

Basic hydrodynamics and Introduction to FLASH

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Outline:

- Astrophysical fluids
- Basic hydrodynamic equations
- Introduction to the FLASH code



Why we need simulations in Astronomy?

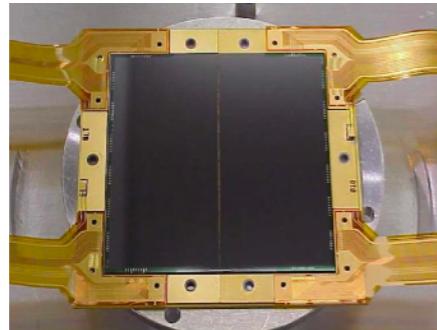
- We **receive** and **record**, but we do not experiment or perturb (too distant)
- Astrophysical systems involve extreme conditions that cannot be replicated in lab (mostly)
- Timescales are >> human lifetime (mostly)



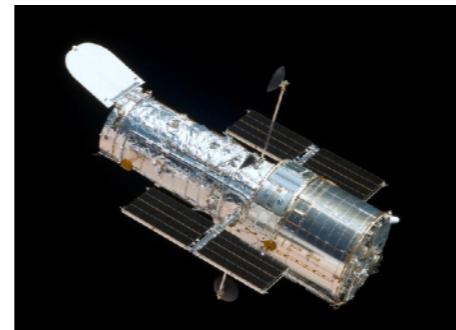
Observation and Computation Analogies



Astrophysical object



Detectors

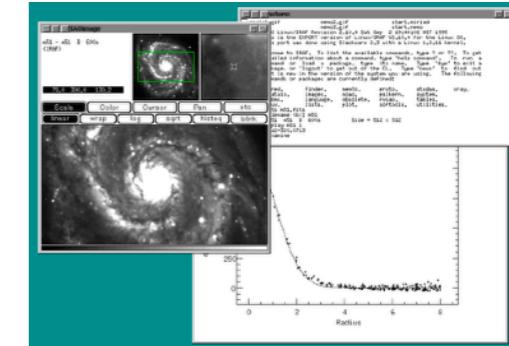


Telescope

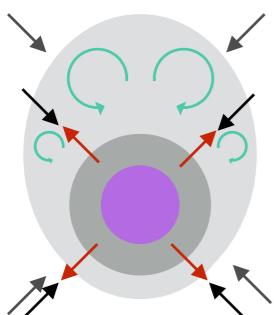
Observation



DATA



Data reduction / Calibration



```
*****  
#include "Neut.h"  
#include "Driver.h"  
#include "Log.h"  
  
using namespace std;  
using namespace Neut;  
using namespace Driver;  
using namespace Log;  
  
int main()  
{  
    Driver driver;  
    Log log;  
    Neut neut;  
  
    driver.start();  
    log.start();  
    neut.start();  
  
    // ...  
    // Your code here  
    // ...  
  
    driver.stop();  
    log.stop();  
    neut.stop();  
}
```

Mathematical Model

Codes

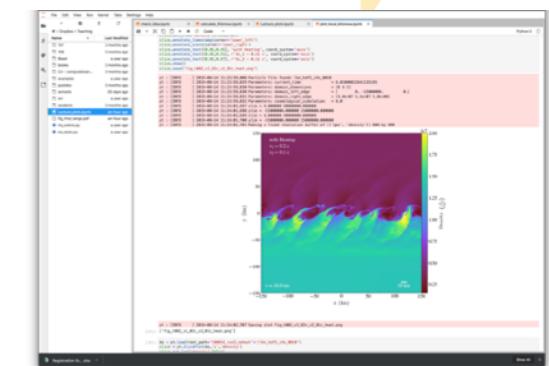


Supercomputer

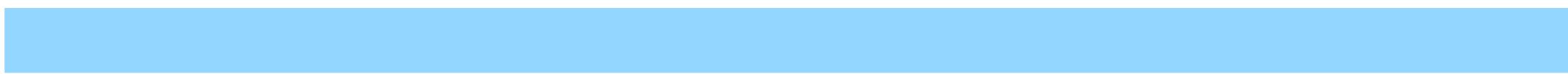
Simulation



DATA

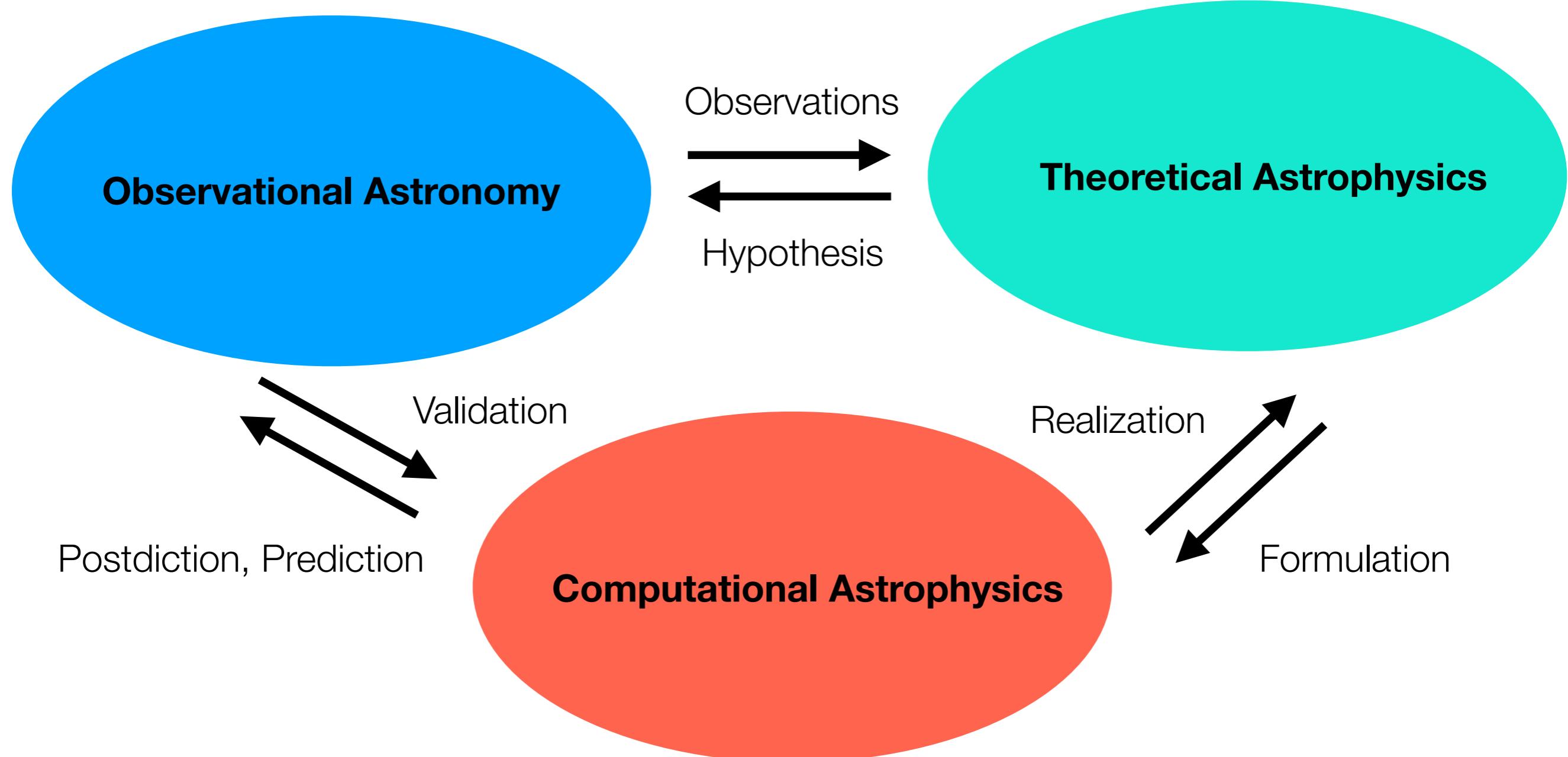


Data analysis / Visualization





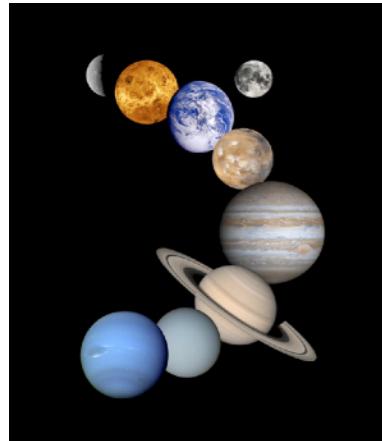
Frontier Astronomy Research



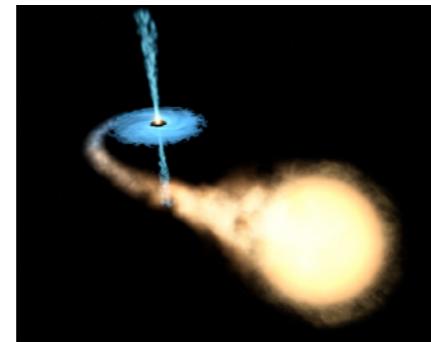
Michael Norman (1997)



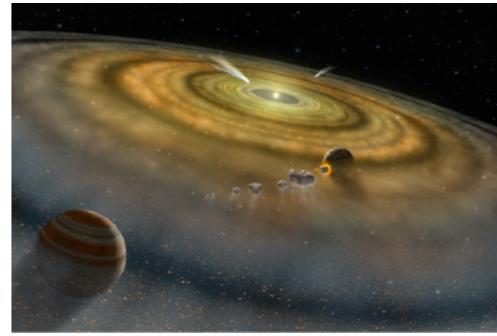
Astrophysical Problems



Planets
(~ Earth radius)



Stars/Binaries
(~0.1-1000 AU)



Planetary disk
(~100-1000 AU)



Interstellar medium
(~10-100 pc)



Galaxy
(~50 kpc)



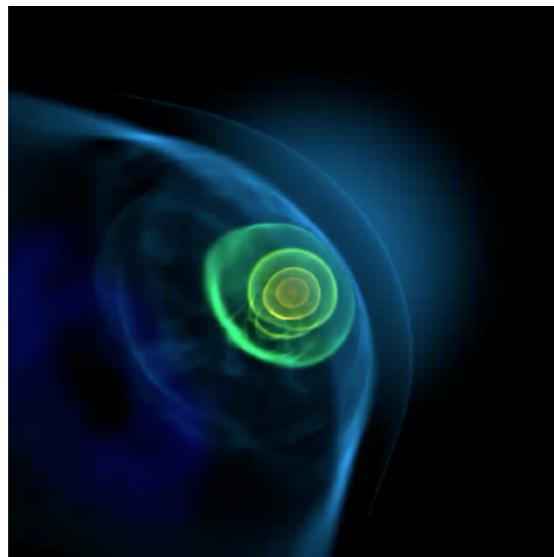
Galaxy clusters
(~1 Mpc)



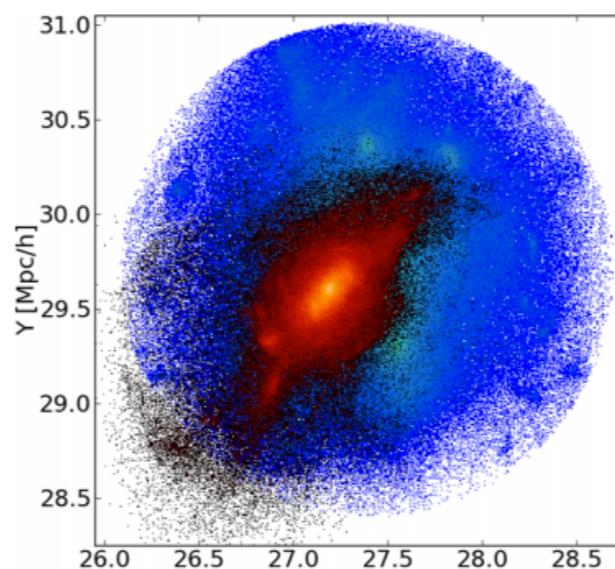
Large scale
structure
(~1 Gpc)



