

#### **ESEIAAT**



# **Cubesat Constellation Astrea**

**ANNEX VI: Matlab Codes** 

Degree: Aerospace Engineering **Course:** Engineering Projects

**Group:** G4 EA-T2016

**Delivery date: 22-12-2016** 

### Students:

Cebrián Galán, Joan Fontanes Molina, Pol Foreman Campins, Lluís Fraixedas Lucea, Roger Fuentes Muñoz, Óscar González García, Sílvia Herrán Albelda, Fernando Kaloyanov Naydenov, Boyan Martínez Viol, Víctor Morata Carranza, David Pla Olea, Laura Pons Daza, Marina Puig Ruiz, Josep Serra Moncunill, Josep Maria

Tarroc Gil, Sergi Tió Malo, Xavier

Customer: Pérez Llera, Luís Manuel

Urbano González, Eva María



## **Contents**

1	Mat	lab Codes	1
	1.1	Satellite Footprint	1
	1.2	Minimum Plane Inclination	3
	1.3	Satellite Number Computation for Polar Orbits	5
	1.4	Orbit Parameters	8
	1.5	Walker Delta Testing	13
	1.6	Orbit Plotter	14
	1.7	Perturbations	21
	1.8	Orbit Decay	27
	1.9	Performance Evaluator	31
	1.10	Thrust	35
	1.11	Satellites Datasheet	40
	1.12	Ground Station Localization	42
	1.13	Received signal strength variation with bandwidth	45
	1.14	Received signal strength for distance variation	47



## **Matlab Codes**

## **Satellite Footprint**

```
1 %----ASTREA CONSTELLATION----
2 %PROJECTS - 220028
3 %Aerospace Engineering Barchelor's Degree
4 %ESEIAAT - UPC
5 %Autumn 2016-2017
7 % ORBIT DESIGN TEAM
8 % COMPUTATION OF A SATELLITE FOOTPRINT
10 clc
11 clear all
12 close all
13
14 %% PHISICAL CONSTANTS AND PARAMETERS
15
16 Re = 6378;
                     %Earth's Radius [Km]
17 Se = 4*pi*Re^2;
                     %Earth's Surface [Km^2]
18 h = 500:100:1000; %Satellte height [Km]
19 eo = 0:5:40;
                     %Elevation angle [deg]
21 %% SOLVER
22
23 for i = 1:length(h)
     for j = 1:length(eo)
24
25
          d(i,j) = Re * (sqrt(((h(i)+Re)/Re)^2-(cosd(eo(j)))^2)-sind(eo(j)));
26
          Bo(i,j) = asind(d(i,j)*cosd(eo(j)) / (h(i)+Re));
27
28
          S(i,j) = 2 * pi * Re^2 * (1-cosd(Bo(i,j)));
29
30
          R(i,j) = sqrt(d(i,j)^2 - h(i)^2);
31
32
          %Satellite coverage expressed as a fraction of Earth's Area (%)
33
```



```
Cov(i,j) = (S(i,j) / Se)*100;
34
35
           %Number of satellites needed
           nsat(i,j) = (Se / S(i,j));
37
38
       end
39
40 end
41
42 %% PLOTS
43
44 figure
45 for k = 1:length(h)
      plot ( eo , Cov(k,:) )
47
      hold on
48 end
49
50 title('Coverage vs Elevation')
51 xlabel('Elevation [deg]') % x-axis label
52 ylabel('Coverage [%]') % y-axis label
legend (num2str(h(1)), num2str(h(2)), num2str(h(3)), ...
       num2str(h(4)), num2str(h(5)), num2str(h(6)))
55
56 figure
for k = 1:length(h)
      plot ( eo , nsat(k,:) )
       legend ('h(j)')
59
      hold on
60
61 end
63 title('Num.Satellites vs Elevation')
64 xlabel('Elevation [deg]') % x-axis label
65 ylabel('Num.Satellites') % y-axis label
legend (num2str(h(1)), num2str(h(2)), num2str(h(3)), ...
       num2str(h(4)), num2str(h(5)), num2str(h(6)))
67
68
69 figure
70 for k = 1:length(h)
      plot ( eo , 2*Bo(k,:) )
      hold on
72
73 end
74
75 title('Angle vs Elevation')
76 xlabel('Elevation [deg]') % x-axis label
77 ylabel('Angle between satellites[deg]') % y-axis label
78 legend(num2str(h(1)), num2str(h(2)), num2str(h(3)), ...
       num2str(h(4)), num2str(h(5)), num2str(h(6)))
```



#### Minimum Plane Inclination

```
1 %----ASTREA CONSTELLATION----
2 %PROJECTS - 220028
3 %Aerospace Engineering Barchelor's Degree
4 %ESEIAAT - UPC
5 %Autumn 2016-2017
7 % ORBIT DESIGN TEAM
9 % MINIMUM INCLINATION - To provide Full coverage %
12 %PROBLEM:
13 % Given: - Minimum Elevation over the horizon to contact the satellite
          - Height of the orbit
15 % Compute the minimum elevation to ensure visual contact
16 clear; clc;
18 %% Key variable
19 min_elevation_pole=5; %Initially set by @lfore as 32;
21 %% Previous Calculation: Visibility(latitude)
22
23 R = 6371;
24 N = 180;
25 \text{ deg} = 180;
26 x = linspace(0, 90, N);
27
28 %%% Minimum angle of elevation due to atmospheric conditions at given
29 %%% latitudes. 0 km represents the poles whereas 6371 km represents the
30 %%% Earth's equator taking as reference point the south pole.
32 elvlat = zeros (1,deg);
33 for i = 1:deg/5
      elvlat(i) = min_elevation_pole-(min_elevation_pole-15)/(2*(90-75))*i;
35 end
36 for i = deg/5+1:4*deg/10
      elvlat(i) = elvlat(deg/5);
37
38 end
39 for i = 4*deg/10+1:6*deg/10
      elvlat(i) = elvlat(2*deg/5)+3/15*(i-4*deg/10);
42 for i = 6*deg/10+1:8*deg/10
      elvlat(i) = elvlat(6*deg/10)-3/15*(i-6*deg/10);
45 for i = 8*deg/10+1:10*deg/10
      elvlat(i) = elvlat(8*deg/10)-1/15*(i-8*deg/10);
```



```
48 elvlat=fliplr(elvlat); %Sets the first value to the value in the equator
                          %The first value matches 0 latitude
51 %% Minimum i computation
52
54 L = 0:5:90;
                     %Latitude angle [deg]
55 e = (interplq(x',elvlat',L'))'; %Minimum elevations interpolation [deg]
56
57 figure (1)
58 subplot (1, 2, 1)
59 plot(x,elvlat)
60 subplot (1, 2, 2)
61 title('Interpolation Verification')
62 plot(L,e)
63
64 inc=zeros(length(h),length(L)); %Inclinations preallocation
65
  for i = 1:length(h)
66
      for j = 1:length(L)
67
68
          A = cosd(e(j))/(1+h(i)/R);
69
          theta=acosd(A)-e(j); %Earth Central Angle
70
           inc(i,j)=L(j)-theta;
71
          if inc(i, j) < 0</pre>
72
               inc(i,j)=0;
73
74
          end
      end
75
76 end
77
78 %% 3. Results Plotting
79
80 figure(2)
81 colors=['r','b','y','g','m','k','c'];
82 for k = 1:length(h)
      hold on
83
      plot ( L , inc(k,:),colors(k))
84
85 end
86
87 grid on
88 title('Minimum inclination required to fulfill global coverage')
89 xlabel('Latitude [deg]')
90 ylabel('Inclination [deg]')
91 legend(num2str(h(1)), num2str(h(2)), num2str(h(3)), ...
      num2str(h(4)), num2str(h(5)), num2str(h(6)), num2str(h(7)))
93
94 %% Bonus!!!
95
96 L=90;
```



```
97 e=5.33;
98 h=500:1:600;
99 inc=zeros(1,length(h));
100 for i = 1:length(h)
           A = cosd(e) / (1 + h(i) / R);
            theta=acosd(A)-e; %Earth Central Angle
102
           inc(i)=L-theta;
104
           if inc(i)<0
                inc(i)=0;
105
           end
106
107 end
108
109 figure (3)
110 plot(h,inc)
111 xlabel('Heights')
112 ylabel('Inclination')
113 grid on
114 axis([min(h) max(h) 45 90])
```

## **Satellite Number Computation for Polar Orbits**

```
1 %----ASTREA CONSTELLATION----
2 %PROJECTS - 220028
3 %Aerospace Engineering Barchelor's Degree
4 %ESEIAAT - UPC
5 %Autumn 2016-2017
7 % ORBIT DESIGN TEAM
8 % POLAR ORBITS Number of satellites computation
10 %PROBLEM:
11 % Given: - Elevation over the horizon to contact the satellite
12 % - Height of the orbit
13 % Compute the final number of satellites
14 % Through the "Streets of coverage" method.
15
16 clear; clc;
17
18 %% 1. Input data
19 % PHISICAL CONSTANTS AND PARAMETERS
21 Re = 6378;
                      %Earth's Radius [Km]
21 Re = 63/8; %Earth's Radius [Km]
22 Se = 4*pi*Re^2; %Earth's Surface [Km^2]
h = 500:20:600; %Satellte height [Km]
24 eo = 20; %Elevation angle [deg]
25
```



```
26 %% 2. Number of satellites computation
27
  d=zeros(length(h),length(eo));
29 nsats_opt=d;
  for i = 1:length(h)
31
       for j = 1:length(eo)
32
33
           d(i,j) = Re*(sqrt(((h(i)+Re)/Re)^2-(cosd(eo(j)))^2)-sind(eo(j)));
34
           Bo = asind( d(i,j)*cosd(eo(j)) / (h(i)*Re) );
35
36
           Sup = 2 * pi * Re^2 * (1-cosd(Bo));
37
           Nspp_min=ceil(360/(2*Bo));
                                               %Number of satellites per plain
38
           Nsatspp=Nspp_min:1:Nspp_min+50; %Array of sats/plain to optimize
39
40
41
           S=360./Nsatspp;
42
43
           Lstreet=acosd(cosd(Bo)./cosd(S/2));
44
           DmaxS=Lstreet+Bo;
45
           DmaxO=2*Lstreet;
46
47
           Nplains=ceil(1+((180-DmaxO)./DmaxS));
48
49
           Nsats=Nplains.*Nsatspp;
           [row, col] = find (min (Nplains));
50
           Nplains_opt(i, j) = Nplains(row, col);
51
           nsats_opt(i,j)=Nsats(row,col);
52
53
           %Number of satellites needed
54
55
           nsat(i,j) = (Se / Sup);
56
       end
57
58 end
59
60 %% 3. Results Plotting
61
62 figure (2)
63 colors=['r','b','y','g','m','k','c'];
64 for k = 1:length(eo)
       subplot(1,2,1)
65
       hold on
66
       semilogy ( h , nsats_opt(:,k),colors(k))
67
       subplot(1,2,2)
68
       hold on
       semilogy ( h , Nplains_opt(:,k),colors(k))
70
71 end
72
73 subplot (1, 2, 1)
74 title('Num.Sats vs Height')
75 xlabel('Height [km]') % x-axis label
```



```
76 ylabel('Num.Satellites') % y-axis label
77 \frac{1}{2} %legend (num2str(eo(1)), num2str(eo(2)), num2str(eo(3)),...
        num2str(eo(4)), num2str(eo(5)), num2str(eo(6)), num2str(eo(7)))
79 axis([400 900 0 1000])
80 grid on
81
82 subplot (1, 2, 2)
83 title('Num.Planes vs Height')
84 xlabel('Height [km]') % x-axis label
85 ylabel('Num of Orbital Planes') % y-axis label
86 % legend(num2str(eo(1)), num2str(eo(2)), num2str(eo(3)),...
         num2str(eo(4)), num2str(eo(5)), num2str(eo(6)), num2str(eo(7)))
88 axis([400 900 0 30])
89 grid on
91 %% Single detailed analysis
92
93 figure (1)
94 plot(h,nsats_opt)
95 title('Num.Sats vs Height - Streets of coverage Method')
96 xlabel('Height [km]') % x-axis label
97 ylabel('Num.Satellites') % y-axis label
98 grid on
100 %% Multianalysis variation with height
101
102 figure (3)
103 colors=['r','b','y','g','m','k'];
104 for k = 1:length(h)
105
       subplot(1,2,1)
106
      hold on
107
      plot ( eo , nsats_opt(k,:),colors(k))
       subplot(1,2,2)
108
109
       hold on
       plot ( eo , Nplains_opt(k,:),colors(k))
110
111 end
112
113 subplot (1, 2, 1)
114 title('Num.Sats vs Elevation - Streets of coverage Method')
115 xlabel('Elevation [deg]') % x-axis label
116 ylabel('Num.Satellites') % y-axis label
legend (num2str(h(1)), num2str(h(2)), num2str(h(3)), ...
       num2str(h(4)), num2str(h(5)), num2str(h(6)))
119 axis([0 40 0 700])
120 grid on
122 subplot (1, 2, 2)
123 title('Num.Planes vs Elevation - Surfaces Method')
124 xlabel('Elevation [deg]') % x-axis label
125 ylabel('Num.Orbital Planes') % y-axis label
```



```
legend(num2str(h(1)), num2str(h(2)), num2str(h(3)),...
num2str(h(4)), num2str(h(5)), num2str(h(6)))
legend(num2str(h(1)), num2str(h(6)))
num2str(h(6)))
legend(num2str(h(1)), num2str(h(6)))
legend(num2str(h(1)), num2str(h(6))), num2str(h(6)))
```

