we get:

$$\{\delta \mathcal{V}_Z^{\mathcal{D}}\}^{\text{SSM}} \sim (4.97 - 2.27\lambda - 2.07 \ \rho - 13.95 \ \kappa) \times 10^{-5} < 0$$

because λ , ρ and κ are all positive parameters for $s_{\alpha} = 0.17$ and w = (4.34, 6.67) v, see Eq. (57).

The black curves show that \mathcal{R}_{VV}^{SSM} exhibits positive (negative) corrections for values of the scalar mass below (above) $m_{\text{scalars}} \sim 750$ GeV. Additional features concerning these curves are visible when $m_S \sim 230$, 330 GeV (new kinks shortly below m_H and m_t) and $m_S \sim 470$ GeV. These behaviors arise from the finite parts of the loop integrals $\mathcal{B}_0(k^2, M^2, M^2)$, with $M = m_Z, m_H, m_t$, which appear in $\delta \alpha^{\text{ios}}$ defined at the singlet scalar momentum $k^{*2} = (m_S^2 + m_H^2)/2$. As a consequence, the kinks are dictated by the condition,

$$k^{*2} = \frac{m_H^2 + m_S^2}{2} = 4M^2 \implies m_S = \sqrt{8M^2 - m_H^2}$$

 $\sim \{230, 330, 470 \,\text{GeV}\}.$ (59)

Finally, in Fig. 5 we summarize our results for the decay widths $\Gamma(S \to ZZ)$ and $\Gamma(S \to W^+W^-(\gamma))$ as a function of m_S for the selected values $s_\alpha = 0.17$, w = 4.34 v. As expected from our previous considerations, the NLO results (solid line) are above the LO behavior (dashed line) in the



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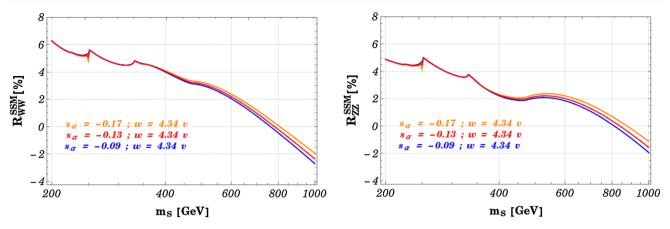


Fig. 6 Ratio $\mathcal{R}_{VV}^{\text{SSM}}$ as a function of the singlet mass m_S for the three fixed values $s_{\alpha} = -0.09, -0.13$ and -0.17

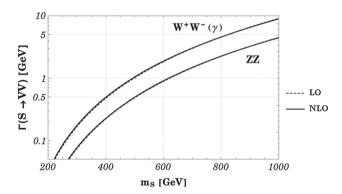


Fig. 7 The same as Fig. 5 but for $s_{\alpha} = -0.17$

small mass region but becomes smaller in the region of larger masses.

Case $s_{\alpha} < 0$ As it was shown in Fig. 2, the ratio $\mathcal{R}_{VV}^{\text{SSM}}$ in the high mass region depends much more on the sign and on the assumed value of $\sin \alpha$ than the case of low-mass, and the variation with $sign(s_{\alpha})$ is more evident when $s_{\alpha} < 0$. To study more in detail the region of negative s_{α} , in Fig. 6 we show $\mathcal{R}_{VV}^{\text{SSM}}$ (V = W on the left, V = Z on the right) as a function of the singlet mass m_S for three fixed values of s_{α} , namely $s_{\alpha} = -0.09, -0.17$ (which are the two extremes of the considered range) and its central value $s_{\alpha} = -0.13$.

