

Physics 105B Lecture Notes, Fall 2024

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The class textbook is *Classical Mechanics of Particles and Systems, 5th edition* by Marion & Thornton. These notes are a supplement for the textbook, not a replacement, so I won't cover everything from the course.

- For midterm 1, study chapters 4, 9, 10
- For midterm 2, study chapters 4, 9, 10, 11, 12
- For the final, study chapters 4, 9, 10, 11, 12, ...

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1 Review of Lagrangian mechanics (physics 105A)

1.1 Calculus of variations

Use catenary and brachistocline as examples

1.2 Least action principle

1.3 Generalized momenta

1.4 Lagrange multipliers, forces from constraints

2 Phase diagrams

For lots of systems, the state can be described by generalized positions (q_1, q_2, \dots) and generalized velocities $(\dot{q}_1, \dot{q}_2, \dots)$. For example, the state of a pendulum at any time can be described by the vector

$$\begin{bmatrix} \theta \\ \dot{\theta} \end{bmatrix},$$

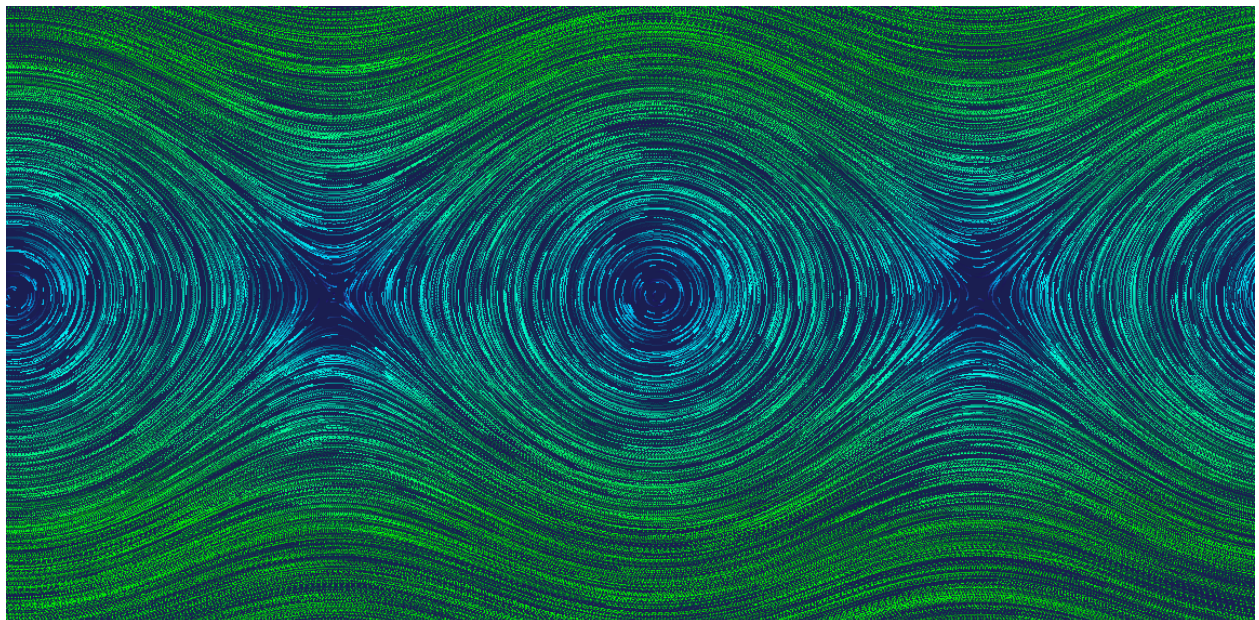
and if we know the state, we can calculate the rate of change of the state. That is,

$$\begin{bmatrix} \dot{\theta} \\ \ddot{\theta} \end{bmatrix}$$

is a function of the state. We can visualize this as a vector field, where $x = \theta$ and $y = \dot{\theta}$. One amazing tool for animating vector fields is <https://anvaka.github.io/fieldplay/>

For the pendulum example, copy this code:

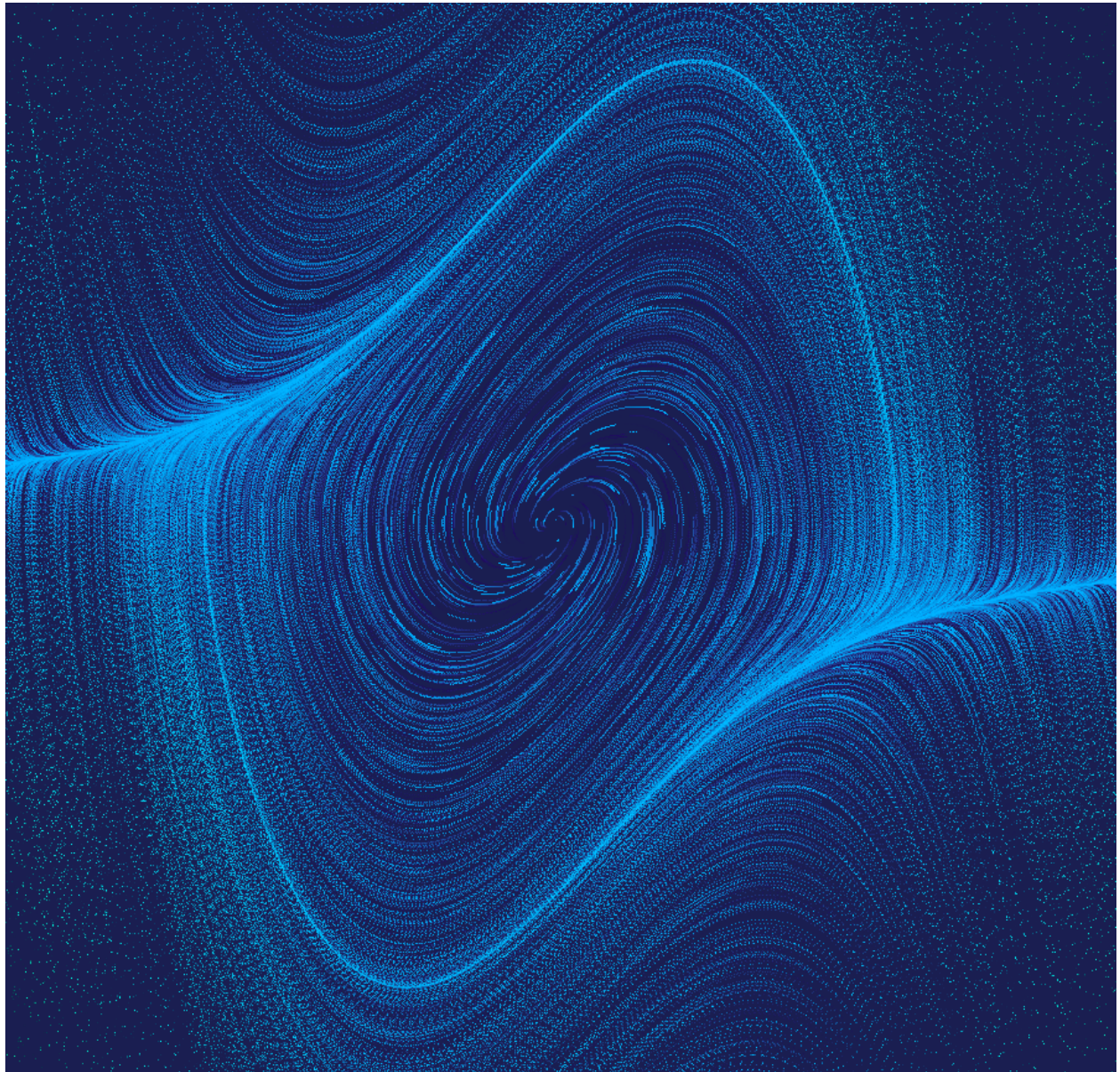
```
// p.x and p.y are current coordinates
// v.x and v.y is a velocity at point p
vec2 get_velocity(vec2 p) {
    vec2 v = vec2(0., 0.);
    v.x = p.y;
    v.y = cos(p.x);
    return v;
}
```