### **Orbit Parameters**

```
1 %----ASTREA CONSTELLATION----
2 %PROJECTS - 220028
3 %Aerospace Engineering Barchelor's Degree
4 %ESEIAAT - UPC
5 %Autumn 2016-2017
7 % ORBIT DESIGN TEAM
8 % ORBIT PARAMETERS
10 % This routine calculates the minimum number of plans and satellites
11 % to guarantee global coverage.
12 % -------
13 % _v2 --> Computes angles with arccosine
14 % _v3 --> Also iterates in f
15
16 clc
17 clear all
18 close all
19
20 tic
21
22 %% DATA INPUT
23
24 Re = 6371;
                         %Earth's Radius [Km]
25 Se = 4*pi*Re^2;
                        %Earth's Surface [Km^2]
26 h = 500:5:600;
                        %Height vector [km]
27 in = 75:3:90;
                        %Inclination vector [deg]
                        %Minimum number of plans
28 pmin = 5; p=12;
29 pmax = 20;
                         %Maximum number of plans
30 sppmin = 10; spp=10; %Minimum number of satellites each plane
31 sppmax = 24;
                        %Maximum number of satellites each plane
32 typeWD = 2;
                         %Defines:
                          % 1- Semi-Walker-Delta
33
                          % 2- 3/4 of Walker-Delta
                         % 3- Full Walker-Delta
35
36 plaunch = 5.76e6;
                         %Launch price
37 psat = 250e3;
                         %Price per satellite
38 eo = 20;
                        %Vison angle
                         %Variable used as counter
q=0;
```



```
40 nplanes_win=0;
  %Boolean, Indicates a successful combination with that number of planes
  %% COMPUTATIONS
  %result=zeros(6,100);
45
  for j=1:length(h)
46
47
       fprintf('\n\nH = %g km', h(j))
       Bo = visibleangle(h(j),eo); %Footprint angle
48
49
       for k=1:length(in)
50
           fprintf('\ni=%g: ',in(k))
51
52
           for p=pmin:pmax
53
               F=0:1:p-1; nF=length(F); anglesp_f=zeros(1,nF);
54
55
               for spp=sppmin:sppmax
56
57
                    angles = 360/spp; %Angle between satellites of same plane
58
                 for kf=1:nF
59
                   % We look for the optimum f for this spp,npp,inc,h
60
                  anglesp_f(kf) = distance_opt_f(in(k),h(j),p,spp,typeWD,F(kf));
61
                   %Angle between satellites of different planes
62
63
                 end
64
                    [anglesp, kf_min] = min(anglesp_f);
65
66
                     if (angles < 2*Bo) && (anglesp < Bo)
67
68
69
                        q=q+1;
                        result(:,q)=[h(j) in(k) spp p spp*p F(kf_min)];
70
                        fprintf(['Solution %g: h=%gkm - i=%g - '...
71
                           '%gspp - %gplanes - %g sats - f=%g'],q,result(:,q))
72
73
                        nplanes_win=1;
                        break %Stop iteration in satellites per plain
74
                     end
76
77
               end
78
                if nplanes_win
                                 %If a combination with that num of planes
79
                    nplanes_win=0; %works, then jump to the next inclination
80
                    break
81
               end
82
           end
83
       end
84
85
  end
86
87 %% Optimum Results
88 % Minimum number of cubesats
89 numsat = result (5,1:q);
```



```
[minsat,index] = min(numsat);
91
   disp(sprintf('\nMINIMUM SATELLITES\nheight: %d', result(1,index)))
93 disp(sprintf('inclination: %d', result(2, index)))
94 disp(sprintf('number of satellites in each plane: %d',result(3,index)))
95 disp(sprintf('number of planes: %d',result(4,index)))
  disp(sprintf('minimum number of satellites: %d',result(5,index)))
97
   % Minimum number of planes
98
99
   numplane = result(4,1:q);
100
   [minplane, index] = min(numplane);
101
102 disp(sprintf('\nMINIMUM PLANES\nheight: %d', result(1,index)))
103 disp(sprintf('inclination: %d',result(2,index)))
104 disp(sprintf('number of satellites in each plane: %d',result(3,index)))
105 disp(sprintf('number of planes: %d',result(4,index)))
   disp(sprintf('minimum number of satellites: %d',result(5,index)))
106
107
108
  % Price Optimization
109 TCost = plaunch*result(4,:) + psat*result(5,:); TCost = TCost/(1e6);
110 [CostOpt,index] = min(TCost);
111 height = result(1,:);
112
disp(sprintf('\nMINIMUM PRICE\nprice[M]: %d', TCost(index)))
114 disp(sprintf('height: %d', result(1,index)))
115 disp(sprintf('inclination: %d', result(2, index)))
116 disp(sprintf('number of satellites in each plane: %d',result(3,index)))
117 disp(sprintf('number of planes: %d',result(4,index)))
   disp(sprintf('minimum number of satellites: %d',result(5,index)))
118
119
   fprintf('\n\n MORE THAN ONE SOLUTION COULD HAVE THE SAME MIN NUMBER\n')
120
121
   %% Plots
122
123
124 x=0:q-1;
plot(x, numplane, x, numsat, x, TCost, x, height)
126 strValues = strtrim(cellstr(num2str([x(:) numplane(:)],'(%d, %d)')));
  text(x, numplane, strValues, 'VerticalAlignment', 'bottom');
128 strValues = strtrim(cellstr(num2str([x(:) numsat(:)],'(%d, %d)')));
129 text(x, numsat, strValues, 'VerticalAlignment', 'bottom');
130 strValues = strtrim(cellstr(num2str([x(:) TCost(:)],'(%d, %.1f)')));
   text(x,TCost,strValues,'VerticalAlignment','bottom');
132 strValues = strtrim(cellstr(num2str([x(:) height(:)],'(%d, %d)')));
133 text(x,height,strValues,'VerticalAlignment','bottom');
134 legend('number of planes', 'number of satellites', 'Total Cost (M)')
135
136
137
138 t.o.c
```



```
1 function [anglesp] = distance_opt_f (i,a,p,s,typeWD,F)
2 %% Data input
4 % i: Inclination angle [deg]
5 % a: Height of the orbit [km]
6 % p: Number of planes
7 % s: Number of satelites per plane
9 %% Walker Delta type of geometry
10 WD=[180 225 360];
11 degreegenerate = WD(typeWD);
12 % Walker Delta 360 deg, SemiWalker Delta 180 deg,
13 % other constellations range between 180 and 360
15 m = degreegenerate/180;
16 % generates Walker Delta Constellation
17 % (m=2 \rightarrow 2*180 generated constellation),
18 % Semi Walker (m=1 -> 1*180) or other 1<m<2.
19
20 % Phasing between adjacent planes
21 %f = F*p*cosd(i); % parameter defined graphically
22 f = F; % Whereas F is a number 0 < f < (p-1)
          % According to Astrodynamics notes
23
24
25 REarth = 6371; % [km]
26 h = (REarth+a)/REarth; %radius of the orbit
27 %(in terms of the sphere radius) R=h*r;
28
29 %% Distribution of Coordinates of the satellites
30 Omega = zeros(1,p);
31 nu = zeros(s,p);
```



```
32 angle = zeros(1,p);
33 X = zeros(3, s, p);
  for j = 1:p
35
36
       Omega(j) = m*pi*(j-1)/p;
       angle(j) = f*2*pi*(j-1)/(s*p); %Phasing due to f between planes
37
38
       for k = 1:s
           % True anomalies of the s satellites
39
           nu(k,j) = 2*pi*(k-1)/s+angle(j);
40
           X(:,k,j) = cartesian(h,0,i,Omega(j),0,nu(k,j));
41
           % h[adim] --> X[adim]
42
       end
43
44 end
45
46 %% Fast comprovation
47 % Does the adjacent phasing exceed pi
48 % THIS MEANS
  % The minimum might be between
49
50
  %% Angle between different satellites
52
53
54 c=zeros(s,p); %General aproach. Where is the minimum angle?)
55
56 for j=2:p
    for i=1:s-1
57
58
       % We need to assess the angles between the last and first plane
       % Specially in Semi-Walker configurations. That's why we add this
60
61
       % auxiliar variables.
62
         p1=j;
63
         if j==1
64
65
             p2=p;
66
         else
             p2=j-1;
67
         end
68
69
       % Angles a2 and a3 could be used to increase de reliability of the
70
       % constellation, allowing greater possibilities of communication
71
       % between planes. (not only between (i,p1) and (i,p2)
72
       % Angle between satellite (i,p1) and satellite (i,p2)
74
       a1=acosd((X(:,i,p1)'*X(:,i,p2))/(norm(X(:,i,p1))*norm(X(:,i,p2))));
76
77
       % Angle between satellite (i,j) and satellite (i+1,j-1)
       a2=a\cos((X(:,i,p1)'*X(:,i+1,p2))/(norm(X(:,i,p1))*norm(X(:,i+1,p2))));
78
79
       % Angle between satellite (i,j) and satellite (i-1,j-1)
80
81
       a3=acosd((X(:,i,p1)'*X(:,i-1,p2))/(norm(X(:,i,p1))*norm(X(:,i-1,p2))));
```



```
82
       % Among the computed angles, choose the minimum
83
       [angle(i,j),c(i,j)]=min([a1 a2]);
85
86
       % Why minimum? Then we know that at least it will be able to
       % cover the angle with one of the two adjacent close nodes.
87
88
    end
89
90 end
91
92 anglesp = max(max(angle));
93 %Maximum angle between satellites of different planes
```

