



Verification of the predictions of SSST in nonlinear simulations

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zonal flows coexist with turbulence

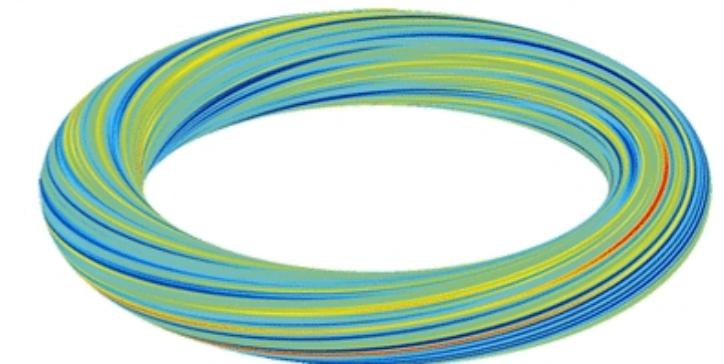
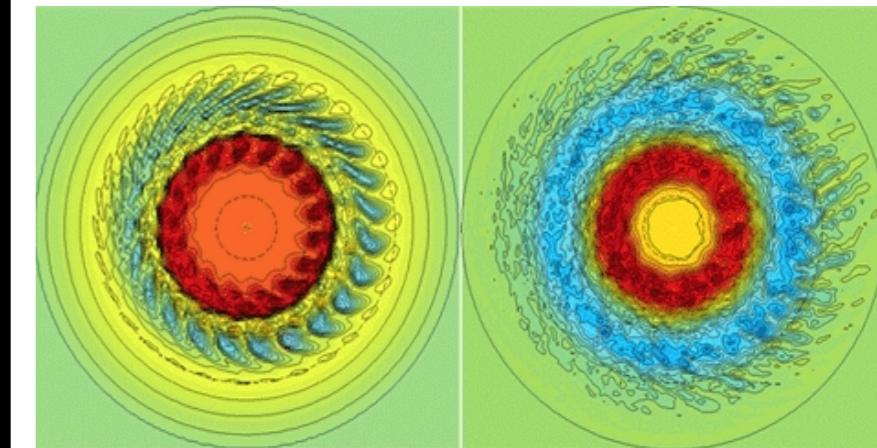


banded Jovian jets

NASA/Cassini Jupiter Images

polar front jet

NASA/Goddard Space Flight Center

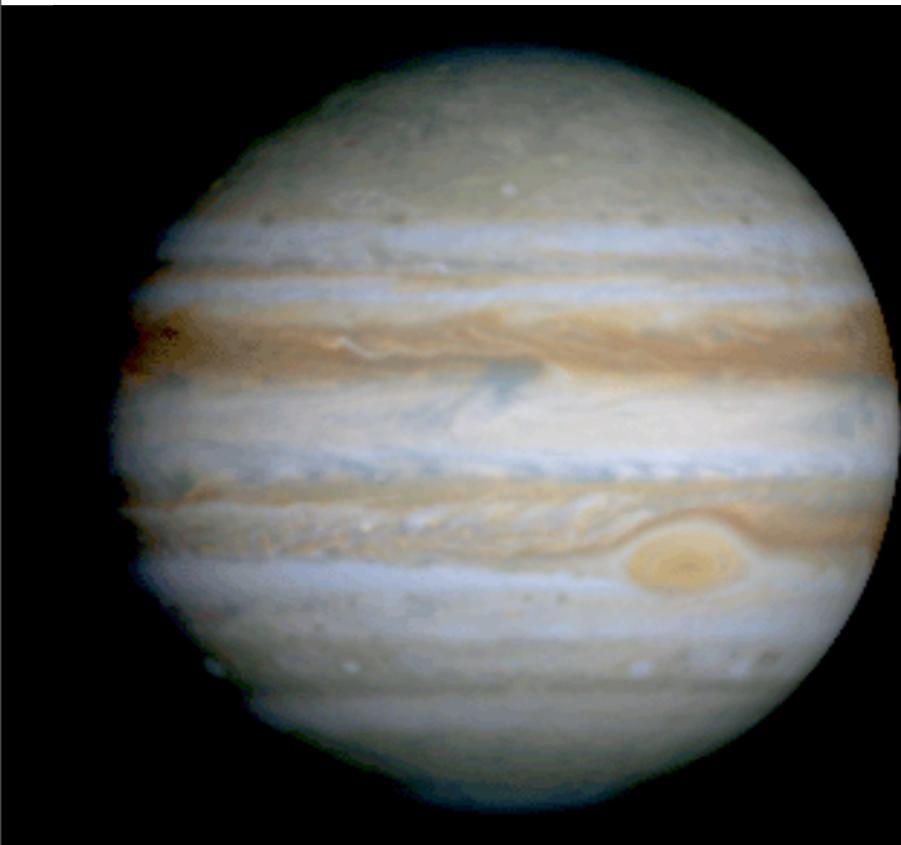


jets in tokamaks

courtesy: L.Villard

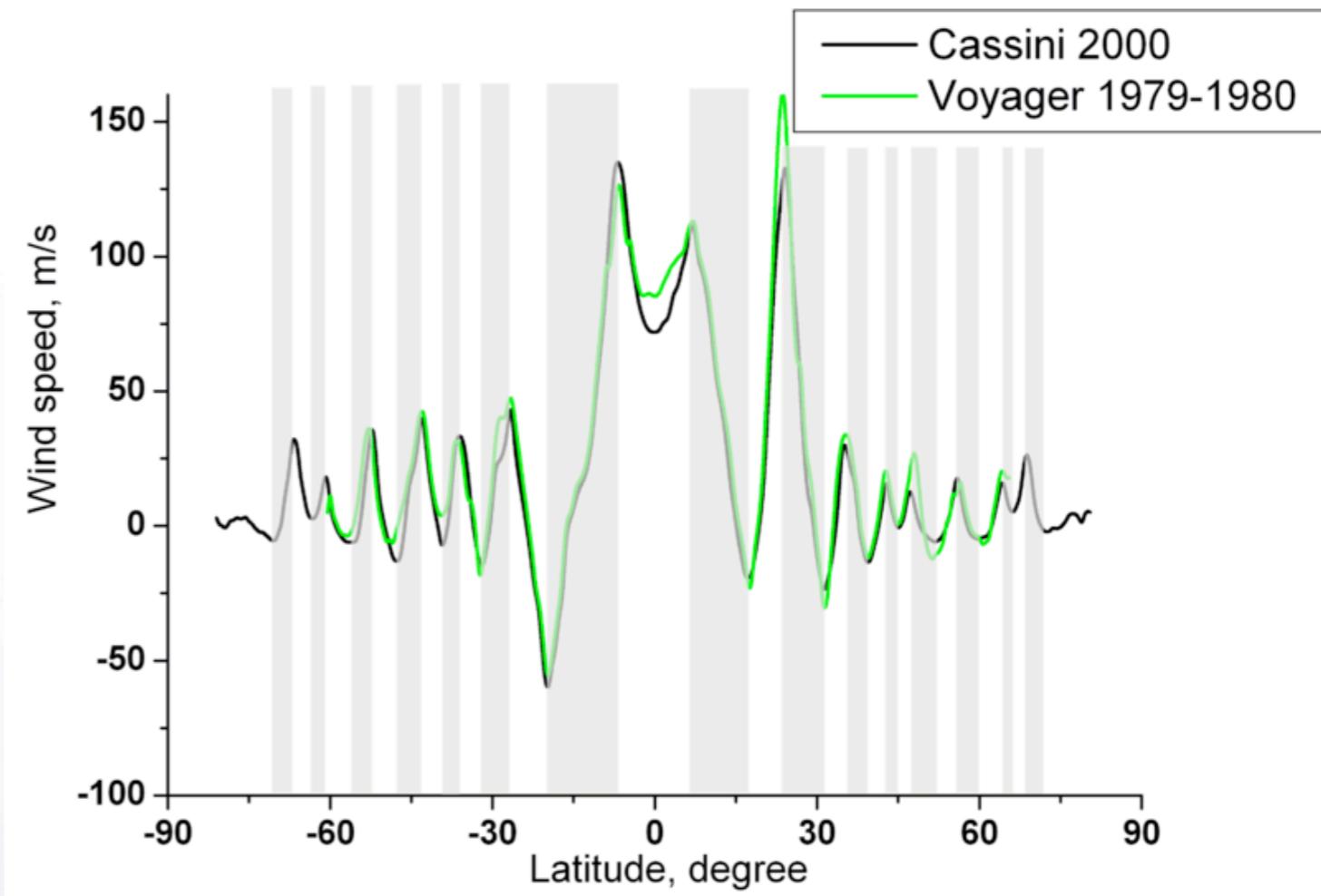


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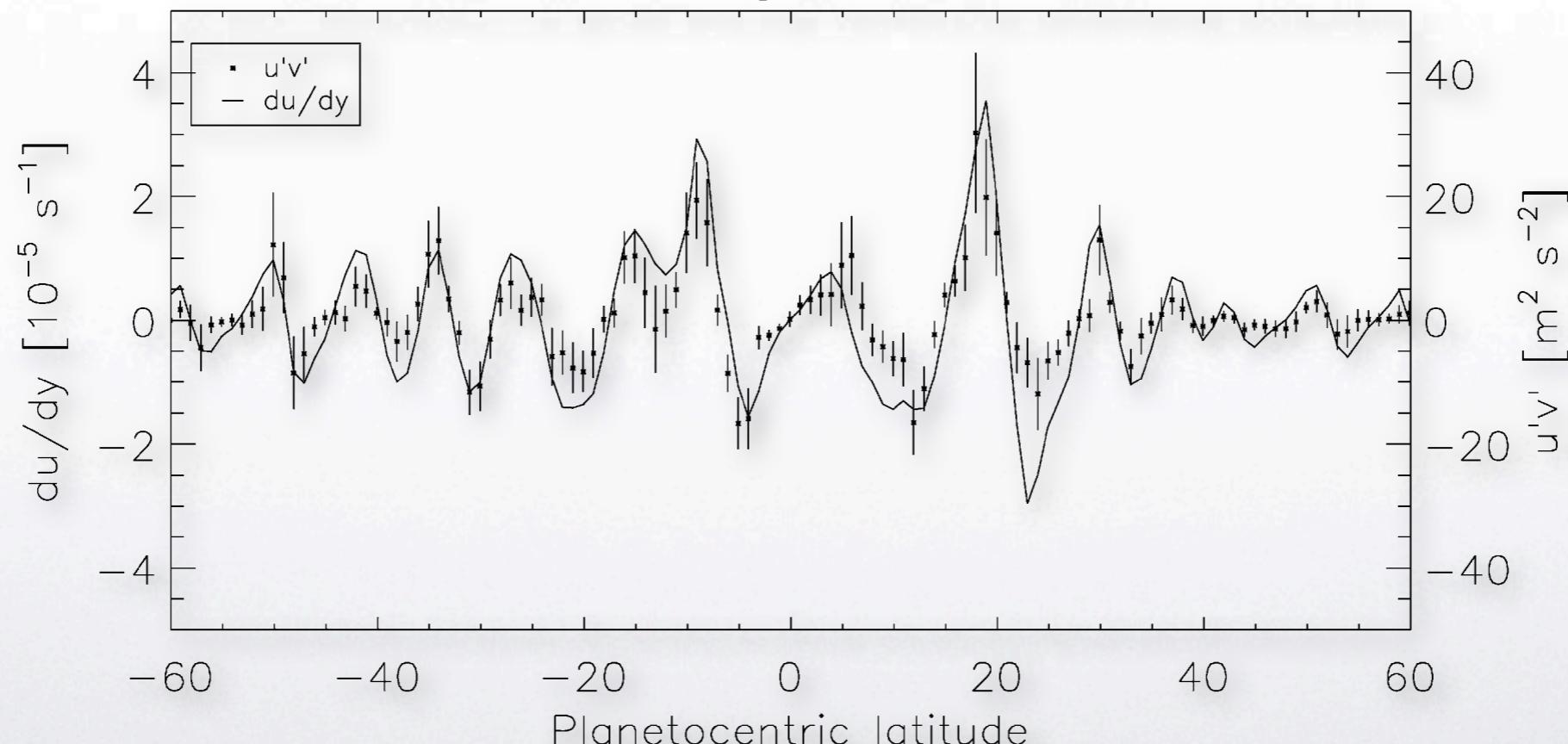


observed Jovian zonal winds
at cloud level
Vasavada & Showman, 2005



zonal flows are maintained by eddies

$$\frac{d}{dt} \int \frac{U^2}{2} dy = \int \frac{dU}{dy} \overline{u'v'} dy - \text{Dissipation}$$



(Salyk et. al. 2006)



Zonal - Eddy field decomposition

$$\varphi(x, y, t) = \Phi(y, t) + \varphi'(x, y, t)$$

where $\Phi(y, t) = \bar{\varphi}(y, t) = \frac{1}{L_x} \int_0^{L_x} \varphi(x', y, t) dx'$



Zonal - Eddy field decomposition

$$\varphi(x, y, t) = \Phi(y, t) + \varphi'(x, y, t)$$

zonal mean

eddy

where

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SSST System

$$\frac{dU}{dt} = \sum_{k=1}^{N_k} \frac{k}{2} \operatorname{Im} [\operatorname{vecd} (\Delta_k^{-1} C_k)] - r_m U$$

$$\frac{dC_k}{dt} = A_k(U) C_k + C_k A_k(U)^\dagger + \epsilon Q_k$$

admits equilibria $(U^E, C_1^E, \dots, C_{N_k}^E) \equiv (U^E, C_k^E)$

homogeneous equilibrium: $(U^E = 0, C_k^E = \epsilon Q_k^E / 2r)$



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$$\overline{v' z'}$$

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admits equilibria (U^E, C_k^E)
homogeneous equilibrium
 N_y discretization points and N_k zonal harmonics \Rightarrow
 $\Rightarrow (U, C_k)$ is a $(2N_k N_y^2 + N_y)$ state vector
(e.g. $N_y = 128, N_k = 10 \Rightarrow (U, C_k) \sim 3.3 \times 10^5$)



SSST System

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Stability of SSST equilibrium

perturbing the SSST equilibrium $(U^E + \delta U, C_k^E + \delta C_k)$

$$\frac{d\delta U}{dt} = \sum_{k=1}^{N_k} \frac{k}{2} \operatorname{Im} [\operatorname{vecd} (\Delta_k^{-1} \delta C_k)] - r_m \delta U$$

$$\frac{d\delta C_k}{dt} = A_k(U^E) \delta C_k + \delta C_k A_k(U^E)^\dagger + \delta A_k C_k^E + C_k^E \delta A_k^\dagger$$

$$\text{with } \delta A_k \equiv A_k(U^E + \delta U) - A_k(U^E)$$



Stability of SSST equilibrium

perturbing the SSST equilibrium $(U^E + \delta U, \mathbf{C}_k^E + \delta \mathbf{C}_k)$

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$$\frac{d\delta \mathbf{C}_k}{dt} = \mathbf{A}_k(U^E) \delta \mathbf{C}_k + \delta \mathbf{C}_k \mathbf{A}_k(U^E)^\dagger + \delta \mathbf{A}_k \mathbf{C}_k^E + \mathbf{C}_k^E \delta \mathbf{A}_k^\dagger$$

$$\text{with } \delta \mathbf{A}_k \equiv \mathbf{A}_k(U^E + \delta U) - \mathbf{A}_k(U^E)$$

split for real/imaginary part and

search for eigensolutions: $(\delta \hat{U}, \delta \hat{\mathbf{C}}_{k,R}, \delta \hat{\mathbf{C}}_{k,I}) e^{\sigma t}$ for $k = 1, \dots, N_k$,

$$\sigma \begin{pmatrix} \delta \hat{U} \\ \delta \hat{\mathbf{C}}_{k,R} \\ \delta \hat{\mathbf{C}}_{k,I} \end{pmatrix} = \mathbb{L} \begin{pmatrix} \delta \hat{U} \\ \delta \hat{\mathbf{C}}_{k,R} \\ \delta \hat{\mathbf{C}}_{k,I} \end{pmatrix}$$



Stability of SSST equilibrium

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$$\frac{d\delta \mathbf{C}_k}{dt} = \mathbf{A}_k(U^E) \delta \mathbf{C}_k + \delta \mathbf{C}_k \mathbf{A}_k(U^E)^\dagger + \delta \mathbf{A}_k \mathbf{C}_k^E + \mathbf{C}_k^E \delta \mathbf{A}_k^\dagger$$

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$\mathbb{L} = \mathbb{L}(U^E, \mathbf{C}_k^E)$ has dimension $(2N_k N_y^2 + N_y)$

Eigenanalysis of stability operator

for $(U^E = 0, C_k^E = \epsilon Q_k^E / 2r) \Rightarrow \delta \hat{U} = e^{iny}$

1. assume $\sigma, \delta \hat{U}_n \Rightarrow \delta \hat{C}_{k,n} \Rightarrow \delta(\bar{v' \zeta'})_n$

2. use mean flow perturbation equation

$$\sigma \delta \hat{U}_n = \delta(\bar{v' \zeta'})_n - r_m \delta \hat{U}_n$$

to calculate new σ

3. repeat until convergence

(this method can be generalized for $U^E \neq 0$ equilibria)

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to calculate new σ

3. repeat until convergence

this method works extremely well and allows us to attack the $(2N_k N_y + N_y) \times (2N_k N_y + N_y)$ eigenvalue problem which would otherwise be impossible

(this method can be generalized for $U^E \neq 0$ equilibria)



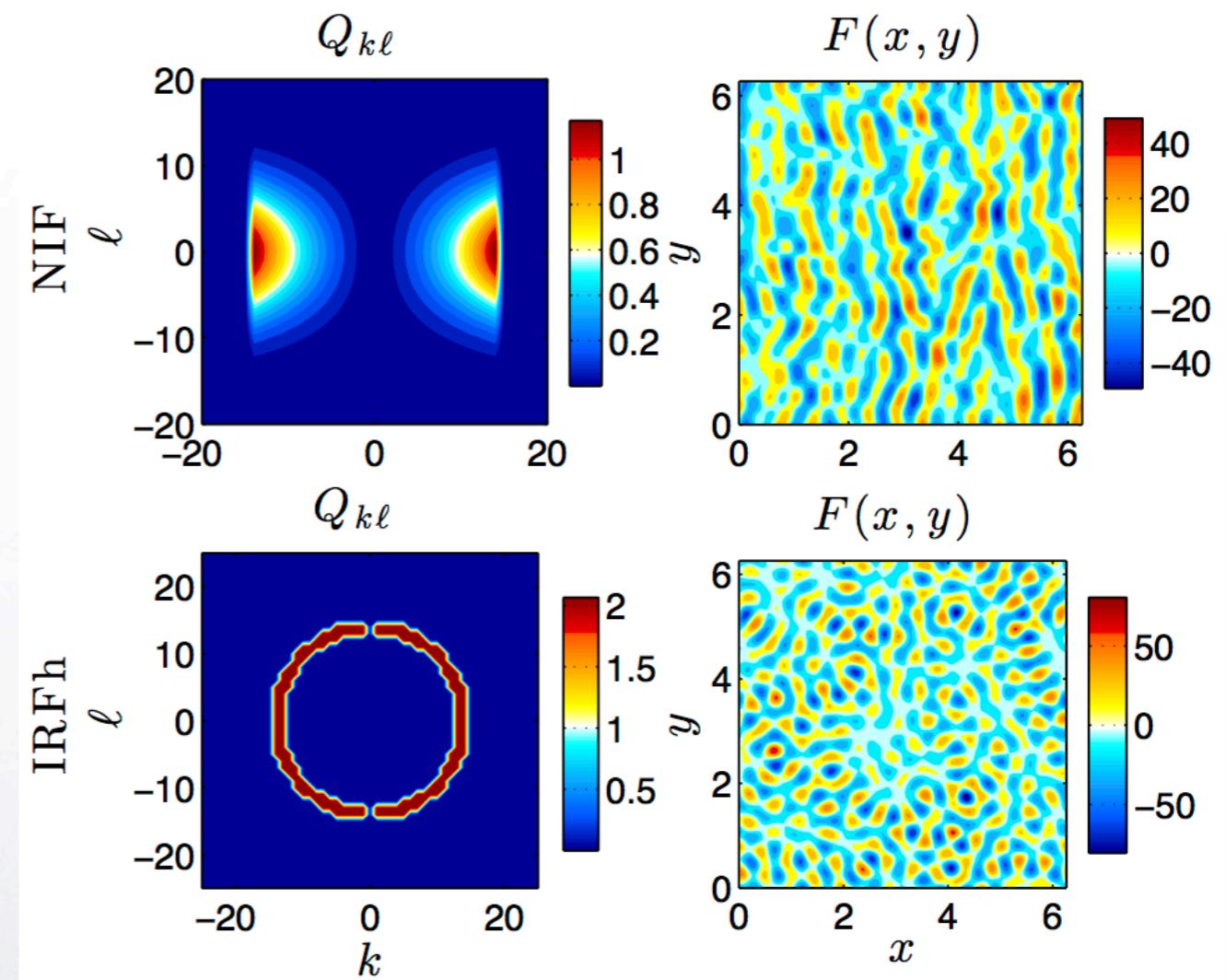
Types of stochastic forcing: NIF & IRFh

Non-isotropic Forcing (NIF)

parametrize excitation by
baroclinic instabilities

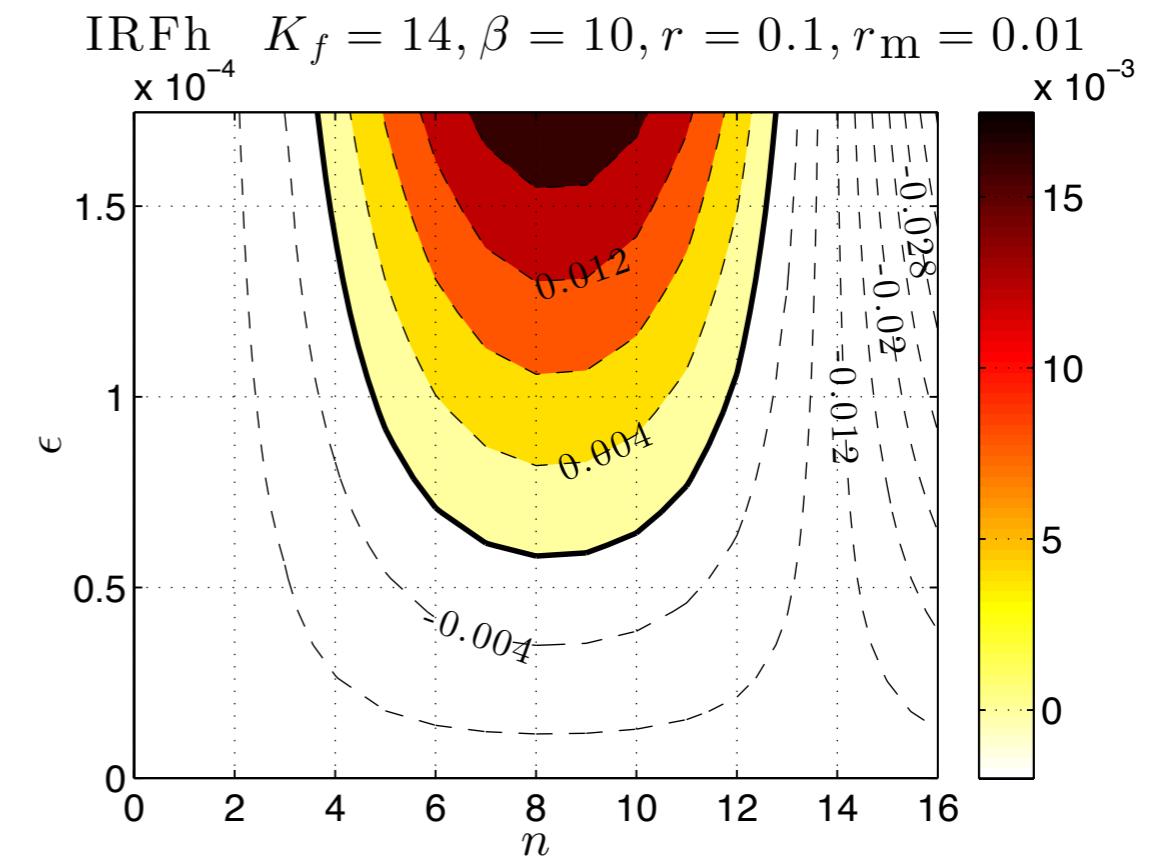
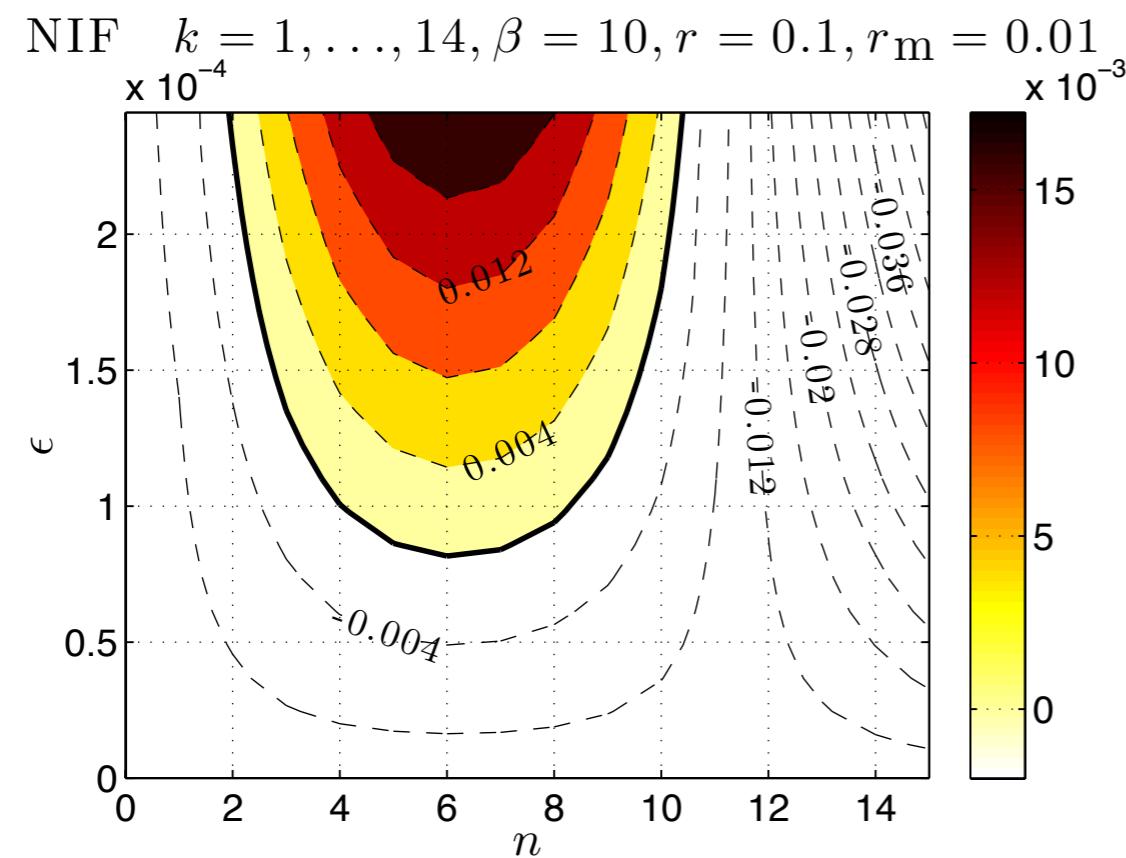
Isotropic Ring Forcing (IRFh)