For a Van der Pol oscillator, copy in this code, and experiment with changing μ (mu):


```
float mu = 1.;
vec2 get_velocity(vec2 p) {
    return vec2(p.y, mu*(1.-p.x*p.x)*p.y-p.x);
}
```

Here is a Van der Pol oscillator with $\mu = 1$:



3 Rotating reference frames

Chapter 10 in the textbook. Copy example 10.5 (Foucault pendulum) to these notes

4 Moment of inertia tensor

If a system of particles is moving, and some point which we will call the origin is moving with velocity V , and r_α is the vector from the origin to some particle α , and that particle is rotating about the origin with angular velocity ω , then the velocity of that particle is $v_\alpha = V + \omega \times r_\alpha$. The kinetic energy of particle α is $m_\alpha v_\alpha^2/2$, and if you sum that over each α , then we will see that the total kinetic energy is $T_{trans} + T_{rot}$, where $T_{trans} = (\sum_\alpha m_\alpha) V^2/2$ and $T_{rot} = \sum_\alpha m_\alpha (\omega \times r_\alpha)^2/2$. It makes sense to choose the origin to be the center of the mass, which minimizes T_{rot} (assuming ω is the same for each particle). The rotational kinetic energy can also be written as

$$T_{rot} = \frac{1}{2} \sum_{i,j} I_{i,j} \omega_i \omega_j$$

where $i, j \in \{x, y, z\}$ and $I_{i,j}$ is the inertia tensor, defined as

$$I_{i,j} := \sum_\alpha m_\alpha \left(\left(\delta_{i,j} \sum_k x_{\alpha,k}^2 \right) - x_{\alpha,i} x_{\alpha,j} \right)$$

The diagonal elements of this matrix (I know it's a tensor, but I'm gonna call it a matrix anyway) are called moments of inertia, and the off-diagonal elements are called products of inertia.

Note that if a mass distribution just lies along the x axis, then $I_{y,y} = I_{z,z} > 0$ and the other 7 entries are all zero, which makes sense because that mass distribution is hardest to rotate about the y axis and about the z axis. If the formula for the off-diagonal entries resembles the formula for Pearson's correlation coefficient, which makes sense because ignoring some normalization factors, r measures how much data lies along the line $y = \pm x$, and the products of inertia also measure how much mass lies along diagonal lines like $y = x, z = 0$ and $z = y, x = 0$ and $x = z, y = 0$.

Also note that I is a symmetric matrix, so we can choose an orthonormal basis of eigenvectors (with real eigenvalues), and I is diagonal in that basis. Those eigenvectors are called principle axes of inertia, and the eigenvalues are called the principle moments of inertia. If all three principle moments of inertia are the same, we call the body a spherical top, if two are the same but the other is different, we call it a symmetric top, and if all three are different, we call it an asymmetric top.

If we're working with a continuous mass distribution, the formula becomes

$$I_{i,j} = \int_{r \in \mathbb{R}^3} \rho(r) \left(\delta_{i,j} \sum_k x_k^2 - x_i x_j \right) d^3 r.$$

If we are working in a reference frame where a is the position of the center of mass and the position of each particle is $R = a + r$ (and the components of that are $X_i = a_i + x_i$), and we let $M := \sum_\alpha m_\alpha$ denote the total mass, then the inertia tensor in the center of mass frame is

$$I_{i,j} = J_{i,j} - M (a^2 \delta_{i,j} - a_i a_j),$$

where J is the inertia tensor in the lab frame. In fact, we don't need a to be the center of mass, we just need J to be the original inertia tensor and I to be the inertia tensor in the frame centered at a . This equation is called Steiner's parallel-axis theorem. It clearly shows that rotational inertia about every axis is minimized iff you are in the COM frame.

4.1 Angular momentum

The momentum of each particle in our system is $p_\alpha = m_\alpha \omega \times r_\alpha$, so the angular momentum of the whole system is $L := \sum_\alpha r_\alpha \times p = \sum_\alpha m_\alpha r_\alpha \times (\omega \times r_\alpha)$. Using the vector identity $A \times (B \times A) = A^2 B - A(A \cdot B)$, this becomes

$$L = \sum_\alpha m_\alpha (r_\alpha^2 \omega - r_\alpha (r_\alpha \cdot \omega)).$$

Each component of L can be written as

$$L_i = \sum_j I_{i,j} \omega_j.$$

so $L = I \cdot \omega$.