## Walker Delta Testing

```
1 %----ASTREA CONSTELLATION----
2 %PROJECTS - 220028
3 %Aerospace Engineering Barchelor's Degree
4 %ESEIAAT - UPC
5 %Autumn 2016-2017
7 % ORBIT DESIGN TEAM
8 % ORBIT SMART WALKER-DELTA
10 % This routine compute the WD possible configurations that give global
11 % coverage by having the distance between planes matches the planes
12 % rotation
14 p=[12 15 18];
15 S=360./p; S=S*pi/180;
re=6378.01e3;
u=3.986012e14;
19 J2=1.0826e-3;
21 \text{ we}=2*\text{pi/(24*3600)};
23 syms a;
24 inc=75; inc=inc*pi/180;
25 a_sol=zeros(1,length(inc));
27 for i=1:length(inc)
       dOmega=-1.5*J2*(re/a)^2*sqrt(u/a^3)*cos(inc(i));
       P0=2*pi*sqrt(a^3/u);
29
      Pn=P0*(1-1.5*J2*(re/a)^2-0.75*J2*(4-5*sin(inc(i))^2)*(re/a)^2);
       eq=S(2)-Pn*(we-dOmega);
31
```



```
sol=vpasolve(eq,re*1.15);
      if imag(sol) == 0
33
         a_sol(i)=sol;
       end
35
36 end
37
38 plot(inc*180/pi,(a_sol-re)/1000)
39 h=(a_sol-re)/1000
40
41 %% Satellites computation
42
43 emin=20*pi/180;
44 Lstreet=5*pi/180;
46 rho=re/(re+h*1000);
47 etha=asin(sin(rho)*cos(emin));
48 Lmax=pi/2-emin-etha;
49 S=2*acos(cos(Lmax)/cos(Lstreet));
N=ceil(2*pi/S)
```

#### **Orbit Plotter**

```
1 %----ASTREA CONSTELLATION----
2 %PROJECTS - 220028
3 %Aerospace Engineering Barchelor's Degree
4 %ESEIAAT - UPC
5 %Autumn 2016-2017
7 % ORBIT DESIGN TEAM
\mathbf{8} % ORBIT PLOTTER - Influential phenomena computation
10 % This matlab routine plots the position of the satellites of a Walker
11 % Delta / SemiWalker Delta or other generated configurations.
12
13 clc, clear, close all;
15 %% O. Data Input
16 %add the subfolder programs to the current path
17 addpath(genpath('./auxiliar'));
18
19 %_____Physical input___
20
21 global h, h = 542; % [km]
22 degreegen = 225; % Walker Delta 360 deg, SemiWalker Delta 180 deg,
                    % other constellations range between 180 and 360 deg
24 in = degtorad(72); % inclination of the planes
```



```
25 s = 21;
                      % satellites per plane
26 p = 9;
                       % number of planes
27 f=8;
                      % f = 1.25*p*cos(i); % parameter defined graphically
28 eo=degtorad(20); % elevation angle [rad]
      __Numerical Input_
30 %
31 r = 1; %radius of the sphere
32 n = 50; %number of cells of the sphere generated
33 N = 50; %number of points used to print the orbit
34 nn = 100; %number of points used to compute the footprint
36 %____Physical Constants_
37 global Re, Re = 6371; %Earth's Radius [Km]
38 %% 1. Preprocessing
39
40 global m, m = degreegen/180;
              % generates Walker Delta Constellation (m=2 -> 2*180)
41
               % generated constellation), Semi Walker (m=1 -> 1*180)
42
              % or other 1<m<2.
43
45 global a, a = (Re+h)/Re;
              %radius of the orbit (in terms of the sphere radius) R=h*r;
46
47
48 %% 2. Main Process
49
50 %compute the sat RAAN and some sat. parameters (Laura)
51 Omega = zeros(1,p);
                      %true anomaly of the satellites
52 \text{ nu} = zeros(s,p);
53 angle = zeros(1,p); %angle between satellites
54 for i = 1:p
      Omega(i) = m*pi*(i-1)/p; %orbits
55
     angle(i) = f*2*pi*(i-1)/(s*p); %sats
57
      for k = 1:s
58
          nu(k,i) = 2*pi*(k-1)/s+angle(i);
59
      end
61 end
62
63 % Compute the satellite coordinates (Laura)
64 X = zeros(3, s, p);
65 for i = 1:p
      for k = 1:s
          X(:,k,i) = cartesian(a,0,in,Omega(i),0,nu(k,i));
67
70 sCOOR=reshape(X,3,p*s)'; %put the data in a more friendly structure
71
72 %compute the radius of the footprint
73 %SATELLITE_FOOTPRINT_COMPUTATION.m
74 d=Re*((((h+Re)/Re)^2-cos(eo)^2)^.5-sin(eo));
```



```
75 Bo=asin(d*cos(eo)/(h+Re));
76 S=2*pi*Re^2*(1-cos(Bo));
77 R=(S/pi)^{.5};
79 %create the footprint of each satellite
80 SATfp=[];
81 for i=1:s*p
       %unitary vector pointing to the sat from the earth's center
82
       u=-sCOOR(i,:)/norm(sCOOR(i,:));
83
       %Generate the circle coordinates (2D)
84
       [satfp(:,1), satfp(:,2), satfp(:,3)] = circle3D(sCOOR(i,:), u, R/Re, nn);
       satfp=satfp+repmat(u*h/Re,nn,1); %move the circle to be tg to Earth
86
87
       for j=1:size(satfp,1)
88
            v=sCOOR(i,:)-satfp(j,:); v=v/norm(v);
89
            %unitary vector sat->footprint
90
91
            %verify if the footprint intersect with the earth's surface
92
93
           tol=satfp(j,1)^2+satfp(j,2)^2+satfp(j,3)^2-r^2;
           while tol>1e-2
94
                satfp(j,:) = satfp(j,:) -0.01*v;
95
                tol=satfp(j,1)^2+satfp(j,2)^2+satfp(j,3)^2-r^2;
96
97
            end
98
       end
       %store the footprint
99
       SATfp=[SATfp; satfp];
100
101 end
102
103
104 %% 3. Plots
105
106 %3D plot
107 [COOR] = Orbitplot 3D(in, r, X, n, SATfp, nn);
109 x=zeros(size(COOR, 1), size(COOR, 2)/3); y=x; z=x;
110 for i=1:size(COOR, 2)/3
       x(:,i) = COOR(:,3*(i-1)+1);
111
       y(:,i) = COOR(:,3*(i-1)+2);
       z(:,i) = COOR(:,3*(i-1)+3);
113
114 end
115 GroundTrackPlot(x, y, z, sCOOR, SATfp, nn);
```

```
1 function [x,y,z]=circle3D(center,normal,radius,n)
2 %This matlab function generates the points of a circle normal to a 3D
3 %vector
4 %In:
5 % center: center of the circle (x, y, z)
6 % normal: vector normal to the circle
7 % radius: radius of the circle
```



```
1 function [COOR] = Orbitplot3D( in,r, X,n, sCOOR, N )
2 %Orbitplot3D is a function that creates a 3D representation of a
3 %constellation and plots the satellites. It can be also configured to plot
4 %their footprint
5 %In:
6 % in: inclination angle [rad]
      r: sphere radius [-]
8 % X: Coordinates of the sats (3,spp,p)
9 % n: number of points used to draw thhe sphere
      ____
     COOR: coordinates of the footprints
12 % N: number of points used to compute the footprint
14 p=size(X,3);
15 s=size(X,2);
16 global a m;
17
18 % Earth
19 %Create the earth instance as an sphere
20 [x,y,z]=sphere(n); %Generate the sphere
21 surf(r*x,r*y,r*z,'FaceColor','interp','EdgeColor','none'); %plot the sphere
22 colormap(linspace(0,1,n)'*[0 0.6 1]); %custom colormap (blue tones);
23 axis equal;
24 box off;
25 hold on;
26 rotate3d on;
28 % Orbits
29 %Generate the circle coordinates (2D)
30 coor=zeros(N+1,3);
31 for i=1:N
      coor(i,:) = [a*r*cos(2*i*pi/N), a*r*sin(2*i*pi/N), 0];
34 coor(N+1,:) = coor(1,:);
36 %Rotation Matrices
37 Ry=[ cos(in) 0 sin(in)
```



```
0 1 0
38
      -sin(in) 0 cos(in)];
39
41 Rz= @(Omega) [-sin(Omega) cos(Omega) 0
                cos(Omega) sin(Omega) 0
                                 0
                                       1]; % he modificat aquesta matriu
43
44
45 coor=coor*Ry; %Rotate the circle in the y axis (inclination)
46
47 %rotate the plane to get the other planes
48 COOR=zeros(N+1, 3*p);
49 for i=1:p
      COOR(:, (1+(i-1)*3):(3+(i-1)*3))=coor*Rz(m*(i-1)*pi/p);
51 end
52 for i=1:p
     plot3 (COOR(:, 1+(i-1)*3), COOR(:, 2+(i-1)*3), COOR(:, 3+(i-1)*3),...
53
          'linewidth',1.5);
54
55 end
56
57 % Satellites (Laura)
58 for i = 1:p
      for k = 1:s
59
        scatter3(X(1,k,i),X(2,k,i),X(3,k,i),25,'square','k','filled');
61
        hold on;
       end
62
63 end
64
65 %Plot also the footprint if the number of input variables is bigger than 4
66 if nargin > 4
67
      for i=1:p*s
          b=1+N*(i-1); c=N+N*(i-1);
68
          fill3(sCOOR(b:c,1),sCOOR(b:c,2),sCOOR(b:c,3),[1 0 0],...
               'FaceAlpha','.3','LineStyle','none');
70
71
       end
72 end
73 end
```

```
function [] = GroundTrackPlot( x, y, z, coor, fp_coor, N)

%GroundTrackPlot is used to plot the ground
ftrack of the constellation orbits

%given their coordinates (cartesian). It has two variants, only introducing
fthe first 3 inputs will plot the orbits and if you introduce all the
finputs it plots the satellites and their footprint.

%In:

% x: matrix containing the x coordinate of the orbits to plot
y: matrix containing the y coordinate of the orbits to plot
x: matrix containing the z coordinate of the orbits to plot
y: matrix containing the z coordinate of the orbits to plot
y: matrix containing the z coordinate of the orbits to plot
y: matrix containing the z coordinate of the orbits to plot
y: matrix containing the z coordinate of the orbits to plot
```



```
coor: satellite coordinates (nx3)
     fp_coor: satellite footprint coordinates ((n*nn)x3)
15 % N: number of points used to compute the footprint
17 p=size(x,2); %n of planes
18 s=size(coor, 1)/size(x, 2); %n of spp
20 %% Set the background image for the ground track plot
21 figure('name','Constellation Ground Track');
22 tit=['Num. of planes: ' num2str(p) ' Num. of spp: ' num2str(s)];
23 ground='World-satellite-map.png';
24 im=imread(ground);
25 imagesc([-180,180],[-90,90],im);
26 grid on;
27 axis([-180 180 -90 90]);
28 title(tit);
29 hold on;
30 %set the plot lines color
31 C = ['r', 'y', 'm', 'g', 'w']; C=[];
32 for i=1:ceil(size(x,2)/5)
      C=[C c];
34 end
35
36 %% Main process
37  for i=1:size(x,2)
       %convert the cartesian coordinates to spherical
38
39
      [lon, lat] = cart2sph(x(:,i),y(:,i),z(:,i));
       %get the coordinates of the orbit in spherical coord.
      lon=radtodeg(lon); %pass to degrees
41
      lat=radtodeg(lat);
      Lon=[lon(end); lon(1:end-1)];
43
44
      %Plot the sat ground track
45
       [\neg, j] = \max(abs(lon-Lon));
46
      plot(lon(1:j-1), lat(1:j-1), C(i), 'linewidth', 1.2);
47
      plot(lon(j:end), lat(j:end), C(i), 'linewidth', 1.2);
49 end
50
if nargin > 3
       [slon, slat] = cart2sph(coor(:,1), coor(:,2), coor(:,3));
52
       %get the coordinates of the sats in spherical coord.
53
       slon=radtodeg(slon); %pass to degrees
54
      slat=radtodeg(slat);
55
       %plot the sats
57
       scatter(slon, slat, 30, 'square', 'w', 'filled');
58
59
       %get the coordinates of the sats footprint in spherical coord.
60
       [fplon, fplat] = cart2sph(fp_coor(:,1), fp_coor(:,2), fp_coor(:,3));
61
       fplon=radtodeg(fplon); %pass to degrees
```



```
fplat=radtodeg(fplat);
63
64
       %Plot their footprint
       for i=1:p*s
66
67
            b=1+N*(i-1); c=N+N*(i-1);
            fpLon=[fplon(c); fplon(b:c-1)];
68
69
            %get the maximum values (plotting purposes)
70
            [sorted, I]=sort(abs(fplon(b:c)-fpLon));
71
            M=sorted(end-1:end);
72
            j=I(end-1:end); j(2)=j(2)-1;
73
74
            if M(1) > 300
75
                if j(1) < j(2)
76
                    auxlon=fplon(b+j(1):b+j(2)-1); %negative lon. part
77
78
                    auxlat=fplat(b+j(1):b+j(2)-1);
79
                    auxlon2=[fplon(b+j(2):c); fplon(b:b+j(1)-2)]; %+ lon. part
80
81
                    auxlat2=[fplat(b+j(2):c); fplat(b:b+j(1)-2)];
82
                    auxlon2=fplon(b+j(2):b+j(1)-2); %negative lon. part
83
                    auxlat2=fplat(b+j(2):b+j(1)-2);
84
85
86
                    auxlon=[fplon(b+j(1):c); fplon(b:b+j(2)-1)];
                     %positive lon. part
87
                    auxlat=[fplat(b+j(1):c); fplat(b:b+j(2)-1)];
88
89
                end
90
             fill(auxlon,auxlat,[1 0 0],'FaceAlpha','.2','LineStyle','none');
91
92
             fill(auxlon2,auxlat2,[1 0 0],'FaceAlpha','.2','LineStyle','none');
            else
93
                fill(fplon(b:c), fplat(b:c), [1 0 0], ...
94
                     'FaceAlpha','.3','LineStyle','none');
95
96
            응응응응응응응응응응응
97
            pause(.05); %
            응응응응응응응응응응응
99
            end
100
101
102
103
       end
104
   end
105 end
```

```
1 % This function converts kepler orbit elements to cartesian coordinates
2
3 % a: semimajor axis
4 % e: eccentricity
5 % i: inclination [rad]
```



```
6 % Omega: longitude of the ascending node [rad]
7 % w: Argument of periapsis [rad]
8 % nu: True anomaly [rad]
9
10 function [X] = cartesian(a,e,i,Omega,w,nu)
11
12 % Position in cylindrical coordinates
13 r = a*(1-e^2)/(1+e*cos(nu));
14
15 % Position components
16 X = [r*(cos(Omega)*cos(w+nu)-sin(Omega)*sin(w+nu)*cos(i));
17 r*(sin(Omega)*cos(w+nu)+cos(Omega)*sin(w+nu)*cos(i));
18 r*sin(i)*sin(w+nu)];
19
20 end
```