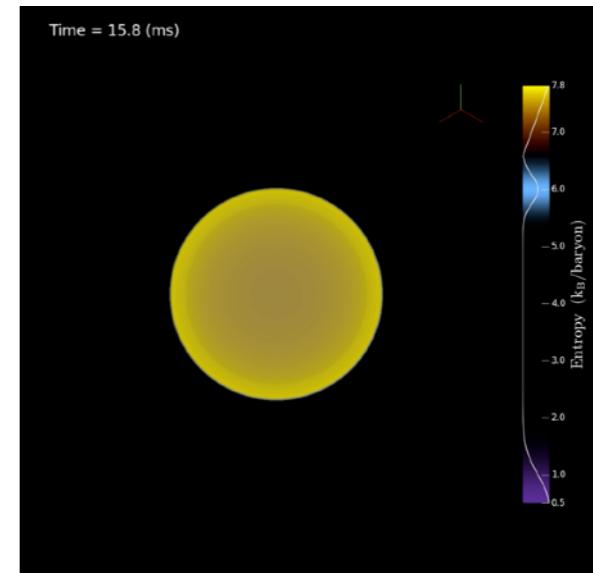
Astrophysical Simulations



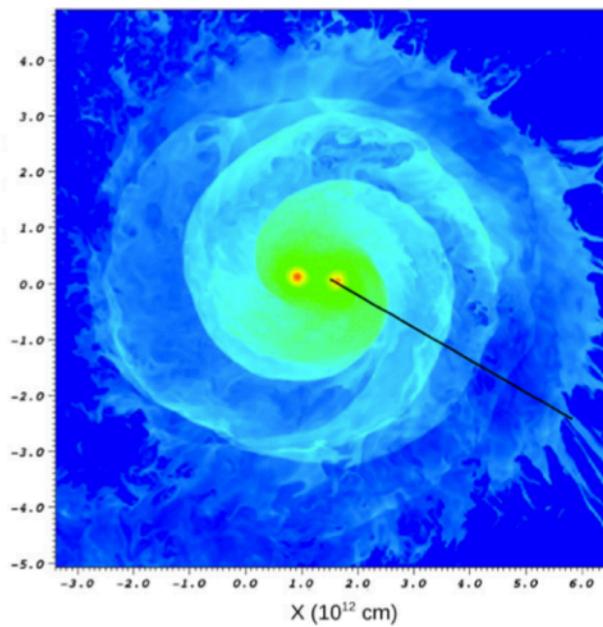
2012 ApJ, 750, 151



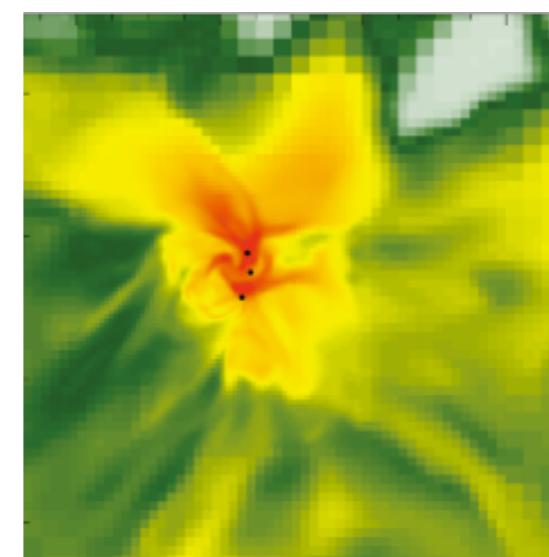
2013 MNRAS



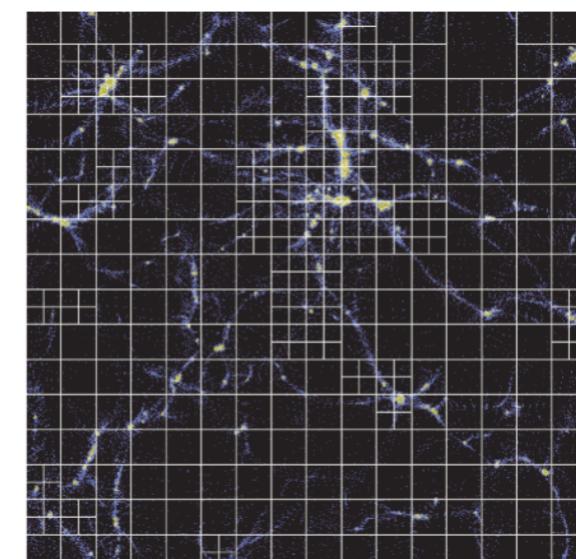
2018 ApJ



2012 ApJ, 746, 74



2009 MNRAS 398, 1082



2005 ApJS 160, 28



Fluid Approximation

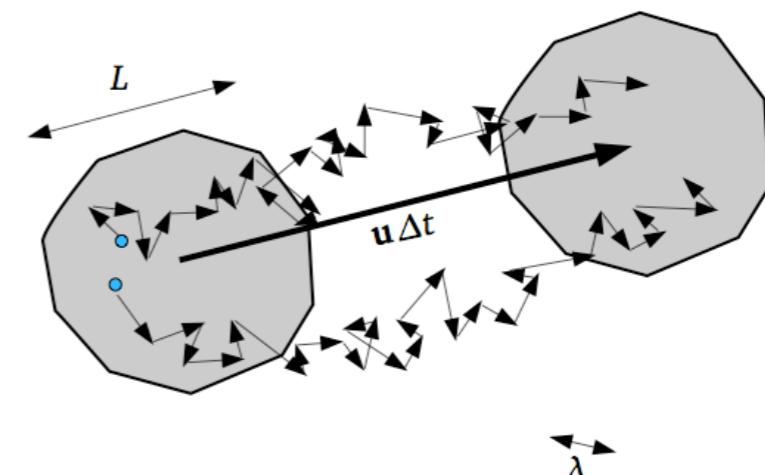
- To solve the gas dynamics equations, one could use **N-body** method if we know the force between two particles.
- However, the number of particles is usually huge ($N_A=10^{23}$). It is impossible to use direct N-body method to study the gas dynamics. -> **Fluid**



Fluid Approximation (continue)

$L \gg \lambda_{\text{mfp}}$, Mean free path

$T \gg \tau_{\text{coll}}$, Collision time scale



Exercise: Consider the air in this room

$$\rho \sim 10^{-3} \text{ g/cm}^3, \mu \sim 28m_p, T \sim 300 \text{ K}$$

$$\text{Cross section } \sigma \sim 3 \times 10^{-15} \text{ cm}^2$$

$$\text{RMS speed of nitrogen } v \sim \sqrt{\frac{8kT}{\pi\mu}} \sim 5 \times 10^4 \text{ cm/s}$$

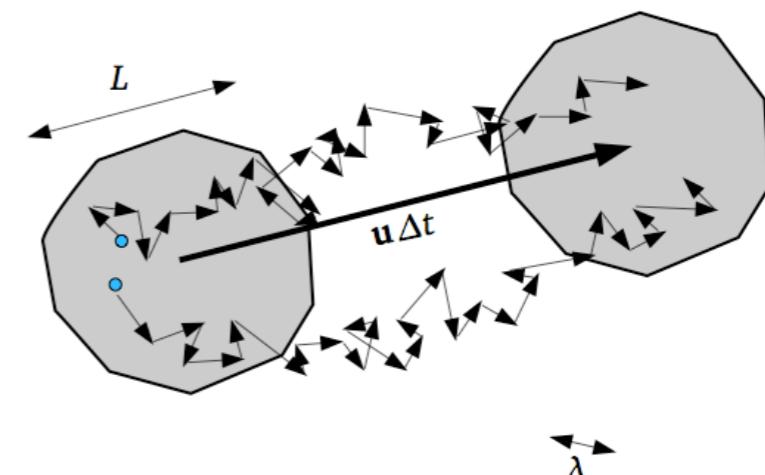
$$\tau_{\text{coll}} \sim (n\sigma v)^{-1} \sim 3 \times 10^{-10} \text{ sec} \quad \lambda_{\text{mfp}} \sim (n\sigma)^{-1} \sim 3 \times 10^{-5} \text{ cm}$$



Fluid Approximation (continue)

$L \gg \lambda_{\text{mfp}}$, Mean free path

$T \gg \tau_{\text{coll}}$, Collision time scale



Exercise: In ISM

Cross section $\sigma \sim?$ cm²

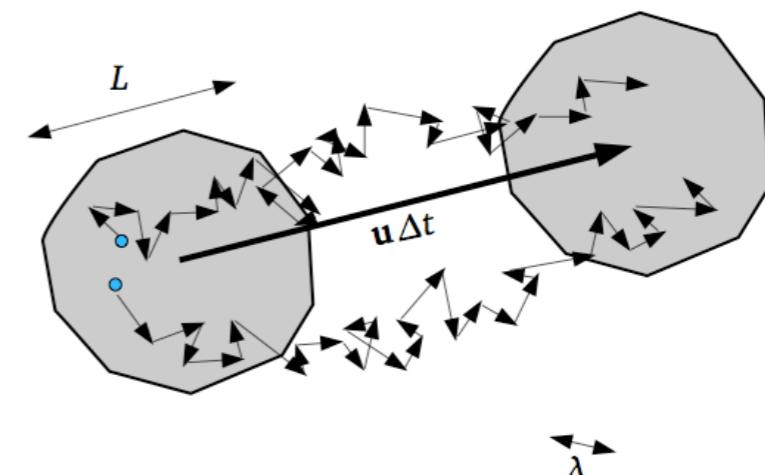
number density $n \sim?$ cm⁻³



Fluid Approximation (continue)