The dependence on the mixing starts to be significant for $m_S \gtrsim 400$ GeV while it can be neglected for smaller masses, for both cases V=W, Z. In addition, \mathcal{R}_{VV}^{SSM} becomes negative when the scalar mass is roughly larger than 800 GeV, as it was the case for $s_\alpha > 0$, see Fig. 4. As before, in Fig. 7 we summarize our results for the decay widths $\Gamma(S \to ZZ)$ and $\Gamma(S \to W^+W^-(\gamma))$ as a function of m_S for the selected values $s_\alpha = -0.17$ and $w = 4.34 \, v$. Also in this case the NLO results (solid line) are very similar to the LO lines (dashed line) and tend to be larger in the region of small masses.

6 Conclusions

In this paper we have studied in details an extension of the SM which involves the presence of a new real scalar field s^0 , singlet under the SM gauge group. Its main effect is to mix with the SM scalar doublet ϕ via a quartic interaction of the form $\kappa (\phi^{\dagger} \phi)(s^0)^2$, giving rise to two mass eigenstates that we call H (the lighter) and S (the heaviest). We have limited our interests here to the study of the decay rates of S to a pair of vector gauge bosons V; as far as we know, the amplitudes of such vertices could be extracted by the one-loop self energies and vertex corrections in Ref. [35] (relevant to the Higgs scalar decays to gauge bosons) but the one-loop corrections to heavy scalar decays into gauge bosons have not been computed explicitly before. In the mass range analysed in this paper, $200 \le m_S \le 1000$ GeV which corresponds to mixing angles in the range $|s_{\alpha}| \in [0.09, 0.17]$ and to singlet vev values $w \ge 4.34v$, the decay $S \to VV$ is kinematically accessible and we estimated that the oneloop corrections to ZZ and $W^+W^-(\gamma)$ channels can be as large as $\mathcal{O}(5\%)$ and $\mathcal{O}(7\%)$ for $m_S \sim 200$ GeV, respectively. Interestingly enough, the sign of the NLO corrections is not fixed a priori: for $m_S \lesssim 700$ GeV, the quantity $\mathcal{R}_{VV}^{\text{SSM}} =$ $[(\Gamma_{VV}^{\rm NLO}/\Gamma_{VV}^{\rm LO})-1]$ is positive for every values of α while for larger masses $\mathcal{R}_{VV}^{\rm SSM}$ can also become negative (the precise turning point depends on the assumed values of α and w). Regarding the dependence on α , $\mathcal{R}_{VV}^{\text{SSM}}$ exhibits a different behavior with respect to the sign of the mixing angle: for $\sin \alpha > 0$ and fixed m_S , it is almost independent on α in the low mass range while a stronger dependence is visible in the high mass range especially for $\sin \alpha < 0$. This dependence is confined in the high mass region for masses somehow larger than 400 GeV. We have also studied the dependence of $\mathcal{R}_{VV}^{\text{SSM}}$ on the singlet vev w; we found that its maximum variation is practically the same for every m_S value; to give an example, when $m_S = 900$ GeV and $s_\alpha \sim 0.2$, the absolute difference



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 $|\mathcal{R}_{VV}^{SSM}(6.67v) - \mathcal{R}_{VV}^{SSM}(4.34v)|\%$ amounts to $\sim 0.3\%$ for

Acknowledgements We are strongly indebted with Giuseppe Degrassi, Roberto Franceschini and Ramona Grober for illuminating discussions and Andrea Di Iura for useful suggestions on technical aspects of the software used in our computations.

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A Feynman rules of the scalar sector

We give the Feynman rules of the trilinear and quartic vertices arising from scalar potential $V_{\rm np}(\phi, s^0)$, shown in Eq. (2). When we expand $V_{\rm np}(\phi, s^0)$ in terms of the physical fields, the scalar trilinear and quartic couplings are generally expressed as

$$V_{np}(\phi, s^{0}) = \dots + C_{S_{1}S_{2}S_{3}} S_{1}S_{2}S_{3} + C_{S_{1}S_{2}S_{3}S_{4}} S_{1}S_{2}S_{3}S_{4},$$
(60)

where S can be H, S, η^3, η^{\pm} and the coefficients $C_{S_1S_2S_3}$, $C_{S_1S_2S_3S_4}$ are given by:

$$\begin{split} \mathbf{C}_{HHH} &= -3ic_{\alpha}s_{\alpha}(s_{\alpha}v - c_{\alpha}w)\kappa - 6i(c_{\alpha}^{3}v\lambda - s_{\alpha}^{3}w\rho), \\ \mathbf{C}_{SSS} &= -3ic_{\alpha}s_{\alpha}(c_{\alpha}v + s_{\alpha}w)\kappa - 6i(s_{\alpha}^{3}v\lambda + c_{\alpha}^{3}w\rho), \\ \mathbf{C}_{HSS} &= -i[c_{\alpha}v(c_{\alpha}^{2} - 2s_{\alpha}^{2}) + s_{\alpha}w(2c_{\alpha}^{2} - s_{\alpha}^{2})] \\ \kappa - 6i(c_{\alpha}s_{\alpha}^{2}v\lambda - c_{\alpha}^{2}s_{\alpha}w\rho), \\ \mathbf{C}_{HHS} &= -i[s_{\alpha}v(s_{\alpha}^{2} - 2c_{\alpha}^{2}) + c_{\alpha}w(c_{\alpha}^{2} - 2s_{\alpha}^{2})] \\ \kappa - 6i(c_{\alpha}^{2}s_{\alpha}v\lambda - c_{\alpha}s_{\alpha}^{2}w\rho), \\ \mathbf{C}_{H\eta^{3}\eta^{3}} &= -i(2c_{\alpha}v\lambda - s_{\alpha}w\kappa), \\ \mathbf{C}_{H\eta^{+}\eta^{-}} &= -i(2c_{\alpha}v\lambda - s_{\alpha}w\kappa), \end{split}$$

$$\begin{split} \mathbf{C}_{S\eta^3\eta^3} &= -i(c_\alpha w\kappa + 2s_\alpha v\lambda), \\ \mathbf{C}_{S\eta^+\eta^-} &= -i(c_\alpha w\kappa + 2s_\alpha v\lambda), \\ \mathbf{C}_{HHHH} &= -6i(c_\alpha^4\lambda + c_\alpha^2 s_\alpha^2 \kappa + s_\alpha^4 \rho), \\ \mathbf{C}_{SSSS} &= -6i(c_\alpha^4\rho + c_\alpha^2 s_\alpha^2 \kappa + s_\alpha^4 \lambda), \\ \mathbf{C}_{HHSS} &= -i(c_\alpha^4 - 4c_\alpha^2 s_\alpha^2 + s_\alpha^4)\kappa - 6ic_\alpha^2 s_\alpha^2 (\lambda + \rho), \\ \mathbf{C}_{HHHS} &= 3ic_\alpha s_\alpha c_{2\alpha} \kappa - 6i(c_\alpha^3 s_\alpha \lambda - c_\alpha s_\alpha^3 \rho), \\ \mathbf{C}_{HSSS} &= -3ic_\alpha s_\alpha c_{2\alpha} \kappa - 6i(c_\alpha s_\alpha^3 \lambda - c_\alpha^3 s_\alpha \rho), \\ \mathbf{C}_{HH\eta^3\eta^3} &= -i(s_\alpha^2 \kappa + 2c_\alpha^2 \lambda), \\ \mathbf{C}_{HH\eta^3\eta^3} &= -i(s_\alpha^2 \kappa + 2c_\alpha^2 \lambda), \\ \mathbf{C}_{HH\eta^3\eta^3} &= -i(s_\alpha^2 \kappa + 2c_\alpha^2 \lambda), \end{split}$$