The time derivative of L is the net torque on the body, $\dot{L} = N := \sum \tau$. The rotational kinetic energy is $T_{rot} = \omega \cdot L/2$.

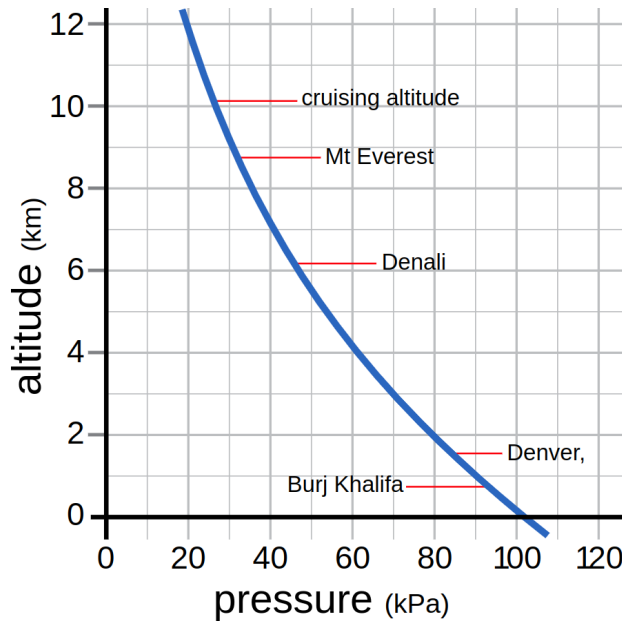
4.2 Stress tensor

A lot of your intuition for the moment of inertial tensor can be applied to the stress tensor σ , which is another contravariant, symmetric, second-order tensor. It represents the forces on each point in a material. Thinking of it as a 3 by 3 matrix, the diagonal elements represent the tensile stress along each axis. So for example, $\sigma_{1,1}$ would be positive if the material is being stretched along the x axis at that point, and negative if it is being squished along the x axis at that point.

The off-diagonal elements of σ represent shear forces. For example, if $\sigma_{1,2} = \sigma_{2,1} > 0$, then there is some shearing force in the x, y plane. Such a force could also be treated as a pulling force in the direction of the unit vector $\pm \frac{1}{\sqrt{2}}(\hat{x} + \hat{y})$. By treating all of the shearing forces as pulling or squishing along some axis other than \hat{x} , \hat{y} , or \hat{z} , we can describe the stress tensor with just the tensile stress along those 3 axes, which we can ensure are orthogonal, because σ is symmetric. This is equivalent to diagonalizing the stress tensor.

To intuitively understand these off-diagonal elements, imagine a square piece of sheet metal lying in the x, y plane, with one edge flushed to the x axis and another flushed to the y axis. Then apply a shear force $F \propto y\hat{x}$ to it, which points along the x axis and is proportional to the y coordinate. This force acts to stretch the square into a rhombus, effectively pulling (that is, applying tensile stress) in the $\frac{1}{\sqrt{2}}(\hat{x} + \hat{y})$ direction.

Stress has units of force per area. One Pascal is a Newton per meter squared, and one atmosphere is about 101.3 kilopascals. If the stress tensor is some scalar $\sigma_{1,1}$ times the identity tensor, then we can say that the pressure is $-\sigma_{1,1}$. Air pressure at sea level is very close to 1 atm, and air pressure pretty much decays exponentially with altitude.



5 Euler angles

For this section, I will call my axes x_1 , x_2 , and x_3 instead of x , y , and z .

Any rotation in 3D space can be decomposed into these three steps (in order):

- Rotate by an angle ϕ around the x'_3 axis. Note that this will move the x'_1 and x'_2 axes. The axes of the new system are now called x''_1, x''_2, x''_3 .

- Rotate by an angle θ about the new x_1'' axis (which will change the directions of the x_2'' and x_3'' axes). The axes of the system after this rotation are called x_1''', x_2''', x_3''' .
- Rotate by an angle ψ about the newest x_3''' axis. After all 3 of these rotations are done, call the transformed axes x_1, x_2, x_3 .