parametrize excitation by
convection



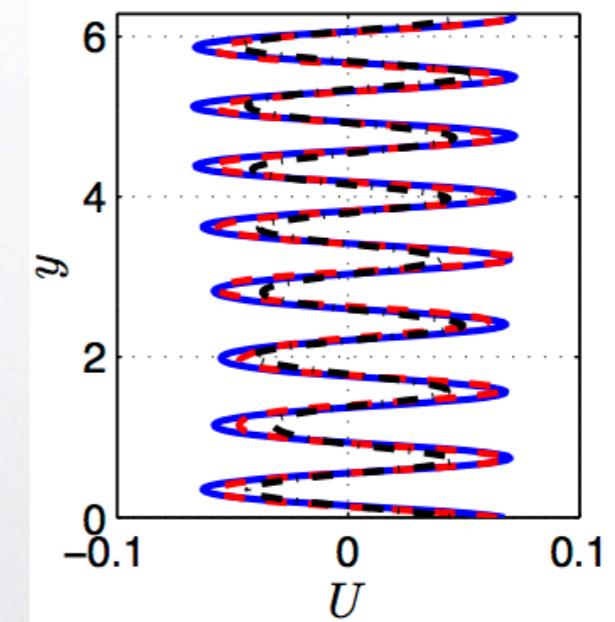
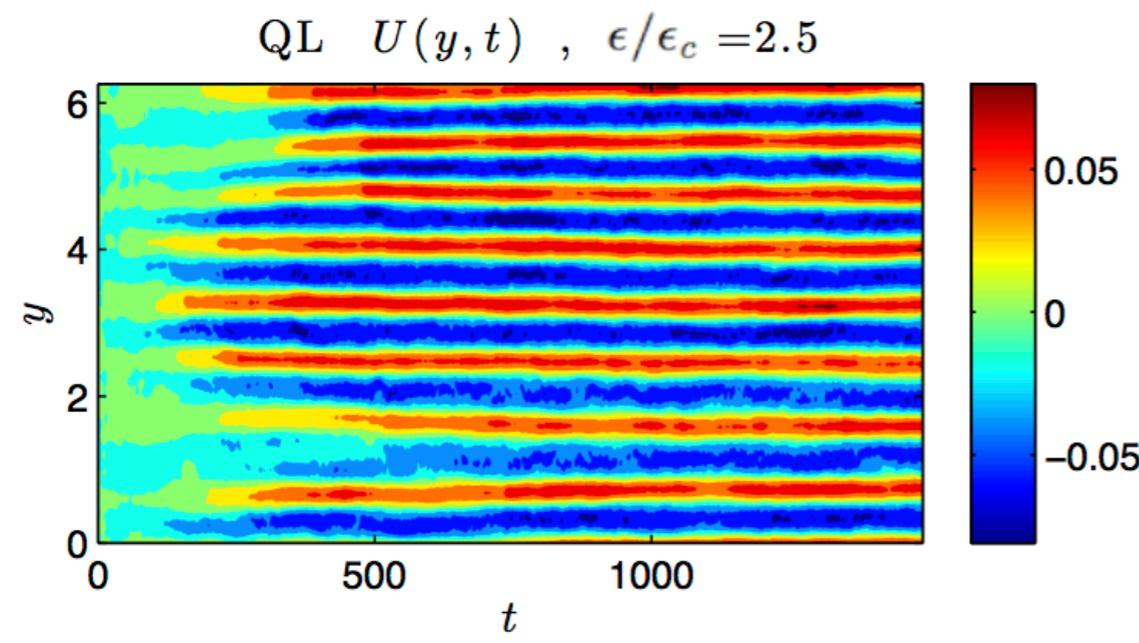
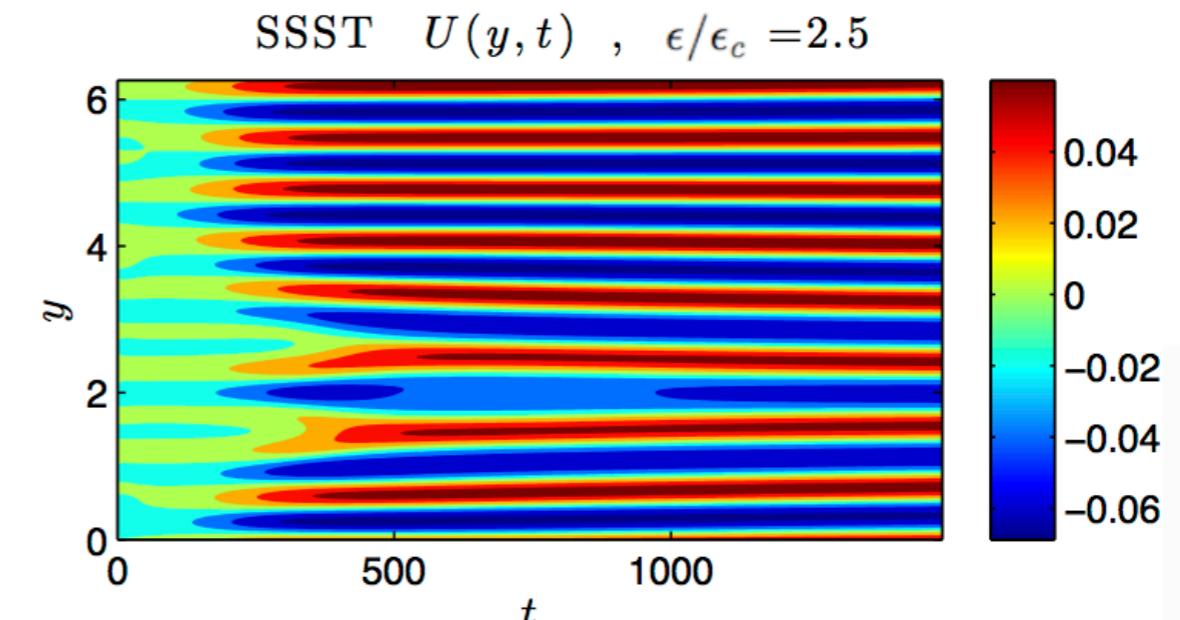
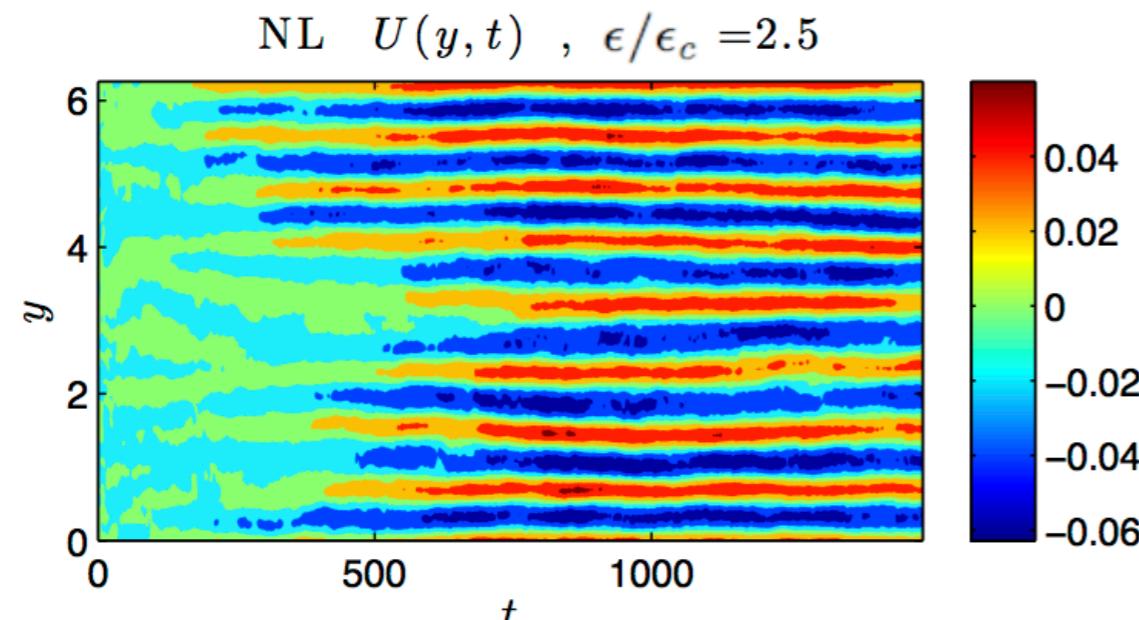


SSST growth rates for homogeneous equilibrium



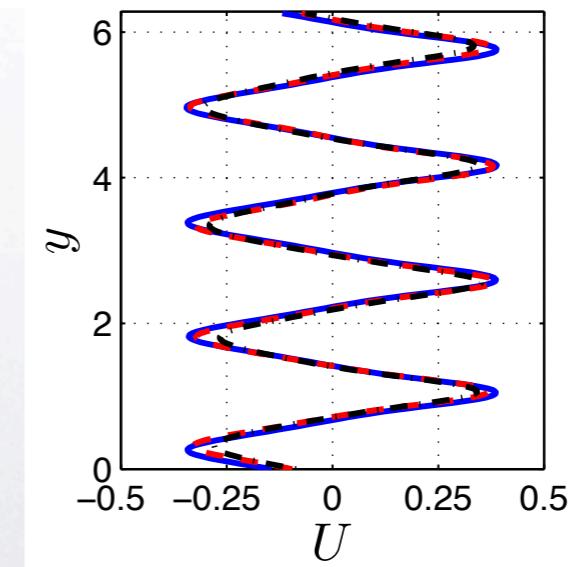
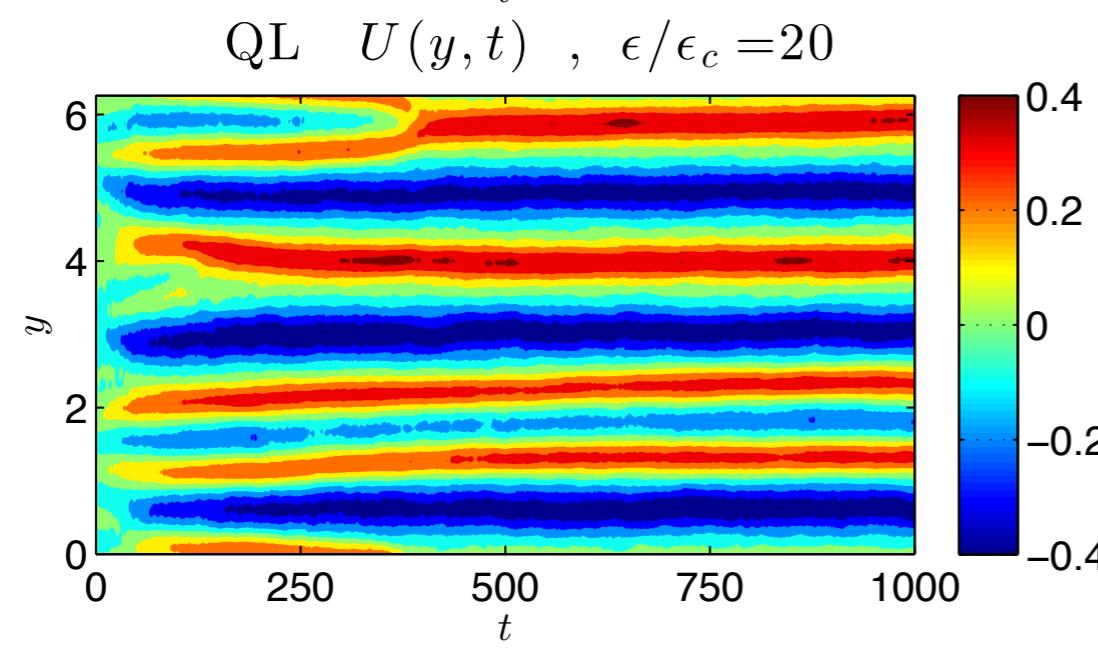
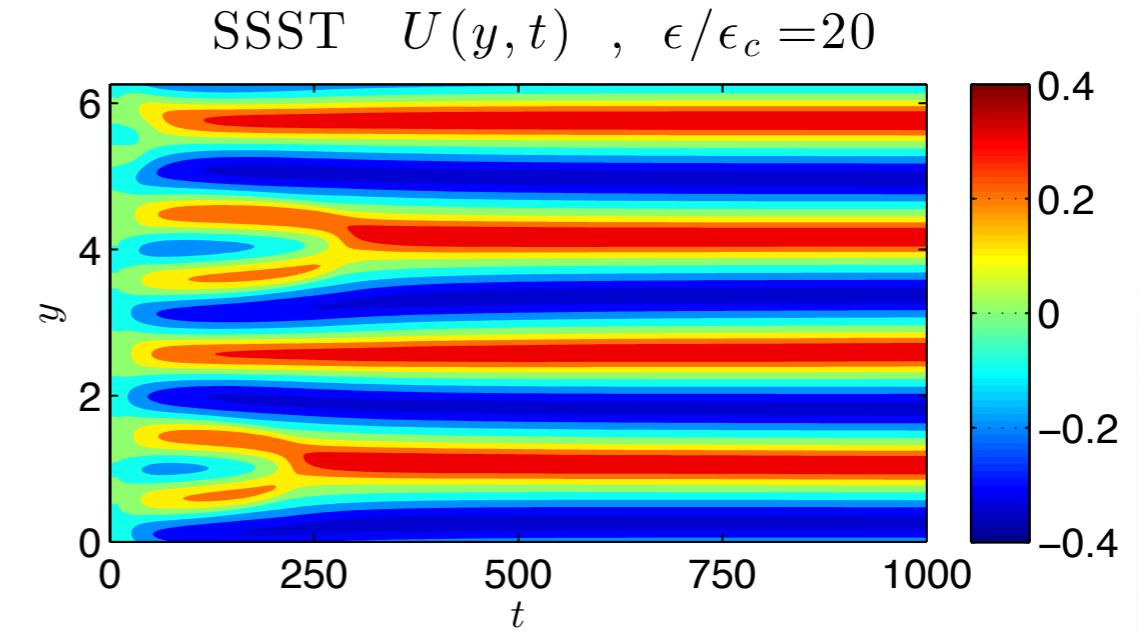
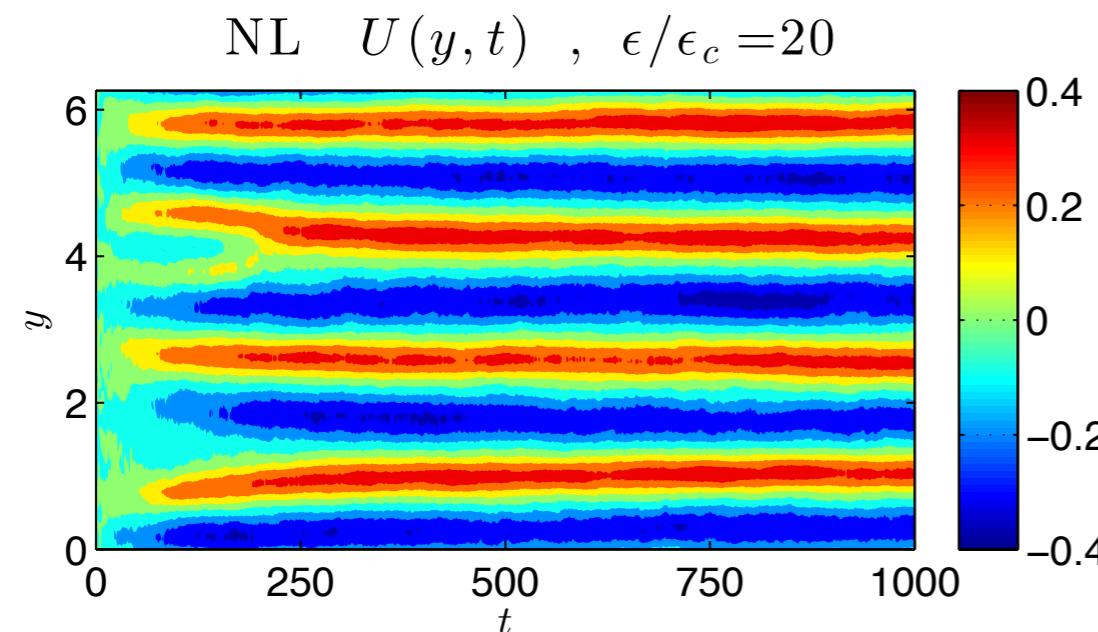


IRFh forcing, $r_m = r/10$ (NL, QL, SSST)





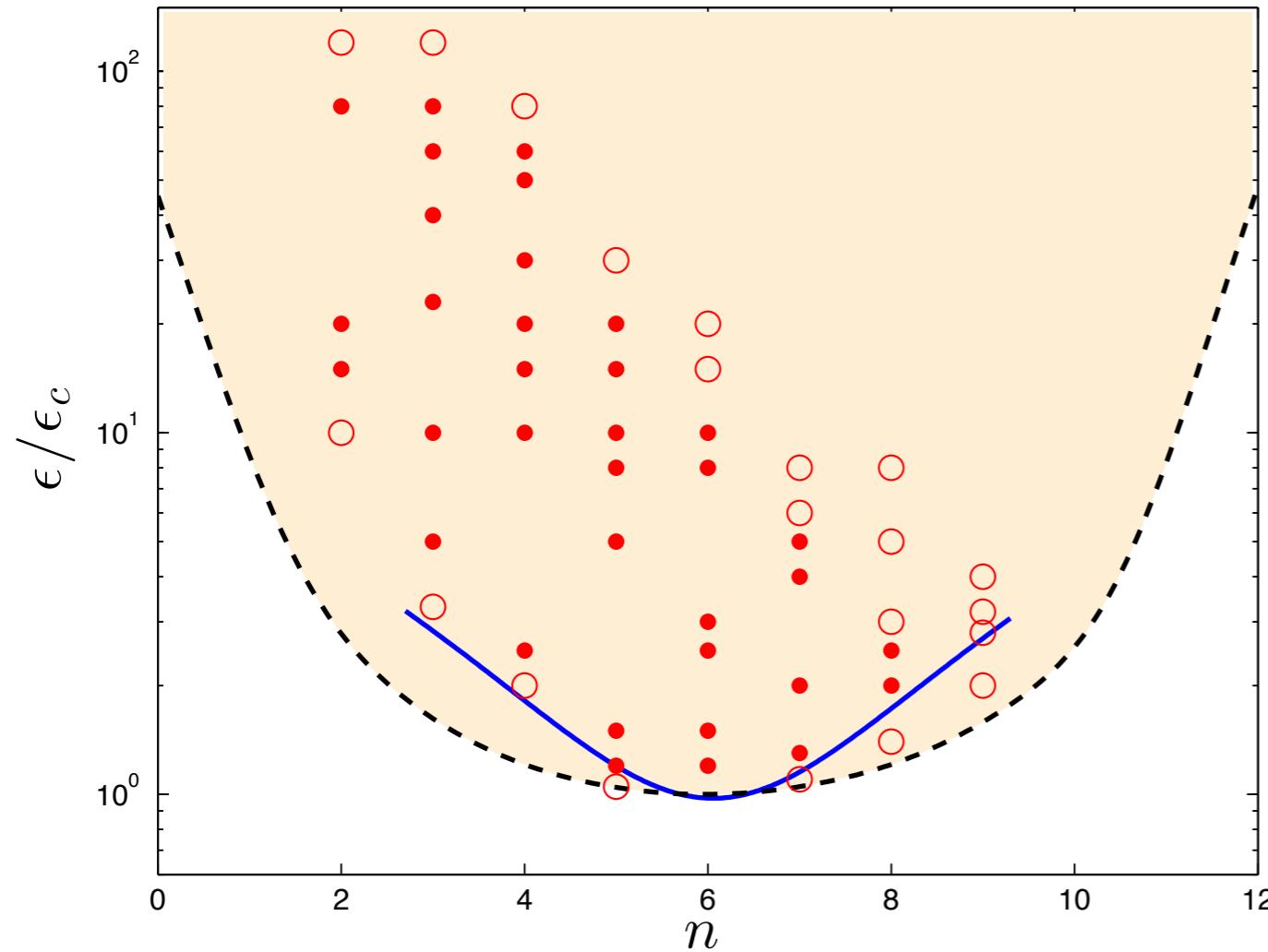
NIF forcing, $r_m = r/10$ (NL, QL, SSST)





SSST equilibria stability diagram

NIF at $k = 1, \dots, 14$, $r = 10^{-1}$, $r_m = 10^{-2}$, $\beta = 10$

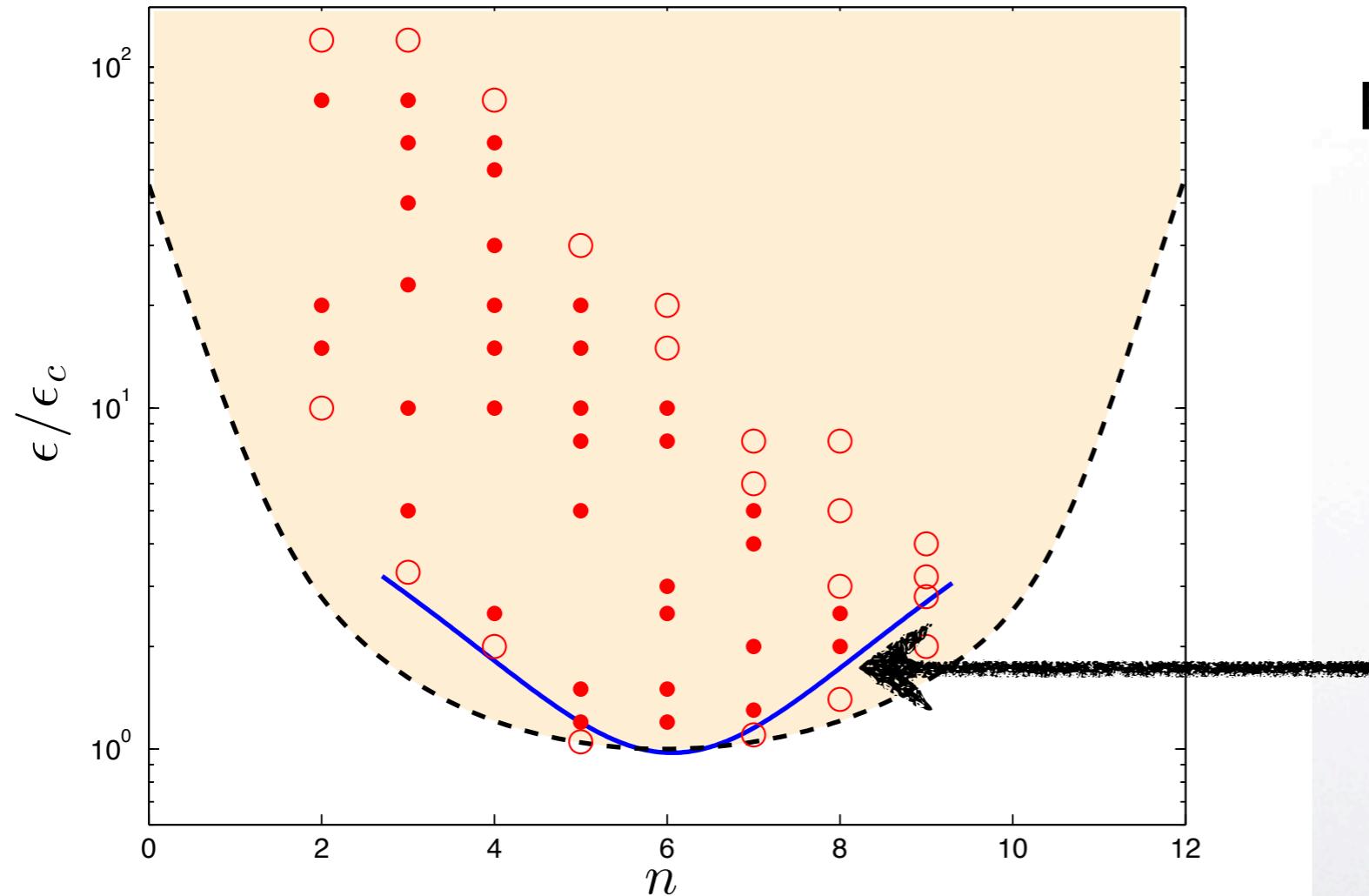


For higher energy input rates
equilibria become SSST
unstable and move towards
the left of the diagram