$L \gg \lambda_{\text{mfp}}$, Mean free path

$T \gg \tau_{\text{coll}}$, Collision time scale



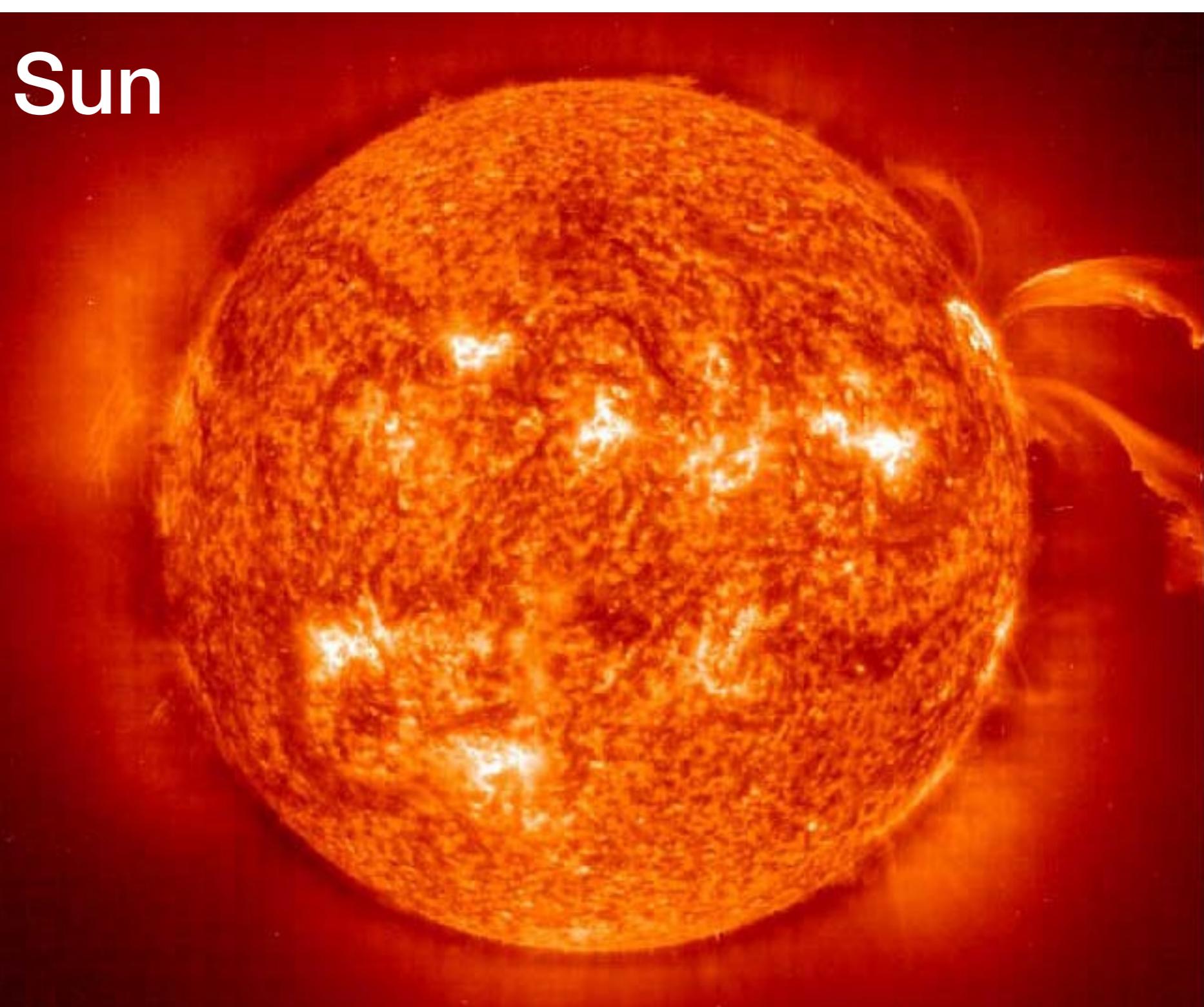
Exercise: In ISM

Cross section $\sigma \sim 3 \times 10^{-15} \text{ cm}^2$

number density $n \sim 1 \text{ cm}^{-3}$

$\lambda_{\text{mfp}} \sim (n\sigma)^{-1} \sim 10^{15} \text{ cm}$

$\tau_{\text{coll}} \sim (n\sigma v)^{-1} \sim 10^9 \text{ sec} \sim 40 \text{ yrs}$





Star forming region (S106)