## **Perturbations**

```
1 %----ASTREA CONSTELLATION----
2 %PROJECTS - 220028
3 %Aerospace Engineering Barchelor's Degree
4 %ESEIAAT - UPC
5 %Autumn 2016-2017
7 % ORBIT DESIGN TEAM
8 % ORBIT PERTURBATIONS - Influential phenomena computation
10 %PROBLEM:
11 % Given: - Initial orbital parameters
12 % Compute the final orbtial parameters after each orbit
13 %clear; clc;
14
15 %% 1.A Input data physical constants
16 % PHISICAL CONSTANTS AND PARAMETERS
17 %Physical data
18 RE=6.378e6; %Earth Radius [m]
u=3.986e14; %GM Earth
20 J2=1082.6e-5; %J2 coefficient;
22 %% 1. Input data Orbit parameters
23 % SPACECRAFT DATA
             % Satellite mass [kg]
24 \text{ m}=4;
           % Drag Coefficient
25 Cd=2.2;
26 A=0.1; % Satellite surface [m^2] --> 3U pointing to Earth
27 Bc=m/(Cd*A); % Ballistic coefficient of the satellite;
28
```



```
29 % ORBIT DATA
               % Altura inicial [m]
30 H0=542e3;
e0=0.0001;
                % Excentricity
32 i0=72*pi/180;
                   % Inclination [deg]
33 a0=RE+H0;
34
35 % INITIAL PARAMETERS
36 P=2*pi*(a0^3/u)^.5; % Orbit Period [s]
37 v=sqrt(u/a0); %satellite velocity [m/s]
38 n=(2*pi/P)*86400; %number of revolutions/day
40
41 % SIMULATION PARAMETERS
42 N=10000000;
44 %% 2. Perturbations propagation
45 %Outputs de la funcio Perturbation:
      Pert: matrix (4x5) with the perturbations causes at the rows (J2, Drag,
     Moon, Sun) and the orbital elements in the columns (a,e,i,omega, Omega)
       a e i w Omega
  용
      [
                           ] J2
49
50
      [
                           ]Drag
51 %
     [
                          ] Moon
52 % [
                          ]Sun as 3rd Body
53
54
55
56 t=zeros(1,N); t(1)=0;
57 a=t; a(1)=a0;
58 e=t; e(1)=e0;
59 i=t; i(1)=i0;
60
61 w_sun=zeros(1,N-1); w_moon=w_sun; w_J2=w_sun;
62 Om_sun=w_sun; Om_moon=w_sun; Om_J2=w_sun;
63
64 w0=0; Omega0=0;
65 w=t; w(1)=w0;
66 Omega=t; Omega(1)=Omega0;
67
68 tic
69 for n=2:N
      a0=a(n-1);
71
      e0=e(n-1);
72
      i0=i(n-1);
73
74
      w0=w(n-1);
75
      Omega0=Omega(n-1);
76
     P=2*pi*(a0^3/u)^.5; % Orbit Period [s]
77
78
```



```
[Perti rho] = Perturbation (a0, e0, i0, Bc);
79
80
        a(n) = a0 + sum(Perti(:,1));
        e(n) = e0 + sum(Perti(:,2));
82
83
        i(n) = i0 + sum(Perti(:,3));
        w(n) = w0 + sum(Perti(:, 4));
84
        Omega(n) = Omega0 + sum(Perti(:, 5));
85
86
        w_sun(n-1) = Perti(4,4);
87
        w_{moon}(n-1) = Perti(3,4);
88
89
        w_J2(n-1) = Perti(1,4);
90
        Om_sun(n-1) = Perti(4,5);
91
92
        Om_{moon(n-1)} = Perti(3,5);
        Om_J2(n-1) = Perti(1,5);
93
94
        % We don't want angles bigger than 360!
95
        w(n) = w(n) - 360 * floor(w(n) / 360);
96
        Omega(n)=Omega(n)-360*floor(Omega(n)/360);
97
98
        if Omega(n)>2*pi
99
100
             Omega(n) = Omega(n) -2*pi;
101
        end
102
        t(n) = t(n-1) + P;
103
104
        if a(n) < (RE+180e3)
105
         fprintf('Your satellite successfully burned in the atmosphere.\n\n')
106
          break
107
108
        end
109
110 end
111
   %% Post processing and results plotting
112
113
fprintf('Time to do %g iterations = %g s.\n ',n,time)
116
117 figure(1)
118 plot (t(1:n)/(3600*24), (a(1:n)-RE)/1000)
119 ylim([100 500])
120 grid on
121 ylabel('Orbit height [km]')
122 xlabel('Time [days]')
123 tfin=floor(t(n)/(3600*24));
124 days = num2str(floor(tfin/365));
125 titulaso=['Orbit decay in ' num2str(tfin) ' days = ' days 'years'];
126 title(titulaso)
127
128
```



```
129 figure (2)
130 subplot (1, 2, 1)
131 plot (t(1:n)/(3600*24), w(1:n))
132 grid on
133 ylabel('Perigee Argument [deg]')
134 xlabel('Time [days]')
135 title('Perigee Argument deviation in 100 days')
136 axis([0 100 0 360])
137
138 subplot (1, 2, 2)
139 plot (t(1:n)/(3600*24), Omega(1:n))
140 grid on
141 ylabel('Ascenent Node Argument [deg]')
142 xlabel('Time [days]')
143 tfin=t(n)/(3600*24);
144 title('Ascendent Node deviation in 100 days')
145 axis([0 100 0 360])
146
147 %% Modulus Analysis
148
149 figure (3)
150
151 subplot (1,2,1)
semilogy(t(2:n+1)/(3600*24), abs(w_sun(1:n)), 'b', 'LineWidth', 2);
153 hold on
semilogy(t(2:n+1)/(3600*24), abs(w_{moon}(1:n)), 'r', 'LineWidth', 2)
155 semilogy(t(2:n+1)/(3600*24),abs(w_J2(1:n)),'k','LineWidth',2)
156 legend('Sun','Moon','J2')
157 xlabel('Time (days)');
158 ylabel('Modulus of Perturbation []');
159 title('Effects on Argument of the Perigee');
160 grid on
161 axis([0 t(n)/(3600*24) 1e-8 1])
163 subplot (1, 2, 2)
semilogy(t(2:n+1)/(3600*24), abs(Om_sun(1:n)),'b','LineWidth',2);
165 hold on
166 semilogy(t(2:n+1)/(3600*24),abs(Om_moon(1:n)),'r','LineWidth',2)
semilogy(t(2:n+1)/(3600*24), abs(Om_J2(1:n)), 'k', 'LineWidth', 2)
168 legend('Sun','Moon','J2')
169 xlabel('Time (days)');
170 ylabel('Modulus of Perturbation []');
171 title('Effects on Argument of Ascendent Node');
axis([0 t(n)/(3600*24) 1e-8 1])
173 grid on
174
175 print -depsc ModulusAngulars
```

```
1 function [pert rho] = Perturbation( a,e,i,Bc)
```



```
2 %% Perturbations
3 % This matlab is used to compte the perturbaons in the classical orbital
4 % elements due to:
5 % J2 (non spherical earth)
6 % Atmospheric drag
7 % Third body (Moon and Sun)
8 %Inputs:
9 % a: semi-major axis [m]
10 % e: eccentricity [-]
     i: inclination [rad]
12 % m: mass of the satellite [kg]
13 % Cd: drag coefficient [-]
     A: Area of the satellite [m^2]
15 %Outputs:
16 % Pert: matrix (4x5) with the perturbations causes at the rows (J2, Drag,
17 % Moon, Sun) and the orbital elements in the columns (a,e,i,omega, Omega)
     incP/revolution
      incv/revolution
20 %TO DO:
21 %Les perturbacions degudes a J2 i 3rd body no se si son en pert/dia o no,
22 %cal mirar-ho per tal de tenir totes les pertrubacions escalades igual
24 %% Previous calculations and PreAllocation
25 %Physical data
26 RE=6.378e6; %Earth Radius [m]
27 u=3.986e14; %GM Earth
28 J2=1082.6e-5; %J2 coefficient;
30 P=2*pi*(a^3/u)^.5; % Orbit Period [s]
31 v=sqrt(u/a); %satellite velocity [m/s]
32 n=(2*pi/P)*86400; %number of revolutions/day
33
34
35 % ATMOSPHERIC MODEL
36 %get the density from MSISE model (extracted as a list from
37 %http://omniweb.gsfc.nasa.gov/vitmo/msis_vitmo.html
       h[km]
               rho[kg/m^3]
39 % filename1='msis_26371.lst'; filename2='msis_2226.lst';
40 % formatSpec='%f %f';
41 % size=[2 Inf];
42 % fileID=fopen(filename1,'r');
43 % rho_h=fscanf(fileID, formatSpec, size);
44 % fclose(fileID);
45 % fileID=fopen(filename2,'r');
46 % rho_h=[rho_h fscanf(fileID,formatSpec,size)];
47 % fclose(fileID);
48 % rho=interp1(rho_h(1,:),rho_h(2,:)',(a-RE)/1000);
     H = (a - RE) *1e-3;
50
51
      % Compute exospheric temperature [K]
```



```
T = 900 + 2.5 * (120-70);
52
       % Compute effective atmospheric molecular mass [km/K], valid 180<H<500
53
       M = 27 - 0.012 * (H - 200);
       % Compute atmospheric scale height [km]
55
       SH = T / M;
       % Compute atmospheric density [kg/m3]
57
      rho = 6E-10 * exp(-(H - 175) / SH);
58
59
60
61 pert=zeros(4,5);
63 %% Computation
64
                        _____PER REVOLUCIO
65 %J2 pertubation___
66 j=1;
67 pert(j,5)=-1.5*n*J2*(RE/a)^2*cos(i)*(1-e^2)^-2;
68 pert(j,4)=0.75*n*J2*(RE/a)^2*(4-5*(\sin(i))^2)*(1-e^2)^-2;
70 %Atmospheric Drag_____
                             ____ PER REVOLUCIO
71 %Perturbations/revolution
72 j=2;
73 pert(j,1)=-2*pi*rho*a^2/Bc;
74
75 % incP=-6*pi^2*a^2*rho/(v*Bc);
76 % incv=pi*a*rho*v/Bc;
77
                         _____/DIA --> /REV
78 %Moon Perturbation___
80 pert(j,5)=-0.00338*\cos(i)/n;
81 pert(j,4)=0.00169*(4-5*(\sin(i))^2)/n;
82
83 % Fins aqui son /dia,
84 % Canviem unitats + apliquem periode actual
85 pert (j, 5) = pert (j, 5) *P/(24*3600);
86 pert(j,4)=pert(j,4)*P/(24*3600);
88 %Sun Perturbation___
                          _____/DIA --> /REV
89 j=4;
90 pert(j,5)=-0.00154*cos(i)/n;
91 pert (j, 4) = 0.00077 * (4-5*(sin(i))^2)/n;
92
93 % Fins aqui son /dia,
94 % Canviem unitats + apliquem periode actual
95 pert(j,5)=pert(j,5)*P/(24*3600);
96 pert(j,4)=pert(j,4)*P/(24*3600);
98
99 end
```