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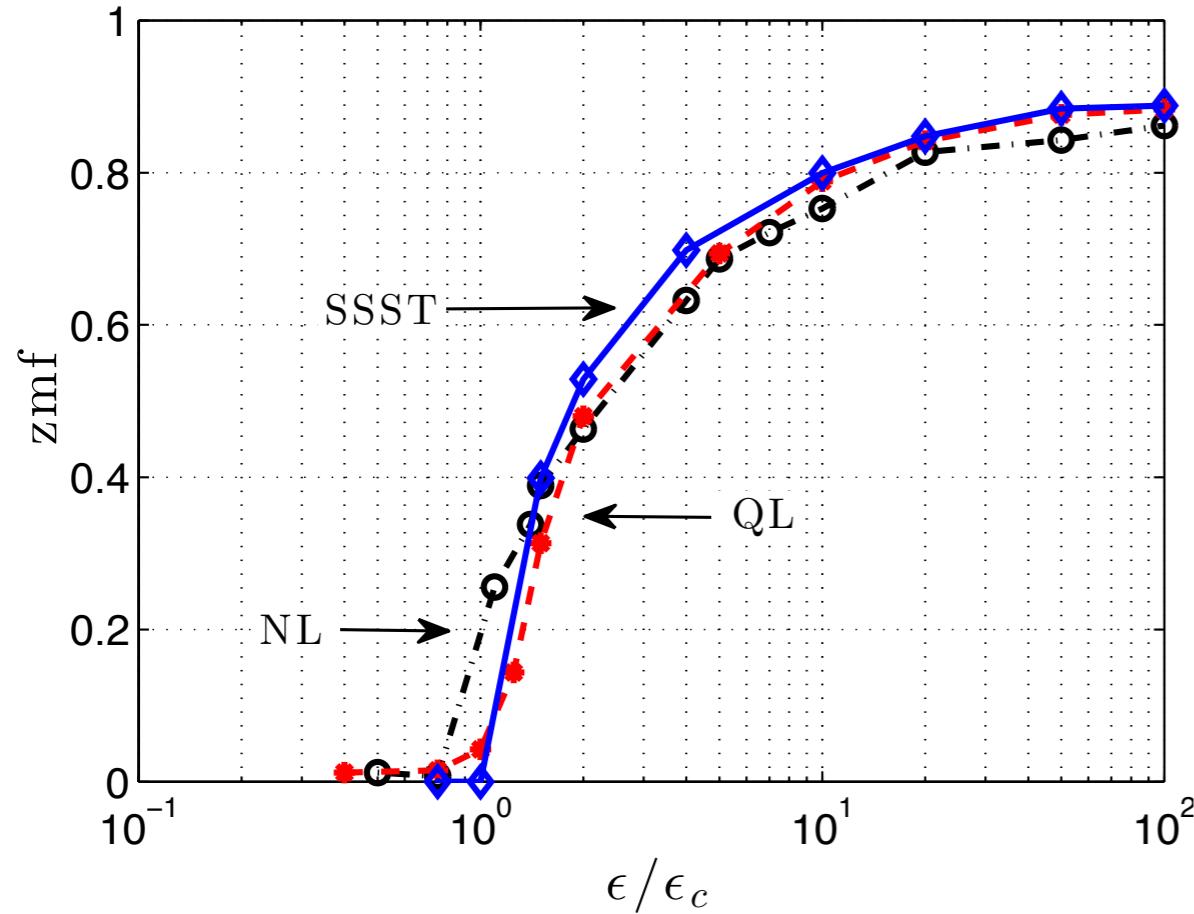
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Eckhaus instability
boundary

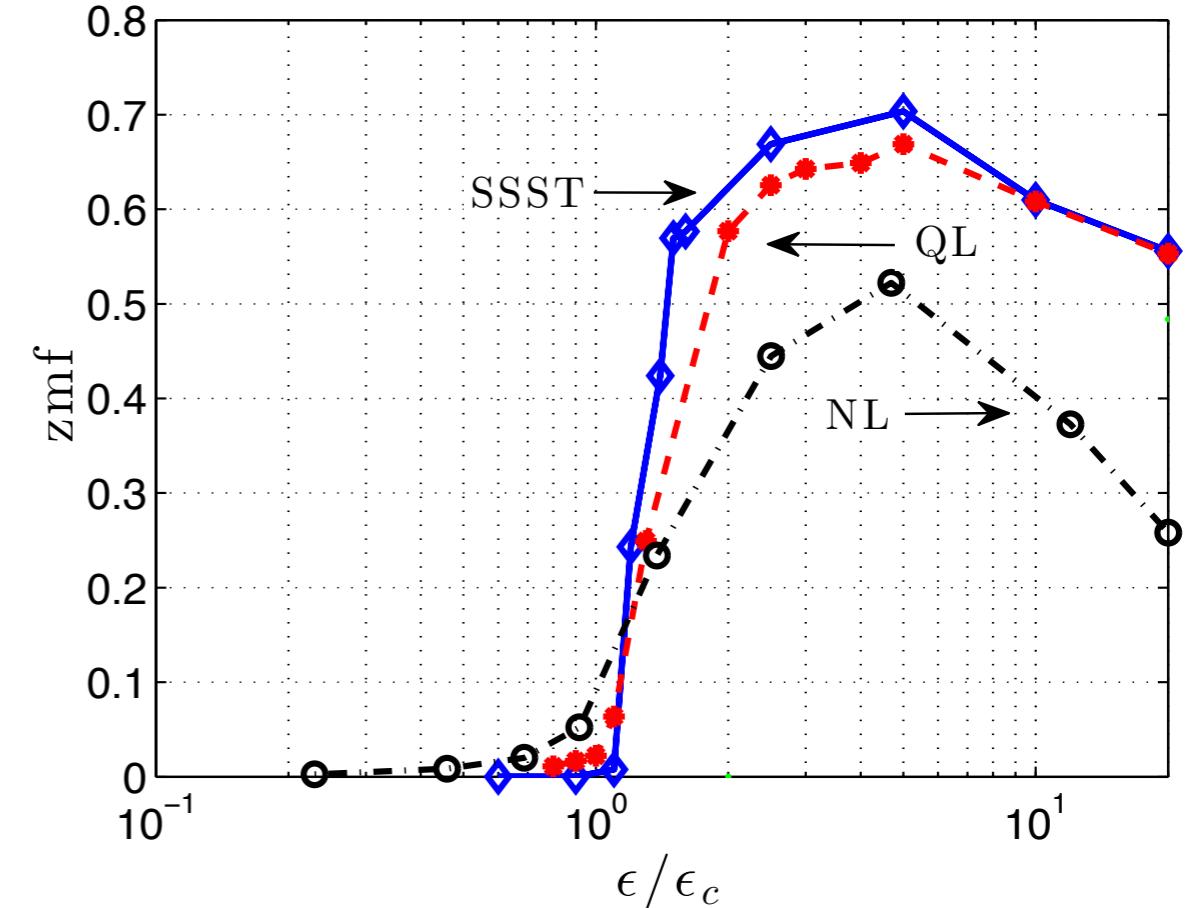


Bifurcation diagrams

NIF, $r = 0.1$, $r_m = r/10$



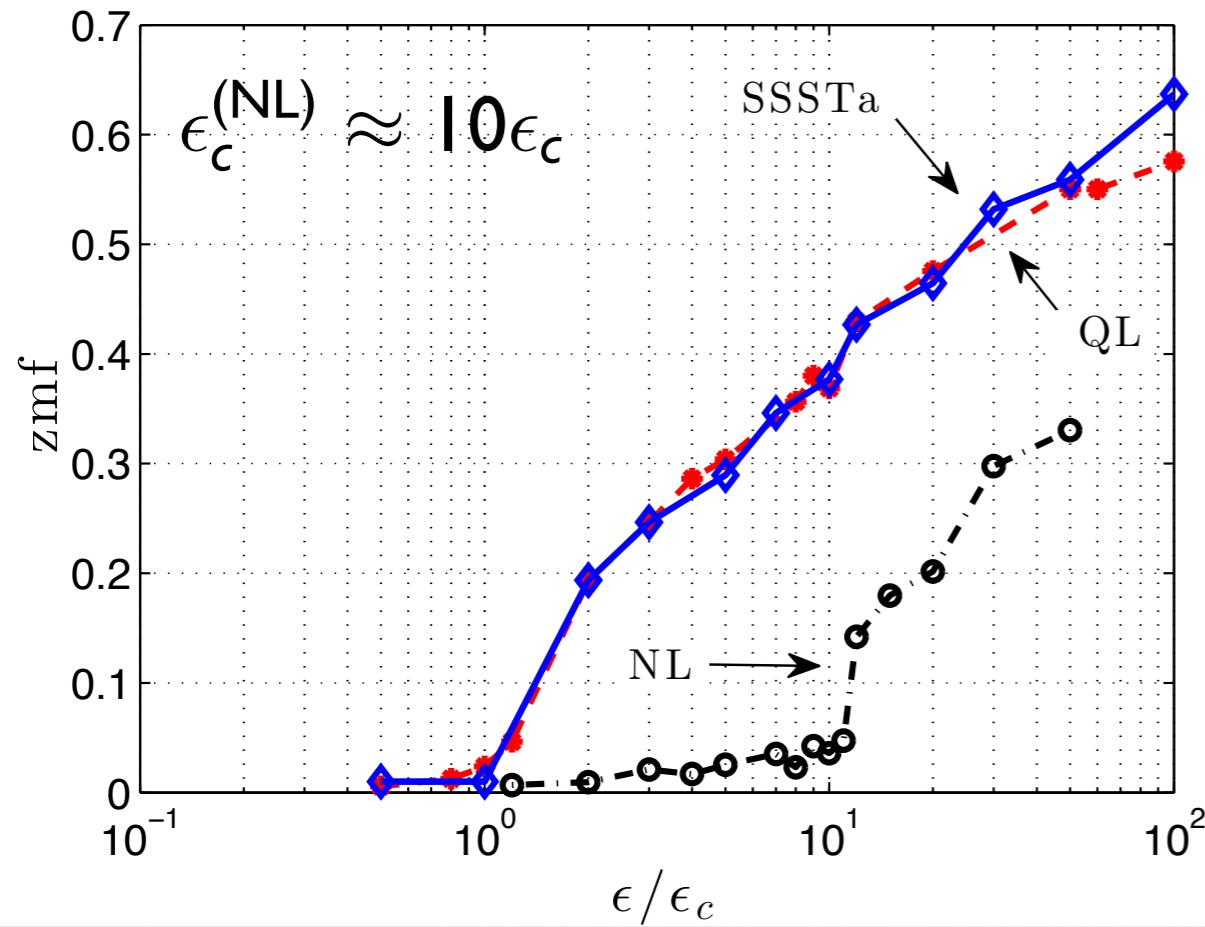
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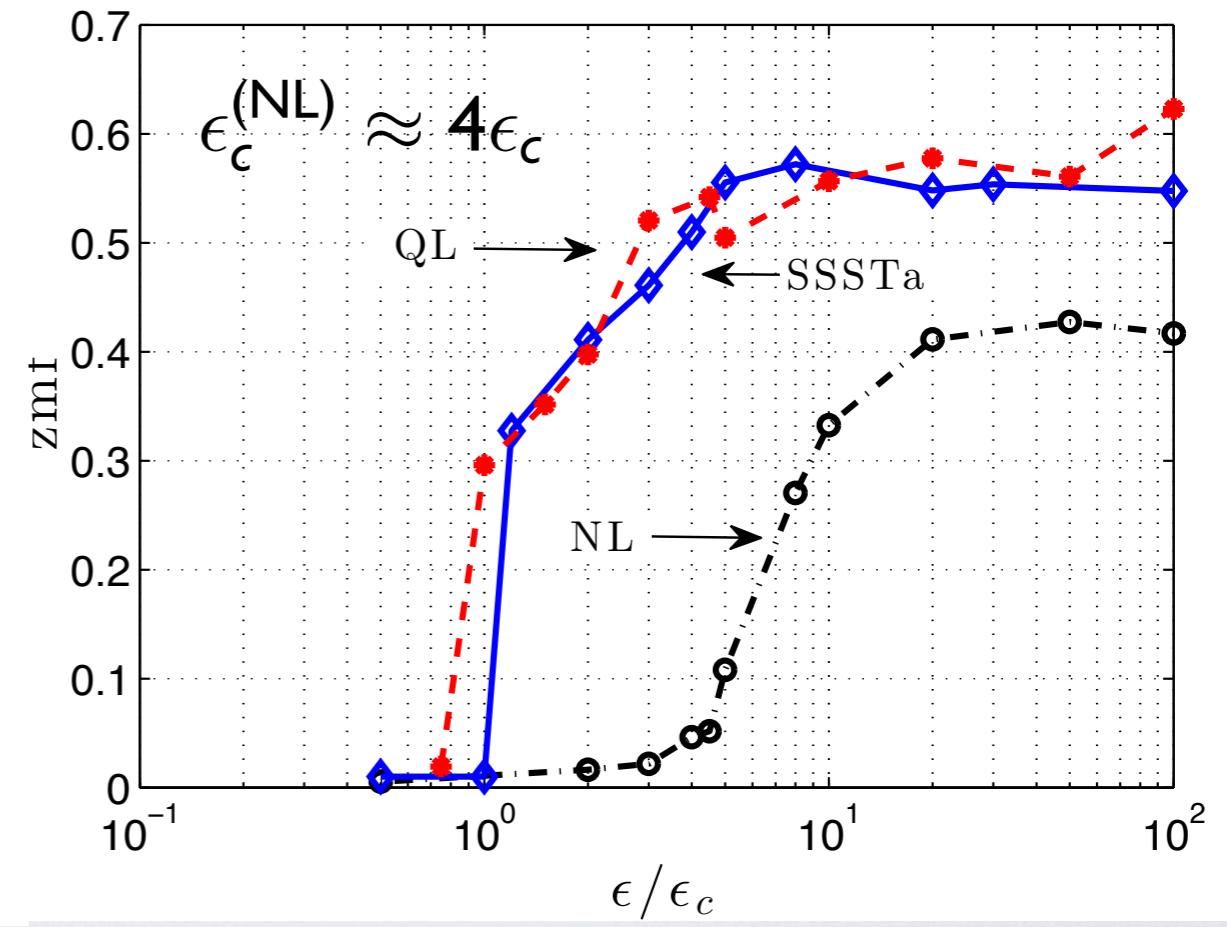


Bifurcation diagrams

NIF, $r = r_m = 10^{-2}$



IRFh, $r = r_m = 10^{-2}$

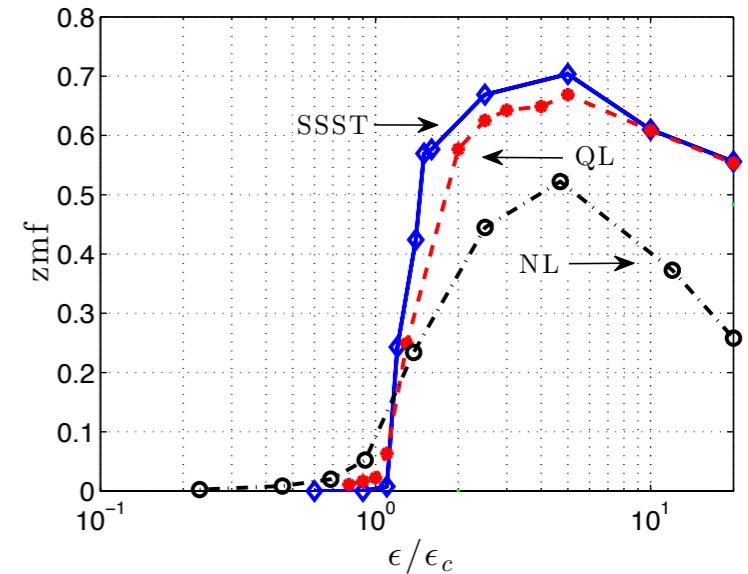




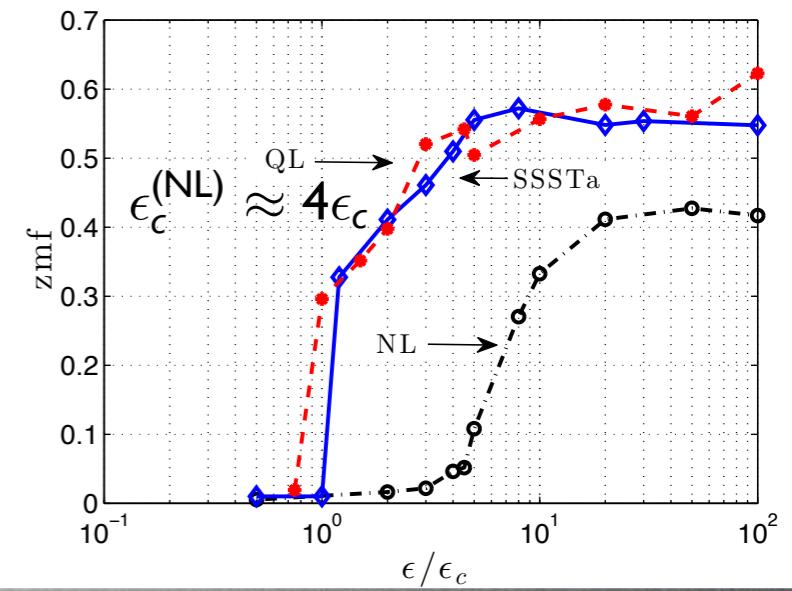
Issues to be clarified

- ▶ What happens below the critical energy input rate, ϵ_c ?
- ▶ When does the predicted ϵ_c agrees with the NL bifurcation point, $\epsilon_c^{(NL)}$ and why they don't always agree?
- ▶ What happens in the range $\epsilon_c < \epsilon < \epsilon_c^{(NL)}$?
- ▶ How can $\epsilon_c^{(NL)}$ be predicted?

IRFh, $r = 0.1$, $r_m = r/10$



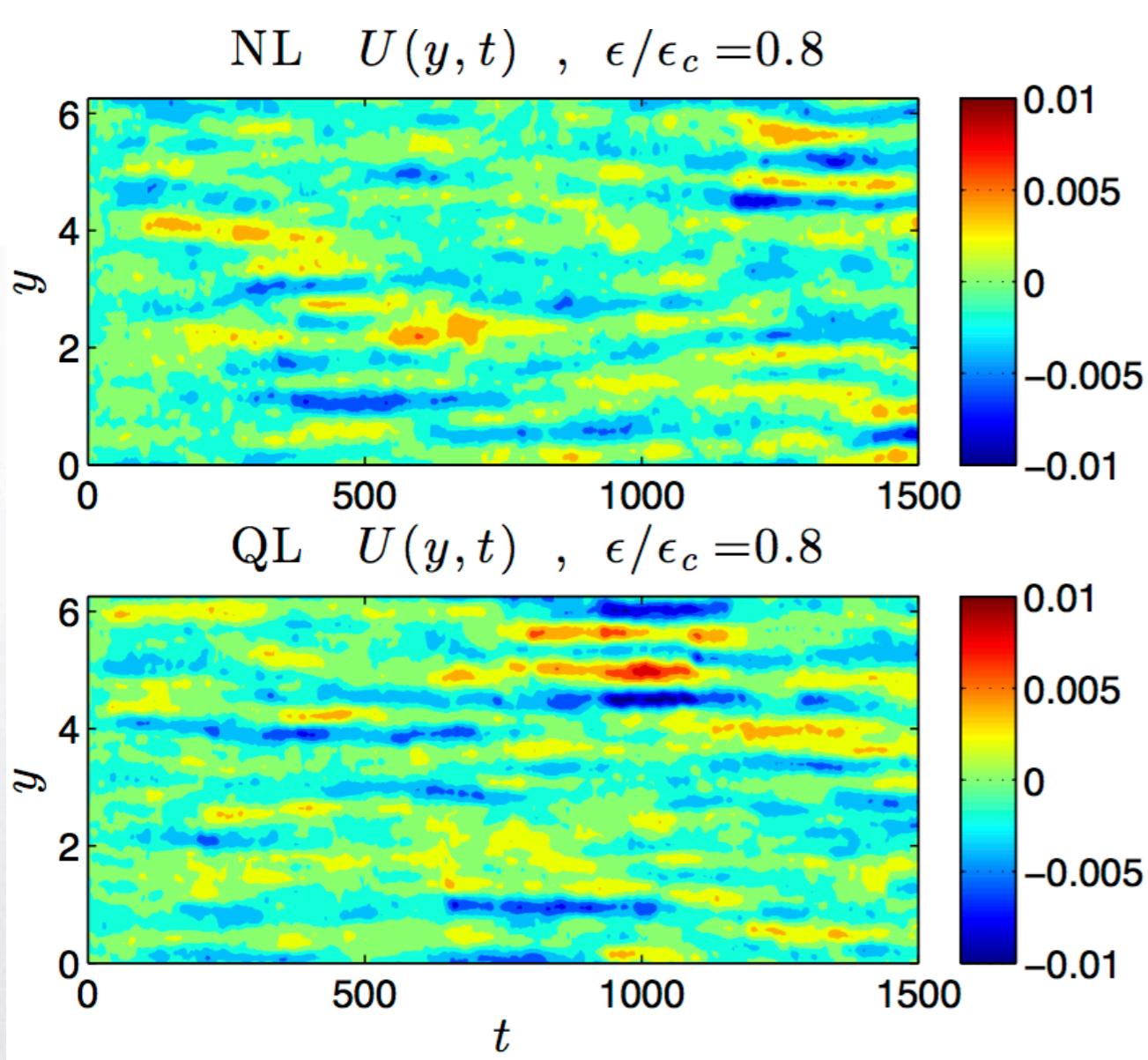
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Stochastically excited stable SSST eigenmodes

(Latent jets in ocean: Berloff et. al. 2009, Berloff et. al. 2011, Cravatte et. al. 2012)



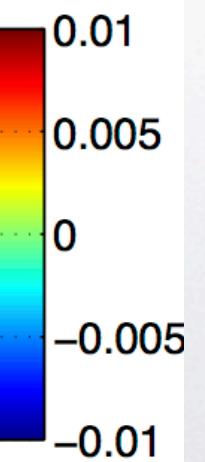
$$U(y, t) = \operatorname{Re} \left[\sum_{n=1}^{15} \alpha_n(t) e^{iny} \right]$$

$$\frac{d\alpha_n}{dt} = \sigma_n \alpha_n + \xi(t)$$

σ_n : SSST growth rates (least stable)