Image credit: NASA

Cas A Supernova Remnant

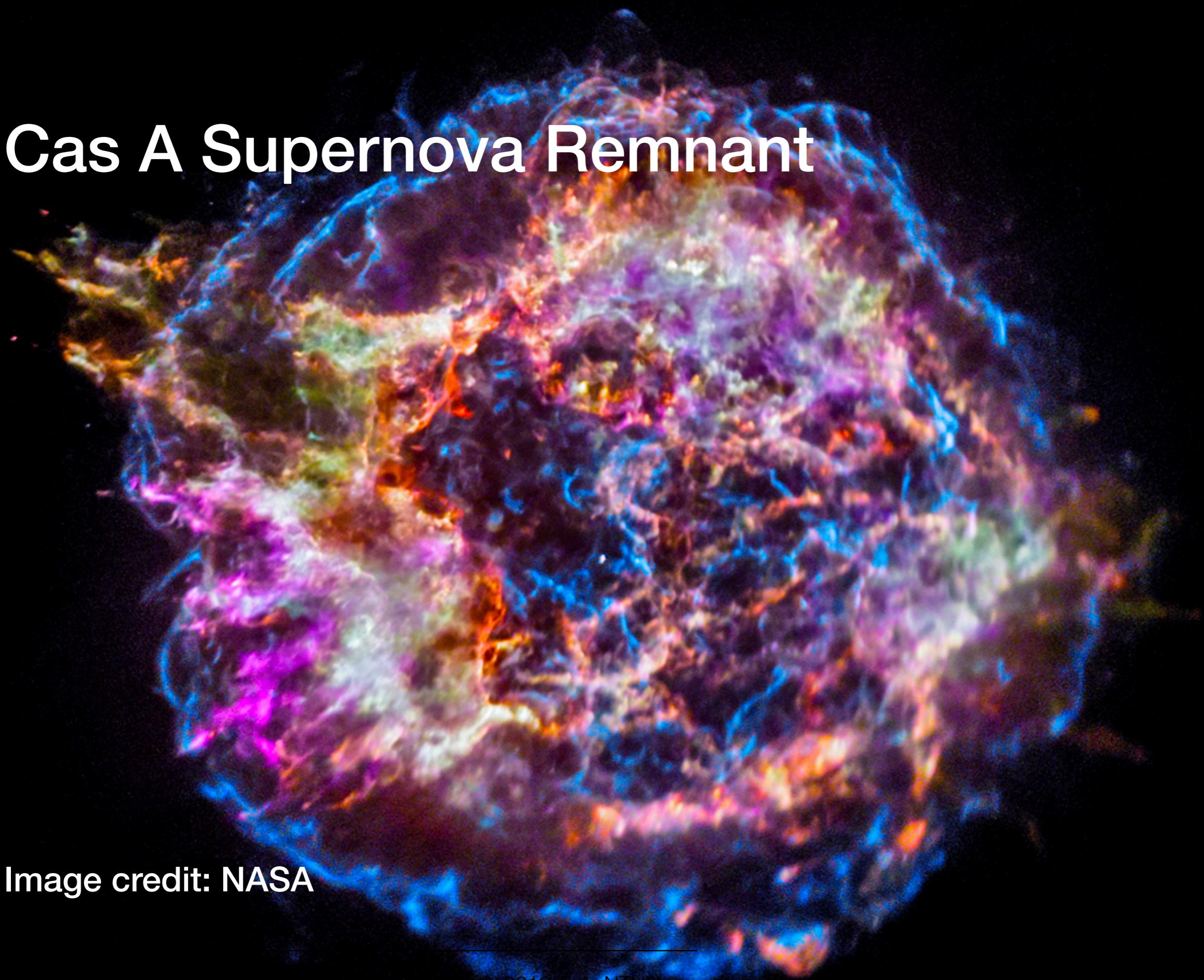


Image credit: NASA

Bullet cluster

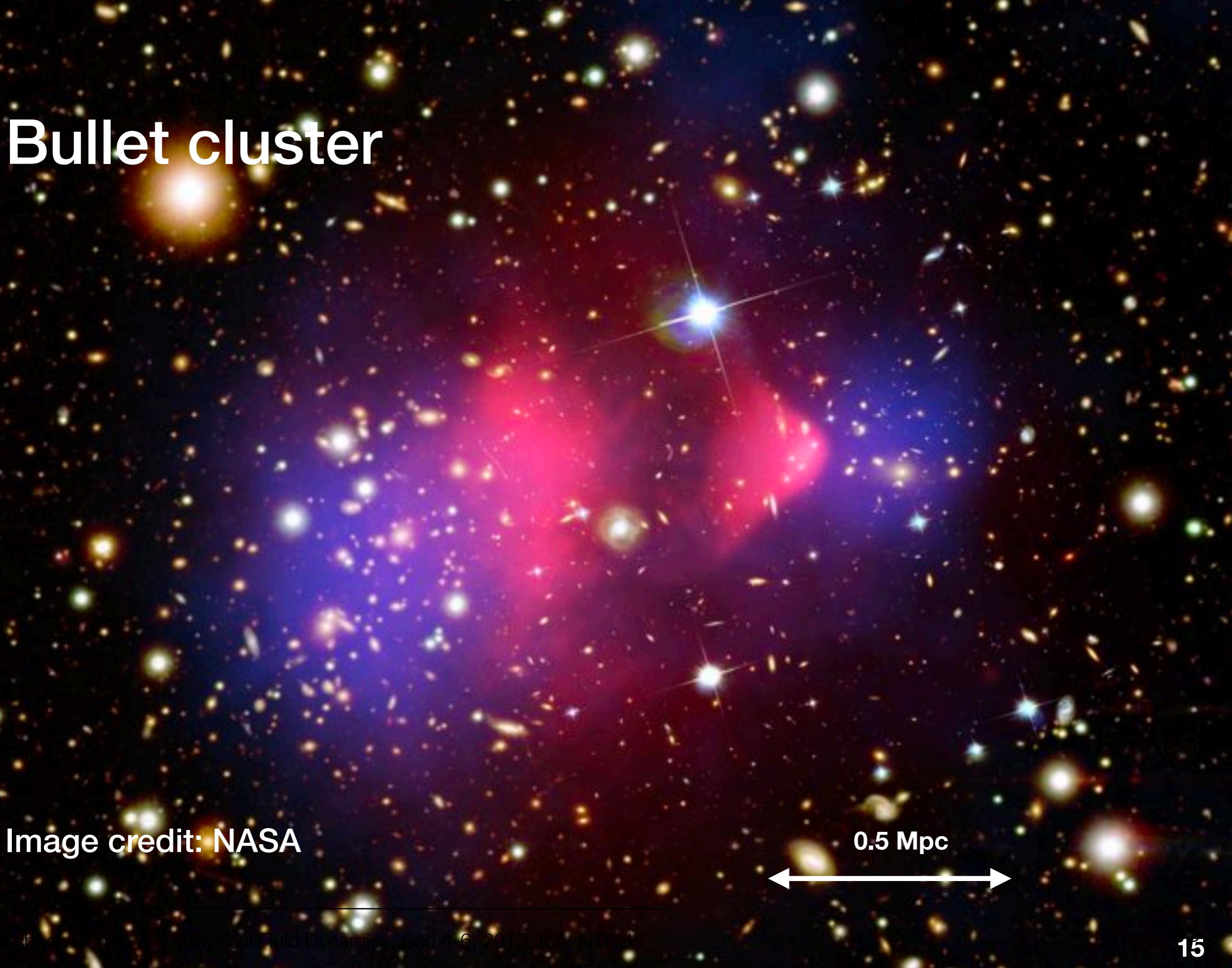


Image credit: NASA

0.5 Mpc

Astrophysical Fluids

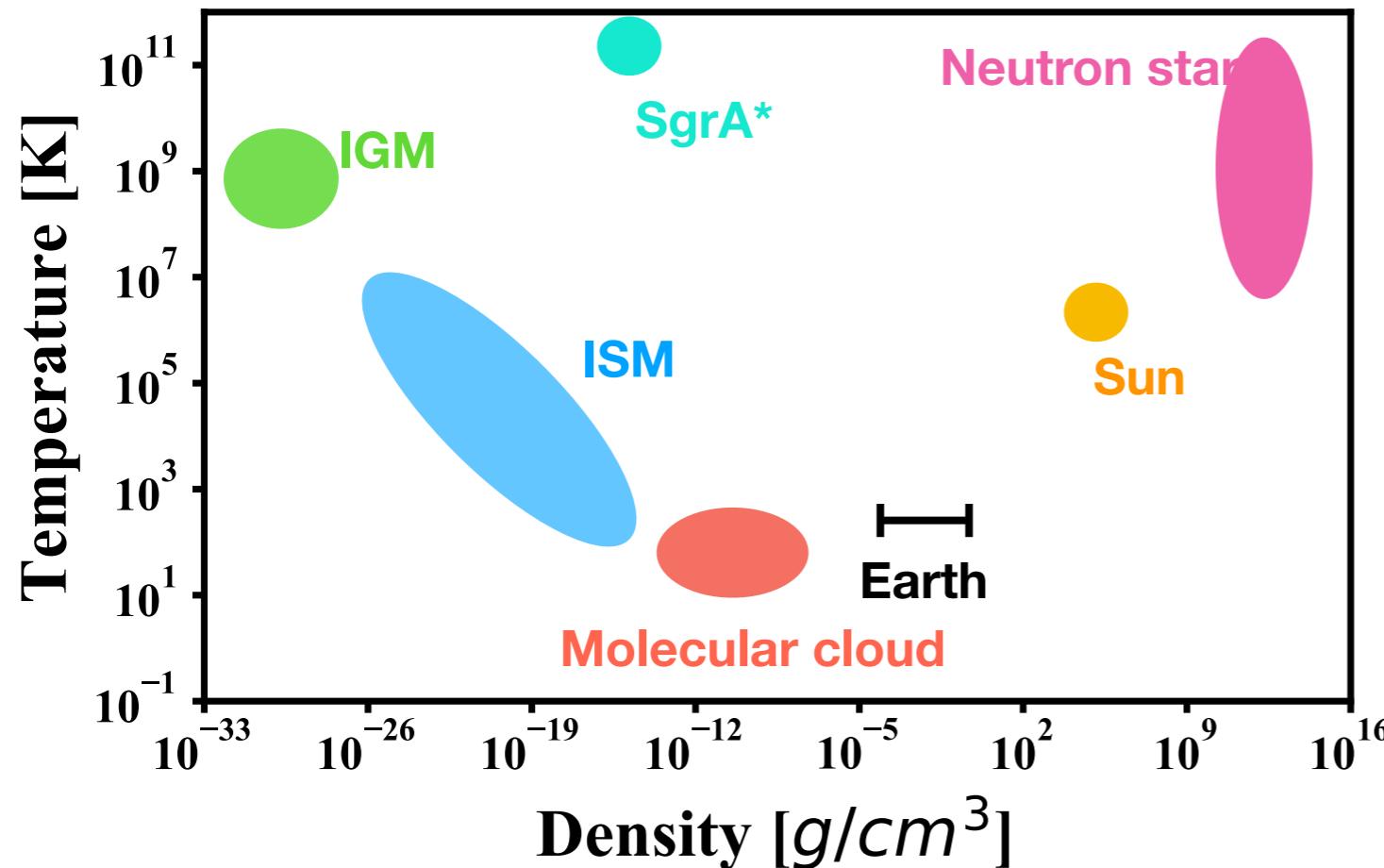


<https://youtu.be/lCx5ekWnUc>

16

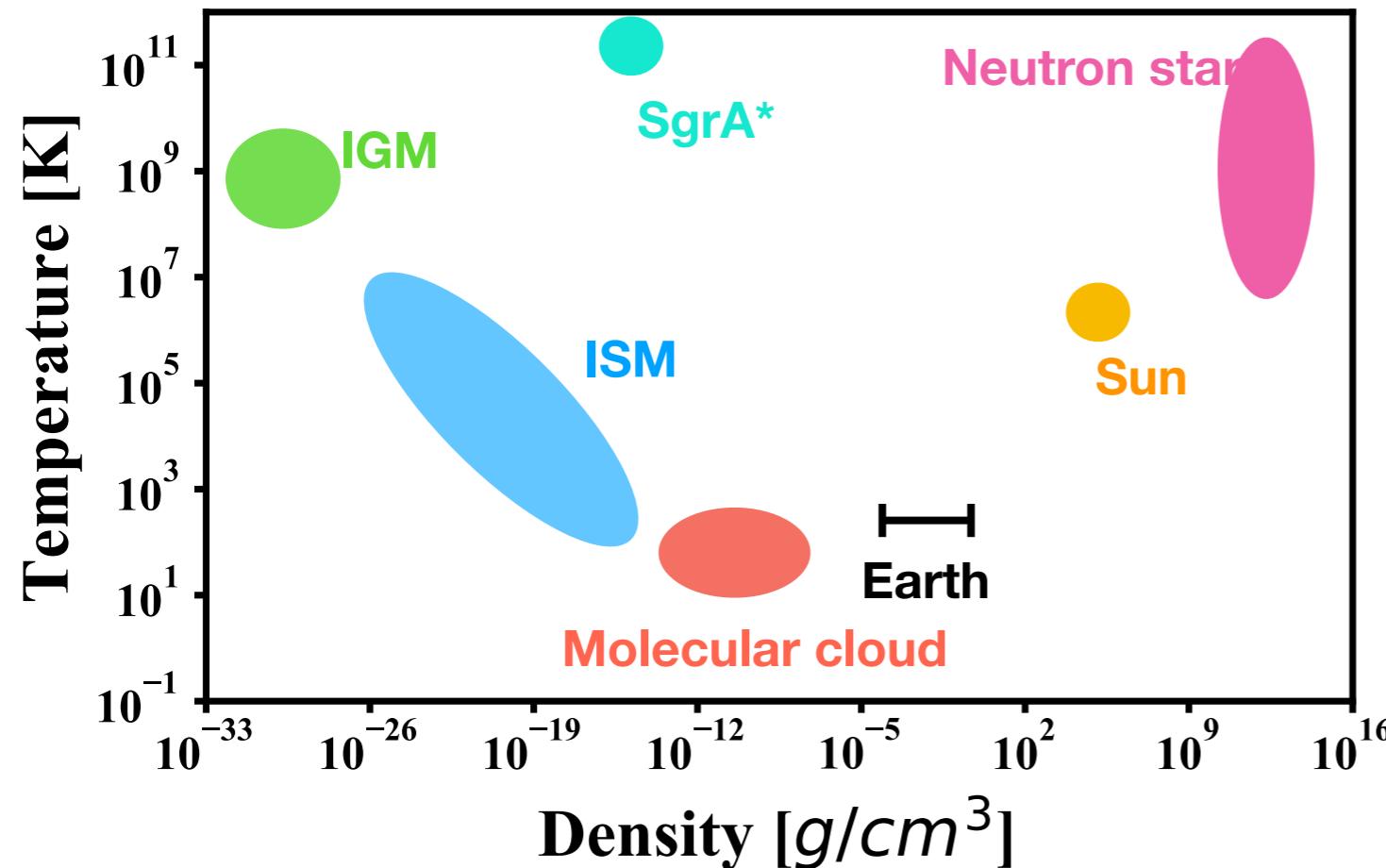


Astrophysical fluids



- Different with Earth-bound fluid
- Wide range of physical Conditions

Astrophysical fluids



- Magnetic fields are often important
- Compressibility is often important
- Occasionally self-gravitating and relativistic
- Non-ideal process are often unimportant
- Sometimes involves interactions with non-thermal process and strong radiation fields



Governing equations

The governing equations for ideal (inviscid) hydrodynamics. The main physical ideas are simple:

- Mass is conserved (1 constraint)
- Momentum is conserved (3 constraints)
- Energy is conserved (1 constraint)

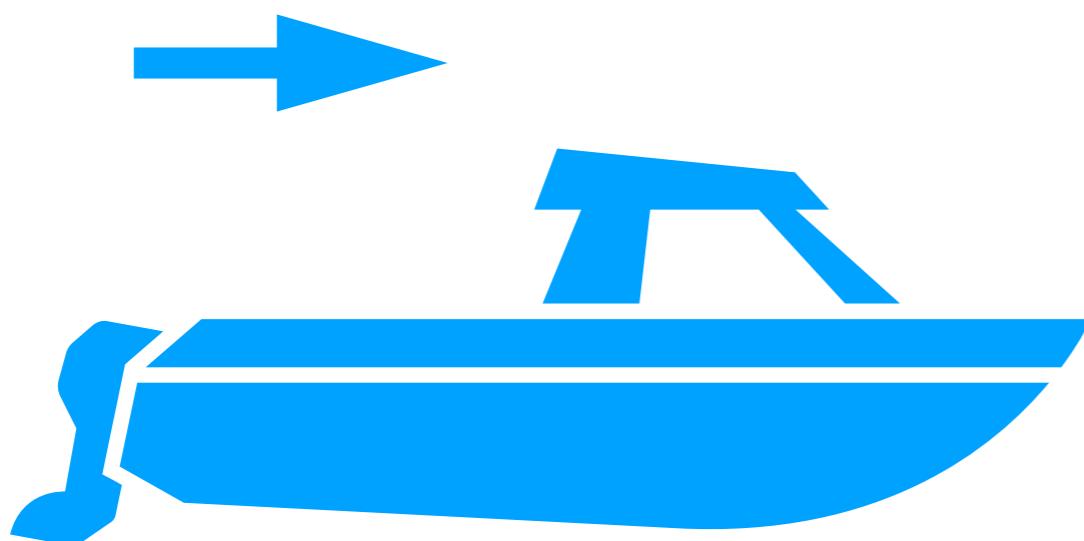
variables: ρ , \mathbf{v} , and u

ρ	Gas density
\mathbf{v}	Gas velocity
u	Gas internal energy



Coordinates

Eulerian



Lagrangian





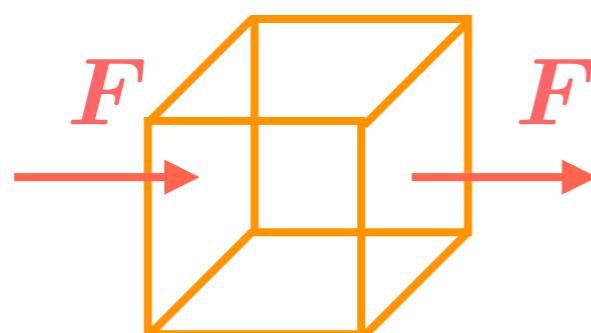
Conservational laws

The conservational laws can be written in what is called “**conservation form**”:

$$\partial_t(U) = -\nabla \cdot \mathbf{F}$$

“Gauss’s Theorem” or “divergence theorem”

$$\int_V (\nabla \cdot \mathbf{F}) dV = \oint_S \mathbf{F} \cdot \mathbf{n} dS$$



U Density of a quantity
 \mathbf{F} Flux density for that quantity



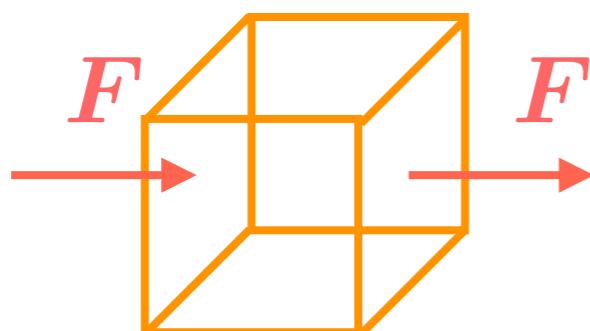
Continuity equation

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{v}) = 0$$

ρ Gas density
 \mathbf{v} Gas velocity

“Gauss's Theorem” or “divergence theorem”

$$\int_V (\nabla \cdot \mathbf{F}) dV = \oint_S \mathbf{F} \cdot \mathbf{n} dS$$





Convective derivative

- Consider a fluid element

$$\mathbf{v} = \frac{d\mathbf{r}}{dt}, \quad \mathbf{a} = \frac{d\mathbf{v}}{dt},$$

- Rewrite continuity equation

$$\begin{aligned}\partial_t \rho &= -\nabla \cdot (\rho \mathbf{v}) \\ &= -(\mathbf{v} \cdot \nabla) \rho - \rho (\nabla \cdot \mathbf{v})\end{aligned}$$

$$\begin{aligned}\mathbf{a} &= \frac{d\mathbf{v}(\mathbf{r}, t)}{dt} \\ &= \frac{\partial \mathbf{v}}{\partial t} + \frac{\partial \mathbf{v}}{\partial \mathbf{r}} \cdot \frac{\partial \mathbf{r}}{\partial t} \\ &= \frac{\partial \mathbf{v}}{\partial t} + (\mathbf{v} \cdot \nabla) \mathbf{v} \\ &\equiv \frac{D\mathbf{v}}{Dt}\end{aligned}$$

Convective derivative

$$\rightarrow \partial_t \rho + (\mathbf{v} \cdot \nabla) \rho = -\rho (\nabla \cdot \mathbf{v})$$

$$\rightarrow \frac{D\rho}{Dt} = -\rho (\nabla \cdot \mathbf{v})$$

Continuity equation



Momentum equation

- What accelerations act on a fluid element?

$$\mathbf{a} = -\frac{\nabla p}{\rho}$$

- This gives us the Euler equations

$$\frac{D\mathbf{v}}{Dt} = -\frac{\nabla p}{\rho}$$

or

$$\frac{\partial \rho \mathbf{v}}{\partial t} + \nabla \cdot (\rho \mathbf{v} \mathbf{v} + P \mathbf{I}) = 0$$



Momentum equation

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or

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Your exercise!