$$C_{SS\eta^3\eta^3} = -i(c_{\alpha}^2\kappa + 2s_{\alpha}^2\lambda),$$

$$C_{HSn^3n^3} = -is_{\alpha}c_{\alpha}(2\lambda - \kappa),$$

$$C_{HH\eta^+\eta^-} = -i(s_\alpha^2\kappa + 2c_\alpha^2\lambda),$$

$$C_{SSn^+n^-} = -i(c_\alpha^2 \kappa + 2s_\alpha^2 \lambda),$$

$$C$$
 $= i\pi \circ (2)$

$$C_{HSn^+n^-} = -is_{\alpha}c_{\alpha}(2\lambda - \kappa).$$

The couplings of H and S with the Standard Model ordinary matter are:

$$C_{H\sigma\bar{\sigma}} = c_{\alpha} C_{h\sigma\bar{\sigma}}, \quad C_{HH\sigma\bar{\sigma}} = c_{\alpha}^{2} C_{hh\sigma\bar{\sigma}},$$

 $C_{S\sigma\bar{\sigma}} = s_{\alpha} C_{h\sigma\bar{\sigma}}, \quad C_{SS\sigma\bar{\sigma}} = s_{\alpha}^{2} C_{hh\sigma\bar{\sigma}},$

where $C_{h\sigma\bar{\sigma}}$, $C_{hh\sigma\bar{\sigma}}$ are the couplings of the Standard Model Higgs field (h) with a pair of bosonic or fermionic fields $(\sigma = W, Z, f).$

B Feynman Diagrams for $S \rightarrow VV$

Contribution to the Self Energies

See Figs. 8 and 9.

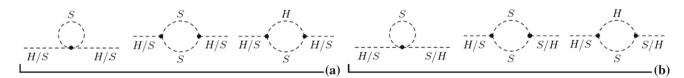


Fig. 8 a S field contributions to the scalar self-energies; b S field contributions to the mixed scalar self-energies

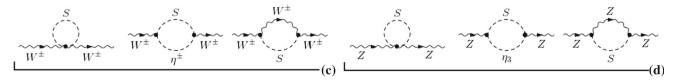


Fig. 9 c S field contributions to the W boson self energies; d S field contributions to the Z boson self energies



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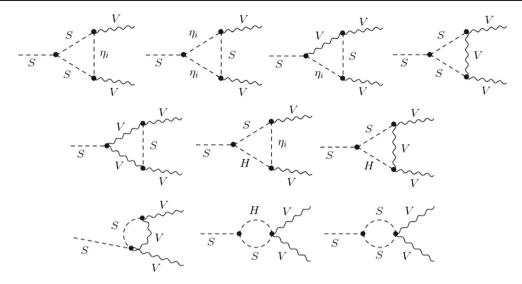


Fig. 10 NLO S field contributions to the SVV vertices

Contribution to the Trilinear Vertices

See Fig. 10.

C: Bremsstrahlung

Here we explicitly report the photon bremsstrahlung contributions to the process $S \to W^+W^-(\gamma)$. As shown in Eq. (53), the total bremsstrahlung decay rate is given by the sum of soft- and hard-photon factors which can be written as,

$$\delta_{W}^{\text{soft}} = \frac{\alpha_{em}}{\pi} \times \left\{ \mathcal{G}(r) \ln \left(\frac{4\Lambda_{\gamma}^{2}}{m_{\gamma}^{2}} \right) + (\mathcal{G}(r) + 1) \left[\frac{\mathcal{N}_{1}}{a_{1}} + \frac{2}{2r - 1} \right] \right\},$$

$$\delta_{W}^{\text{hard}} = \frac{\alpha_{em}}{\pi} \times \left\{ \mathcal{G}(r) \ln \left(\frac{m_{S}^{2}}{4\Lambda_{\gamma}^{2}} \right) + \frac{\mathcal{N}_{2} (\mathcal{G}(r) + 1)}{a_{1}} + \frac{14}{3} \left(1 - \frac{t}{r} \sqrt{1 - \frac{1}{t}} \right) + \frac{1}{\sqrt{3}} \left[\frac{a_{2} - a_{1}}{2r^{2}} \left(2 - \frac{1}{r} \right) + \frac{t}{3r} \sqrt{1 - \frac{1}{t}} \left(1 - \frac{t}{r} \right) \left(2 - \frac{4t - 1}{r} \right) \right] - 2 \ln \left(\frac{1 - (c_{-}d_{+})^{2}}{1 - (c_{-}d_{-})^{2}} \right) + 4 \left(\frac{t}{r} a_{2} - b_{-} \right) \right\},$$
(62)