$\xi(t)$: delta-correlated noise

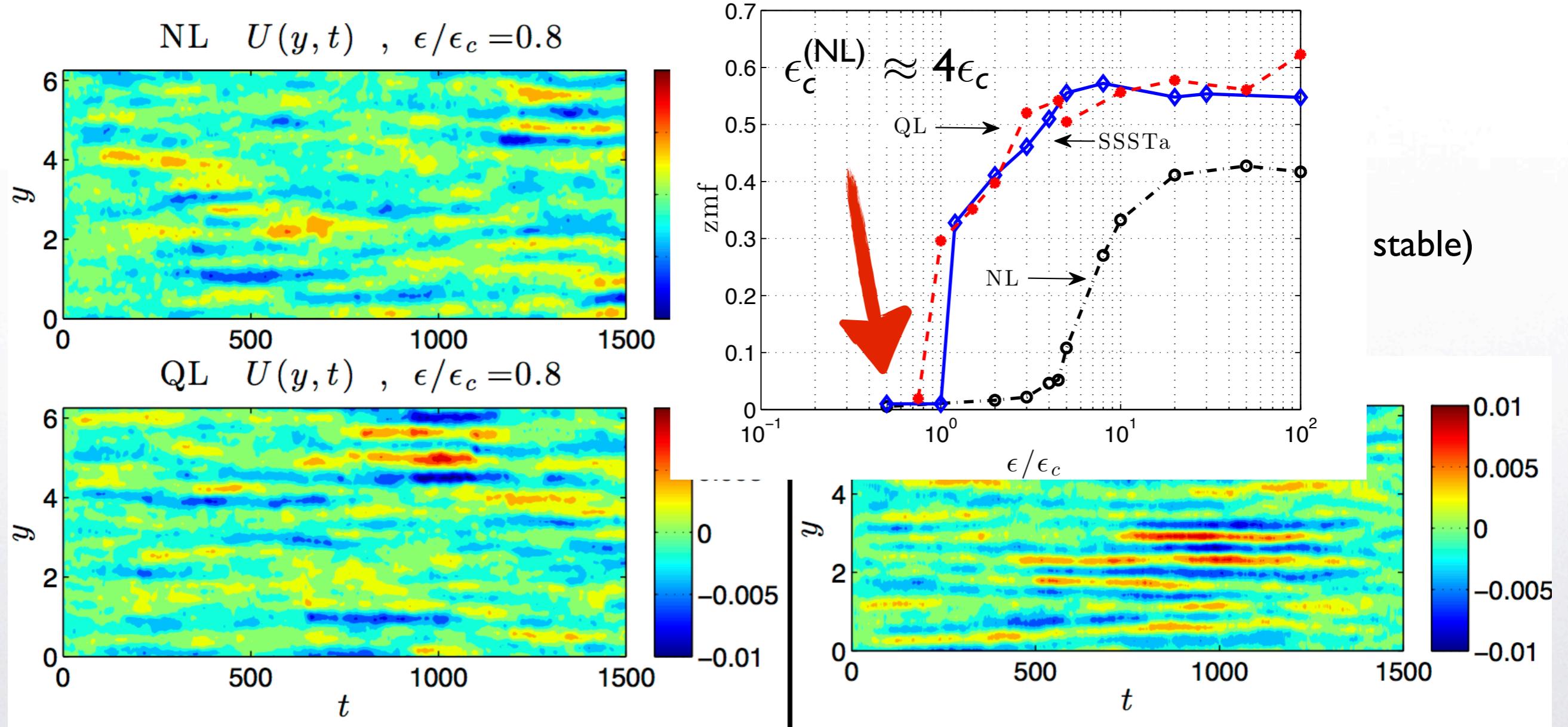
SSST $U(y, t)$, $\epsilon/\epsilon_c = 0.8$





Stochastically excited stable SSST eigenmodes

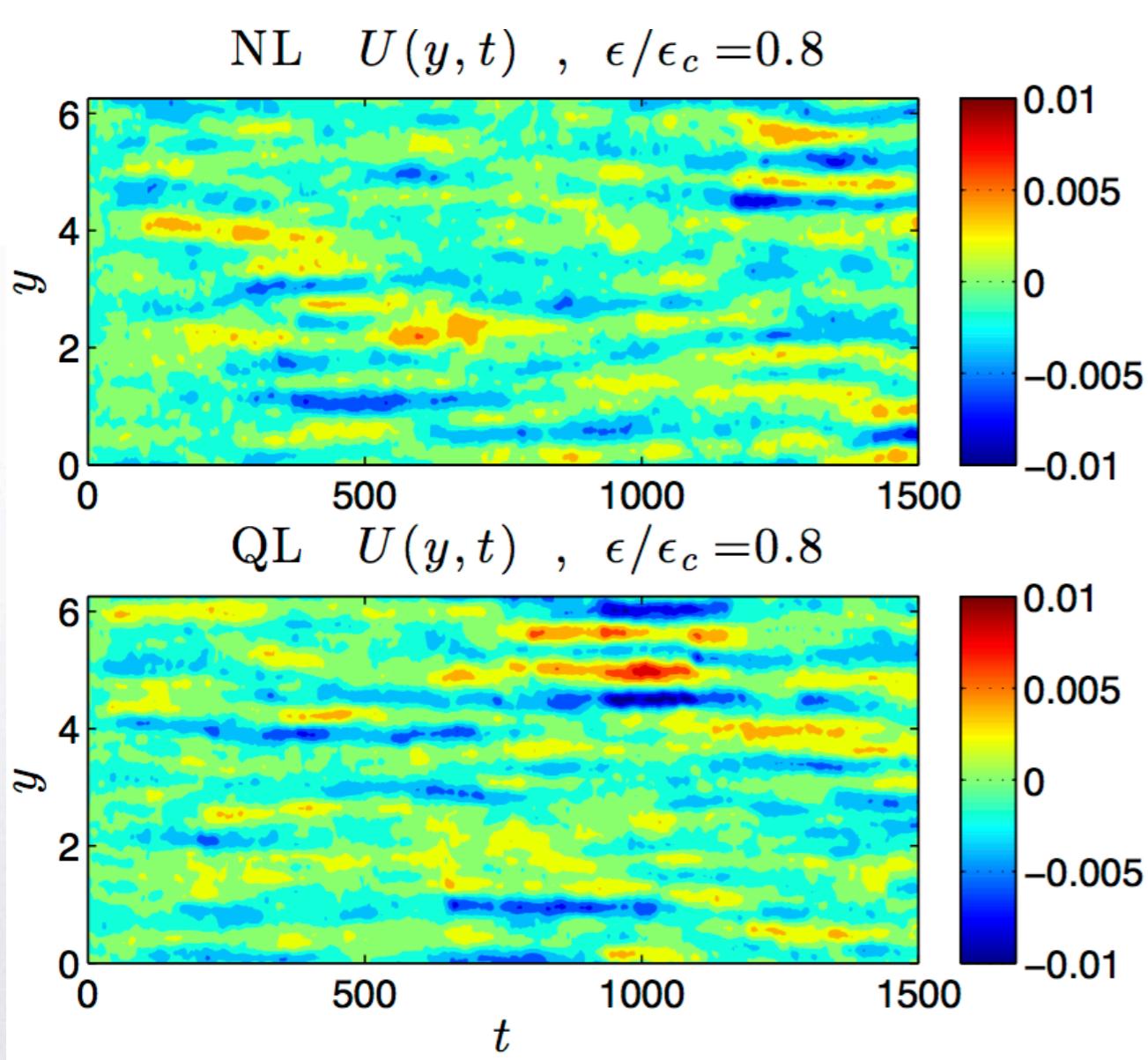
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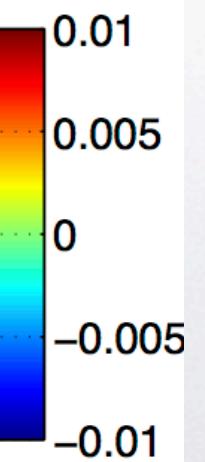
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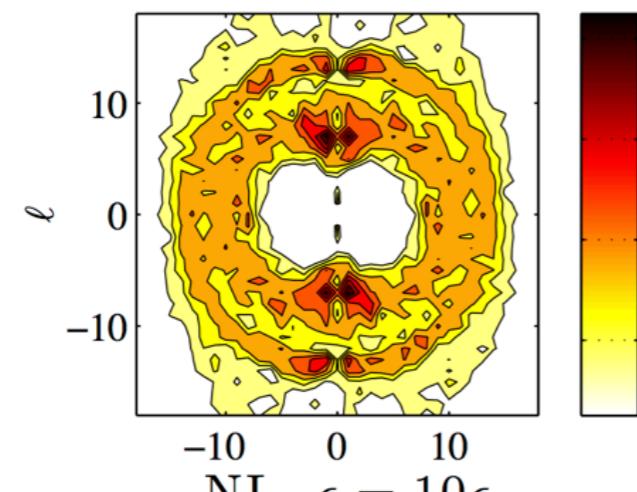


Zonons at NL

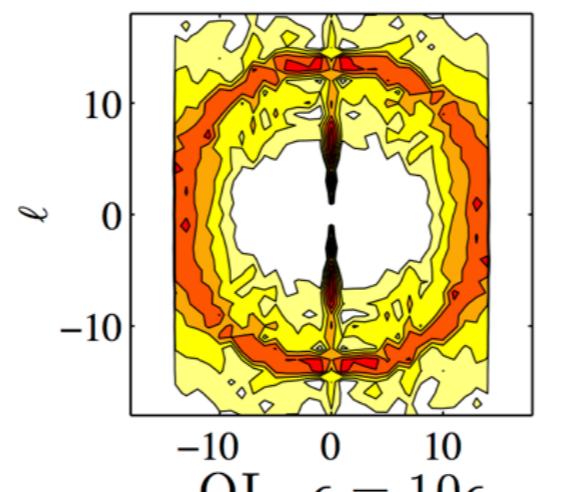
energy spectrum

IRFh, $r = r_m = 10^{-2}$

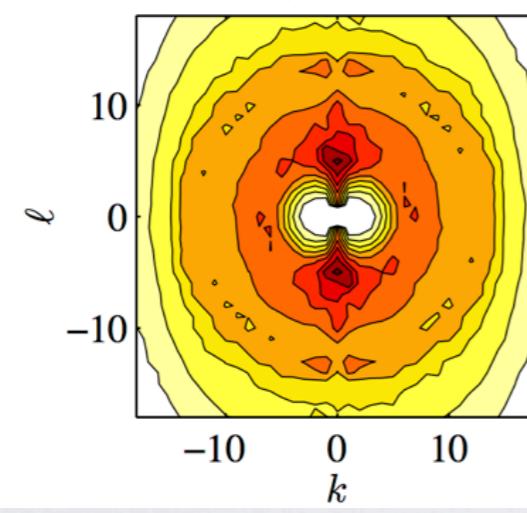
NL, $\epsilon = 2\epsilon_c$



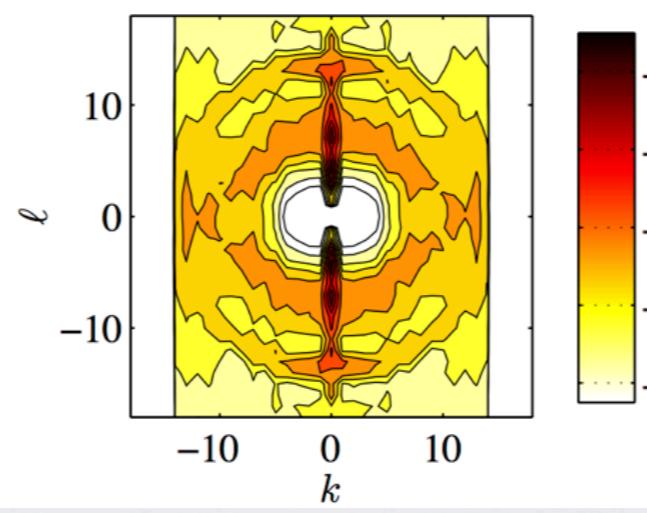
QL, $\epsilon = 2\epsilon_c$



NL, $\epsilon = 10\epsilon_c$



QL, $\epsilon = 10\epsilon_c$



QL / SSST do not allow the interaction of the turbulence with coherent non-zonal structures

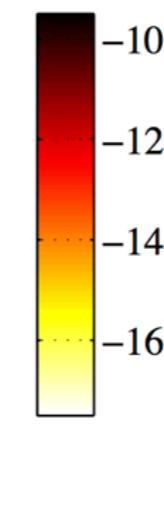
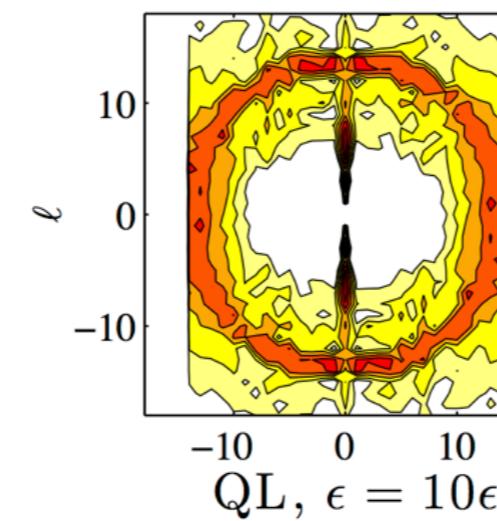
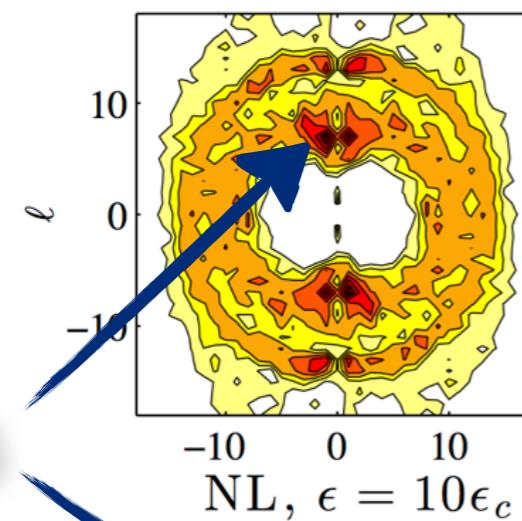


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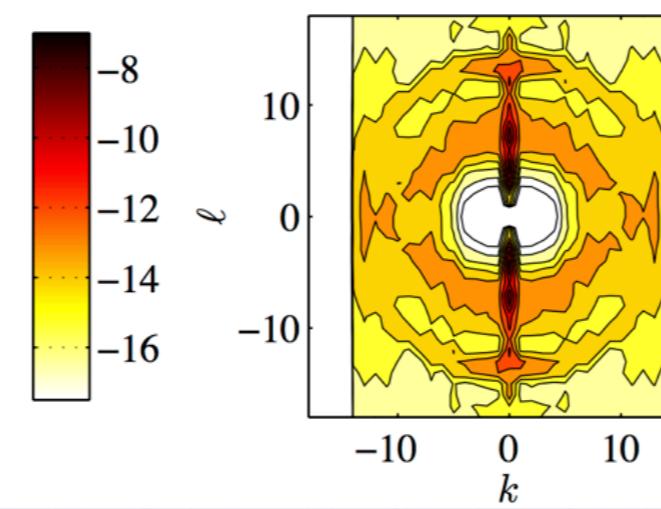
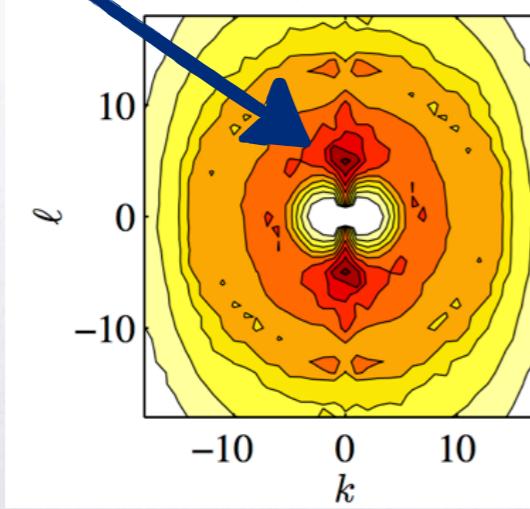
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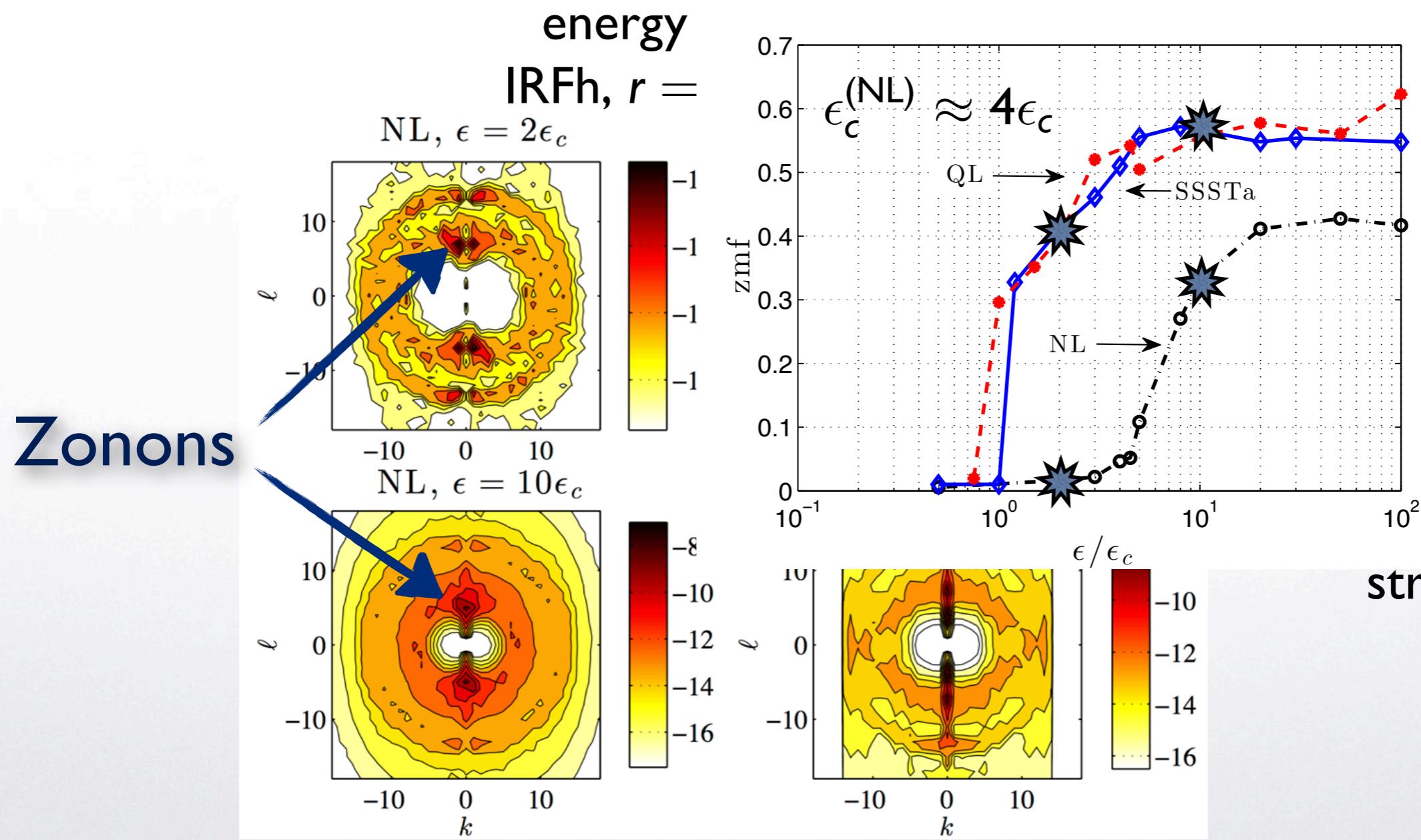
Zonons



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Zonons at NL



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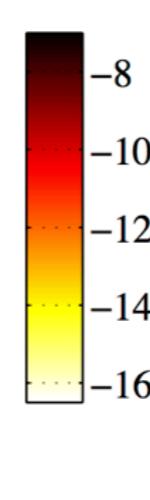
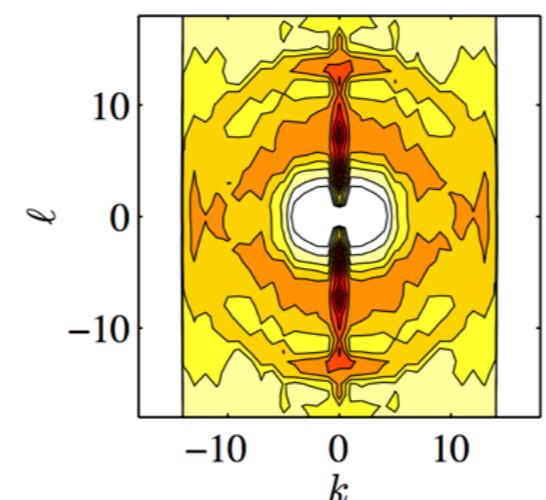
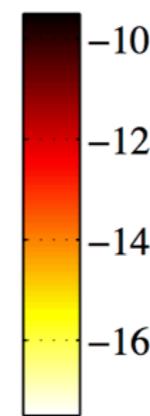
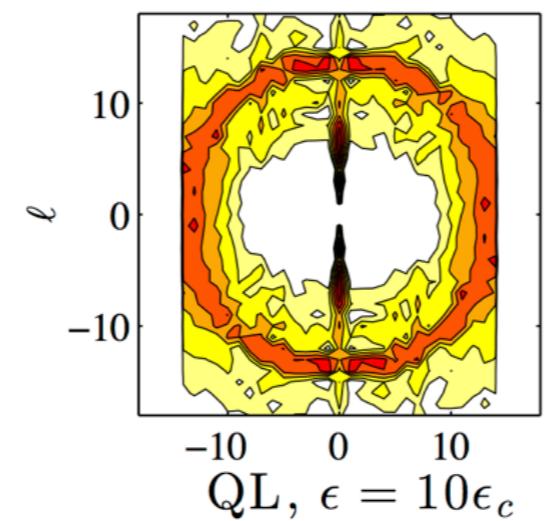
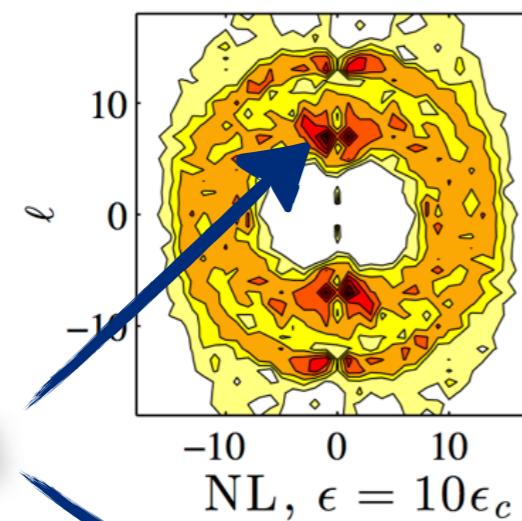


Zonons at NL

energy spectrum

IRFh, $r = r_m = 10^{-2}$

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Zonons

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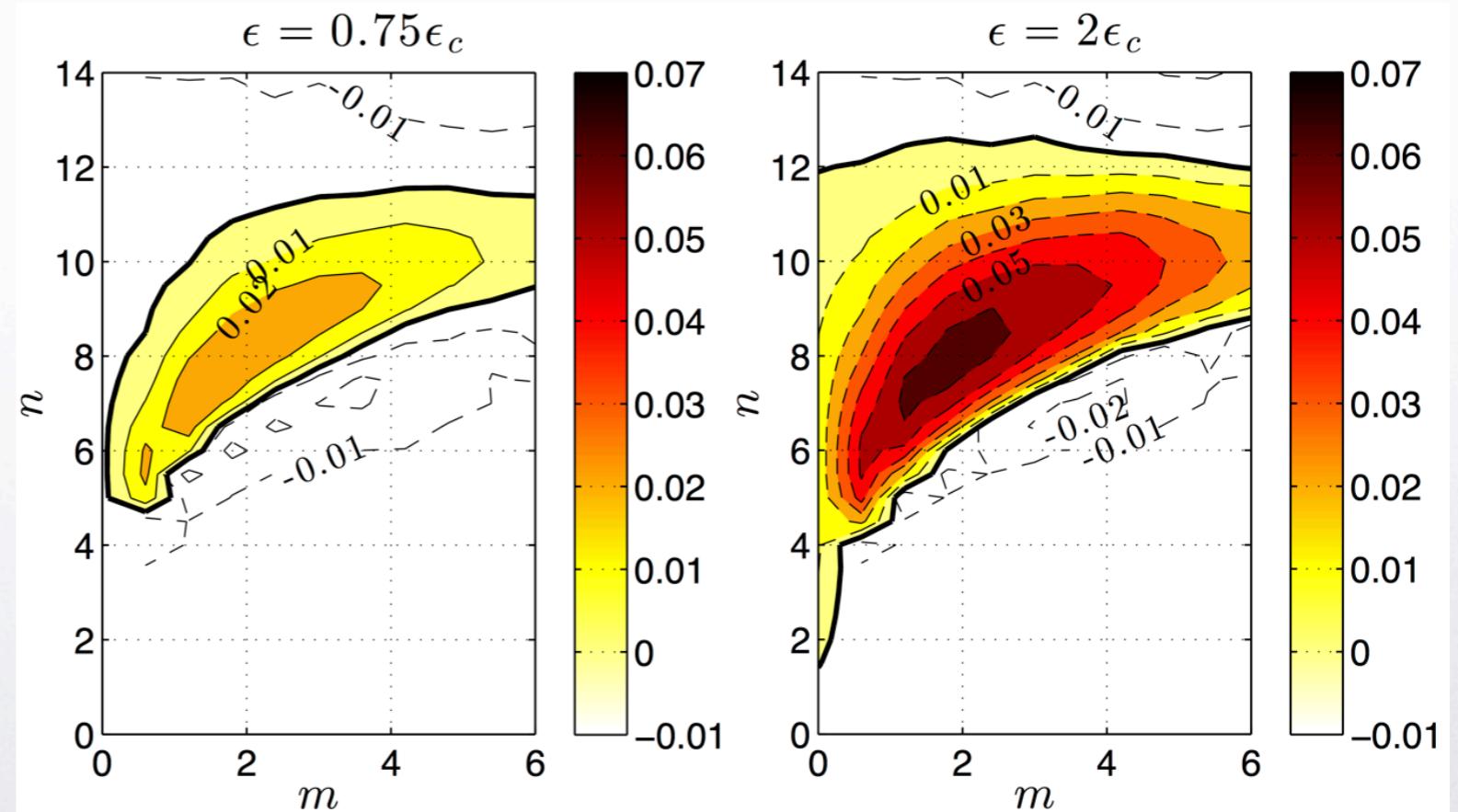