Internal Energy equation

- From first law of thermodynamics,

$$d\epsilon = \cancel{Tds}^0 - pdV, \quad dV = d\left(\frac{1}{\rho}\right) = \frac{-d\rho}{\rho^2}$$

Volume per unit mass

$$d\epsilon \equiv d\left(\frac{u}{\rho}\right) = \underline{\frac{du}{\rho}} - u\underline{\frac{d\rho}{\rho^2}} = -pdV = p\frac{d\rho}{\rho^2}$$

$$\frac{Du}{Dt} = \frac{u + p}{\rho} \frac{D\rho}{Dt} = \frac{\partial u}{\partial t} + (v \cdot \nabla)u = -(u + p) \nabla \cdot v$$

- Combine internal energy equation with the momentum equation

$$\frac{\partial \rho E}{\partial t} + \nabla \cdot [(\rho E + P)\mathbf{v}] = 0$$

Total Energy equation

$$E = \frac{1}{2}v^2 + \epsilon$$



Summary: Euler equations

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{v}) = 0$$

Continuity equation

$$\frac{\partial \rho \mathbf{v}}{\partial t} + \nabla \cdot (\rho \mathbf{v} \mathbf{v} + P \cdot \mathbf{I}) = 0$$

Momentum equation

$$\frac{\partial \rho E}{\partial t} + \nabla \cdot [(\rho E + P) \mathbf{v}] = 0$$

Energy equation

ρ Gas density

\mathbf{v} Gas velocity

p Gas pressure

ϵ Gas specific internal energy

E Gas specific total energy

u Gas internal energy

$$u = \rho \epsilon$$

$$E = \frac{1}{2} v^2 + \epsilon$$



Summary: Euler equations + Gravity

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{v}) = 0$$

Continuity equation

$$\frac{\partial \rho \mathbf{v}}{\partial t} + \nabla \cdot (\rho \mathbf{v} \mathbf{v} + P \cdot \mathbf{I}) = -\rho \nabla \Phi$$

Momentum equation

$$\frac{\partial \rho E}{\partial t} + \nabla \cdot [(\rho E + P) \mathbf{v}] = -\rho \mathbf{v} \cdot \nabla \Phi$$

Energy equation

$$\nabla^2 \Phi = 4\pi G \rho$$

Poisson equation

ρ Gas density

\mathbf{v} Gas velocity

p Gas pressure

ϵ Gas specific internal energy

E Gas specific total energy

u Gas internal energy

Φ Gravitational Potential

$$u = \rho \epsilon$$

$$E = \frac{1}{2} v^2 + \epsilon$$



Summary: Lagrangian form

$$\frac{D\rho}{Dt} = -\rho \nabla \cdot v$$

Continuity equation

$$\frac{Dv}{Dt} = -\frac{\nabla p}{\rho} - \nabla \Phi$$

Momentum equation

$$\frac{Du}{Dt} = -(u + p) \nabla \cdot v$$

Energy equation

ρ	Gas density
v	Gas velocity
p	Gas pressure
ϵ	Gas specific internal energy
E	Gas specific total energy
u	Gas internal energy
Φ	Gravitational Potential

$$u = \rho \epsilon$$

$$E = \frac{1}{2} v^2 + \epsilon$$



Equation of State (EoS)

Need an Equation of State to close system of equations.

For ideal gas,

$$P = (\gamma - 1)u = nk_B T$$

$n = \rho/m$ Number density

T Temperature

k_B Boltzmann constant

$\gamma = C_v/C_p$ Adiabatic index

- Isothermal EoS: $\gamma = 1$ $P = \rho c_s^2$
- Polytropic EoS: $P = K\rho^\gamma$



Finite Volume Method

		$U_{i,j+1}^n$		
	$U_{i-1,j}^n$	U_{ij}^n	$U_{i+1,j}^n$	
		$U_{i,j-1}^n$		

- Initial conditions of all hydrodynamics variables at time step n=0 (a given EoS)
- Boundary conditions at simulation boundaries
- Evolve hydrodynamics variables with a time step dt

$$\frac{\partial}{\partial t} \underbrace{\begin{pmatrix} \rho \\ \rho U \\ E \end{pmatrix}}_{W(x,t)} + \frac{\partial}{\partial x} \underbrace{\begin{pmatrix} \rho U \\ \rho U^2 + p \\ (E + p)U \end{pmatrix}}_{F(W)} = 0,$$

See Hsi-Yu Schive's lecture tomorrow



FLASH Code



FLASH Code: summary

- 3D AMR HD/MHD multi-physics code
- Mainly developed at FLASH center at University of Chicago (<http://flash.uchicago.edu>)
- Contribution from many individual groups
- Written principally in Fortran, with some C and Python (> 1 million lines; 25% are comments)
- Current release version 4.6.1
- > 700 scientists around the world
- > 1000 publications



FLASH Code: Requirements

- Fortran compiler
- MPI - Message Passing Interface
- HDF5 - Hierarchical Data Format



FLASH Code: research applications

- Thermonuclear flashes
- High energy density physics (HEDP)
- Fluid—structure interaction
- Star formation
- Star-star & Star-planetets interactions
- Core-collapse supernova
- Galaxy-galaxy cluster
- Magnetic field amplification
- Turbulence
- Cosmology
- ... (more)



FLASH Code: Physic solvers

- **Hydrodynamics**: unspilt PPM, WENO; split PPM; 2T + Radiation
- **Magneto-hydrodynamics (MHD)**: unspilt staggered mesh; split 8-wave
- **Equation of State (EoS)**: idea gas; degenerate ionized plasma; nuclear,..
- **Radiation Transfer**: multi-group flux-limited diffusion
- **Diffusion and Conduction**: implicit with AMR
- **Laser Energy Deposition**: geometric optics with inverse Bremsstrahlung
- **Opacity**: constant; multi-material tabular
- **Particles**: tracer; massive; sink; charged
- **Nuclear Burning**
- **Gravity**: constant; point mass; planar; self-gravity
- **Cosmology**
- **Heating and Cooling** as source terms
- **Magnetic Resistivity; Conductivity**
- **Primordial Chemistry**



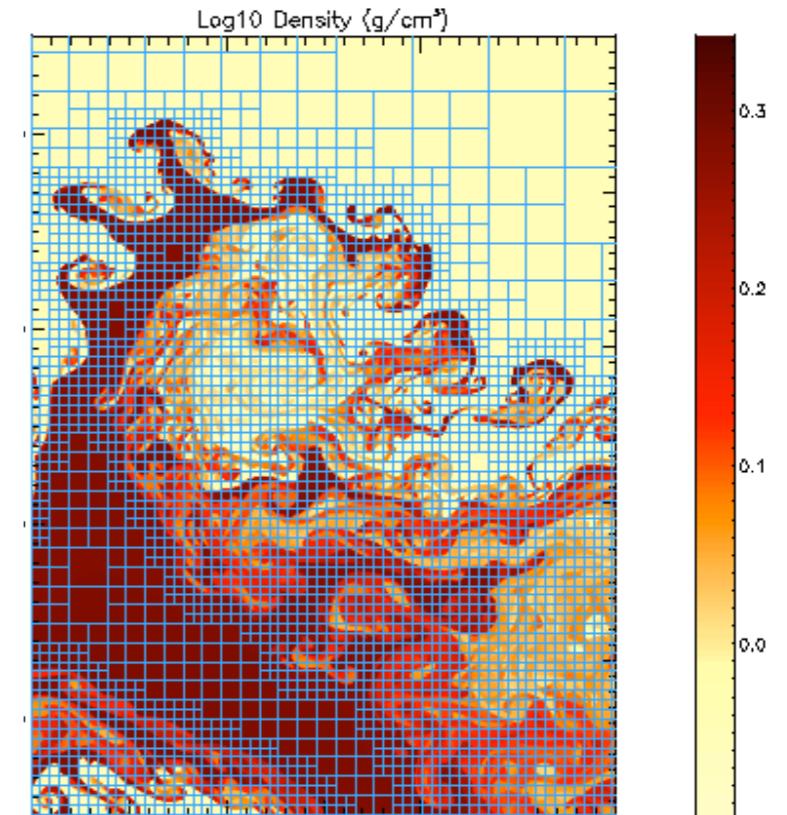
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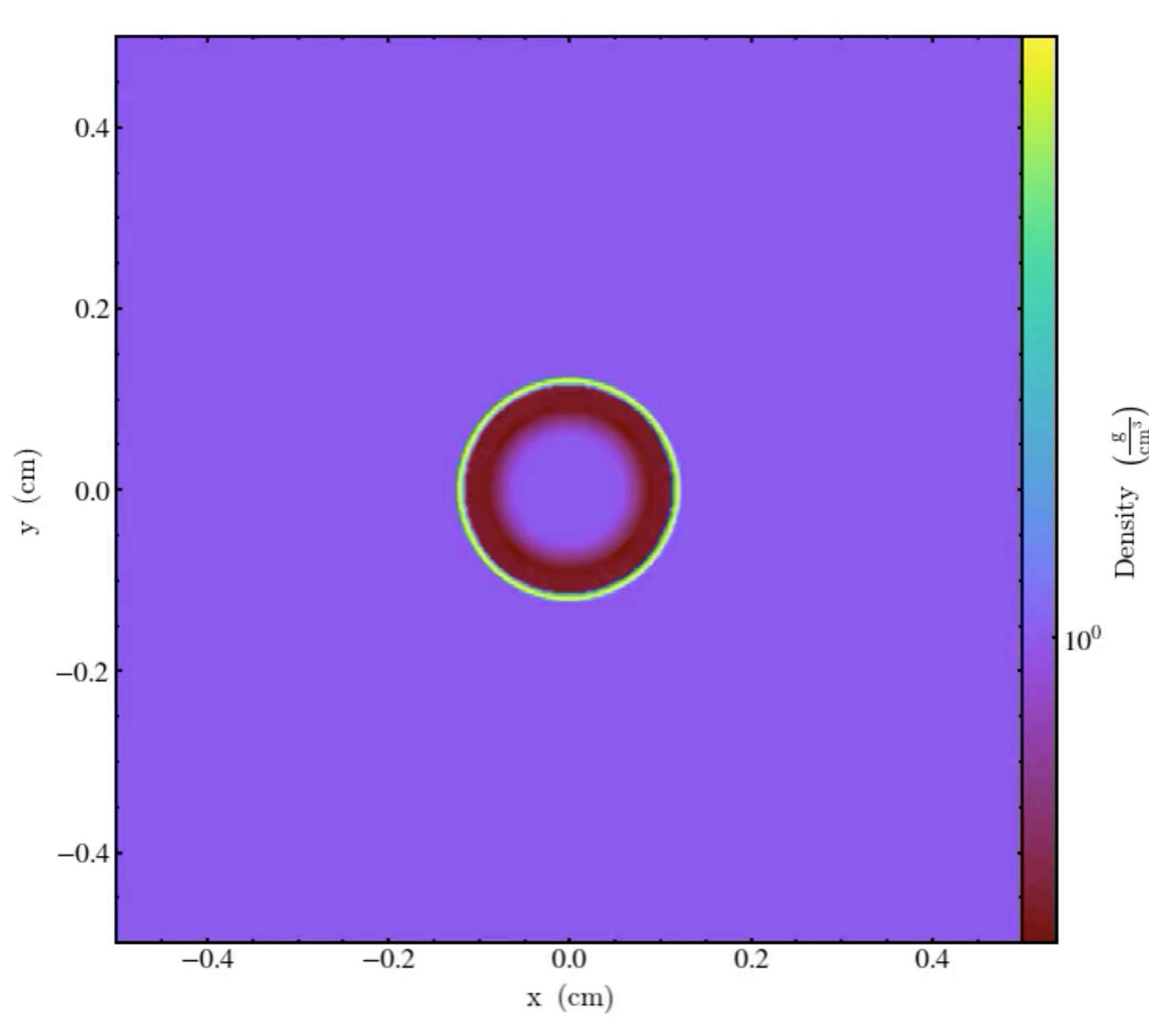