where,

$$\begin{split} \mathcal{G}(r) &= \left[\, a_1 \, (2 - 1/r) / (\sqrt{1 - 1/r}) \, \right] - 1, \\ \mathcal{N}_1 &= \, \operatorname{Li}_2(c_-^2) + a_1 \, (a_1 - 2b_-) - (\pi^2/6), \\ \mathcal{N}_2 &= \, \operatorname{Li}_2((c_-d_+)^2) - \operatorname{Li}_2((c_-d_-)^2) + \operatorname{Li}_2(c_-)^4 \\ &+ 4a_1 \, (a_1 - b_+) \\ &+ 2a_2 \ln[(1 - (c_-d_+)^2)(1 - (c_-d_-)^2)] - (\pi^2/6), \\ \mathcal{N}_3 &= \, \sqrt{1 - 1/r} \, \Big[\, 1 - (1/r) + (3/4r^2) \, \Big], \\ a_1 &= \, \ln(c_+), \\ a_2 &= \, \ln(d_+), \\ b_{\pm} &= \, \ln(c_+ \pm c_-), \\ c_{\pm} &= \, \sqrt{r} \pm \sqrt{r - 1}, \\ d_{\pm} &= \, \sqrt{t} \pm \sqrt{t - 1}, \end{split}$$

with $r = m_S^2/4m_W^2$ and $t = r(1 - 2q_\gamma/m_S)$. For $q_\gamma = q_\gamma^{\rm max}$, Eq. (62) is reduced to,

$$\delta_W^{\text{hard}} = \frac{\alpha_{em}}{\pi} \left\{ \mathcal{G}(r) \ln \left(\frac{m_S^2}{4\Lambda_\gamma^2} \right) - 4b_- + \frac{14}{3} + \frac{(\mathcal{G}(r) + 1)}{a_1} \left[\text{Li}_2(c_-)^4 - \frac{\pi^2}{6} + 4a_1 (a_1 - b_+) - \frac{2a_1}{4r^2 - 4r + 3} \right] \right\}.$$
(63)

References

 G. Aad et al. (ATLAS Collaboration), Phys. Lett. B 716, 1 (2012). https://doi.org/10.1016/j.physletb.2012.08.020. arXiv:1207.7214 [hep-ex]