## **Orbit Decay**

```
1 %----ASTREA CONSTELLATION----
2 %PROJECTS - 220028
3 %Aerospace Engineering Barchelor's Degree
4 %ESEIAAT - UPC
5 %Autumn 2016-2017
7 % ORBIT DESIGN TEAM
8 % ORBIT DECAY
10 % This routine computes the orbital decay of an spacecraft with time
11 % using the cowell's method.
12 % The differential equations are integrated using "ode45" (Runge-Kutta 4th,
13 % 5th order)
14
17 응응
18 clear all;
19 clc;
20 close all;
21
23 global mu mmu smu re rs rm omega J2;
24
        _Physical Variables_
26 mu=3.986004418e14; % earth gravitational constant [m^3/s^2]
27 mmu=4902.800076e9; %moon gravitational constant [m^3/s^2]
28 smu=132712440040.944e9; %sun gravitational constant [m^3/s^2]
29
30 re=6378136.3; %earth radius [m]
31 rs=696e6; %sun radius [m]
32 rm=1730e3; %moon radius [m]
34 omega=7.292115e-5; %earth angular velocity [rad/s]
35 J2=0.001081874; %earth oblateness gravity coeff []
36
37 au=149597870691; %astronomical unit [m]
38 c=2.99792458e8; %[m/s] speed of light
39
40 %____Numerical variables_
41 hlim=180; %minimum height for considering the decay [km]
42 dt=15*60; %timestep of the simulation [s]
43 t0=0; %initial time of the sim
44 tf=dt;
45
```



```
47 %__Orbital elements___
48 a=re+542e3; %[m] semimajor axis
49 e=0.0001; % eccentricity
50 i=degtorad(72); %[rad] inclination
51 w=0; %[rad] argumetn f periapsis
52 Omega=0; %[rad] RAAN
53 nu=0; %true anomaly
54
55 date=datetime(2017,2,4); %datetime array [Y, M, D]
56
57 %__Oblateness parameters___
58 znls=0; %zonals [0-18]
59 tssrls=0; %tesserals [0-18]
60
61 %__Drag perturbation inputs_
62 Cd=2.2; %Drag coefficient (Typical for sats, various ref.)
63 Adrag=.09; %[m^2] Wet area for the drag
64 m=4.1; %[kg] Cubesat mass
65
66 Bc=Adrag*Cd/m; %Ballistic coeff.
67
68 %__SRP perturbation inputs__
69 Cr=2; %reflectivity ct.
70 Asrp=.5; %[m^2] Area for the SRP
71
72 % Ws (@Dap) = 1361/(1+0.0334*\cos(2*pi*Dap/365);
73 Ws=1361; %W/m^2 approximation for the solar irradiance
75 srp=Ws/c*Cr*Asrp/m*au^2; %Solar radiation pressure ct.
78 display('Computing...');
79 f=0(q) q*((3+3*q+q^2)/(1+(1+q)^3/2)); %richard Battin's function
80
and aobl=zeros(3,1);
82 adrag=aobl; asun=aobl; amoon=aobl; asrp=aobl;
y=cartesian2(a,e,i,Omega,w,nu); %obtain the state vector (r,V)
86 jdate0=juliandate(date); %compute the julian date
87
88 tic
89 iter=0; tt=[]; xx=[];
91 h = (norm(y(1:3)) - re) *1e-3;
92 while h>hlim
      jdate=jdate0+t0/86400;
93
94 %
       GMST=JD2GMST(jdate);
95 %
       %Acceleration due to oblateness
```



```
aobl=Oblat((t0+tf)/2,y,znls,tssrls,GMST);
97
   응
98
   응
         %Acceleration due to the 3rd body pert.
         rsun=planetEphemeris(jdate, 'Earth', 'Sun');
100
101
          %geocentric pos of the Sun
         rmoon=planetEphemeris(jdate, 'Earth', 'Moon');
102
         %geocentric pos of the Moon
103
104 %
         rsun=rsun'; rmoon=rmoon';
105
         rs2s=y(1:3)-rsun; %vector sun-sat
106
   응
107
   응
         rm2s=y(1:3)-rmoon; %vector moon-sat
108
109
         qs=dot(y(1:3),(y(1:3)-2*rsun))/dot(rsun,rsun);
110
         qm=dot(y(1:3),(y(1:3)-2*rmoon))/dot(rmoon,rmoon);
111 %
112 %
         asun=-smu/norm(rs2s)^3*(y(1:3)+f(qs)*rsun);
   양
         amoon=-mmu/norm(rm2s)^3*(y(1:3)+f(qm)*rmoon);
113
114
         %Acceleration due to the solar radiation presure.
115
   읒
         usun=rsun/norm(rsun); %unit vector of the pos of the sun
116
         us=y(1:3)/norm(y(1:3)); %unit vector of the pos of the sat
117
118
         shadow=random('Normal', 0.5, .1); %shadow factor
119
   응
120
   응
   9
         asrp=srp*rs2s/norm(rs2s)^3;
121
122
       %Acceleration due to the drag force
123
       % Compute exospheric temperature [K]
124
       T = 900 + 2.5*(100-70);
125
126
        % Compute effective atmospheric molecular mass [km/K], valid 180<H<500
       M = 27 - 0.012 * (h - 200);
127
       % Compute atmospheric scale height [km]
128
       SH = T / M;
129
        % Compute atmospheric density [kg/m3]
130
       rho = 6E-10 * exp(-(h - 175) / SH);
131
132
       atmosV=y(4:6) + omega*[y(2); -y(1); 0];
133
       %relative vel between sat and atmos
134
135
       adrag=-.5*rho*Bc*atmosV*norm(atmosV);
136
137
       138
       acc=aobl+asun+amoon+asrp+adrag;
139
140
       ode=@(t,y)[y(4);
141
142
                y(5);
                y(6);
143
                acc(1) - mu/norm(y(1:3))^3*y(1);
144
                acc(2) - mu/norm(y(1:3))^3 * y(2);
145
                acc(3) - mu/norm(y(1:3))^3 * y(3)
146
```



```
]; %function to integrate (6eqn)
147
148
149
       [t,x] = ode45 (ode, [t0 tf], y);
150
       y=x(size(x,1),:);
151
       %save one value each 12 hours
152
       if mod(t(length(t)), 12*3600) == 0
153
154
          tt=[tt; t(length(t))]; xx=[xx; y];
       end
155
      y=y';
156
157
158
       %End of the timestep
159
160
       t0=tf;
      tf=tf+dt;
161
162
      h = (norm(y(1:3)) - re) *1e-3;
163
164
165
      iter=iter+1;
       if mod(iter, 300) == 0
166
          fprintf('Days elapsed: %0.1f \nH=%0.3f km\n', (jdate-jdate0),h);
167
168
       end
169 end
170 toc
172 fprintf('The satellite decays in %0.0f days',tf/86400);
173
174 h=zeros(size(xx,1),1);
175    for i=1:size(xx,1)
      h(i) = (norm(xx(i,1:3)) - re) *1e-3;
177 end
179 plot(tt/86400,h);
```

```
1 % This function converts kepler orbit elements to cartesian coordinates
2
3 % a: semimajor axis
4 % e: eccentricity
5 % i: inclination [rad]
6 % Omega: longitude of the ascending node [rad]
7 % w: Argument of periapsis [rad]
8 % nu: True anomaly [rad]
9
10 function [y] = cartesian2(a,e,i,Omega,w,nu)
11 global mu;
12 % Position in cylindrical coordinates
13 r = a*(1-e^2)/(1+e*cos(nu));
14
15 % Position components
```