FLASH Code: Infrastructure

- **Driver**: split; unsplit
- **Grid**: uniform Grid; Adaptive Mesh Refinement (AMR)
- **GridParticles**: Lagrangian framework
- **GridSolvers**: multigrid; multipole; Barnes-Hut Tree; PFFT; direct solvers
- **IO**: Hierarchical Data Format 5 (HDF5); PnetCDF
- **Multispecies**
- **Runtime parameters**
- **Monitor**: MPI Timers; Hooks for TAU



Block structure PARAMESH AMR



See the afternoon lectures for data analysis and visualization



FLASH Code: research applications

- Highly modular (`setup.py`)

<code>Driver/</code>	<code>Particles/</code>	<code>diagnostics/</code>	<code>numericalTools/</code>
<code>Grid/</code>	<code>PhysicalConstants/</code>	<code>flashUtilities/</code>	<code>physics/</code>
<code>IO/</code>	<code>RuntimeParameters/</code>	<code>monitors/</code>	
<code>Multispecies/</code>	<code>Simulation/</code>	<code>multiprocessorTools/</code>	

<code>Cosmology/</code>	<code>Hydro/</code>	<code>RayTrace/</code>	<code>sourceTerms/</code>
<code>Diffuse/</code>	<code>ImBound/</code>	<code>SolidMechanics/</code>	<code>utilities/</code>
<code>Eos/</code>	<code>IncompNS/</code>	<code>TreeRay/</code>	
<code>Gravity/</code>	<code>RadTrans/</code>	<code>materialProperties/</code>	

<code>Burn/</code>	<code>EnergyDeposition/</code>	<code>Heatexchange/</code>	<code>PrimordialChemistry/</code>
<code>Cool/</code>	<code>Flame/</code>	<code>Ionize/</code>	<code>Stir/</code>
<code>Deleptonize/</code>	<code>Heat/</code>	<code>Polytrope/</code>	<code>Turb/</code>



FLASH Code: Simulation Domain

- AMR

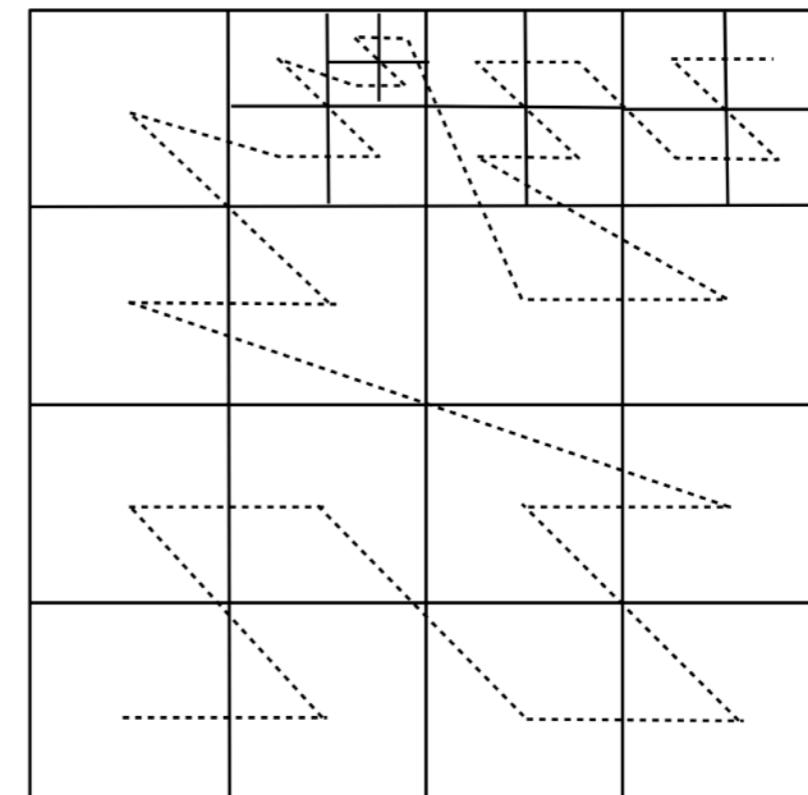
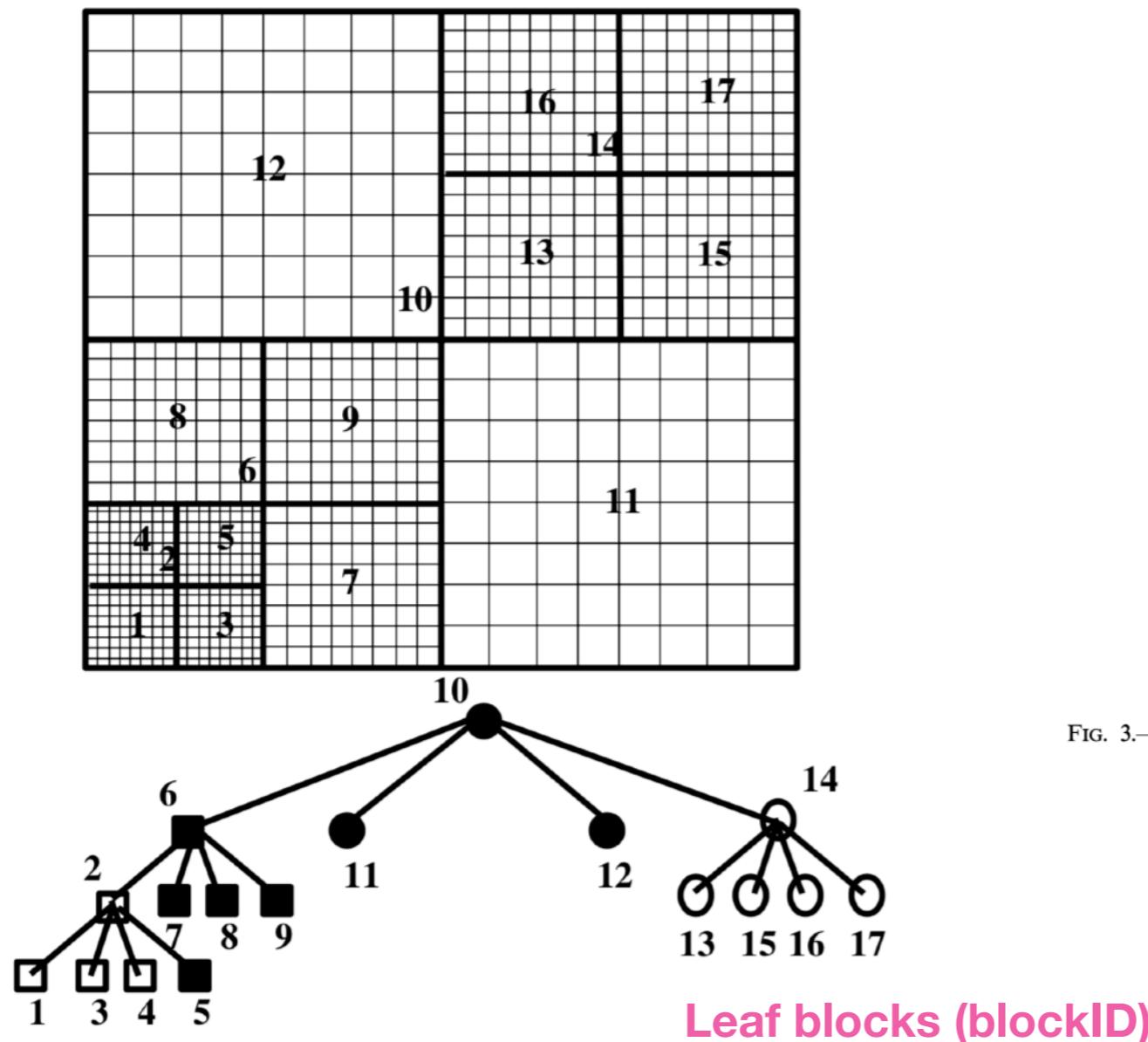
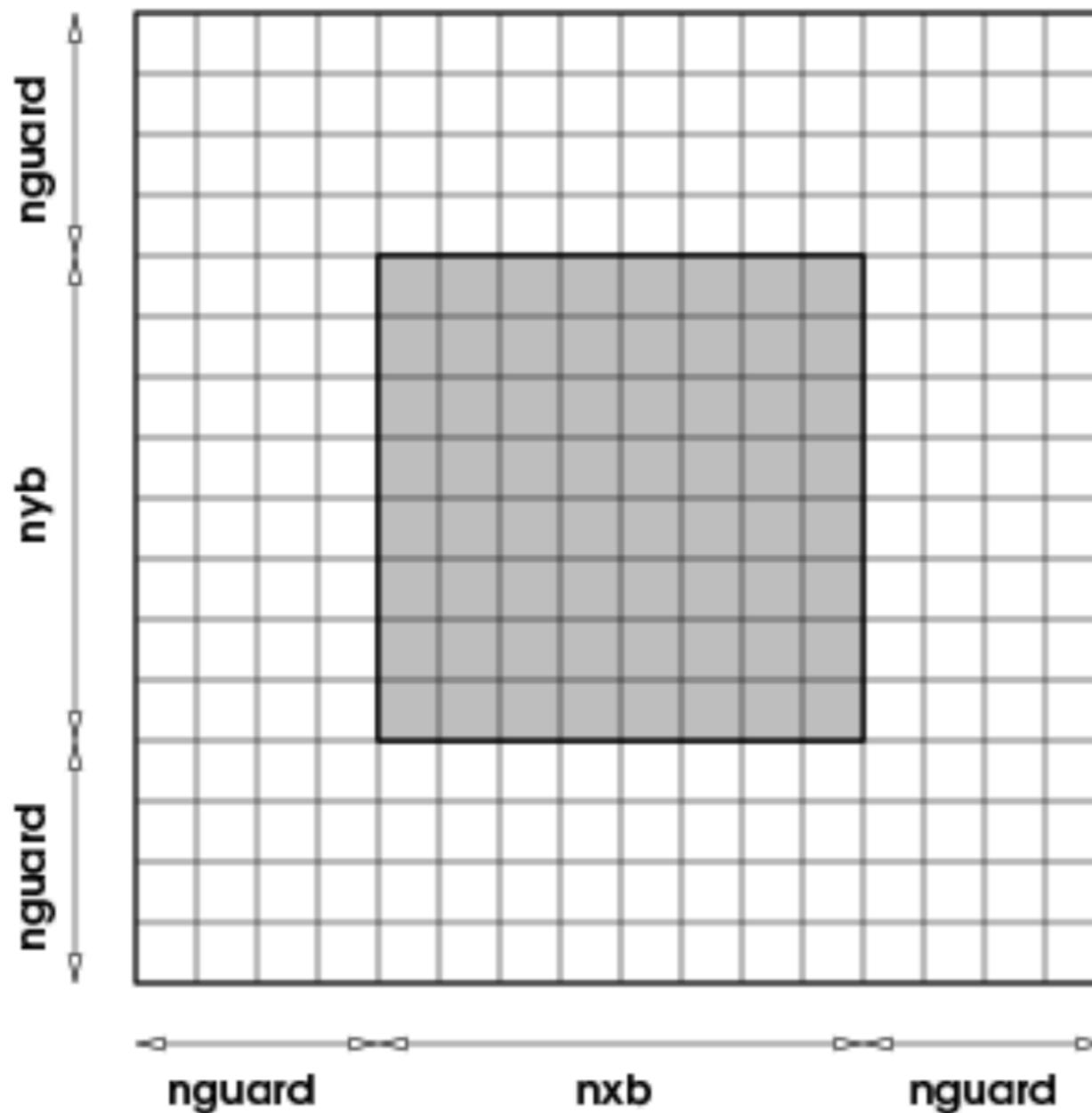