Eur. Phys. J. C (2018) 78:270 Page 13 of 13 270

- S. Chatrchyan et al. (CMS Collaboration), Phys. Lett. B 716, 30 (2012). https://doi.org/10.1016/j.physletb.2012.08.021. arXiv:1207.7235 [hep-ex]
- P.W. Higgs, Phys. Rev. Lett. 13, 508 (1964). https://doi.org/10. 1103/PhysRevLett.13.508
- 4. P.W. Higgs, Phys. Lett. **12**, 132 (1964). https://doi.org/10.1016/ 0031-9163(64)91136-9
- F. Englert, R. Brout, Phys. Rev. Lett. 13, 321 (1964). https://doi. org/10.1103/PhysRevLett.13.321
- G.S. Guralnik, C.R. Hagen, T.W.B. Kibble, Phys. Rev. Lett. 13, 585 (1964). https://doi.org/10.1103/PhysRevLett.13.585
- R.M. Schabinger, J.D. Wells, Phys. Rev. D 72, 093007 (2005). https://doi.org/10.1103/PhysRevD.72.093007. arXiv:hep-ph/0509209
- G. M. Pruna, T. Robens, Phys. Rev. D 88(11), 115012 (2013). https://doi.org/10.1103/PhysRevD.88.115012. arXiv:1303.1150 [hep-ph]
- C. Englert, T. Plehn, D. Zerwas, P. M. Zerwas, Phys. Lett. B 703, 298 (2011). https://doi.org/10.1016/j.physletb.2011.08.002. arXiv:1106.3097 [hep-ph]
- B. Batell, S. Gori, L. T. Wang, JHEP 1206, 172 (2012). https://doi. org/10.1007/JHEP06(2012)172. arXiv:1112.5180 [hep-ph]
- D. Buttazzo, F. Sala, A. Tesi, JHEP 1511, 158 (2015). https://doi. org/10.1007/JHEP11(2015)158. arXiv:1505.05488 [hep-ph]
- T. Robens, T. Stefaniak, Eur. Phys. J. C 75, 104 (2015). https://doi. org/10.1140/epjc/s10052-015-3323-y. arXiv:1501.02234 [hep-ph]
- T. Robens, T. Stefaniak, Eur. Phys. J. C 76(5), 268 (2016). https://doi.org/10.1140/epjc/s10052-016-4115-8. arXiv:1601.07880 [hep-ph]
- 14. I.M. Lewis, M. Sullivan. arXiv:1701.08774 [hep-ph]
- S. Dawson, I.M. Lewis, Phys. Rev. D 92(9), 094023 (2015). https://doi.org/10.1103/PhysRevD.92.094023. arXiv:1508.05397 [hep-ph]
- C. Y. Chen, S. Dawson, I. M. Lewis, Phys. Rev. D 91(3), 035015 (2015). https://doi.org/10.1103/PhysRevD.91.035015. arXiv:1410.5488 [hep-ph]
- F. Bezrukov, M. Y. Kalmykov, B. A. Kniehl, M. Shaposhnikov, JHEP 1210, 140 (2012). https://doi.org/10.1007/ JHEP10(2012)140. arXiv:1205.2893 [hep-ph]
- S. Alekhin, A. Djouadi, S. Moch, Phys. Lett. B 716, 214 (2012). https://doi.org/10.1016/j.physletb.2012.08.024. arXiv:1207.0980 [hep-ph]
- D. Buttazzo, G. Degrassi, P. P. Giardino, G. F. Giudice, F. Sala,
 A. Salvio, A. Strumia, JHEP 1312, 089 (2013). https://doi.org/10. 1007/JHEP12(2013)089. arXiv:1307.3536 [hep-ph]
- J. Elias-Miro, J. R. Espinosa, G. F. Giudice, H. M. Lee,
 A. Strumia, JHEP 1206, 031 (2012). https://doi.org/10.1007/ JHEP06(2012)031. arXiv:1203.0237 [hep-ph]
- A. Falkowski, C. Gross, O. Lebedev, JHEP 1505, 057 (2015). https://doi.org/10.1007/JHEP05(2015)057. arXiv:1502.01361 [hep-ph]
- 22. B. Patt, F. Wilczek. arXiv:hep-ph/0605188
- G. Aad et al. (ATLAS Collaboration), Phys. Rev. D 92(1), 012006 (2015). https://doi.org/10.1103/PhysRevD.92.012006 arXiv:1412.2641 [hep-ex]
- V. Khachatryan et al. (CMS Collaboration), JHEP 1510,
 144 (2015). https://doi.org/10.1007/JHEP10(2015)144
 arXiv:1504.00936 [hep-ex]