Non-zonal instabilities to homogeneous equilibrium

perturbations of the homogeneous equilibrium in form $\delta U = e^{i(mx+ny)}$

for these parameters non-zonal perturbations are more SSST unstable than jets ($m = 0$)

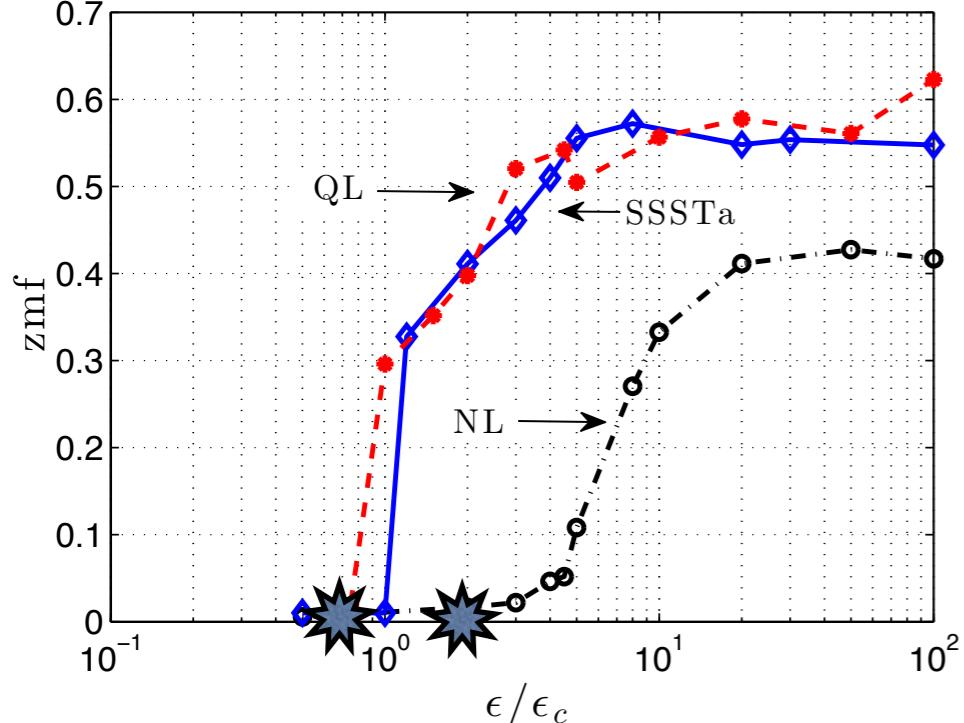
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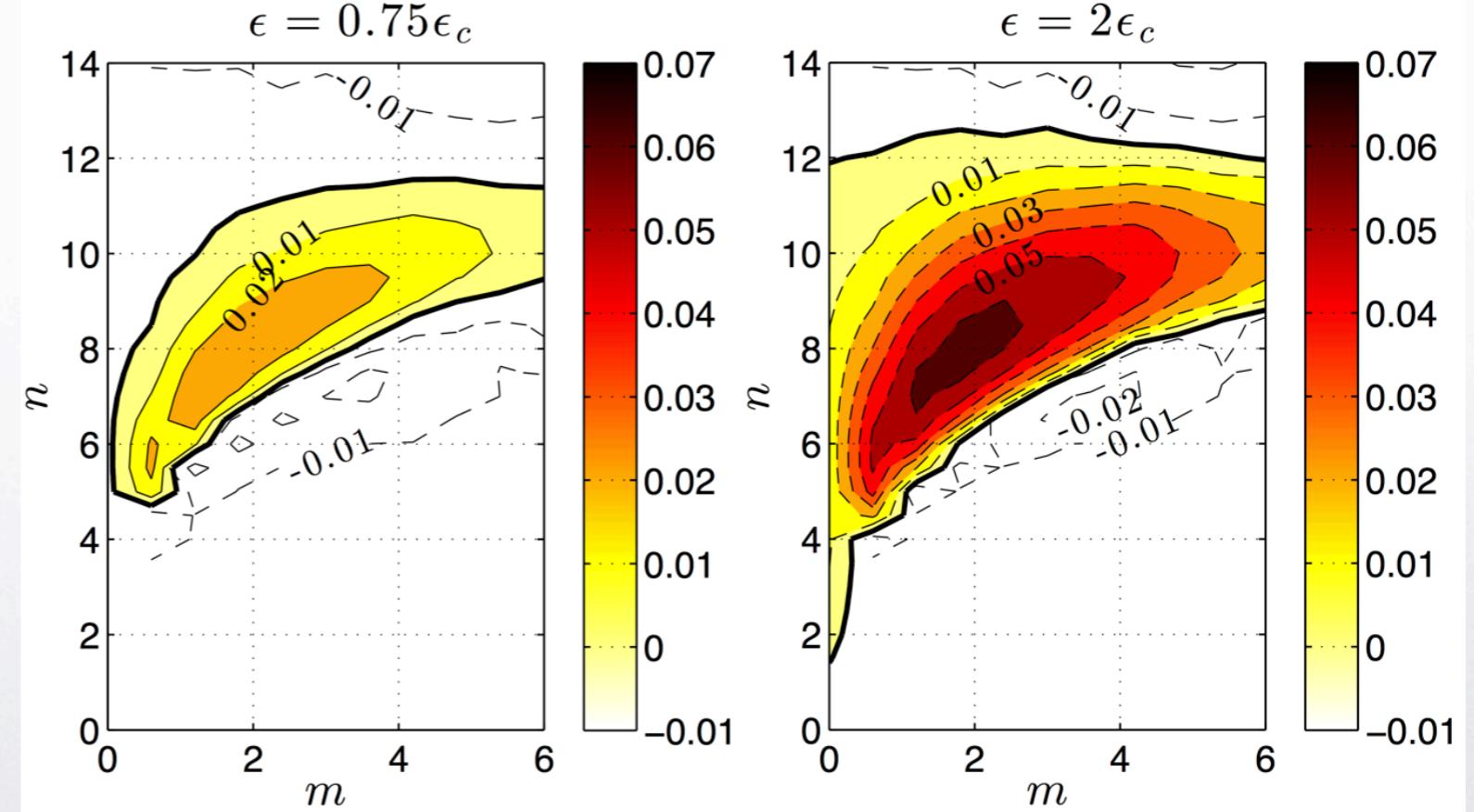


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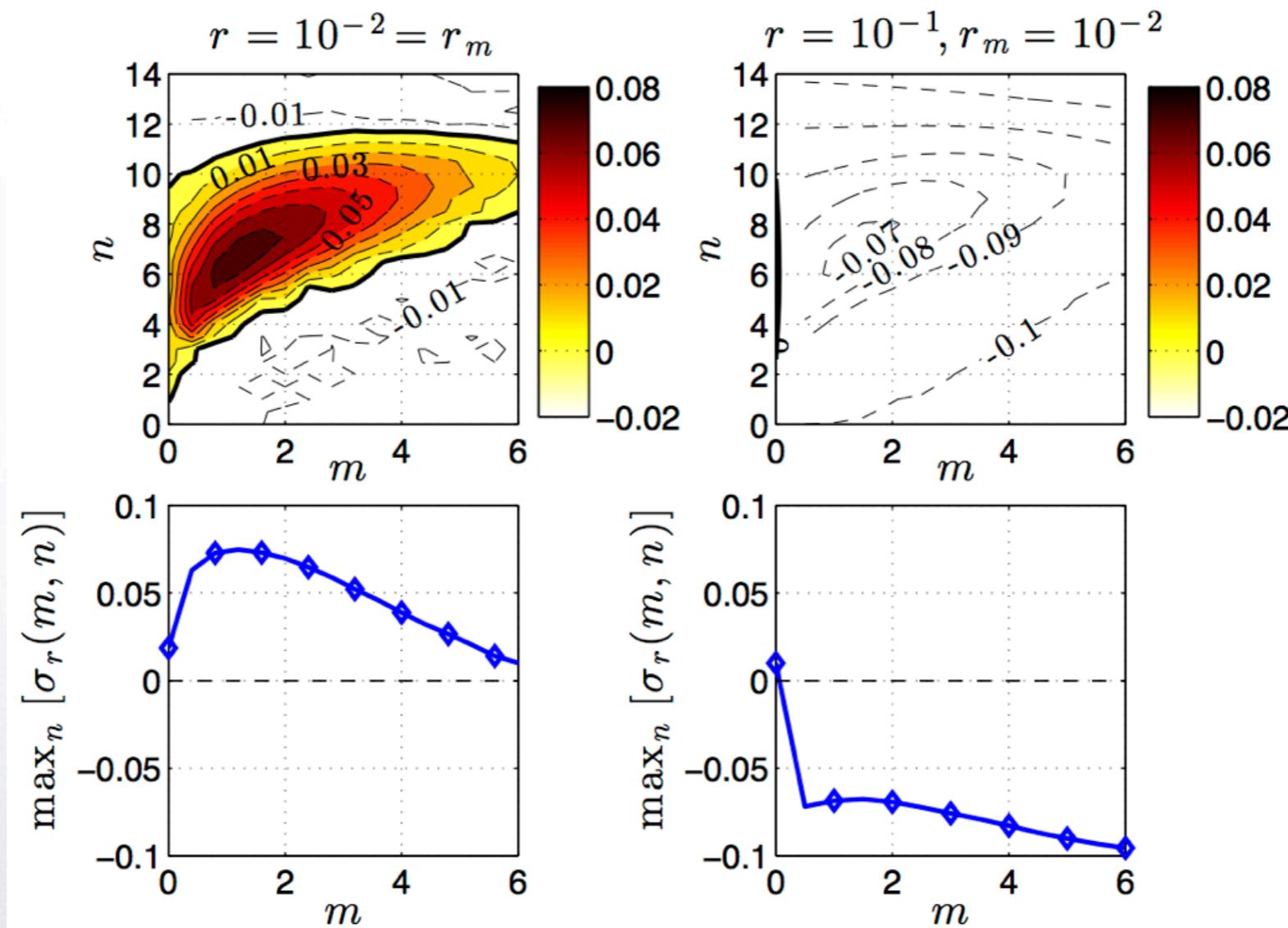


IRFh, $r = r_m = 10^{-2}$





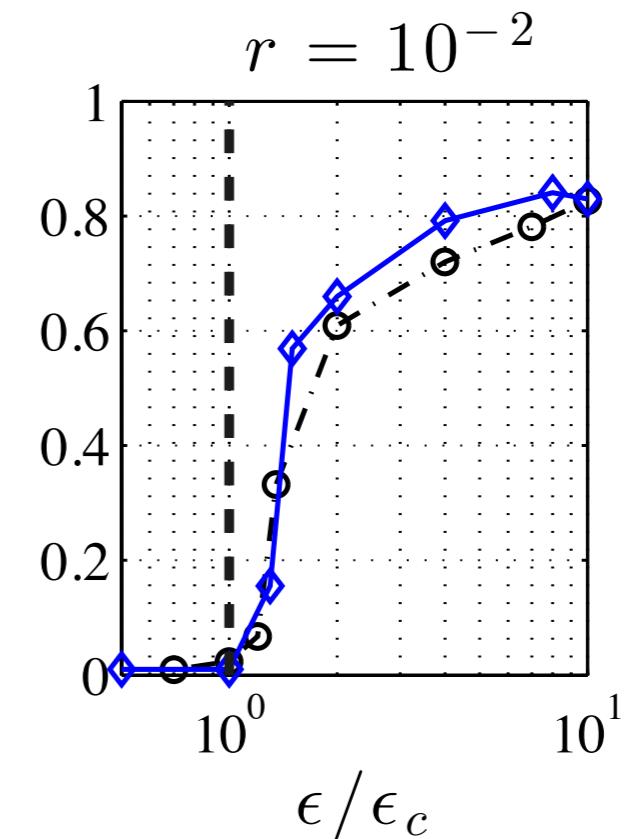
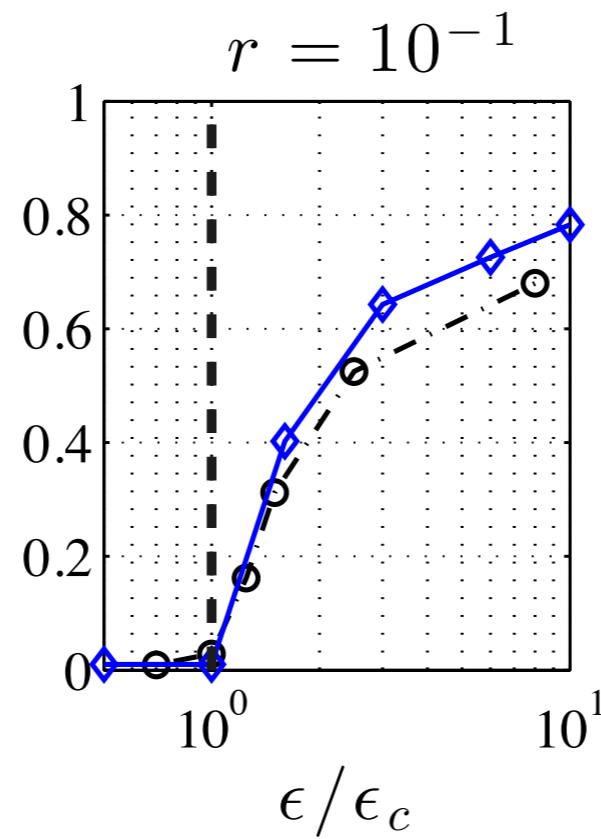
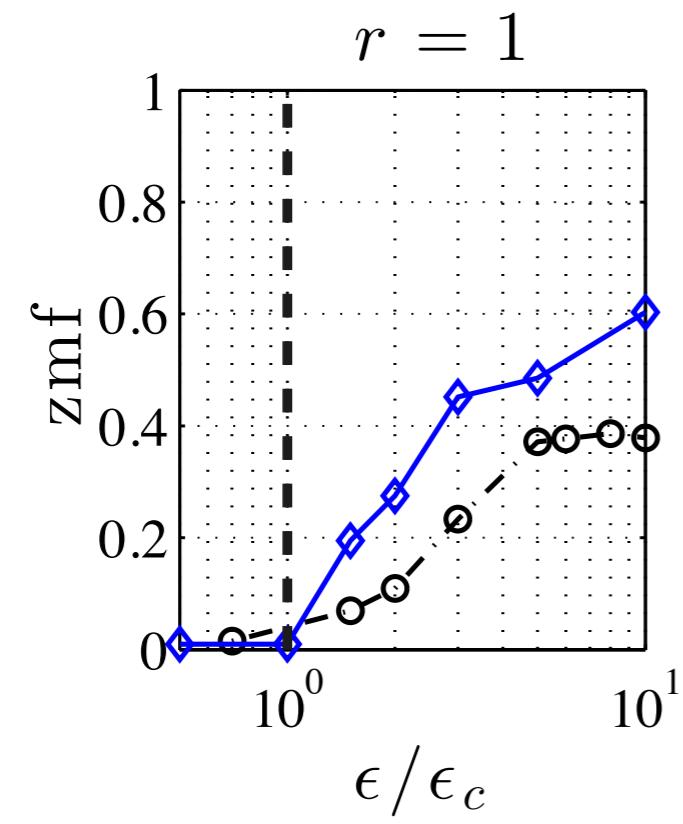
Depression of non-zonal instabilities using $r_m = r/10$





SSST captures the NL bifurcation

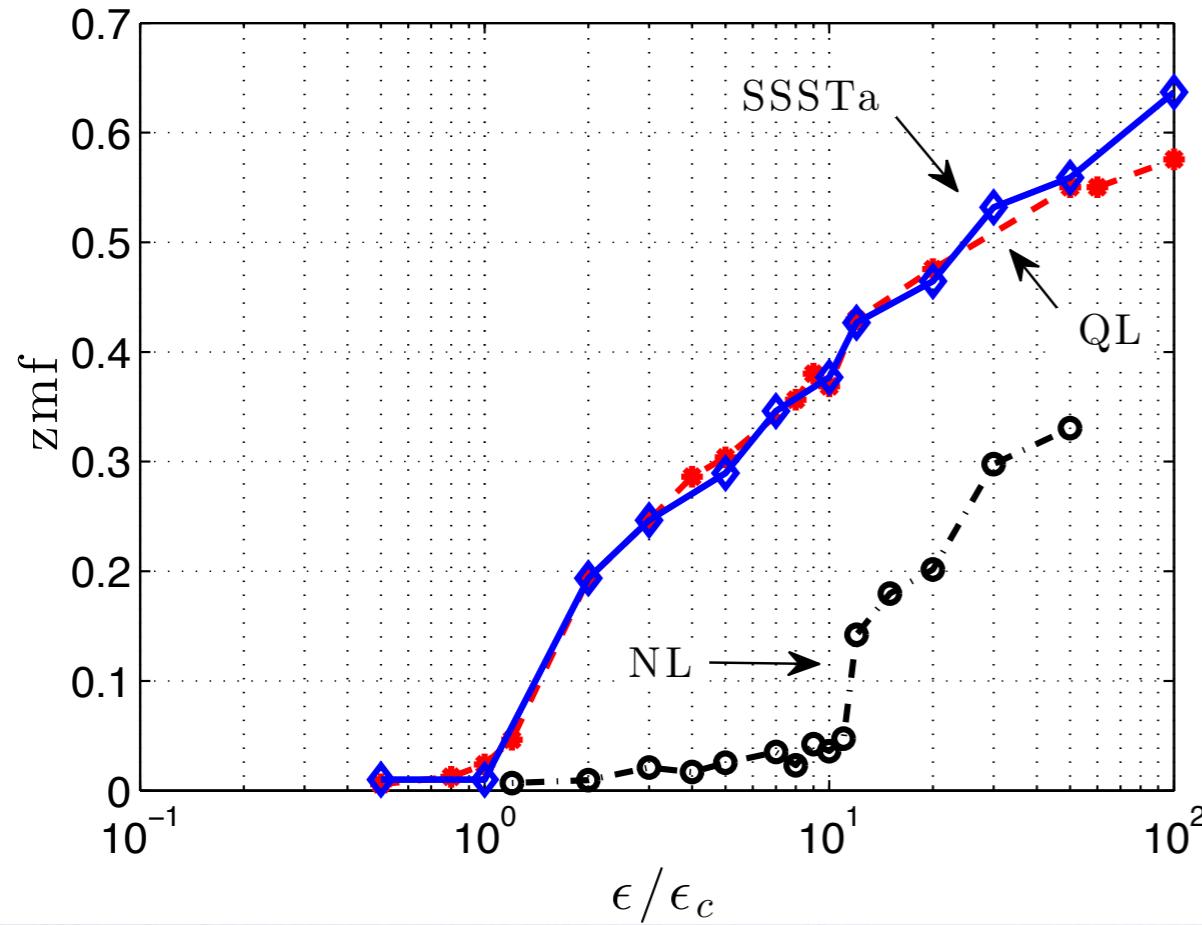
NIF, $r_m = r/10$



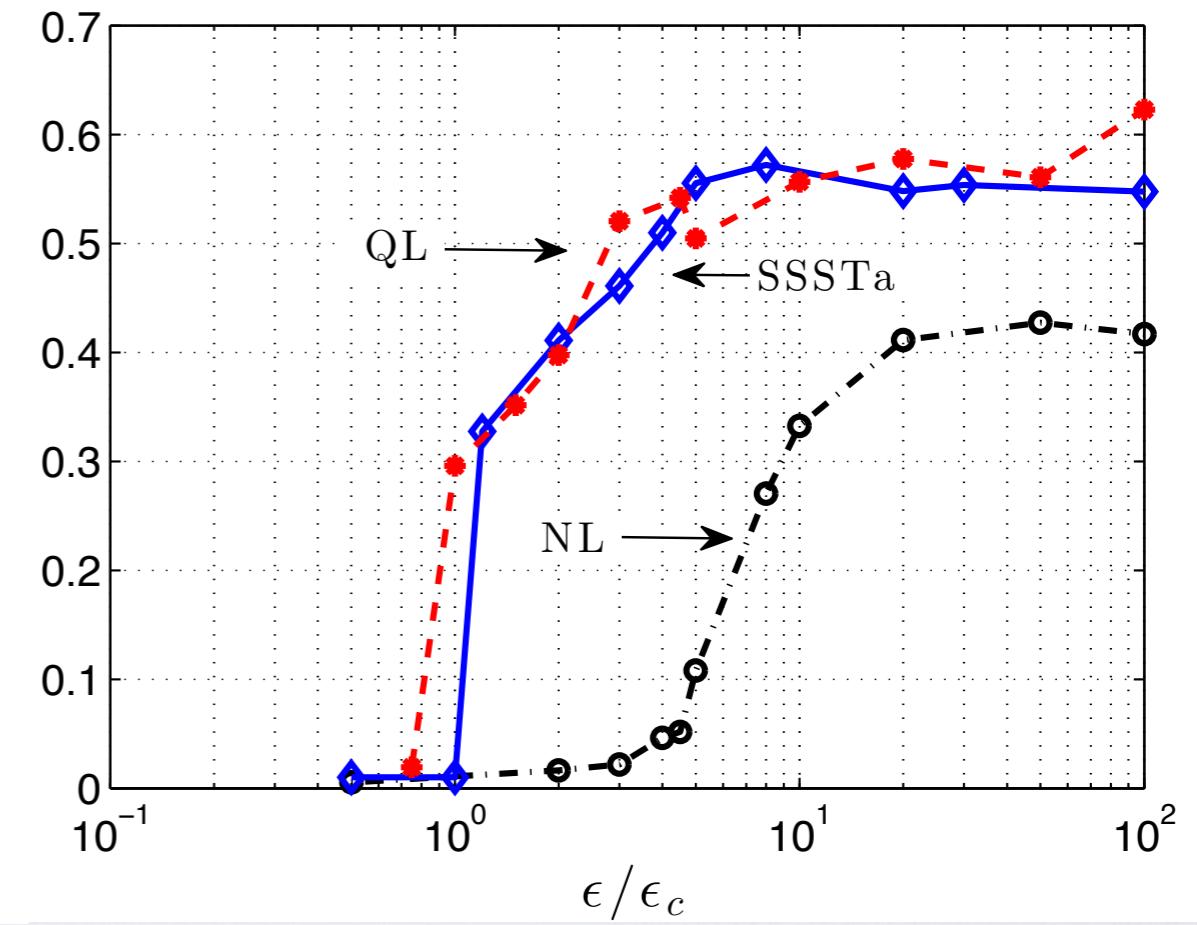


Bifurcation diagrams for $r = r_m$ (revisited)

NIF, $r = r_m = 10^{-2}$



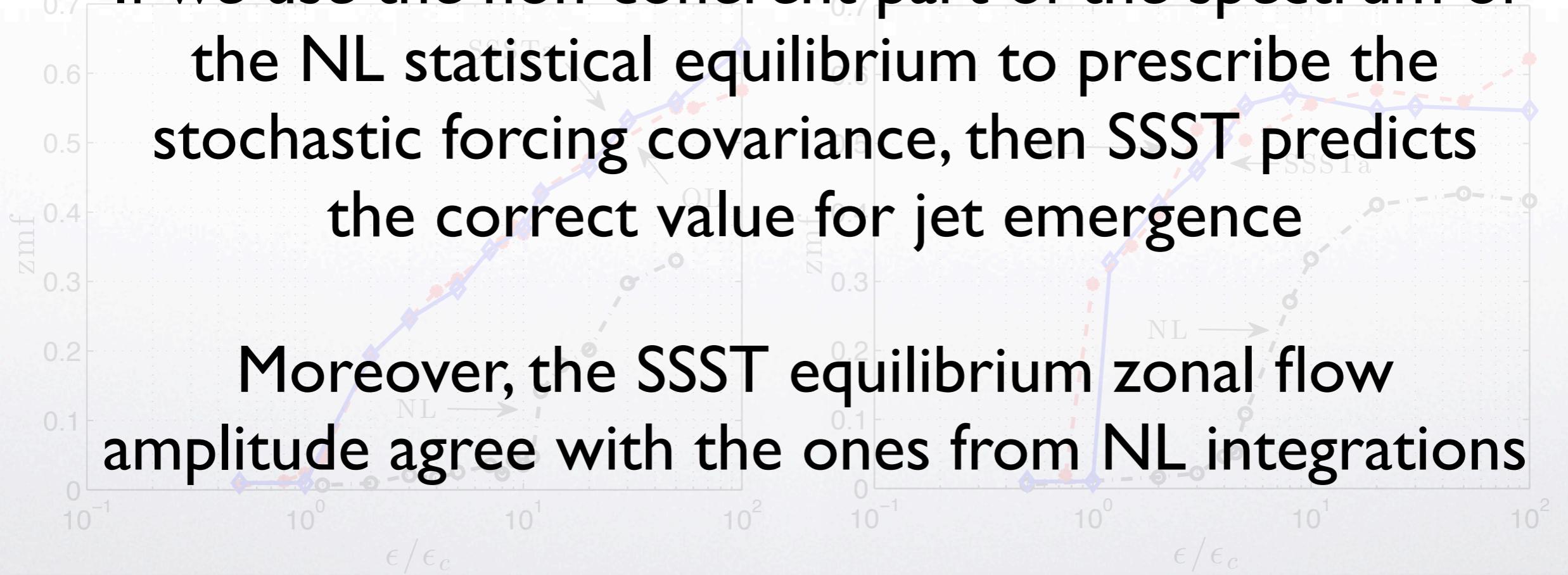
IRFh, $r = r_m = 10^{-2}$





Bifurcation diagrams for $r = r_m$ (revisited)