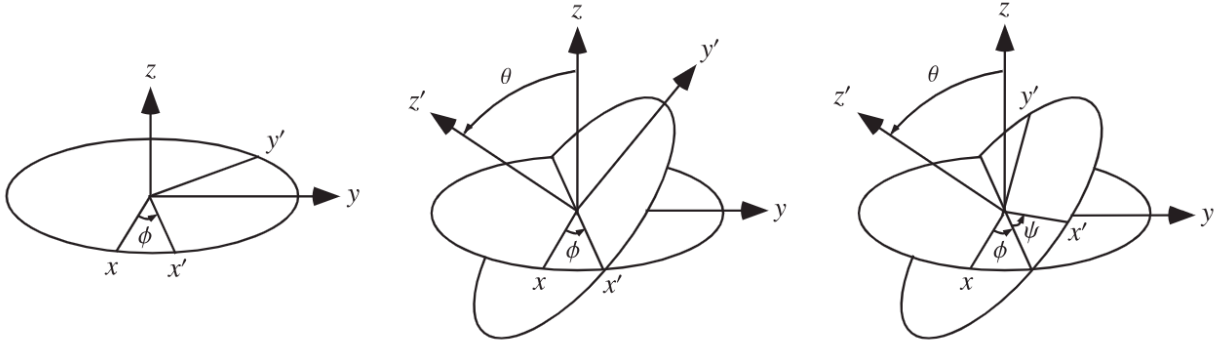
Call those rotations λ_ϕ , λ_θ , and λ_ψ , respectively. Their composition, $\lambda = \lambda_\psi \lambda_\theta \lambda_\phi$, will map a vector x' to a vector x .

$$\lambda_\phi = \begin{bmatrix} \cos \phi & \sin \phi & 0 \\ -\sin \phi & \cos \phi & 0 \\ 0 & 0 & 1 \end{bmatrix}, \lambda_\theta = \begin{bmatrix} 1 & 0 & 0 \\ 0 & \cos \theta & \sin \theta \\ 0 & -\sin \theta & \cos \theta \end{bmatrix}, \lambda_\psi = \begin{bmatrix} \cos \psi & \sin \psi & 0 \\ -\sin \psi & \cos \psi & 0 \\ 0 & 0 & 1 \end{bmatrix}.$$

$$x = \lambda_\psi x''', x''' = \lambda_\theta x'', x'' = \lambda_\phi x'.$$

If we define $\omega_\phi = \dot{\phi}$, $\omega_\theta = \dot{\theta}$, $\omega_\psi = \dot{\psi}$, then observe that $\dot{\phi}$ points along the fixed x_3' axis, $\dot{\theta}$ points along the line of nodes – that is, the line formed by the intersection of the x_1, x_2 plane and the x_1', x_2' plane – and $\dot{\psi}$ points along body's x_3 axis.

The labels in this picture do not match the variable names I am using, but it illustrates the idea well:



$$\begin{aligned} \omega_1 &= \dot{\phi}_1 + \dot{\theta}_1 + \dot{\psi}_1 = \dot{\phi} \sin \theta \sin \psi + \dot{\theta} \cos \psi \\ \omega_2 &= \dot{\phi}_2 + \dot{\theta}_2 + \dot{\psi}_2 = \dot{\phi} \sin \theta \cos \psi - \dot{\theta} \sin \psi \\ \omega_3 &= \dot{\phi}_3 + \dot{\theta}_3 + \dot{\psi}_3 = \dot{\phi} \cos \theta + \dot{\psi}. \end{aligned}$$

If we let the body axes x_1, x_2, x_3 be the principal axes of a rigid body, then treating the Eulerian angles ϕ, θ, ψ as generalized coordinates and plugging them into the Euler-Lagrange equation gives us the Euler equations:

$$\begin{aligned} (I_2 - I_3)\omega_2\omega_3 - I_1\dot{\omega}_1 &= 0 \\ (I_3 - I_1)\omega_3\omega_1 - I_2\dot{\omega}_2 &= 0 \\ (I_1 - I_2)\omega_1\omega_2 - I_3\dot{\omega}_3 &= 0. \end{aligned}$$

These equations are useful for describing precession.