FIG. 3.—Morton space-filling curve for an arbitrary set of blocks of differing spatial resolution



FLASH Code: Simulation Domain

- Blocks



- The grid is composed of blocks
- FLASH4: all blocks are of same size
- May cover different fractions of the physical domain, depending on a block's resolution
- Data storage area for each block reserves space for some layers of guard cells
- Size of guard cells can be changed by putting a line "GUARDCELLS 6" in the Config file. Default is 4.



FLASH Code: Variables

- Density: DENS_VAR
- Velocities: VELX_VAR, VELY_VAR, VELZ_VAR
- Temperature: TEMP_VAR
- Pressure: PRES_VAR
- Specific internal energy: EINT_VAR



Simulations

- All sample simulations can be found in
`FLASH/source/Simulation/SimulationMain/`
- A simulation requires the following 5 files:
 - `Config`
 - `Makefile`:
 - `Simulation_data.F90`
 - `Simulation_init.F90`
 - `Simulation_initBlock.F90`

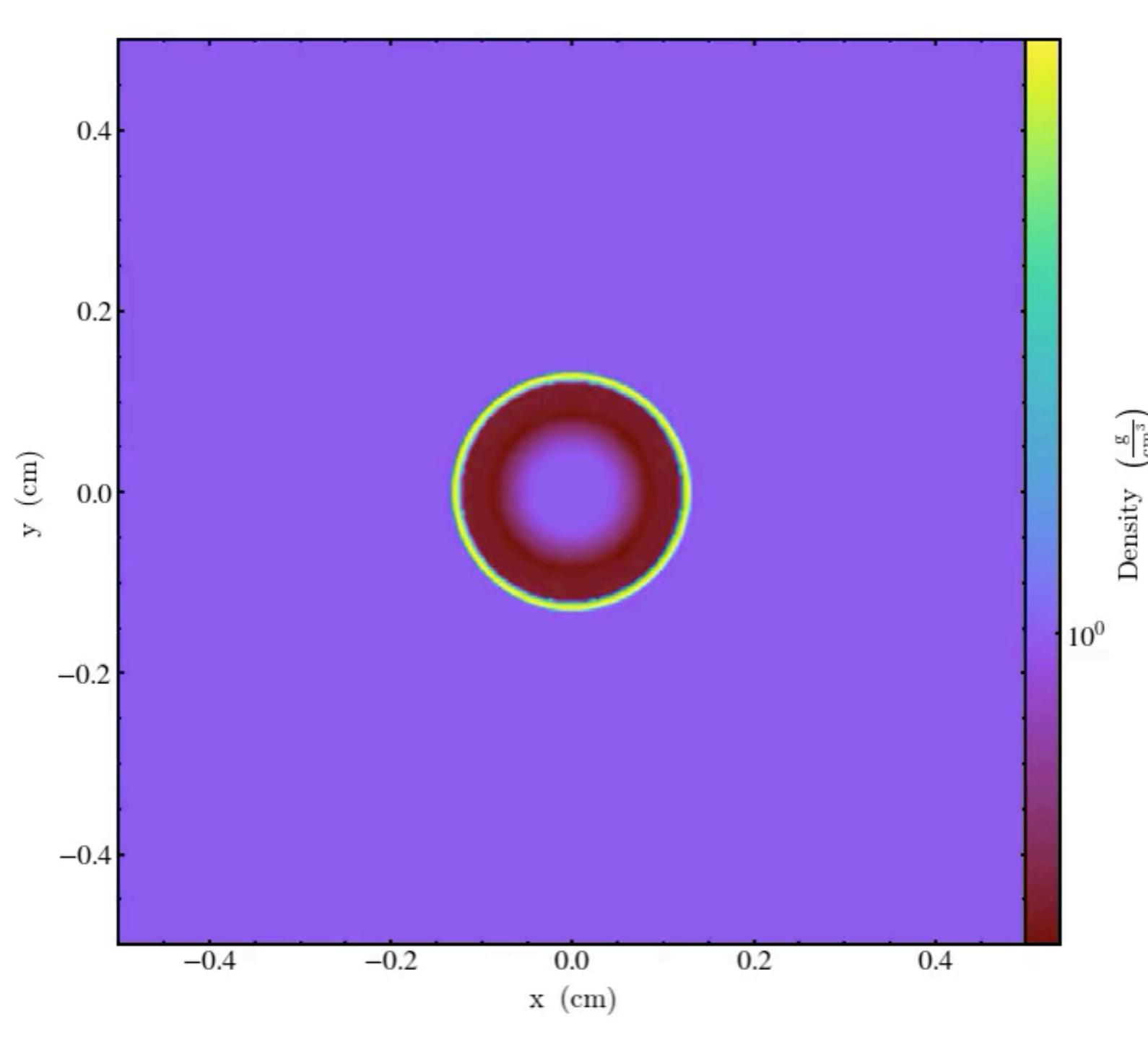


Design a FLASH simulation

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`FLASH/source/Simulation/SimulationMain/`
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 - `Config`
 - `Makefile`:
 - `Simulation_data.F90`
 - `Simulation_init.F90`
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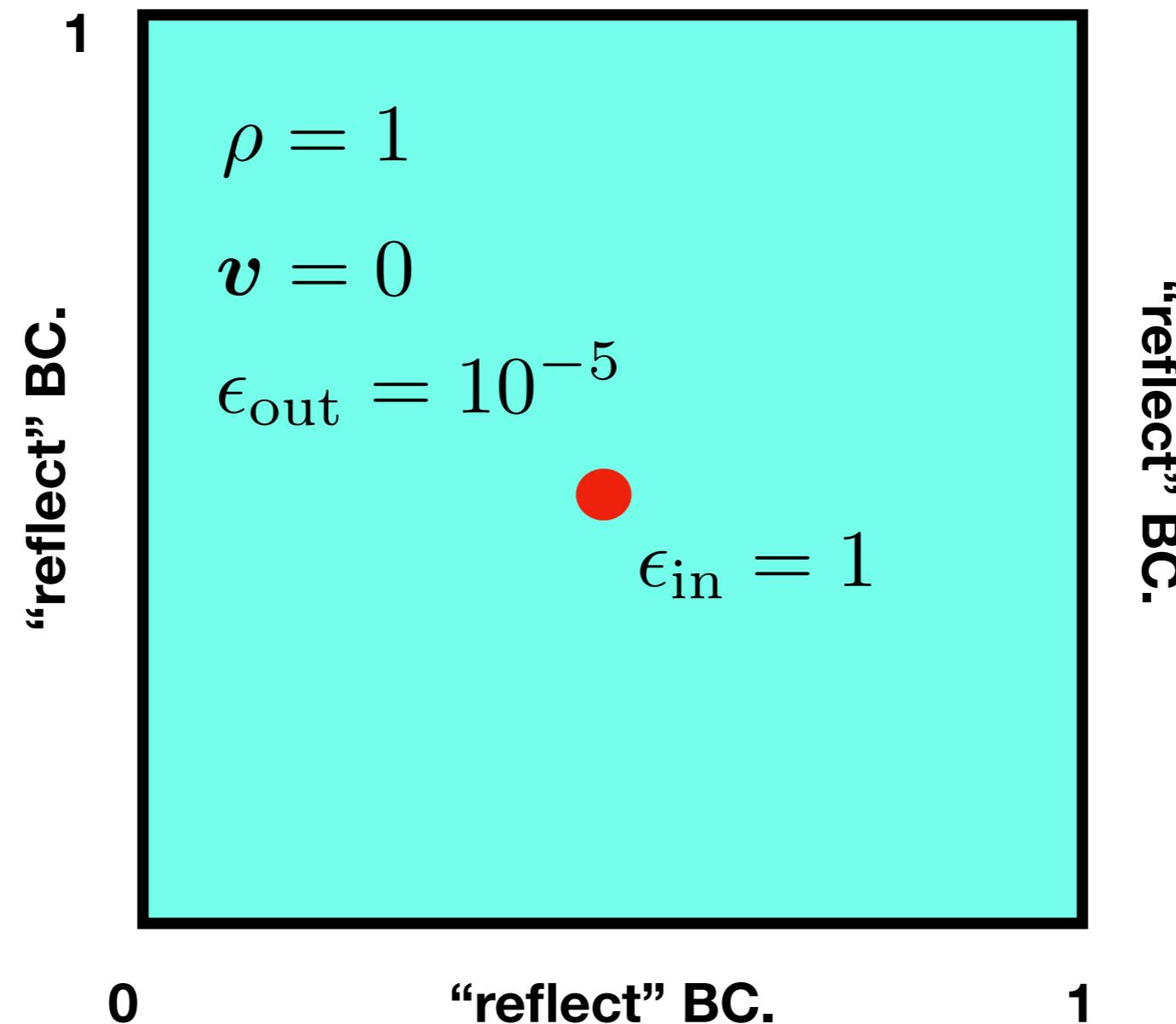
A simple explosion simulation





The initial condition

“reflect” BC.