#### **Performance Evaluator**

```
1 %----ASTREA CONSTELLATION----
2 %PROJECTS - 220028
3 %Aerospace Engineering Barchelor's Degree
4 %ESEIAAT - UPC
5 %Autumn 2016-2017
7 % ORBIT DESIGN TEAM
8 % PERFORMANCE EVALUATOR
10 clear all;
11
12 %% Input Data
13
a=6378.01e3;
u=3.986012e14;
h = (a+546.5101e3)/a;
17 ap=h*a;
18 J2=1.0826e-3;% J2=0;
19
20 ws=sqrt(u/ap^3); % Satellite mean motion
21 we=2*pi/(24*3600); % Earth mean motion
22 dOmega=-1.5*J2*(a/ap)^2*sqrt(u/ap^3)*cosd(75);
23
24 At=1/6; % [min] % Every 10 seconds
25 t=0:At:60*24; t=t*60; % Time array in seconds
i=72*pi/180;
29 p=9;
```



```
30 spp=21;
31 N=p*spp;
32 %f=floor(0.25*spp);
34 fprintf('\ni=%gdeg, p=%g, spp=%g, N=%g\n',i*180/pi,p,spp,N);
35
36 latGS=57.5*pi/180;
37 longGS=(17.73-60)*pi/180; longGS=0;
38
39 fprintf('GS Coordinates: Lat=%g, Long=%g \n',latGS*180/pi, longGS*180/pi)
40
41 f1=pi/2-latGS; f2=longGS;
42 R1=[\cos(f1) \ 0 \ -\sin(f1); 0 \ 1 \ 0; \sin(f1) \ 0 \ \cos(f1)];
43 R2 = [\cos(f2) \sin(f2) 0; -\sin(f2) \cos(f2) 0; 0 0 1];
44 RotGS=R1*R2;
45
46 % Rotation Matrix:
47 % From ECEF - To SEZ
48 % Where:
49 % ECEF = Earth Centered Earth Fix - Coordinate System
50 % SEZ = Topocentric Horizon - Coordinate System
51
52 emin=20*pi/180;
53
54 %% Le simulasion
55 %Omega = 0:30:240; Omega=Omega*pi/180;
omega = 0:225/(p-1):225; Omega=Omega*pi/180;
57 %Omega=0:360/p:360-360/p; Omega=Omega*pi/180;
58
nu = zeros(spp,p);
60 angle = zeros(1,p);
61 X = zeros(3, spp*p);
62
63 flight_time=zeros(1,1000); % Length of the flyby of a satellite
64 links_at_time=zeros(1,1000); % Number of links when one sat has finished
65 time_end_flyby=zeros(1,1000); % When did this flyby end?
67 contact=zeros(1,length(t)); % Number of links
68
69 quality_time=zeros(1,length(t));
70 %If a flyby lasts longer than 3 minutes,
71 % then the previous 3 minutes were successfuly covered.
73 time_record=zeros(1,N);
                               % Accumulates Timesteps being on the GS
74 before_tracking=zeros(1,N);
75 % Boolean to know which ones are already passing by
76 now_tracking=zeros(1,N); % Boolean to know which ones are now passing by
                              % Number of flight paths computed
77 Nflyby=0;
78
79 for n=1:length(t)
```



```
80
       % J2 deviation
81
       Omegat=Omega+dOmega*t(n);
83
84
       % Ground Station Coordinates
       f1=we*t(n)+longGS;
85
       f2=latGS;
86
       RGS=cos(f2);
87
       XGS=[RGS*cos(f1);RGS*sin(f1);sin(f2)];
88
89
       R2 = [\cos(f1) \sin(f1) 0; -\sin(f1) \cos(f1) 0; 0 0 1];
90
       RotGS=R1*R2;
91
92
93
       % Constellation Coordinates
       nu_t=ws*t(n); % True anomaly due to time passing by
94
       for j = 1:p
95
           angle(j) = f*2*pi*(j-1)/(spp*p); %Phasing due to f between planes
96
           for k = 1:spp
97
                % True anomalies of the s satellites
98
               A = spp*(j-1)+k; % Number of the satellite
99
               nu(k,j) = 2*pi*(k-1)/spp+angle(j);
100
               X(:,A) = cartesian(h,0,i,Omegat(j),0,nu(k,j)+nu_t);
101
102
           end
103
       end
104
       % Contact Evaluation
105
       win=0; % Counter to know number of links
106
       now_tracking=zeros(1,N);
107
       for sat=1:N
108
109
           X_hor=RotGS*(X(:,sat)-XGS);
      fprintf('%g,%g,%g --> |r|=%g\n',X_hor(1),X_hor(2),X_hor(3),norm(X_hor))
110
           ang=asin(X_hor(3)/norm(X_hor));
111
           if ang \geq emin
112
                win=win+1;
113
               time_record(sat) = time_record(sat) + 1;
114
               now_tracking(sat)=1;
115
           end
116
117
           118
            if before_tracking(sat) == 1 && now_tracking(sat) == 0
119
120
               Nflyby=Nflyby+1;
                flight_time(Nflyby) = At *time_record(sat);
121
122
               links_at_time(Nflyby) = contact(n-1);
123
               time_end_flyby(Nflyby)=t(n);
124
125
                % THE KEY OUESTION
126
127
                % Was this flyby useful?
               if flight_time(Nflyby)>3
128
129
                    start=n-time_record(sat);
```



```
130
                    finish=n;
131
                    quality_time(start:finish) = quality_time(start:finish)+1;
132
                end
                time_record(sat)=0;
133
134
            end
       end
135
136
       before_tracking=now_tracking;
137
       contact(n)=win;
138
   end
139
140 %% Post-Processing
141 % PLOT 1: NUMER OF LINKS
142 figure(1)
143 title('Links vs Time')
144 plot(t/(3600),contact)
145 axis([min(t)/(3600) max(t)/(3600) 0 max(contact)+1]);
   fails=length(find(contact==0));
146
147
    ratio=fails/length(contact);
148
   fprintf('Links on GS
                                  : %g percent of the time\n', (1-ratio) *100)
149
150 % PLOT 2: LENGTH OF THE LINKS
151 figure (2)
152 plot(flight_time(1:Nflyby));
153 titola=['Length of the ' num2str(Nflyby) ' flyby s. Mean time = '...
       num2str(mean(flight_time(1:Nflyby))) 'min'];
154
155 title(titola)
156 ylabel('Length (minutes)')
157 xlabel('Contact')
158 %xlim([1 Nflyby])
160 % PLOT 3: ANALYSIS OF THE FLYBYS
161 figure (3)
index=1:Nflyby; % X variable to the following plots
   plot(time_end_flyby(index)/3600,flight_time(index),...
163
         time_end_flyby(index)/3600,links_at_time(index),...
164
        t/3600, quality_time);
165
166
   titola='Flybys Analysis';
167
168 title(titola)
169
170 legend('Length of flybys', 'Links by end of flyby',...
        'Num of sats @flybys longer than 3 min')
172 xlabel('Time (h)')
173
174 epic_wins=length(find(quality_time≥1));
175 ratio_covered=epic_wins/length(t);
176 fprintf('Quality flybys on GS: %g percent of the time\n',ratio_covered*100)
177 fprintf('Mean flyby time = %g minutes\n', mean(flight_time(1:Nflyby)))
178
179
```



```
180 %% MAXIMUM GAP SEARCH
181 gap=0;
182 at_gap=0;
183 gap_record=zeros(1,100);
184 ngap=0;
185
   for n=1:length(t)
186
187
        if quality_time(n) == 0
188
            at_gap=1;
            gap=gap+1;
189
190
        end
191
        if quality_time(n)>0 && at_gap==1
192
            at_gap=0;
193
            ngap=ngap+1;
194
195
            gap_record(ngap) = gap * At;
            gap=0;
196
197
        end
198 end
199
200 max_gap=max(gap_record);
201 fprintf('Maximum gap = %g minutes\n', max_gap)
202 fprintf('Number of gaps = gnn', ngap)
```