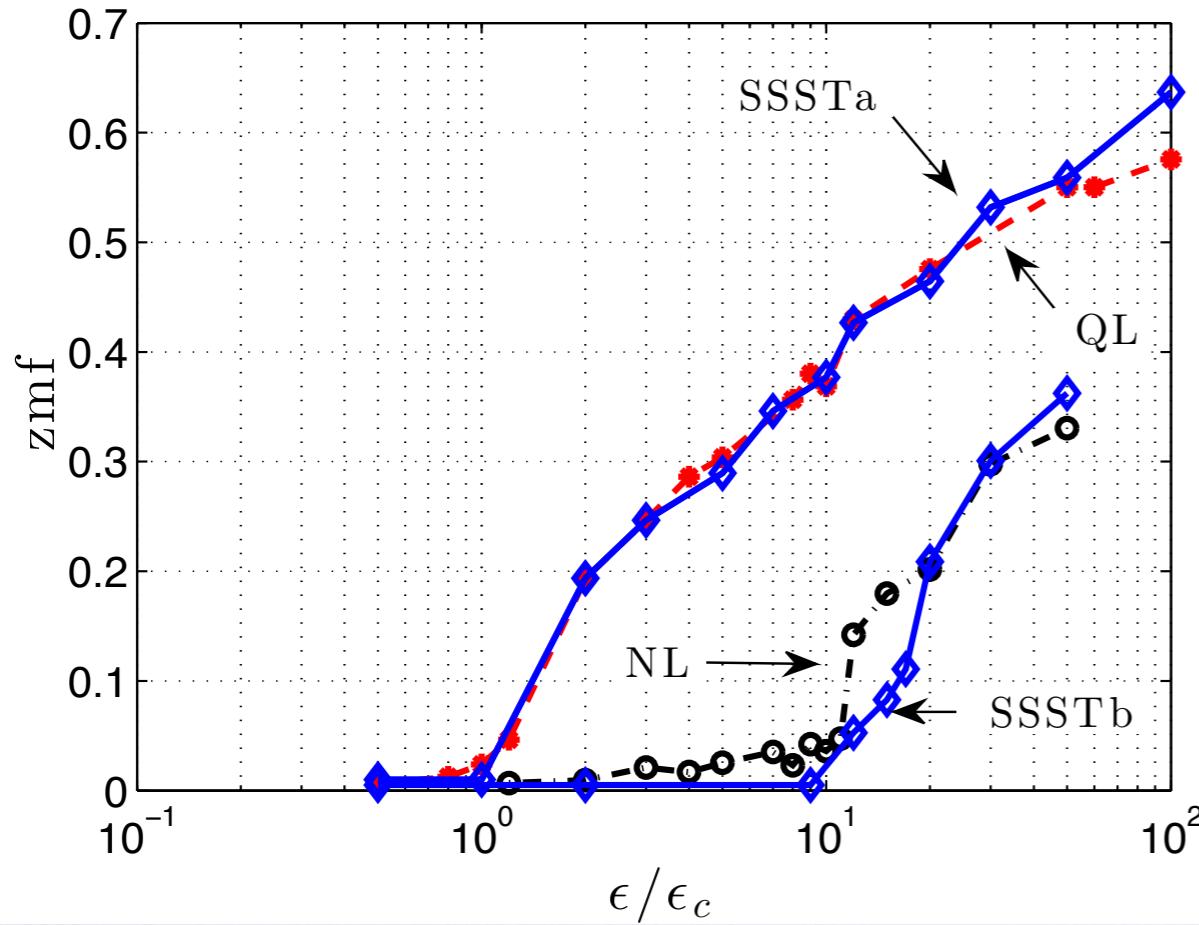
If we use the non-coherent part of the spectrum of the NL statistical equilibrium to prescribe the stochastic forcing covariance, then SSST predicts the correct value for jet emergence



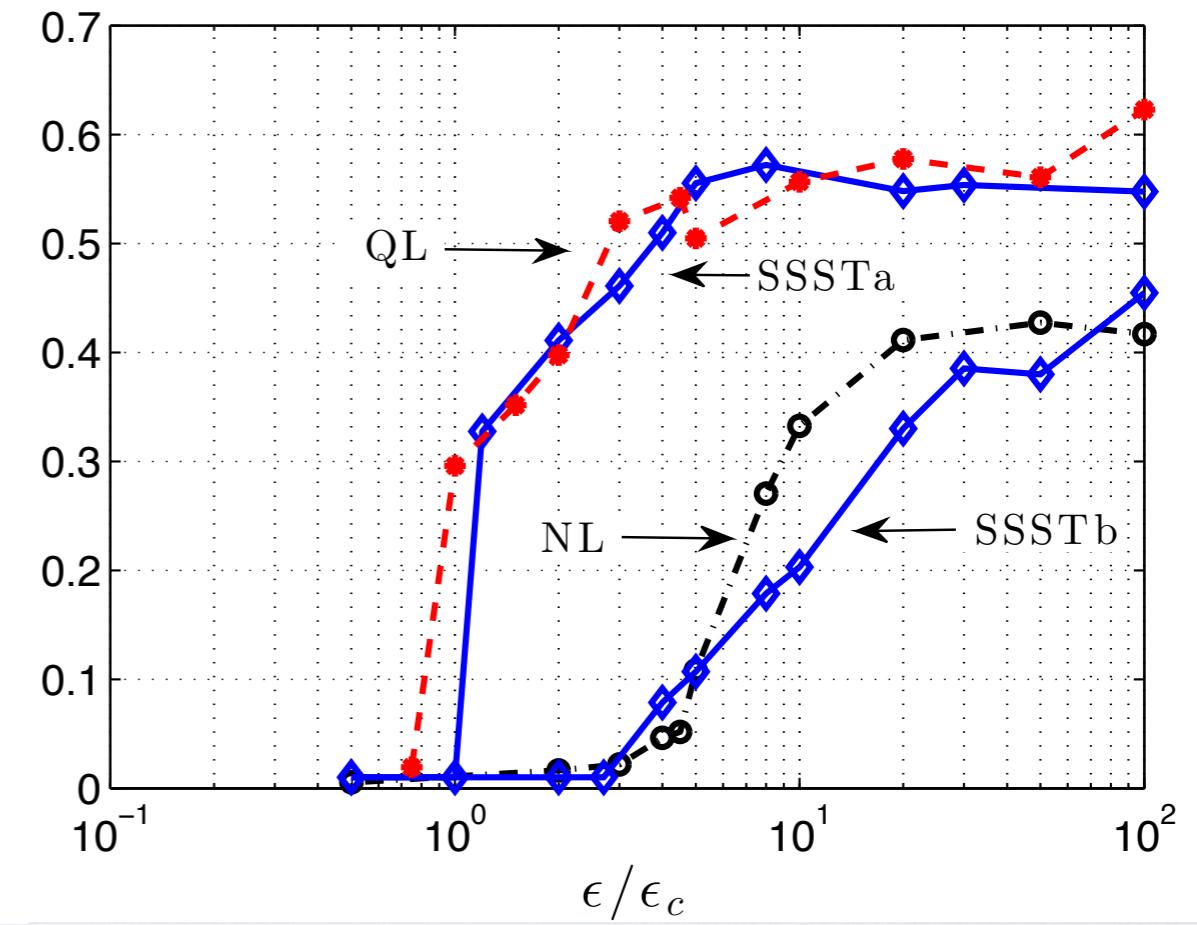


Bifurcation diagrams for $r = r_m$ (revisited)

NIF, $r = r_m = 10^{-2}$



IRFh, $r = r_m = 10^{-2}$



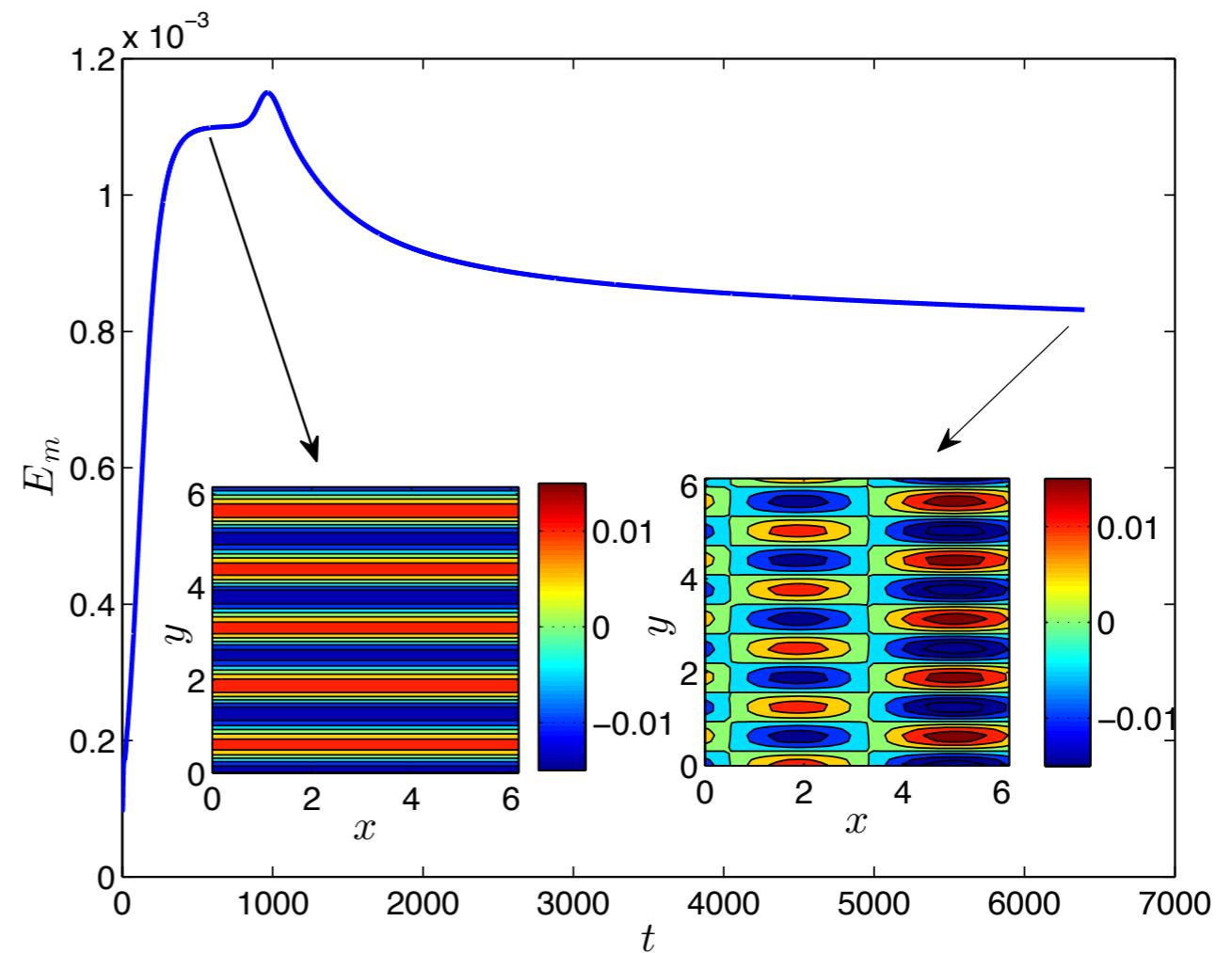


Generalized SSST admits non-zonal equilibria

IRFh at $\epsilon = 2\epsilon_c$

SSST at $\epsilon = 2\epsilon_c$ forms jets

generalized SSST initially goes to a jet equilibrium, which is SSST unstable and switches to a NZCS



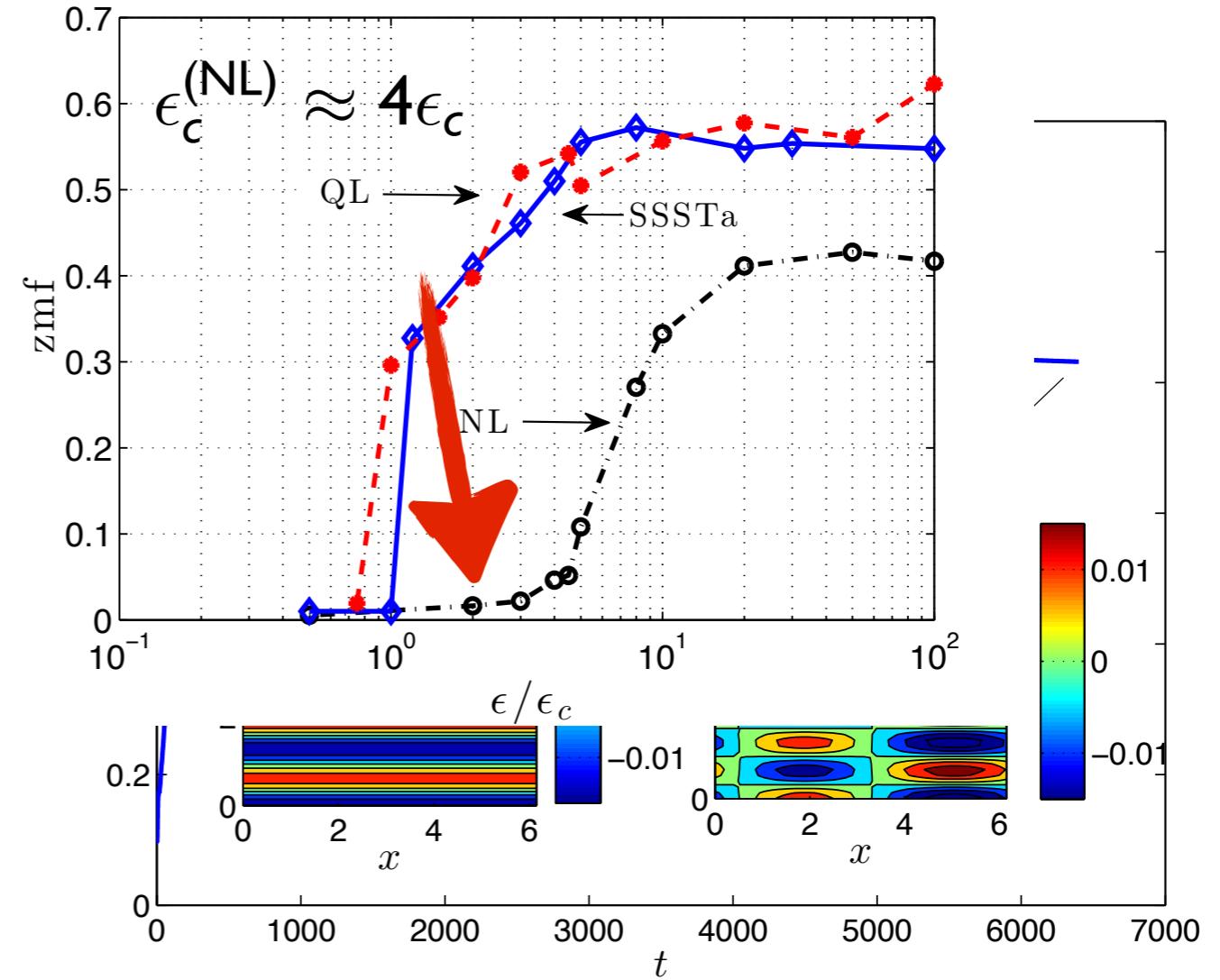


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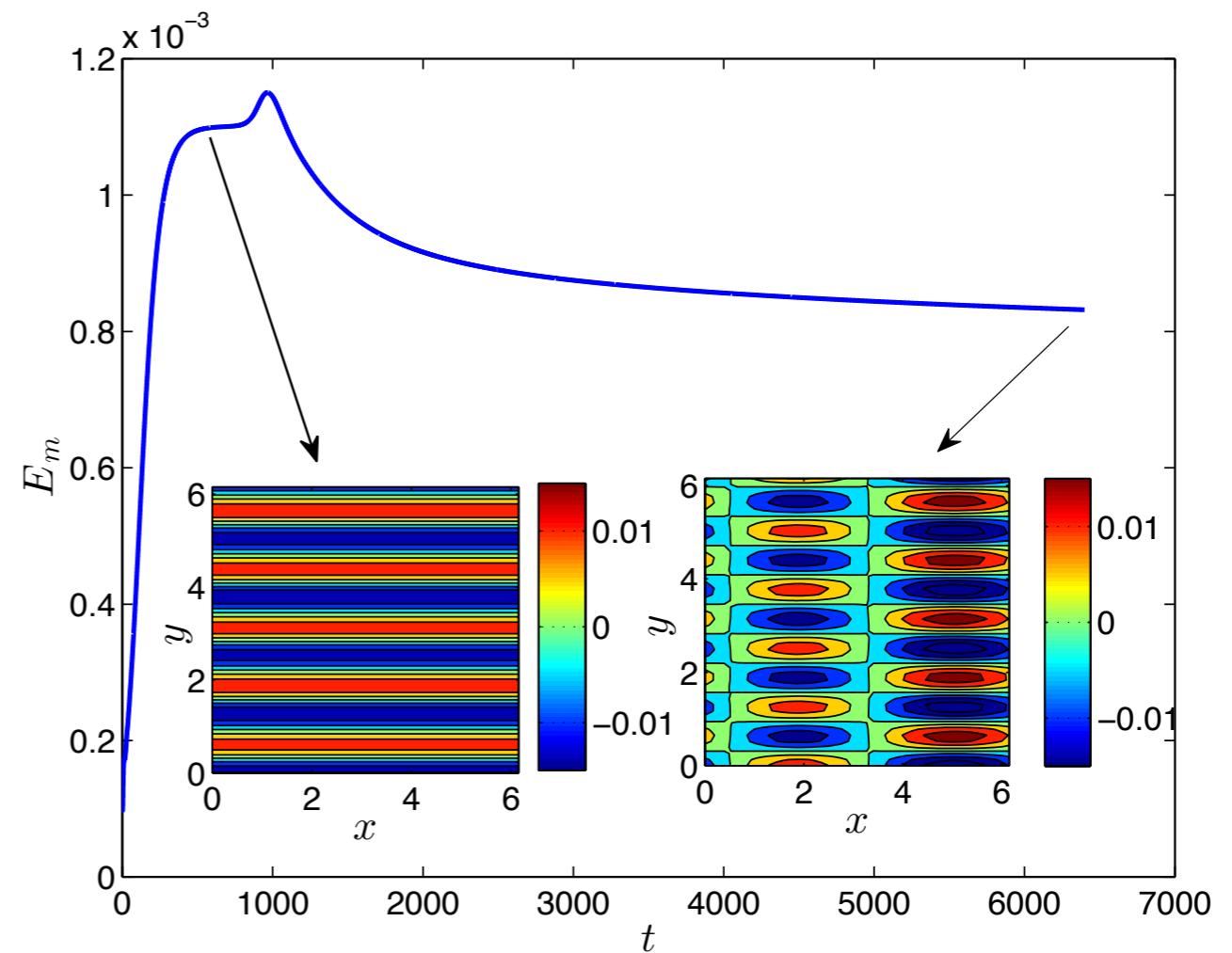


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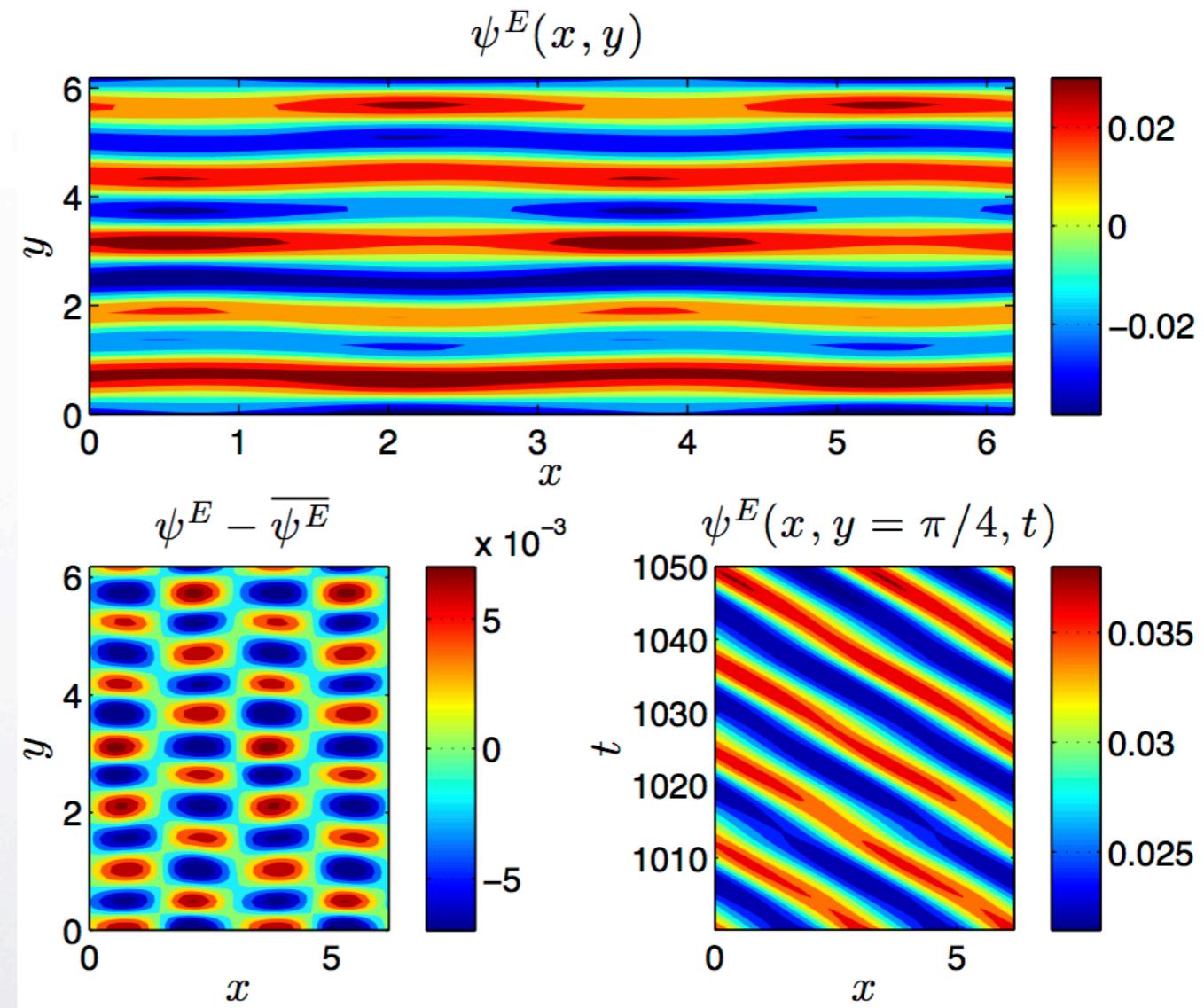




Generalized SSST equilibria show both zonal and non-zonal spectral peaks

IRFh at $\epsilon = 6\epsilon_c$

generalized SSST admits equilibria with zonal as well as non-zonal spectral components