- G. Aad et al. (ATLAS Collaboration), Eur. Phys. J. C 75(9),
 412 (2015). https://doi.org/10.1140/epjc/s10052-015-3628-x
 arXiv:1506.00285 [hep-ex]
- G. Aad et al. (ATLAS Collaboration), Phys. Rev. Lett. 114(8), 081802 (2015).https://doi.org/10.1103/PhysRevLett.114.081802 arXiv:1406.5053 [hep-ex]
- V. Khachatryan et al. (CMS Collaboration), Phys. Lett. B 749, 560 (2015). https://doi.org/10.1016/j.physletb.2015.08.047 arXiv:1503.04114 [hep-ex]
- V. Barger, P. Langacker, M. McCaskey, M.J. Ramsey-Musolf, G. Shaughnessy, Phys. Rev. D 77, 035005 (2008). https://doi.org/10.1103/PhysRevD.77.035005. arXiv:0706.4311 [hep-ph]
- W.L. Guo, Y.L. Wu, JHEP 1010, 083 (2010). https://doi.org/10. 1007/JHEP10(2010)083. arXiv:1006.2518 [hep-ph]
- A. Biswas, D. Majumdar, Pramana 80, 539 (2013). https://doi.org/ 10.1007/s12043-012-0478-z. arXiv:1102.3024 [hep-ph]
- J.A. Casas, D.G. Cerdeño, J.M. Moreno, J. Quilis. arXiv:1701.08134 [hep-ph]
- D. López-Val, T. Robens, Phys. Rev. D 90, 114018 (2014). https://doi.org/10.1103/PhysRevD.90.114018. arXiv:1406.1043 [hep-ph]
- 33. G. Chalons, D. Lopez-Val, T. Robens, T. Stefaniak, arXiv:1611.03007 [hep-ph]
- F. Bojarski, G. Chalons, D. Lopez-Val, T. Robens, JHEP 1602, 147 (2016). https://doi.org/10.1007/JHEP02(2016)147. arXiv:1511.08120 [hep-ph]
- S. Kanemura, M. Kikuchi, K. Yagyu, Nucl. Phys. B 907, 286 (2016). https://doi.org/10.1016/j.nuclphysb.2016.04.005. arXiv:1511.06211 [hep-ph]
- S. Kanemura, M. Kikuchi, K. Yagyu, Nucl. Phys. B 917, 154 (2017). https://doi.org/10.1016/j.nuclphysb.2017.02.004. arXiv:1608.01582 [hep-ph]
- S. Kanemura, M. Kikuchi, K. Sakurai, K. Yagyu, arXiv:1705.05399 [hep-ph]
- A. Denner, Fortsch. Phys. 41, 307 (1993). https://doi.org/10.1002/ prop.2190410402. arXiv:0709.1075 [hep-ph]
- J.R. Espinosa, Y. Yamada, Phys. Rev. D 67, 036003 (2003). https://doi.org/10.1103/PhysRevD.67.036003. [hep-ph/0207351]
- D. Binosi, J. Papavassiliou, Phys. Rept. 479, 1 (2009). https://doi. org/10.1016/j.physrep.2009.05.001. arXiv:0909.2536 [hep-ph]
- S. Kanemura, M. Kikuchi, K. Sakurai, K. Yagyu, Phys. Rev. D 96(3), 035014 (2017). https://doi.org/10.1103/PhysRevD.96. 035014. arXiv:1705.05399 [hep-ph]
- A. Alloul, N.D. Christensen, C. Degrande, C. Duhr, B. Fuks, Comput. Phys. Commun. 185, 2250 (2014). https://doi.org/10.1016/j.cpc.2014.04.012. arXiv:1310.1921 [hep-ph]
- 43. T. Hahn, Comput. Phys. Commun. **140**, 418 (2001). https://doi.org/10.1016/S0010-4655(01)00290-9. [hep-ph/0012260]
- H.H. Patel, Comput. Phys. Commun. 197, 276 (2015). https://doi. org/10.1016/j.cpc.2015.08.017. arXiv:1503.01469 [hep-ph]
- B.A. Kniehl, Nucl. Phys. B 352, 1 (1991). https://doi.org/10.1016/ 0550-3213(91)90126-I
- A. Sirlin, Phys. Rev. D 22, 971 (1980). https://doi.org/10.1103/ PhysRevD.22.971
- B.A. Kniehl, Nucl. Phys. B 357, 439 (1991). https://doi.org/10. 1016/0550-3213(91)90476-E