### **Thrust**

```
1 %----ASTREA CONSTELLATION----
2 %PROJECTS - 220028
3 %Aerospace Engineering Barchelor's Degree
4 %ESEIAAT - UPC
5 %Autumn 2016-2017
7 % ORBIT DESIGN TEAM
8 % THRUST
10 % This function computes the \Delta V and propellant mass necessary to
11 % maintain an orbit between to heights
12
13 %function [mp,\Delta V,tHoh] = thrust(hmax,hmin,ms,Isp,Thr)
14
15 % This function computes the \Delta V and propellant mass necessary to
16 % maintain an orbit between to heights
17
18 % Input variables:
19 % - hmax: maximum height [m]
20 % - hmin: minimum height [m]
```



```
21 % - ms: dry mass of the spacecraft [kg]
22 % - Isp: specific impulse of the spacecraft [s]
23 % - Thr: thrust of the spacecraft [N]
25 % Output variables:
26 % - mp: array of propellant mass necessary for every Hohmann transfer [kg]
27 % - \Delta V: array of \Delta Vs necessary for every Hohmann transfer [m/s]
28 % - tHoh: array of time necessary to do a Hohmann transfer [s]
30 % Proposed values:
n = 550e3;
32 hmin = hmax-8;
ms = 3.95;
                   % Dry mass [kg]
34
   Isp = 2150;
35 Thr = 100e-6;
36
37 %% Data
38
39 % PHYSICAL CONSTANTS AND PARAMETERS
40 RE = 6.378e6; %Earth Radius [m]
41 mu = 3.986e14; %GM Earth
42 	 g0 = 9.81;
43
44 % SPACECRAFT DATA
45 m = 4;
               % Satellite mass [kg]
46 Cd = 2.2; % Drag Coefficient
47 A = 0.1*0.3; % Satellite surface [m^2] --> 3U pointing to Earth
49 % ORBIT DATA
52 IO = 80*pi/180; % Inclination [deg]
53 A0 = RE + H0;
55 % SIMULATION PARAMETERS
56 N=100000;
57 M = 10000;
59 %% 2. Perturbations propagation
61 % Creation of the matrices
62 temp=zeros(M,N);
63 temp(1,1)=0;
64 H = zeros(M,N);
65 H(1,1) = H0;
66 % \Delta V = zeros(2,N);
68 % Asign initial values
69 W0=0; OMEGA0=0;
70 a=zeros(1,N); a(1)=A0;
```



```
71 e=a; e(1)=E0;
72 i=a; i(1)=I0;
73 t=a; t(1)=0;
74 w=a; w(1)=W0;
75 Omega=a; Omega(1)=OMEGA0;
76
   for j = 1:M
77
78
        n = 1;
79
        while a(n) \ge (RE + hmin)
80
81
             n = n+1;
82
83
84
             a0=a(n-1);
             e0=e(n-1);
85
             i0=i(n-1);
86
             w0=w(n-1);
87
             Omega0=Omega(n-1);
88
89
             P=2*pi*(a0^3/mu)^.5; % Orbit Period [s]
90
91
92
             Bc = m/(Cd*A); Ballistic coefficient of the satellite
             Perti=Perturbation(a0,e0,i0,Bc);
93
94
             a(n) = a0 + sum(Perti(:,1));
95
             e(n) = e0 + sum(Perti(:, 2));
96
             i(n) = i0 + sum(Perti(:,3));
97
             w(n) = w0 + sum(Perti(:, 4));
98
             Omega(n) = Omega0 + sum(Perti(:,5));
99
100
             % We don't want angles bigger than 360!
101
             w(n) = w(n) - 360 * floor(w(n) / 360);
102
             Omega(n) = Omega(n) -360 \times floor(Omega(n)/360);
103
104
             if Omega(n)>2*pi
105
                  Omega(n) = Omega(n) -2*pi;
106
             end
107
108
             t(n) = t(n-1) + P;
109
110
        end
111
        % Hohmann
113
114
        [\Delta V1, \Delta V2, mpit, tHohit] = Hohmann(a(n)-RE, hmax, m, Isp);
        m = m-mpit;
115
116
        if m<ms
117
118
            break
        end
119
120
```



```
\Delta V (1, j) = \Delta V1; % first row of the column -> \Delta V1
121
        \Delta V(2,j) = \Delta V2; % second row of the column -> \Delta V2
122
123
        mp(j) = mpit;
        tHoh(j) = tHohit;
124
125
126
       a(1) = A0;
       e(1) = E0;
127
       i(1) = I0;
128
       w(1) = W0;
129
       Omega(1)=OMEGA0;
130
131
       altura = a-RE;
132
      H(j,:) = altura;
133
134
        temp(j,:) = t;
135
136 end
137
138 %% Post process
139
140 mfr = Thr/(g0*Isp);
141
142 for i = 1:N
      if H(1,i) \leq 0
143
        break
144
145
        end
146 end
147 for j = 1:M
    if H(j,1) \leq 0
148
            break
149
150
        end
151 end
152 H = H(1:(j-1), 1:(i-1));
153
154 for i = 2:N
    if temp(1,i) \le 0
155
            break
156
      end
157
158 end
159 for j = 2:M
    if temp(j,2) \le 0
160
            break
161
162
        end
163 end
164 temp = temp(1:(j-1),1:(i-1));
165 for j = 2:size(temp,1)
        temp(j,:) = temp(j,:)+tHoh(j-1)+temp(j-1,size(temp,2));
167 end
169 % Convert matrix to vector
170 H = reshape(H.', 1, size(H, 1) * size(H, 2));
```



```
171 temp = reshape(temp.',1,size(temp,1)*size(temp,2));
172
173 figure (1)
174 plot (temp/(3600*24), H/1000)
175 grid on
176 ylabel('Orbit height [km]')
177 xlabel('Time [days]')
tfin=floor(temp(length(temp))/(3600\times24));
179 days=num2str(floor(tfin/365));
180 titulaso=['Orbit decay in ' num2str(tfin) ' days = ' days ' years'];
181 title(titulaso)
182
183 %% Final calculations
184
185 % % Perimeter of the ellipse
186 % r1 = RE+hmax; % semimajor axis
187 % r2 = RE+hmin; % semiminor axis
188 % h = (r1-r2)^2/(r1+r2)^2;
189 % % Ramanujan approximation
190 % C = pi*(r1+r2)*(1+3*h/(10+sqrt(4-3*h)));
191 %
192 % % Check if \Delta V is possible for the given thruster
193 % OK = true;
194 % for j = 1:length(mp)
        for i = 1:2
             mfrnecessary = \Delta V(i,j) * mp(j) / (C/2);
196 %
             if mfrnecessary>mfr
                 OK = false;
199 응
             end
         end
201 % end
203 % if OK ==true
      fprintf('The trajectory is possible\n\n');
205 % else
         fprintf('Trajectory not possible\n\n');
207 % end
208 if length(mp)<M
       fprintf('There is still propellant left ._.\n\n')
210 end
```

```
1 function [ΔV1,ΔV2,mp,t] = Hohmann(hinicial,hfinal,m,Isp)
2
3 % Hohmann transfer orbit between two circular orbits
4 % - hinicial: height of the first orbit [m]
5 % - hfinal: height of the second orbit [m]
6 % - m: mass of the satellite (total) [kg]
7 % - Isp: Specific impulse [s]
8 % - mp: fuel mass [kg]
```



```
9 % - t: time needed to do the maneuver [s]
nu mu = 3.986004418e14; % Standard gravitational parameter (Earth)
12 REarth = 6.371e6; % [m]
13 g0 = 9.81; % Earth's gravity [m/s^2]
14
15 % First orbit
16 r1 = REarth+hinicial; % [m]
v1 = sqrt(mu/r1); % [m/s]
18
19 % Second orbit
20 r2 = REarth+hfinal; % [m]
v2 = sqrt(mu/r2); % [m/s]
23 % Transfer orbit (ellipse)
24 vp = sqrt(2*mu*r2/(r1*(r1+r2))); % [m/s]
va = sqrt(2*mu*r1/(r2*(r1+r2))); % [m/s]
a = (r1+r2)/2;
27 T = sqrt(2*pi^2*a^3/mu);
29 % \DeltaV 1 -> p(transfer orbit)
30 \DeltaV1 = vp-v1;
31 % \Delta V a(transfer orbit) -> 2
32 \Delta V2 = v2-va;
33 % Total
34 \Delta V = \Delta V1 + \Delta V2;
36 % Fuel mass
37 mp = m*(1-exp(-\Delta V/(g0*Isp)));
39 % Time
40 t = T/2;
41
42 end
```

## **Satellites Datasheet**

```
1 %----ASTREA CONSTELLATION----
2 %PROJECTS - 220028
3 %Aerospace Engineering Barchelor's Degree
4 %ESEIAAT - UPC
5 %Autumn 2016-2017
6
7 % ORBIT DESIGN TEAM
8 % SATELLITES DATASHHET
9
```



```
10 % Writing the constellation caracteristics
11
a=6378.01e3;
u=3.986012e14;
h = (a+542e3)/a;
15 ap=h*a;
16 J2=1.0826e-3;% J2=0;
17
18 ws=sqrt(u/ap^3); % Satellite mean motion
i=72*pi/180;
20
21 p=9;
22 spp=21;
23 N=p*spp;
24 %f=floor(0.25*spp);
26 fprintf('\ni=%gdeg, p=%g, spp=%g, N=%g\n',i*180/pi,p,spp,N);
27
28
  %Omega = 0:30:240; Omega=Omega*pi/180;
30 Omega = 0:225/(p-1):225; %Omega=Omega*pi/180;
31
   %Omega=0:360/p:360-360/p; Omega=Omega*pi/180;
32
33 P=2*pi/ws/60;
34
35 %% Excel IMPORT
36
37 D=zeros(N,9);
38 for sat=1:N
      Om=ceil(sat/spp); ID_p=sat-(Om-1)*spp;
40
      ph = (ID_p-1) *360/spp;
41
      %[ID Plane h P i e Omega phase per]
42
       D(\text{sat}, 1:9) = [\text{sat Om (ap-a})/1000 P i*180/pi 0 Omega(Om) ph 0];
43
44
46 NameCells=['B2:I' num2str(N)];
47 xlswrite('SatsDatasheet.xls',D,NameCells)
48
49 %% Latex IMPORT
50
  for sat=1:N
51
52
       Om=ceil(sat/spp); ID_p=sat-(Om-1)*spp;
53
       ph = (ID_p-1) *360/spp;
54
      %[ID Plane h P i e Omega phase per]
       D(\text{sat}, 1:9) = [\text{sat Om } (\text{ap-a})/1000 \text{ P } i*180/\text{pi } 0 \text{ Omega}(\text{Om}) \text{ ph } 0];
56
58 end
59
```



```
60 [rows, cols] = size(D);
61 file=fopen('Autotable.txt','w');
63 for i=1:rows
      Text=['AstreaSAT ' num2str(i) ' $ '];
      for j=2:cols
65
          if j≠cols
               Text=[Text num2str(D(i, j-1)) ' $ '];
67
68
           else
               Text=[Text num2str(D(i, j-1)) ' \\ '];
69
70
           end
      end
71
       fprintf(file,[Text '\n']);
72
73 end
75 fclose(file);
```

#### **Ground Station Localization**

```
1 %----ASTREA CONSTELLATION----
2 %PROJECTS - 220028
3 %Aerospace Engineering Barchelor's Degree
4 %ESEIAAT - UPC
5 %Autumn 2016-2017
7 % COMMUNICATION TEAM
8 % GROUND STATIONS LOCALIZATION
10 clc; clear;
11
12 %% Input Data
13 N_sat=21; %number of sats in a plane
14 N_planes=9; %number of orbital planes
15 h=542;
            %high of the sats in km
             %inclination of the orbits in degrees
16 I=72;
17 phase=210/(N_planes-1); %Angle between planes in the equator
18 At=0.1; %time step in minutes
19 T=48;
             %time to simulate in hours
20 t=0:At:T*60; %array of time in minutes
                 %minumum elevation in degrees
21 e_min=7.5;
22 lambda=57.5; %latitudes to simulate in degrees
23 lat=length(lambda);
24 mu=0;
                  %longitudes to simulate in degrees
25 long=length(mu);
26 l=length(t);
27
```



```
1 function [X] = Orbital_position(N_sat, N_planes, h, I, mu, t)
2 %N_sat: number of sats per plane
3 %N_planes: number of planes
4 %h: Orbit high in km
5 %I: inclination of the orbital plane in degrees
6\, %mu: Angle between two adjacent orbital planes at the equatiorial plane in
7 %degrees
8 %t: array of time in minutes
10 I=I*pi/180; %I in rad
11 mu=mu*pi/180; %mu in rad
12 Omega=0:mu:(N_planes-1)*mu;
13 %longitude of the ascending node respect the inertial X axe [rad]
14
15 %for defining the orbits are defined a system of
16 %local axes X_o in which the y_o is the rotation axe and x_o is
17 %contained in the equatiorial plane. Thetransformation for the inertial
18 %system X_I to the local system 2 rotations are defined.
19 %FIRST ROTATION: it defines a intermediate system X_1. The inertial system
20 %rotates arround z_I an Omega angle.
21 %SECOND ROTAtion: The X_1 system is rotated (90deg-I) arround x_1
23 L_10=[1 \ 0 \ 0; \ 0 \ \sin(I) \ \cos(I); \ 0 \ -\cos(I) \ \sin(I)]; %1st rotation matrix.
24 %It transform a vector from local coordinates X_o to intemediate
25 %coordinates X_1. This matrix is the same for every orbital plane
26 %since all orbits have the same inclination
28 Re=6371; %Earth radius km.
29 %The distances are going to be refered to the earth radius
30 h=h/Re;
            %high in earth radius
31 GM=6.474e-11*5.972e24*3600/Re^3;
32 %Gravitational constant*earth mass [(Re^3)/min^2]
33 Ro=1+h; %orbital radius in earth radius
w=sqrt(GM/(Ro^3)); %sat's angular speed [rad/min]
```



```
36 phi=0:2*pi/N_sat:(N_sat-1)*2*pi/N_sat;
37 %relative angle between every sat of a plane respect the first one
38 f=8;
39 %factor to module the relative desfase of the first sat of one plane
40 %to the first sat of the next plane
41 l=length(t);
42
43 for p=1:N_planes
                       %going through each plane
44
       phi_0(p) = f*2*pi*(p-1)/(N_planes*N_sat);
45
       %relative desfase of the first sat of one plane to the
46
       %first sat of the next plane
47
48
       One = cos(Omega(p));
49
       Two = sin(Omega(p));
50
       L_I1=[One -Two 0; Two One 0; 0 0 1];
51
       %2nd rotation matrix. It transforms a vector from
52
53
       %intermediate coordinates X_1
       \mbox{\ensuremath{\mbox{$^{\circ}$}}} to inertial coordinates X_I. This matrix is characteristic of every
54
       %plane since evey plane has a diferent longitude of the ascending node
55
       %(Omega)
56
57
       L_eo=L_I1*L_1o;
58
59
       %Global rotation matrix form local coordinates to nertial coordinates
60
                       %going through every sat of the plane
61
       for s=1:N_sat
62
           for k=1:1 %going through every time step
63
64
65
                First = Ro*cos(w*t(k)+phi(s)+phi_0(p));
                Second = -\text{Ro}*\sin(w*t(k)+\text{phi}(s)+\text{phi}_0(p));
66
                X_loc=[First;0;Second];
67
                %Coordinates of the sat at the given time in
68
                %local coordinates X_o.
69
                %It defines a circumference at the plane x_o-z_o with a angular
70
                %velocity w, a initial phase respect the first sat and a
71
                %initial phase respect the first sat of the previous orbit
72
                %first sat.
73
74
                X(:, (p-1)*N_sat+s, k) = L_eo*X_loc;
75
                %Coordinates of the sat in inertial system
76
                %(Coordinate [x y z], Sat number, instant(time)]
77
           end
78
       end
80 end
  end
```

```
1 function [X] = Ground_position(lambda,long,t)
2 %Gorund_position calculates the position of a place of the earth, with a
```



```
3 %wiven longitude and latitude, in an inertial system of coordiantes X_I in
4 %a given period of time
6 %lamda: is the latitude of the place in degrees[-90,90]
7 %long: is the longitude of the place iin degrees [0,360)
8 %t: is a array of time in minutes
10 long=long*pi/180; %longin radians
11 lambda=lambda*pi/180; %lambda in radians
12
13 Rg=cos(lambda);
14 %distance between the Earth point to the rotation axis of the Earth [0,1]
15 hg=sin(lambda);
16 %distance between th Earth point to the equatorial plane [0,1]
17 %this 2 distances are measured in earth radius.
19 w=2*pi/(24*60); %angular velocity of the Earth [rad/min]
20
21 X=[Rg*cos(w*t+long);Rg*sin(w*t+long);hg*ones(1,length(t))];
22 %position of the point in X_I
23 %for the given time. The point descrives a cicumference
24 %with a Rg radius in the horizontal
25 %plane x_I-y_I with an angular velocity w and a initial phase long. The
26 %coordinate z_I is allways hg. It is expresed in earth radius.
27 %X(cooridinates, instant (time)
29 end
```