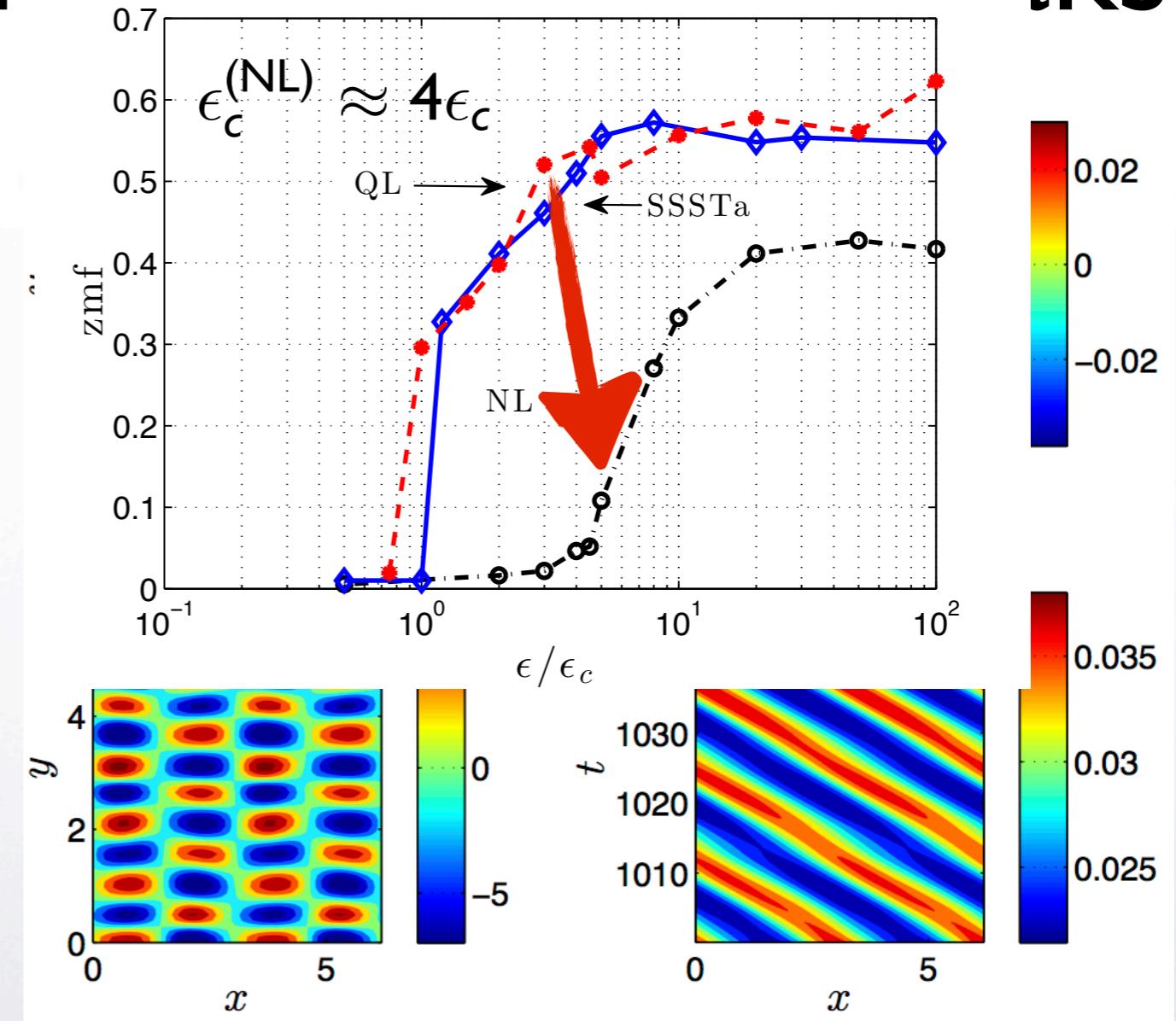




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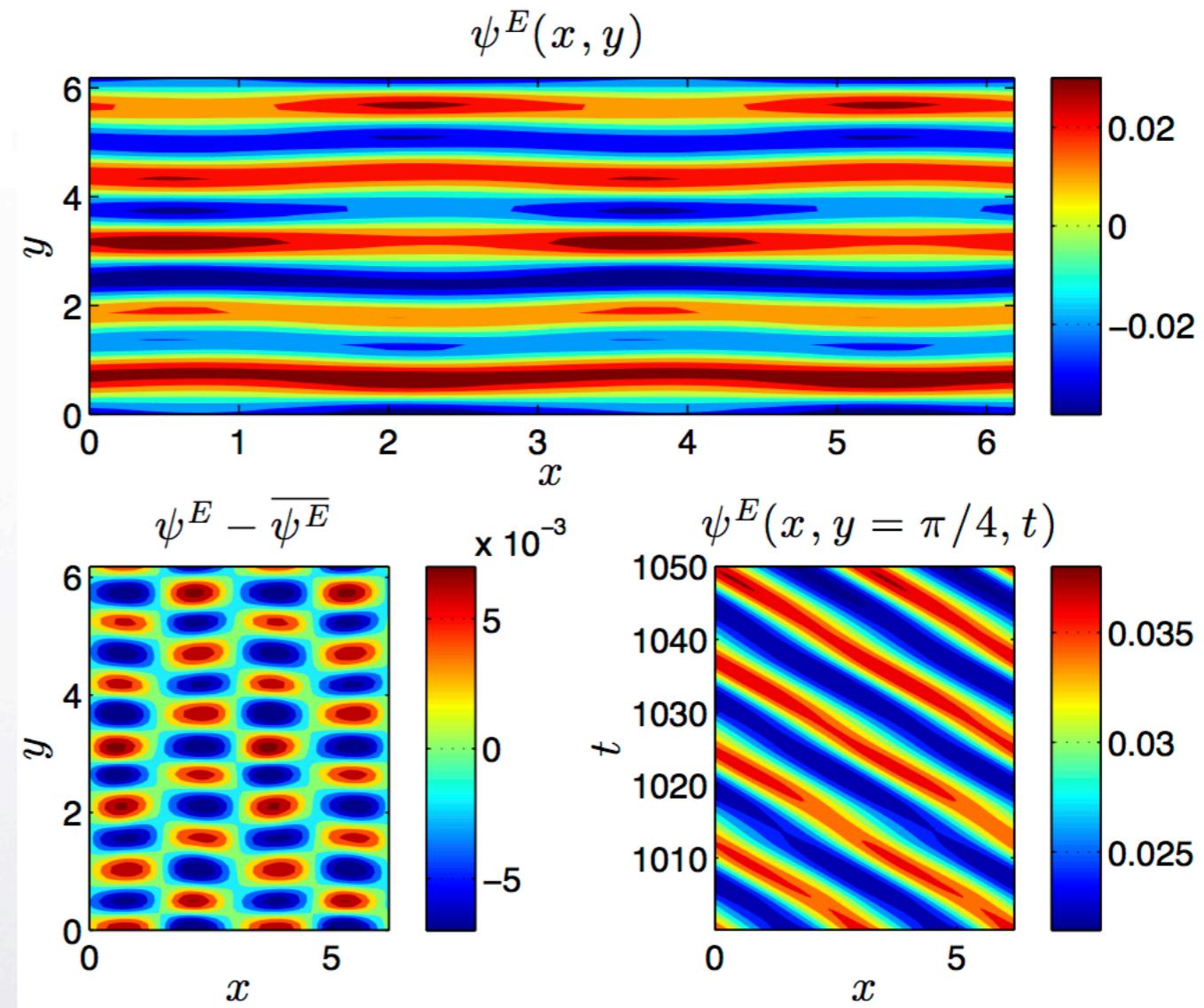




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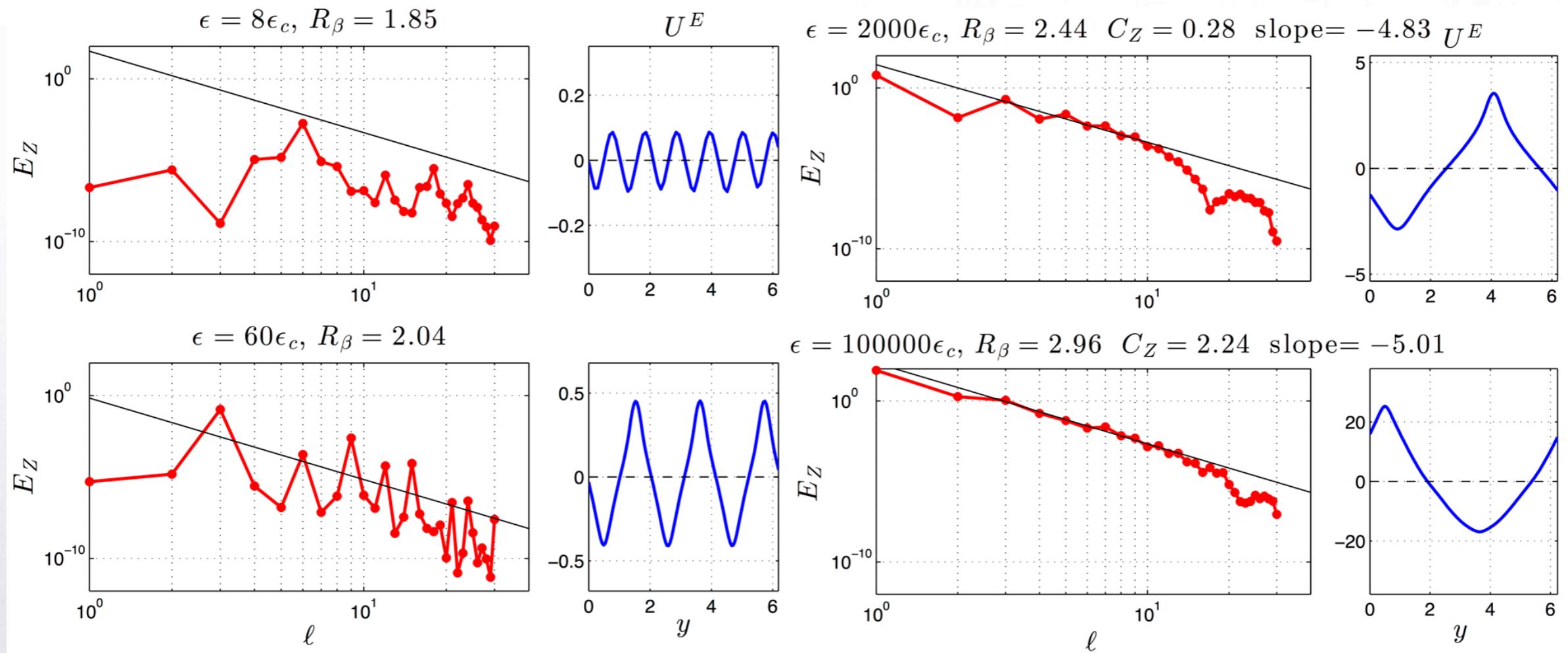
Things to discuss

- ▶ What is the shape of the SSST mean flow equilibria? Does it capture the NL jet shape?
- ▶ Jet mergings / why do they occur?
(is it because of U becomes hydrodynamically unstable?)
- ▶ PV staircases (?)



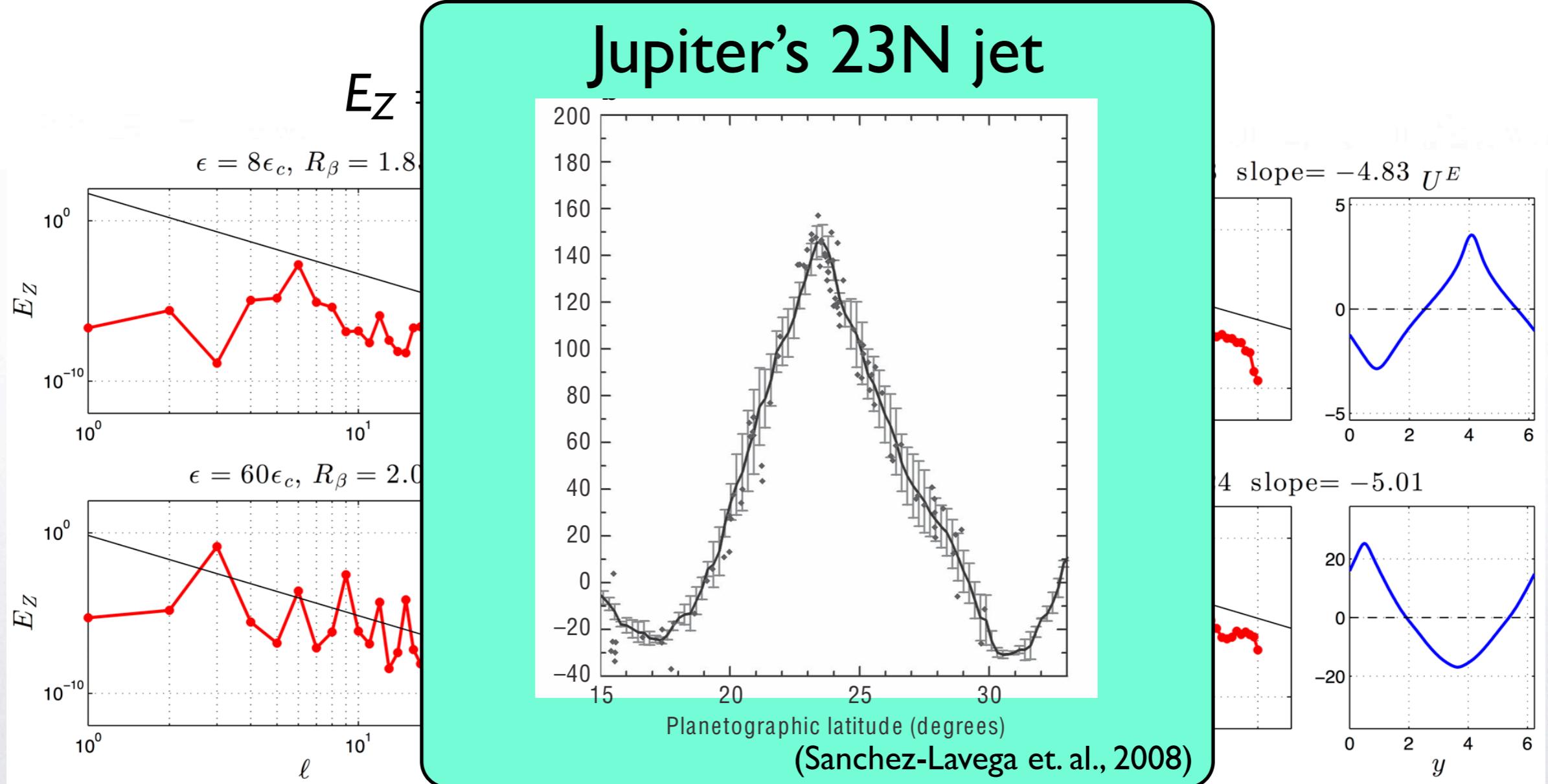
SSST equilibrium zonal flow spectral structure

$$E_Z = E(k=0, \ell) = C_Z \beta^2 \ell^{-5}, C_Z \approx 0.5$$





SSST equilibrium zonal flow spectral structure



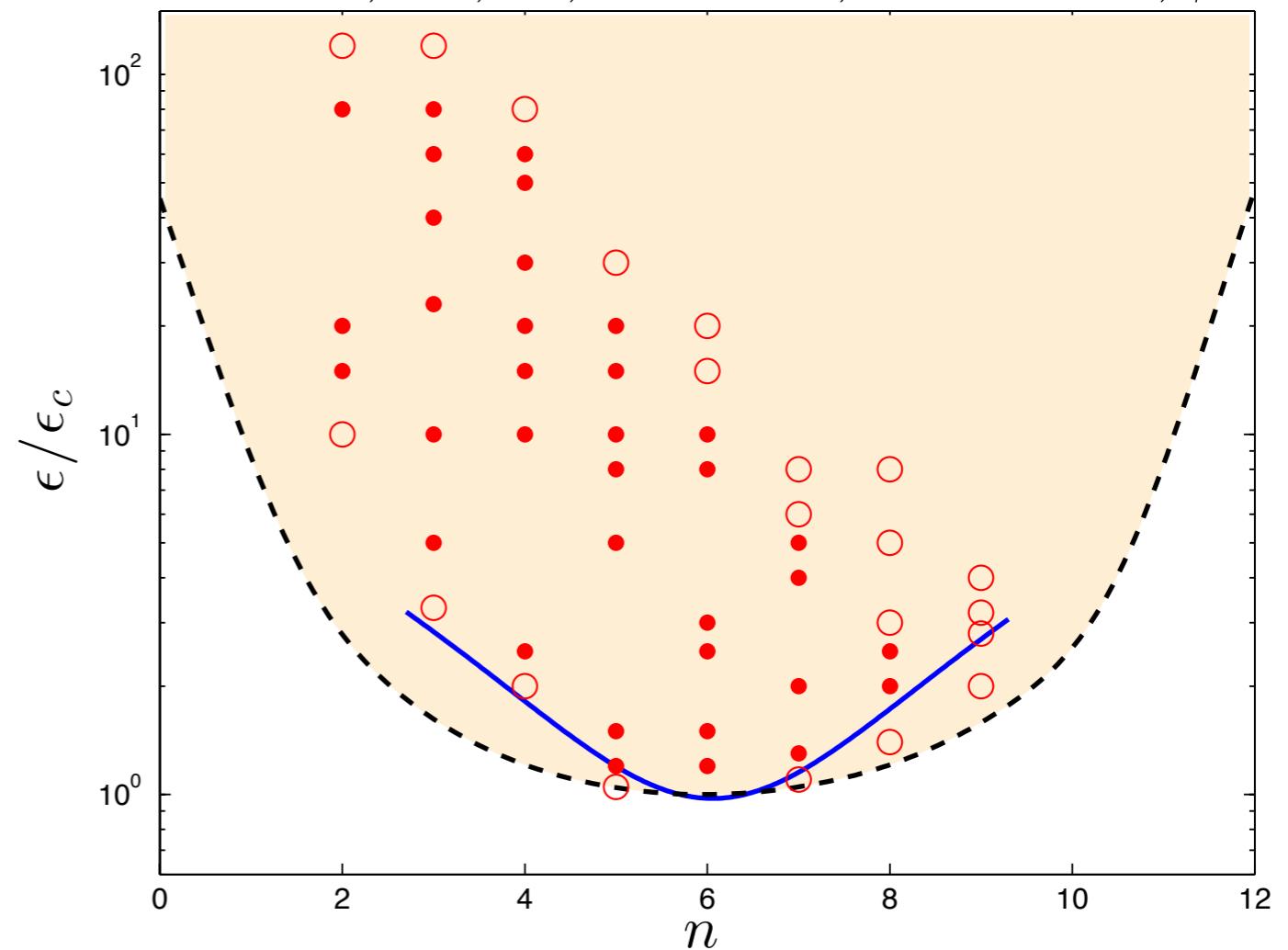


Jet merging is due to SSST instability

SSST equilibria
(stable & unstable)
are hydrodynamically stable

SSST instability comes prior
to hydrodynamic instability
as we move upwards in a
constant n axis

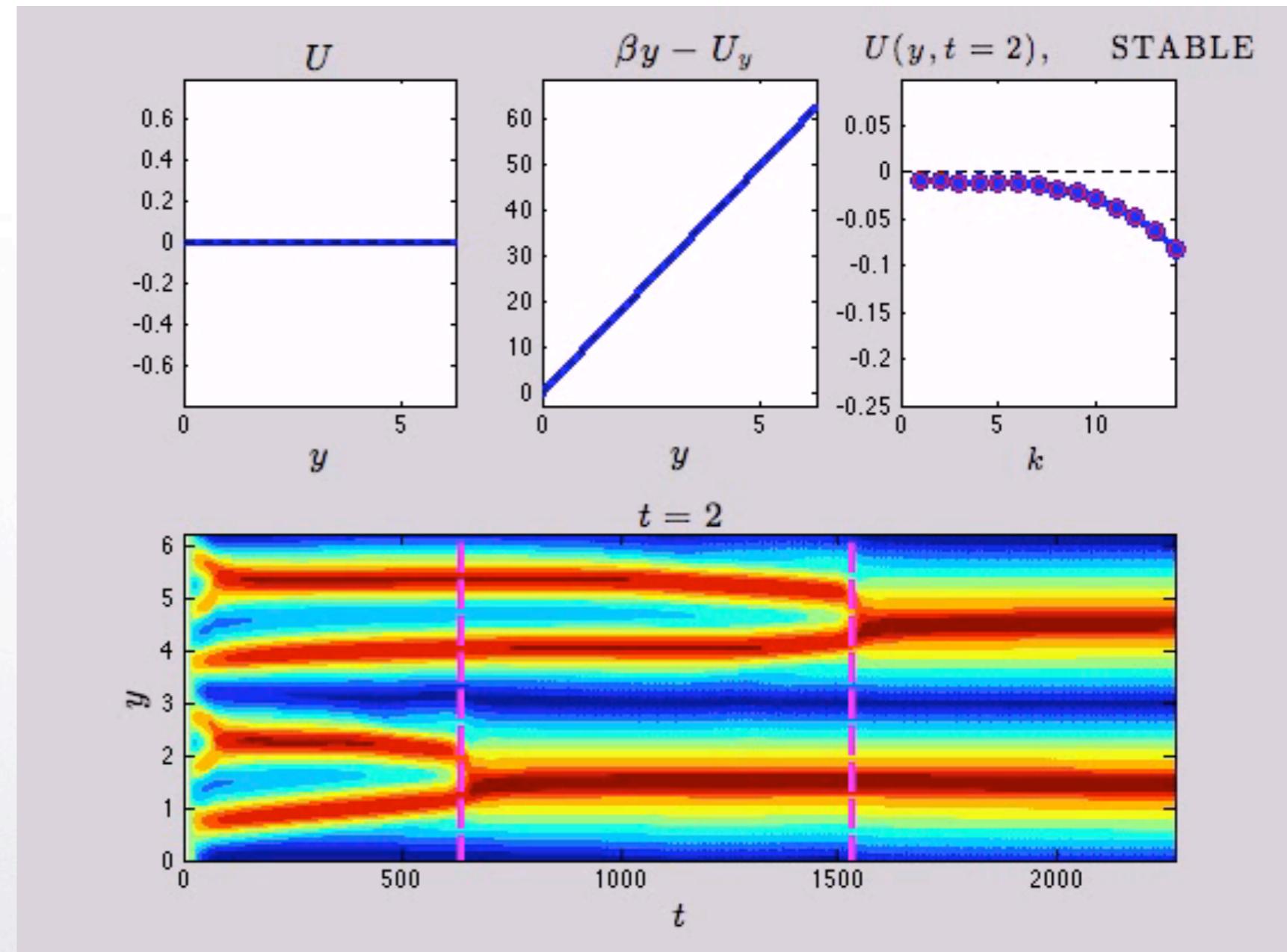
NIF at $k = 1, \dots, 14$, $r = 10^{-1}$, $r_m = 10^{-2}$, $\beta = 10$





Jet merging is due to SSST instability

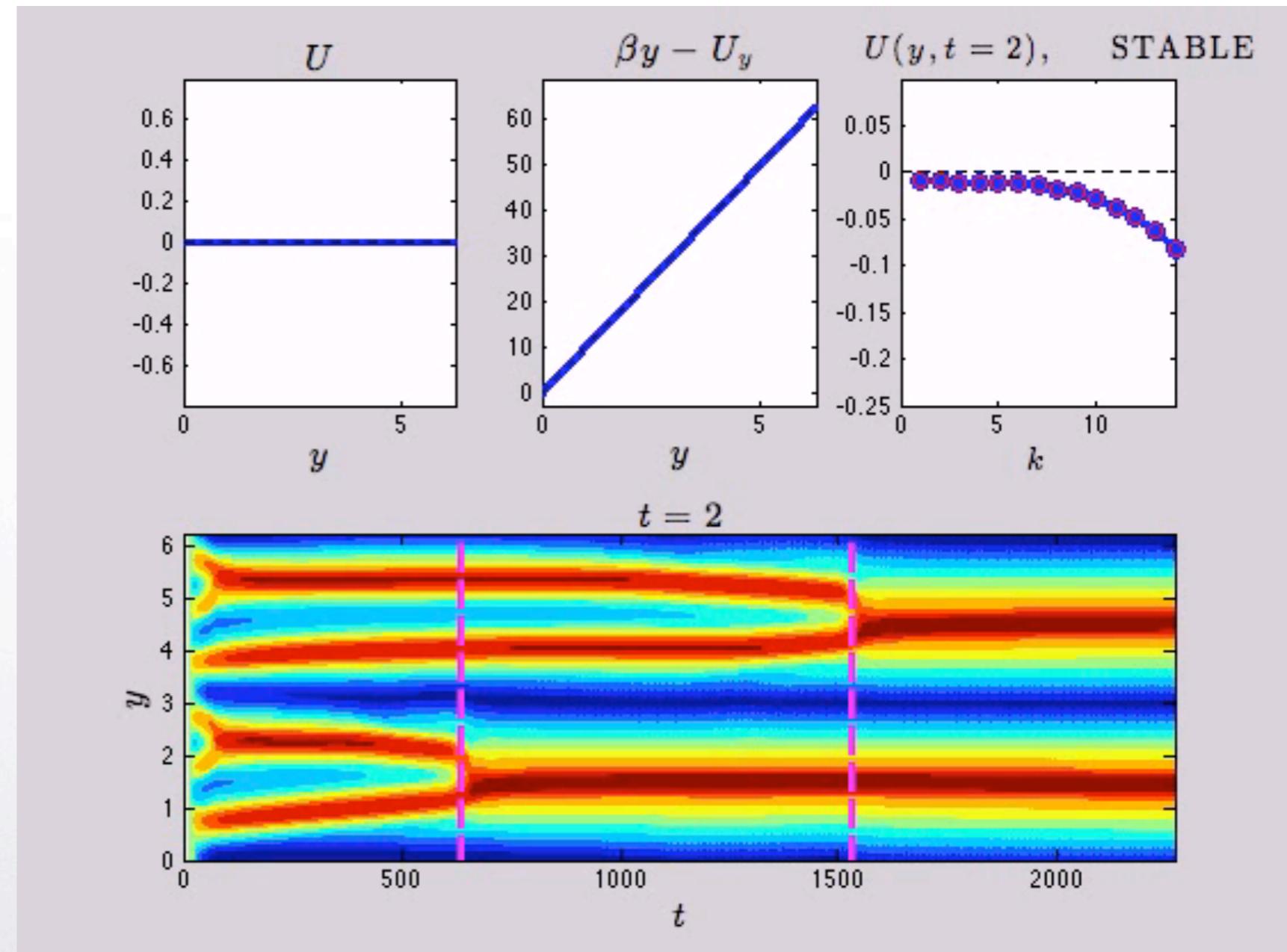
NIF at $k = 1, \dots, 14$
 $r = r_m = 10^{-2}$, $\beta = 10$
 $\epsilon = 100\epsilon_c$