6 Coupled oscillators

If you have a bunch of atoms in a crystal, and you look at a one dimensional piece of it (along the right axis), it will look like a bunch of atoms in a line, each one bonded to its two neighbors. We can treat this as a bunch of masses joined by massless springs. This system is not chaotic, because the differential equation for the positions of each mass is linear. Instead of representing the state of the system by describing the position of each atom, we can change basis, and represent it in terms of normal modes, The dimension of

the state space doesn't change – if we had n degrees of freedom before changing basis (we could say there are n/d masses which are free to move in d dimensions), there will be n normal modes.

If we have two masses of mass M joined by a spring with spring constant κ_{12} , and each mass is joined to a wall by a spring of mass κ , then the equations of motion look something like:

$$\begin{aligned} M\ddot{x}_1 + (\kappa + \kappa_{12})x_1 - \kappa_{12}x_2 &= 0 \\ M\ddot{x}_2 + (\kappa + \kappa_{12})x_2 - \kappa_{12}x_1 &= 0. \end{aligned}$$

We can guess that the solutions will be

$$\begin{aligned} x_1(t) &= B_1 e^{i\omega t} \\ x_2(t) &= B_2 e^{i\omega t}. \end{aligned}$$

for some complex numbers B_1, B_2 . When we solve for ω , we get

$$\omega = \sqrt{\frac{\kappa + \kappa_{12} \pm \kappa_{12}}{M}}$$

which means the characteristic frequencies are

$$\omega_1 = \sqrt{\frac{\kappa + 2\kappa_{12}}{M}}, \omega_2 = \sqrt{\frac{\kappa}{M}}.$$

Since there are two characteristic frequencies, the number of constants in the general solution has doubled. There exists η_1, η_2 which are linear combinations of x_1 and x_2 (in this case, $\eta_1 = x_1 - x_2$ and $\eta_2 = x_1 + x_2$) such that

$$M\ddot{\eta}_i + \kappa_i \eta_i = 0,$$

where κ_i is some frequency that corresponds to the normal mode η_i . By finding these normal modes, we have uncoupled our equations.

When we have several coupled oscillators, the equations of motion will have the form

$$\sum_j (A_{j,k} q_j + m_{j,k} \ddot{q}_j) = 0$$

where $A_{j,k}$ and $m_{j,k}$ are the constants such that

$$\begin{aligned} T &= \frac{1}{2} \sum_{j,k} m_{j,k} \dot{q}_j \dot{q}_k \\ U &= \frac{1}{2} \sum_{j,k} A_{j,k} q_j q_k. \end{aligned}$$

This becomes

$$\sum_j (A_{j,k} - \omega^2 m_{j,k}) a_j = 0$$

when we guess that $q_j(t)$ is an imaginary exponential function of t . Then the normal modes are the eigenvectors of the matrix $A_{j,k}/m_{j,k}$ (that's elementwise division) and the eigenvalues ω are the characteristic frequencies. Once again, we have a symmetric matrix, so the eigenvectors can be made orthonormal.

7 Waves on a 1D string

8 Special relativity

Note. When studying special relativity, we use the Einstein dot notation for time derivatives ($\dot{x} = \frac{\partial}{\partial t}x$), and the apostrophe (e.g. “x prime”) to mean a change of coordinate system. For example, if we have an event whose coordinates are (x, y, z, t) and we apply a Lorentz transform so the new coordinates are (x', y', z', t') , then those apostrophes do not mean to take the time derivative.

9 Final review

The final exam will be almost entirely on chapter 13 (continuous systems & waves) and chapter 14 (special theory of relativity). There may be some questions on the content from midterm 2 (moment of inertia tensor & coupled oscillators), but we do not need to review content from midterm 1.

It is definitely worthwhile to review the discussion section notes on Canvas. If you have spare time, also review all the other notes and sample problem solutions on Canvas.

Definitely review:

- Reflection and transmission coefficients
- Boundary conditions for waves
- “Completing the square” in order to solve Gaussian integrals
- Using Lorentz transformations to derive other formulas, like length contraction and time dilation