# Received signal strength variation with bandwidth

```
1 %ASTREA CONSTELLATION
2 %PROJECTS - 220028
3 %Aerospace Engineering Barchelors Degree
4 %ESEIAAT - UPC
5 %Autumn 2016-2017
6
7 % SATELLITE DESIGN TEAM
8 % SENSITIVITY vs. BANDWITH
9
10 %{
11 This matlab routine calculates the Bandwidth to received signal Strength
12 correlation.
13 %}
14 clc; clear; close all;
15 %
16 %% Sband calculus
```



```
C = 25*8*1e6; %bits/s
18 B = 100; %Hz
19 N = 1.38e-23*87.79*B; %W
20 Binf=5e6; Bsup=56e6; Bstep=1e4;
22 i=1;
23 S=zeros(size(Binf:Bstep:Bsup));
24 for B = Binf:Bstep:Bsup
      N = 1.38e - 23 * 87.79 * B;
       S(i) = 10 * log 10 (100 * (N * (2^(C/B) - 1)));
      SdB(i) = 10*log10*(1000*S/1);
27
      i=i+1;
28
29 end
30 Bs = Binf:Bstep:Bsup;
31 Ss = S;
32
33 %% Xband calculus
34 Binf=5e6; Bsup=100e6; Bstep=1e4;
35
37 S=zeros(size(Binf:Bstep:Bsup));
38 for B = Binf:Bstep:Bsup
     N = 1.38e - 23 * 87.79 * B;
      S(i) = 10 * log10 (100 * (N * (2^{(C/B)}-1)));
40
       SdB(i) = 10*log10*(1000*S/1);
41
       i=i+1;
42
43 end
44 Sx=S; Bx = Binf:Bstep:Bsup;
45 %% PLOT results
46 singleFig(Bx, Sx);
```

```
function singleFig(X1, Y1)
    %CREATEFIGURE1(X1, Y1)
    % X1: vector of x data
4 % Y1: vector of y data

6 % Auto-generated by MATLAB on 15-Dec-2016 02:41:15

7
    % Create figure
9 figure1 = figure('PaperSize',[20.98404194812 29.67743169791]);

10
11 % Create axes
12 axes1 = axes('Parent',figure1,'YGrid','on','XGrid','on','FontSize',12);
13 box(axes1,'on');
14 hold(axes1,'all');
15
16 % Create plot
17 plot(X1,Y1,'LineWidth',2,'Color',[0 0 1]);
18
```



```
19 % Create xlabel
20 xlabel('Bandwidth [Hz]','Interpreter','latex','FontSize',14);
21
22 % Create ylabel
23 ylabel('S [dBm]','Interpreter','latex','FontSize',14);
24
25 % Create title
26 title('Received power with bandwith variation','Interpreter','latex',...
27 'FontSize',16);
28 legend1 = legend('S vs. B','show');
29 set(legend1,'Location','Best');
```

# Received signal strength for distance variation

```
1 %----ASTREA CONSTELLATION----
2 %PROJECTS - 220028
3 %Aerospace Engineering Barchelors Degree
4 %ESEIAAT - UPC
5 %Autumn 2016-2017
7 % SATELLITE DESIGN TEAM
8 % FRIIS EQUATION RANGE CALCULATOR
10 % {
11 This matlab routine calculates the strength of the received signal for
12 distance variation. Taking into account losses considered on the Link
13 Budget Annex. Calculus are based on Friis equation and Bandwidth-Strength
14 correlation.
16 clc; clear; close all;
18 %% Initial variables definition
19 r_start = 100; %km
20 \text{ r\_end} = 1000; \%km
21 n=1; %step
22 Lprop=0;
24 %% Sband Space to Ground station calculus
25 Gt=10; %dBi
26 Gr=35.4; %dBi
27 f=2228; %MHz
28 Pw=2:
29 L_cables = 2; Labs=0.58; Laml=1; Lph=20;
30 Lprop = L_cables+Labs+Laml+Lph;
31 [ PrS, radiusS, FSL_S]=friis_eq( Gt, Gr, f, Pw, r_start, r_end, n, Lprop);
32
```



```
33 %% Sband intersatellite
34 Gt=10; %dBi
35 Gr=10; %dBi
36 f=2228; %MHz
37 Pw=2;
38 %L_cables = 2; Labs=0.58; Laml=1; Lph=20;
39 Lprop = 0;%L_cables+Labs+Laml+Lph;
40 [PrSi, radiusSi, FSL_Si]=friis_eq( Gt, Gr, f, Pw,r_start, r_end, n,Lprop);
42 %% Xband signal strength between satellites calculation
43 Gt=16.5; %dB
44 Gr=16.5; %dB
45 f=9e3; %MHz
46 Pw=12; %W
47 [PrX1, radiusX1, FSL_X1] = friis_eq( Gt, Gr, f, Pw,r_start, r_end, n,Lprop);
49 %% Xband between sat to ground
_{50} r_start = 100; %km
r_end = 1000; %km
52 n=1; %step
53 Gt=16.5; %dB
54 Gr=36; %dB
55 f=10e3; %MHz
56 Pw=12; %W
57 L_cables = 2; Labs=0.58; Laml=1; Lph=20;
58 Lprop = L_cables+Labs+Laml+Lph;
59 [PrX2, radiusX2, FSL_X2] = friis_eq( Gt, Gr, f, Pw, r_start, r_end, n, Lprop);
61 %% PLOT results
62 X1=radiusS;
63 YMatrix1=[PrS' PrSi' PrX1' PrX2'];
64 llegenda1='Sband Sat to GND';
65 llegenda2='Sband Sat to Sat';
66 llegenda3='Xband Sat to Sat ';
67 llegenda4='Xband Sat to GND';
68 eixX='Distance [Km]';
69 eixY='Power [dB]';
70 titol='Change in received power';
71 createfigure (X1, YMatrix1, llegenda1, llegenda2, ...
      llegenda3, llegenda4, titol, eixX, eixY)
73 legend(llegenda1, llegenda2, llegenda3, llegenda4)
```

```
1 function [Pr,radius,FSL]=friis_eq(Gt, Gr, f, Pw, r_start, r_end, n,Lprop)
2 %FRIIS_EQ: Calculate the received on ground power in dBm.
3 % Gt, Gr in dBi
4 % f in Mhz
5 % Pw in watts
6
7 %initiate
```



```
8 i=1;
9 Pr=zeros(size(r_start:n:r_end));
10 radius=zeros(size(Pr));
11 FSL=zeros(size(Pr));
12 P = 10 * log10 (Pw/1e-3);
13 for r=r_start:n:r_end
14
      Ls = 32.45+20*\log 10(r)+20*\log 10(f); %r-->Km; f-->Mhz%Spatial losses
15
      Pr(i) = P + Gt + Gr - Ls - Lprop;
16
       radius(i) = r;
17
       FSL(i) = Ls;
       %fprintf('Received power %fdBm \n',Pr(i));
19
20
       i=i+1;
21 end
22
23
24 end
```

```
1 function createfigure(X1,YMatrix1,llegenda1,llegenda2,llegenda3,...
      llegenda4,titol,eixX, eixY)
4 %CREATEFIGURE(X1, YMATRIX1)
5 % X1: vector of x data
6 % YMATRIX1: matrix of y data
8 % Auto-generated by MATLAB on 30-Nov-2016 11:29:05
10 % Create figure
figure1 = figure('Name', titol, 'NumberTitle', 'off');
12 colormap('hsv');
14 % Create axes
axes1 = axes('Parent', figure1, 'XMinorTick', 'on', 'CLim', [0 1]);
16 box(axes1, 'on');
17 hold(axes1, 'all');
18
19 % Create multiple lines using matrix input to plot
20 plot1 = plot(X1, YMatrix1, 'Parent', axes1, 'LineWidth', 2);
21 set(plot1(1), 'Color', [1 0 0], 'DisplayName', llegenda1);
22 set(plot1(2), 'Color', [1 0.694117665290833 0.39215686917305],...
       'DisplayName', llegenda2);
24 set(plot1(3),...
       'Color',[0.0431372560560703 0.517647087574005 0.780392169952393],...
25
       'DisplayName', llegenda3);
26
28 % Create title
29 title(titol,'Interpreter','latex','FontSize',20);
31 % Create xlabel
```



```
32 xlabel(eixX,'Interpreter','latex','FontSize',16);
33
34 % Create ylabel
35 ylabel(eixY,'Interpreter','latex','FontSize',16);
36
37 % Create legend
38 legend1 = legend(axes1,'show');
39 set(legend1,'Location','Best');
```