Jet merging is due to SSST instability

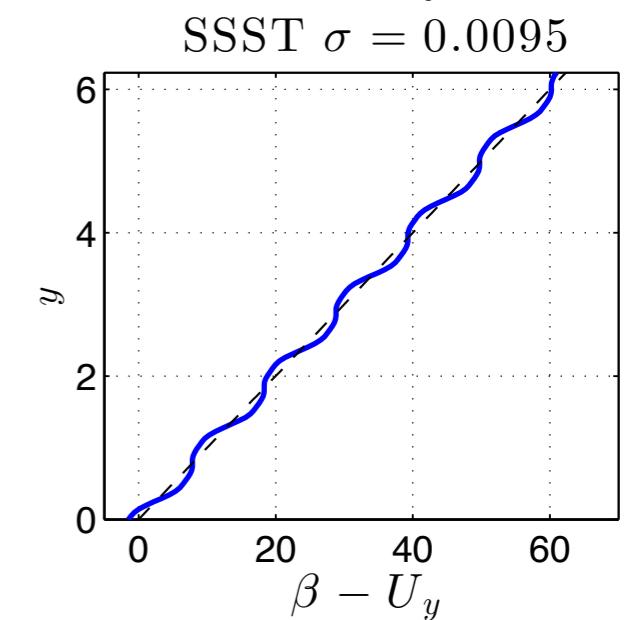
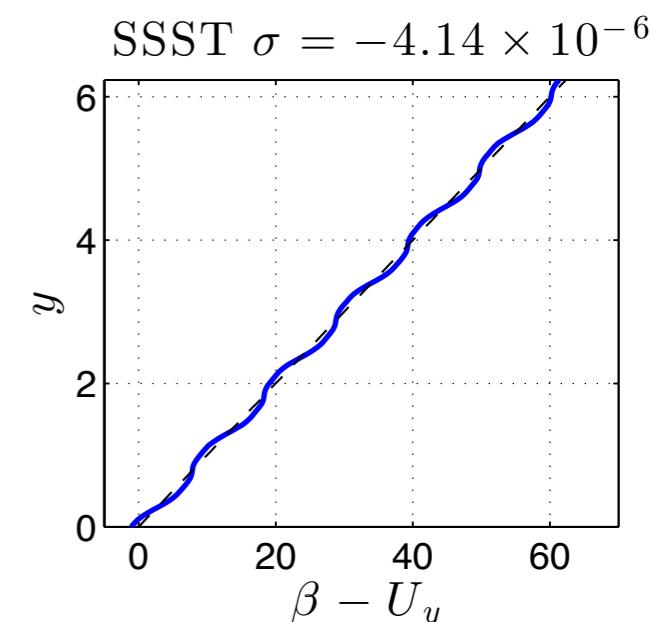
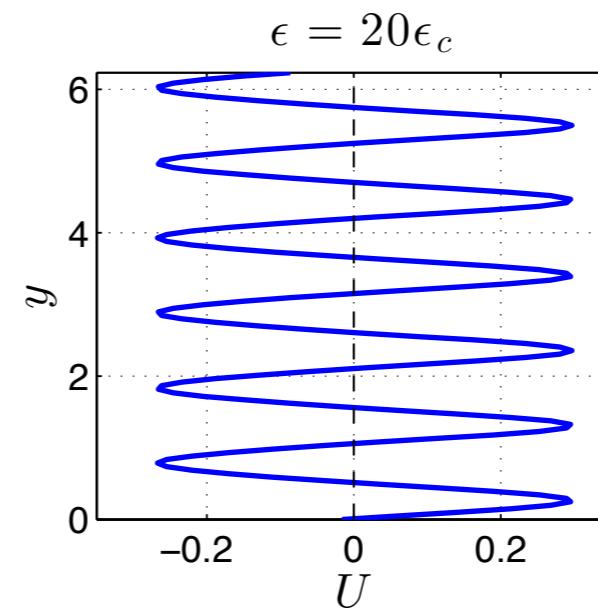
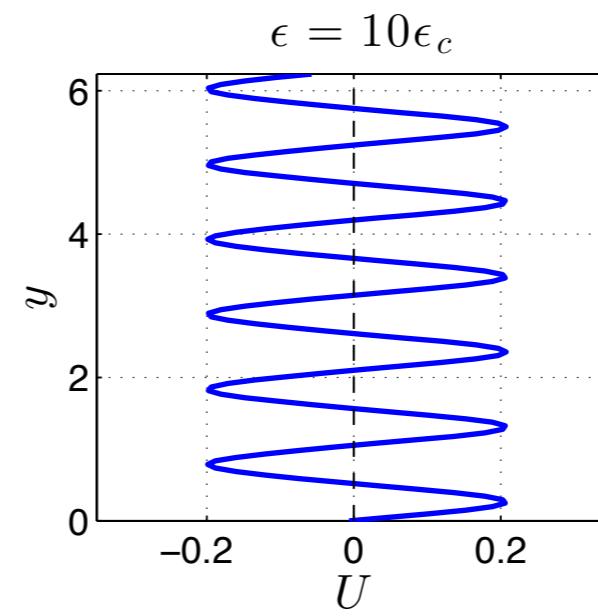
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 $\epsilon = 100\epsilon_c$





PV Staircases

NIF, $k = 1, \dots, 14$
 $r = 0.1, r_m = r/10, \beta = 10$

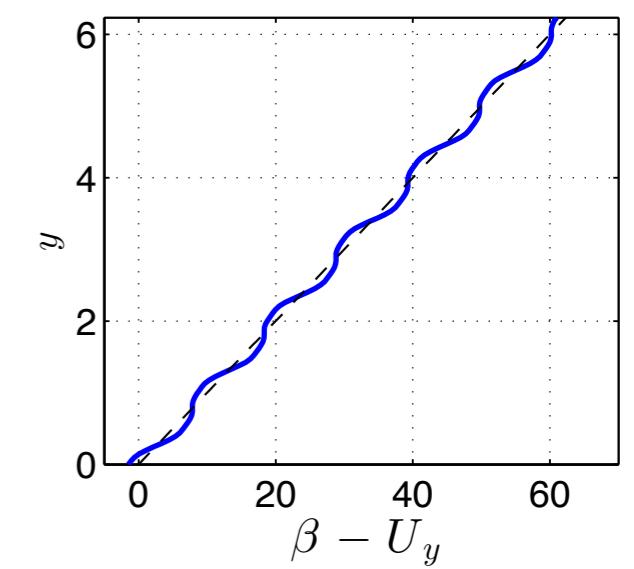
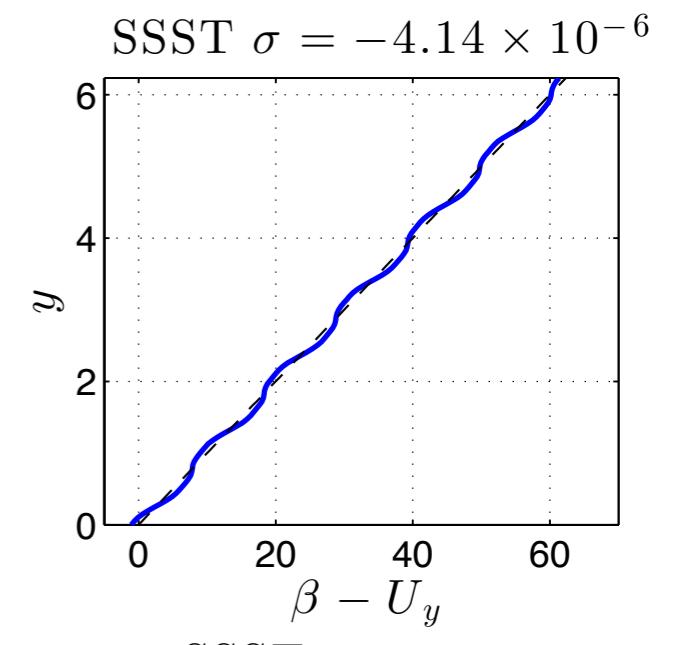
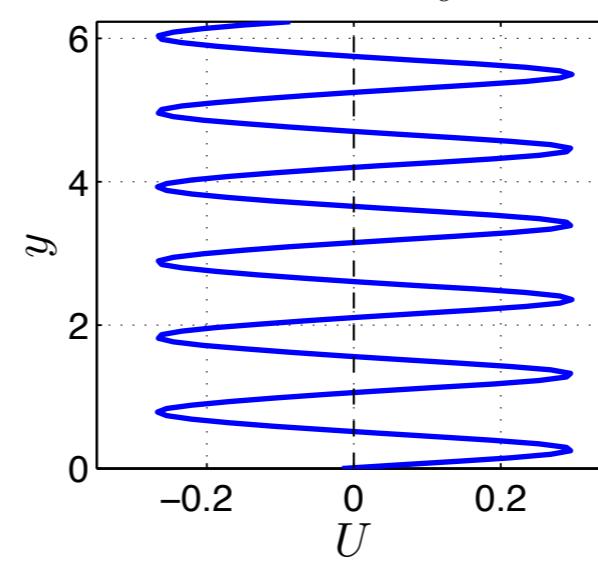
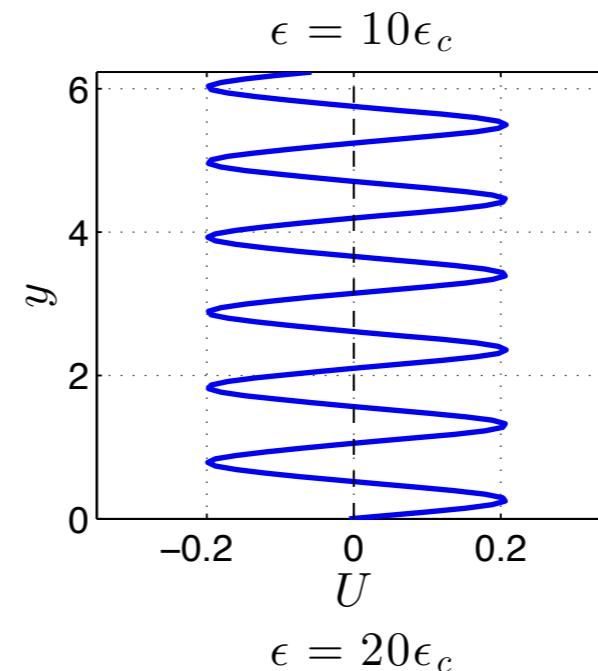




PV Staircases

NIF, $k = 1, \dots, 14$
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SSST stable

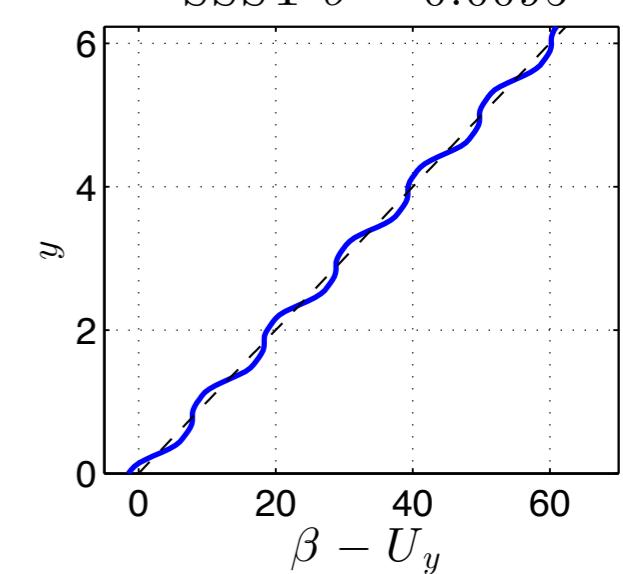
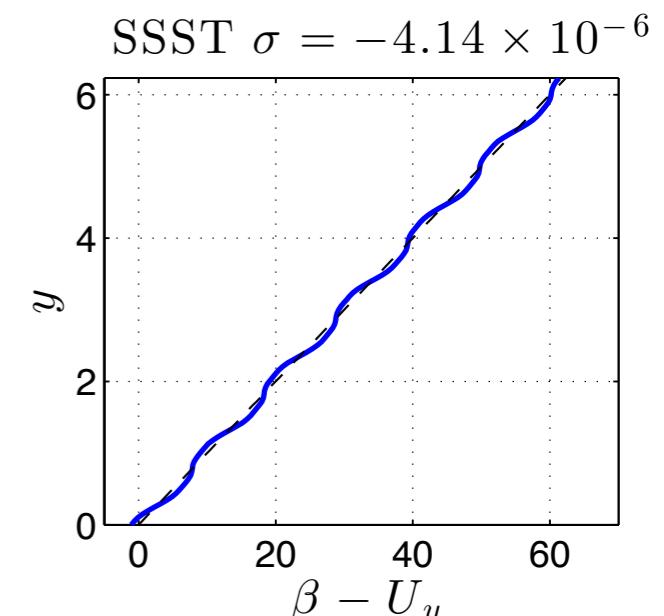
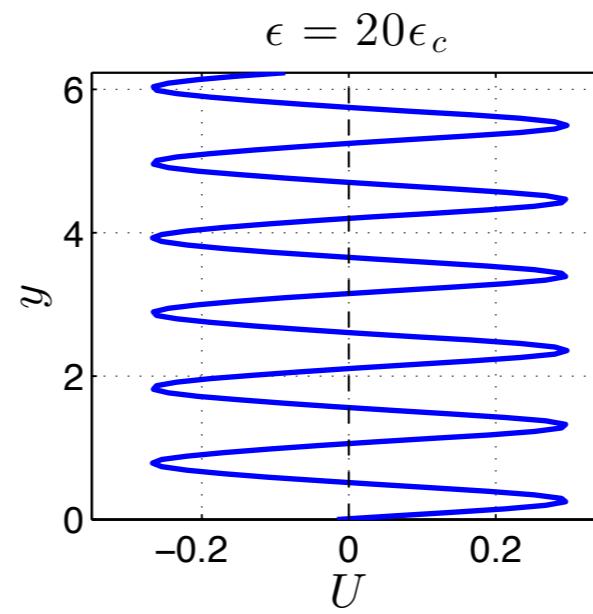
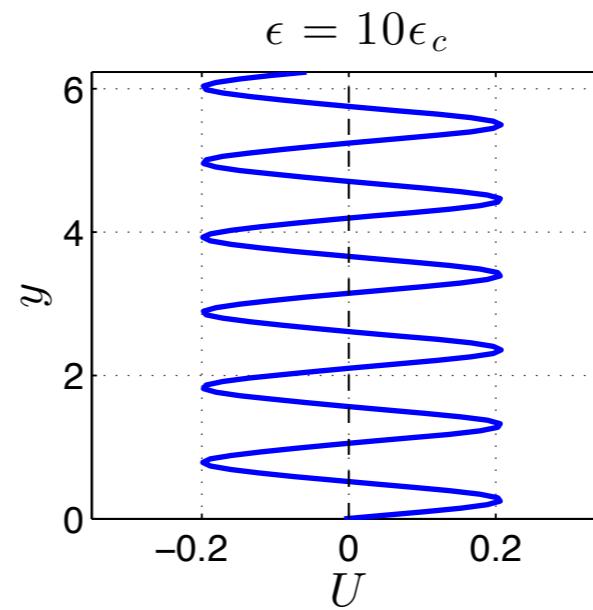




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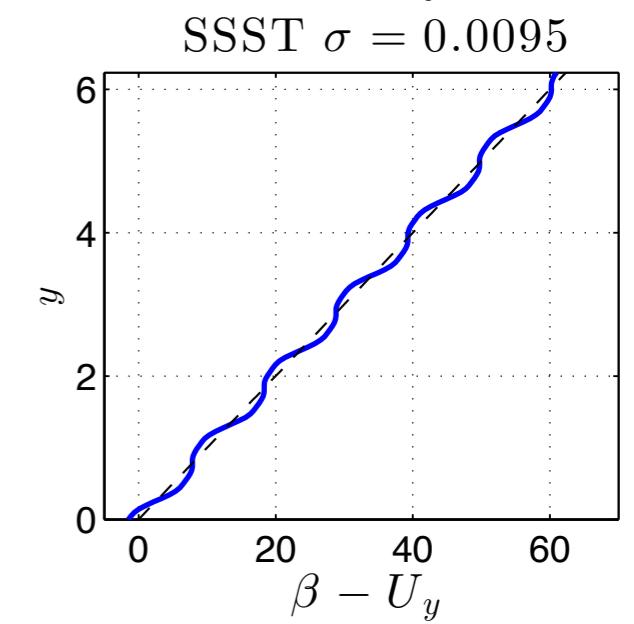
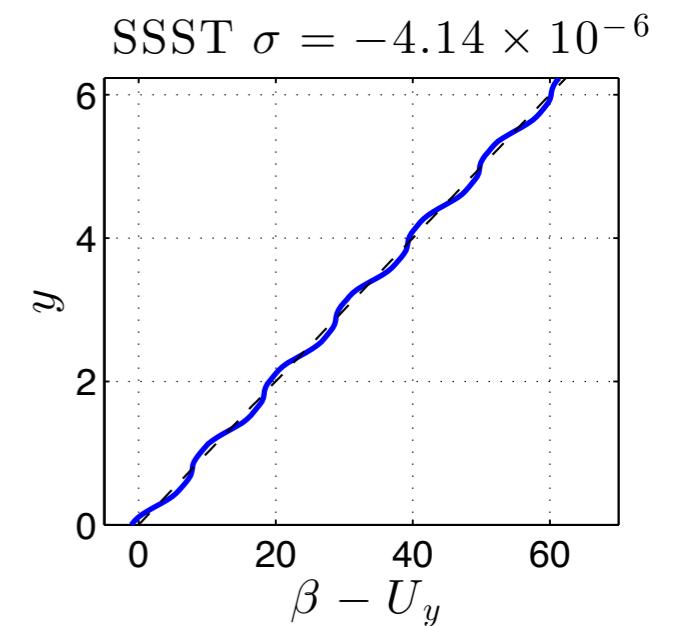
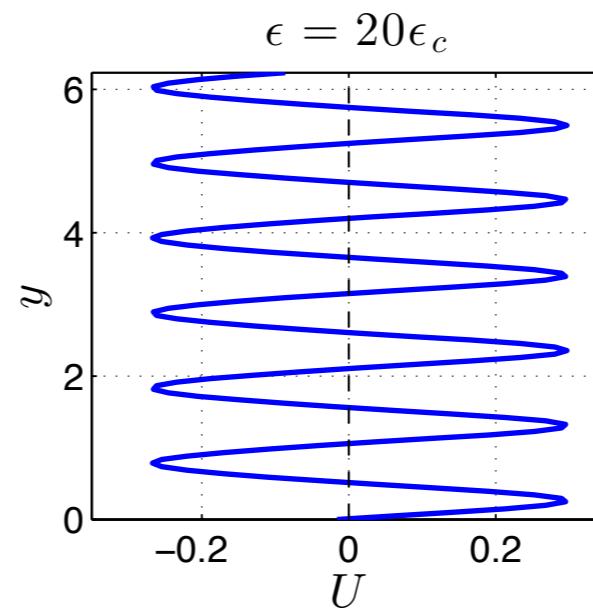
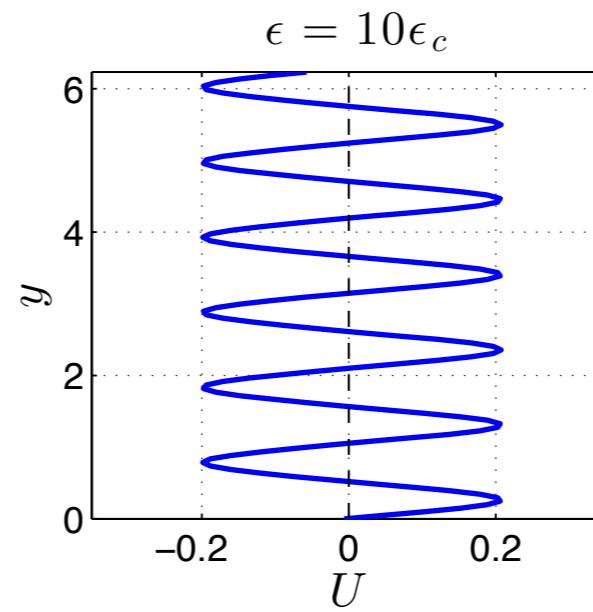
SSST unstable





PV Staircases

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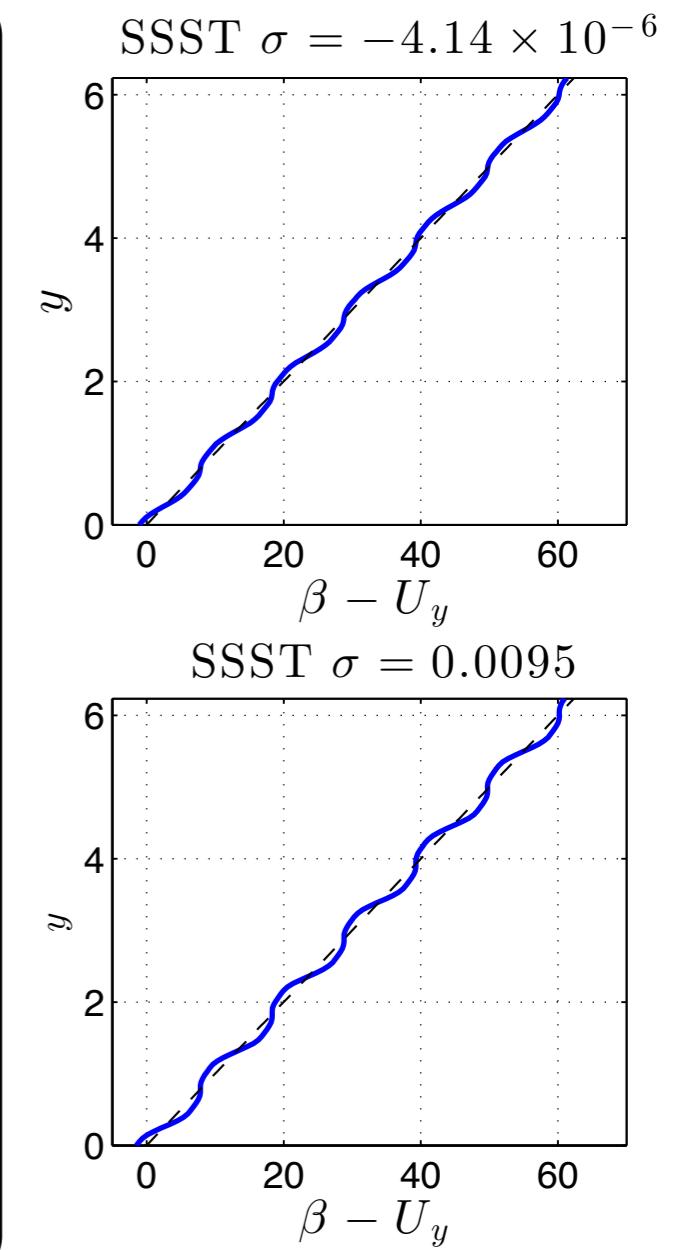
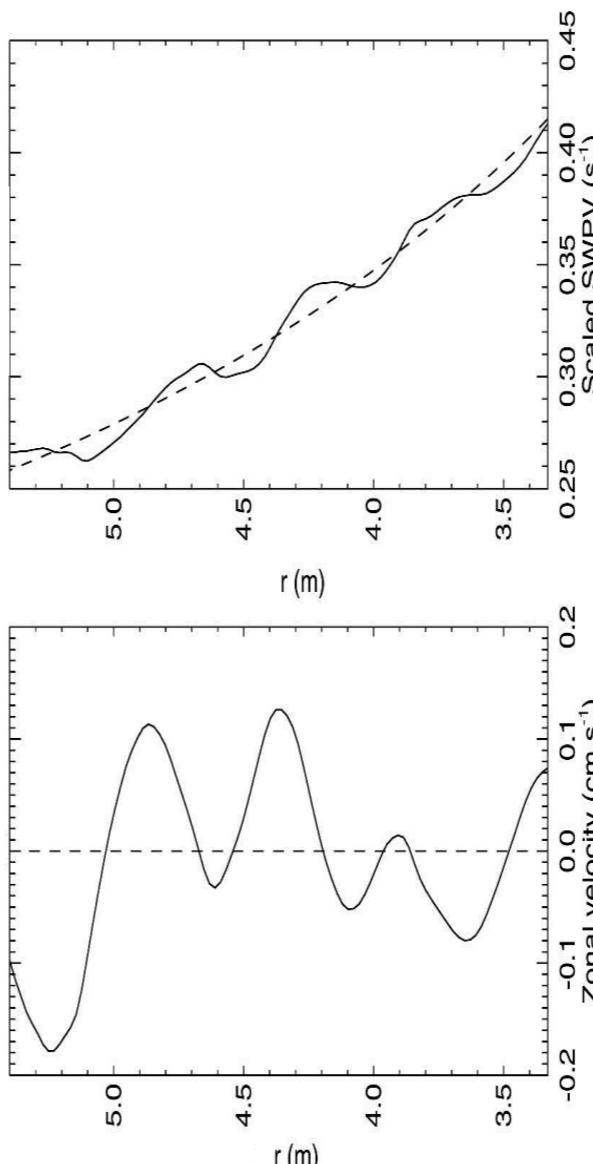


PV Staircases

NIF, $k = 1$,
 $r = 0.1$, $r_m = r/$

PV staircases in
rotating tank
experiment

(Read et. al., 2007)





Conclusions

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...
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Conclusions

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- with the modification of the turbulent spectrum due to the emergence of NZCS, SSST captures the bifurcation for jet emergence
- stochastic excitation of SSST eigenmodes reveals that these modes underlie in the dynamics even in subcritical cases



Thank you

This work has been
supported by



Constantinou, N.C, Ioannou, P.J. and Farrell, B.F., 2012:
Emergence and equilibration of jets in beta-plane turbulence.
(arXiv:1208.5665 [physics.flu-dyn])