Config

```
● ● ● 退出 Config + (~codes/FLASH/s...ationMain/Template) - VIM
1 # The Configuration file for a test problem
2
3 REQUIRES Driver
4 REQUIRES physics/Hydro
5 REQUIRES physics/Eos
6
7 D sim_rho0 The ambient density [g/cc]
8 PARAMETER sim_rho0 REAL 1. [0 to ]
9
10
11
```

~

~

~

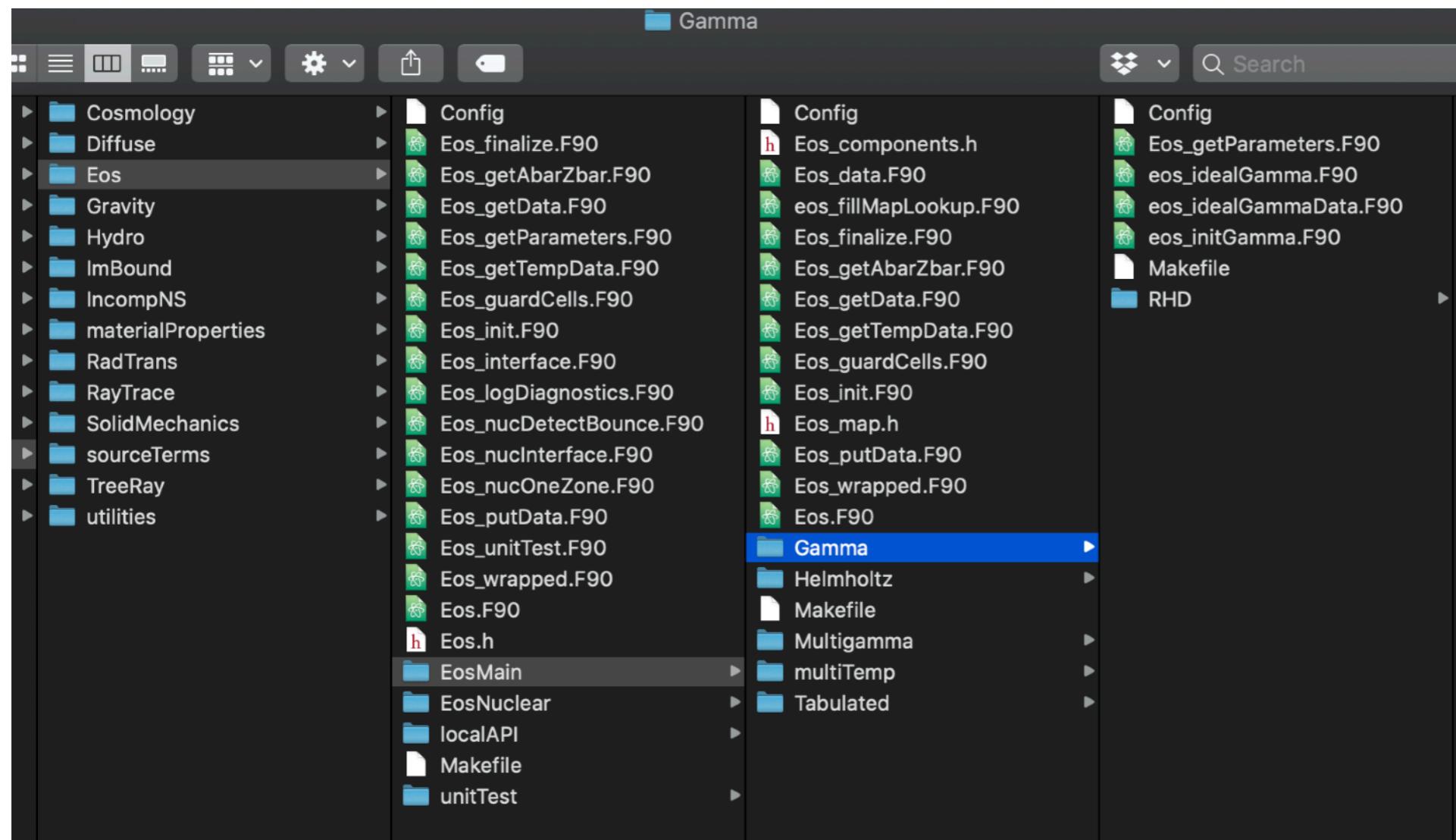
Config [+] 11,0-1 All

pan vega.astr.nthu.edu.tw ~c/F/s/S/S/Template vim - zsh 8/28, 3:16 PM



Using a Flash Unit

- REQUIRES physics/EoS
- Default: physics/Eos/EosMain/Gamma





Makefile

Makefile (~/codes/FLASH/s...ationMain/Template) - VIM

```
1 Simulation += Simulation_data.o
2 Simulation_init.o : Simulation_data.o
3 Simulation_initBlock.o : Simulation_data.o
4
5
```

~
~
~
~
~
~
~
~
~
~
~
~
~

Makefile 5,0-1 All
"Makefile" 5L, 116C written

pan vega.astr.nthu.edu.tw ~c/F/s/S/S/Template vim zsh 8/28, 3:17 PM



Simulation_data.F90

```
25 module Simulation_data
26 #include "Flash.h"
27 implicit none
28
29 !! *** Runtime Parameters ***
30 real, save :: sim_rho0
31
32 !! *** Variables pertaining to Simulation
33 logical, save :: sim_gCell
34
35 integer, save :: sim_meshMe
36 end module Simulation_data
37
```



Simulation_init.F90

```
28 subroutine Simulation_init()
29
30     use Simulation_data
31     use Driver_interface, ONLY : Driver_getMype, Driver_abortFlash
32     use RuntimeParameters_interface, ONLY : RuntimeParameters_get
33     use Logfile_interface, ONLY : Logfile_stamp
34     implicit none
35 #include "constants.h"
36 #include "Flash.h"
37 #include "Multispecies.h"
38
39     call Driver_getMype(MESH_COMM, sim_meshMe)
40
41     call RuntimeParameters_get('sim_rho0', sim_rho0)
42
43     call Logfile_stamp( "initializing the Template problem", &
44                         "[Simulation_init]")
45
46 end subroutine Simulation_init
```



Simulation_init.F90

```
28 subroutine Simulation_init()
29
30     use Simulation_data
31     use Driver_interface, ONLY : Driver_getMype, Driver_abortFlash
32     use RuntimeParameters_interface, ONLY : RuntimeParameters_get
33     use Logfile_interface, ONLY : Logfile_stamp
34     implicit none
35 #include "constants.h"
36 #include "Flash.h"
37 #include "Multispecies.h"
38
39     call Driver_getMype(MESH_COMM, sim_meshMe)
40
41     call RuntimeParameters_get('sim_rho0', sim_rho0)
42
43     call Logfile_stamp( "initializing the Template problem", &
44                         "[Simulation_init]")
45
46 end subroutine Simulation_init
```



Simulation_initBlock.F90

```
30 !***  
31  
32 subroutine Simulation_initBlock(blockID)  
33 !  
34 #include "constants.h"  
35 #include "Flash.h"  
36 #include "Eos.h"  
37  
38 use Simulation_data, ONLY : sim_rho0  
39  
40 use Grid_interface, ONLY : Grid_getBlkIndexLimits, &  
41     Grid_getCellCoords, Grid_getBlkPtr, Grid_releaseBlkPtr  
42 use Eos_interface, ONLY : Eos, Eos_wrapped  
43  
44  
45 implicit none  
46  
47 integer, intent(in) :: blockID  
48  
49  
50 !! Handling AMR blocks and pointer  
51  
52 integer :: i, j, k, n  
53 real, dimension(:,:,:,:),pointer :: solnData  
54 integer, dimension(2,MDIM) :: blkLimits, blkLimitsGC  
55 integer :: sizeX,sizeY,sizeZ  
56 logical :: gcell = .true.  
57  
58 !! Physical coordinates  
59 real :: xx, yy, zz  
60 real :: radius  
61 real,allocatable, dimension(:) ::yCenter, xCoord, zCoord  
62  
63 !! physical parameters  
64 real :: rho, pres, velx, vely  
65 --
```

```
66 ! get the integer index information for the current block  
67 call Grid_getBlkIndexLimits(blockId,blkLimits,blkLimitsGC)  
68 call Grid_getBlkPtr(blockId,solnData)  
69  
70  
71 sizeX = blkLimitsGC(HIGH,IAXIS)  
72 sizeY = blkLimitsGC(HIGH,JAXIS)  
73 sizeZ = blkLimitsGC(HIGH,KAXIS)  
74 allocate(yCenter(sizeY))  
75 allocate(xCoord(sizeX))  
76 allocate(zCoord(sizeZ))  
77 yCenter = 0.0  
78 xCoord = 0.0  
79 zCoord = 0.0  
80  
81 if (NDIM == 3) call Grid_getCellCoords&  
     (KAXIS, blockId, CENTER,gcell, zCoord, sizeZ)  
82 if (NDIM >= 2) call Grid_getCellCoords&  
     (IAXIS, blockId, CENTER,gcell, xCoord, sizeX)  
83  
84 call Grid_getCellCoords(JAXIS, blockId, CENTER, gcell, yCenter, sizeY)  
85  
86  
87
```



Simulation_initBlock.F90

```
95  do k = blkLimits(LOW,KAXIS),blkLimits(HIGH,KAXIS)
96
97    ! get the coordinates of the cell center in the z-direction
98    zz = zCoord(k)
99    do j = blkLimits(LOW,JAXIS),blkLimits(HIGH,JAXIS)
100      ! get the coordinates of the cell center in the y-direction
101      yy = yCenter(j)
102      ! The position in the current yz-row.
103      do i = blkLimits(LOW,IAXIS),blkLimits(HIGH,IAXIS)
104
105        ! get the cell center, left, and right positions in x
106        xx = xCoord(i)
107
108        radius = sqrt(xx**2 + yy**2 + zz**2)
109
110        if (radius .le. 0.1) then
111          pres = (4./3. - 1.)*1.0
112        else
113          pres = 1.e-5
114        endif
115        rho = sim_rho0
116        velx = 0.0
117        vely = 0.0
118
119        solnData(DENS_VAR, i,j,k) = rho
120        solnData(PRES_VAR, i,j,k) = pres
121        solnData(VELX_VAR, i,j,k) = velx
122        solnData(VELY_VAR, i,j,k) = vely
123        solnData(VELZ_VAR, i,j,k) = 0.0
124
125      enddo
126    enddo
127  enddo
128
129  call Grid_releaseBlkPtr(blockID, solnData)
130  call Eos_wrapped(MODE_DENS_PRES,blkLimits,blockId)
131
132  deallocate(yCenter)
133  deallocate(xCoord)
134  deallocate(zCoord)
135
136  return
137 end subroutine Simulation_initBlock
```



Setup the Simulation

'FLASH/source/Simulation/SimulationMain/Template

- A machine depend Makefile.h

```
~/codes/FLASH/sites(sim_school*) » ls
Aliases          buckbeak
FomalhautIMP    cetus.asci.uchicago.edu
FomalhautMP4     clark.asci.uchicago.edu
P640278.nist.gov clogin1
Prototypes       code.uchicago.edu
SEAS10926.gwu.edu coyote.lanl.gov
SEAS10927.gwu.edu crash2.umd.edu
SEAS10927UB.gwu.edu cthinkpad
SEAS12982.gwu.edu ctsv.astro.sunysb.edu
SEAS13033.gwu.edu cube.uchicago.edu
absoft-mpi2      datastar.sdsc.edu
alc.llnl.gov     deepthought.umd.edu_tfitz
amazon.nswccd   duce.gsfc.nasa.gov
animal5          edison.nersc.gov
archimedes.uchicago.edu elan.uchicago.edu
asterix.asci.uchicago.edu eldorado.astro.sunysb.edu
babbage.nersc.gov eldorado.uchicago.edu
bassi.nersc.gov  ellipse.uchicago.edu
beagle.ci.uchicago.edu ellipse02.uchicago.edu
bgl.llnl.gov     ellipse03.uchicago.edu
bgl.mcs.anl.gov  ellipse04.uchicago.edu
bgl.sdsc.edu     ellipse05.uchicago.edu
bonsai.cfa.harvard.edu ellipse06.uchicago.edu
brassica.asci.uchicago.edu ellipse08.uchicago.edu

ellipse09.uchicago.edu
ellipse10.uchicago.edu
ellipse11.uchicago.edu
ellipse_pgf.uchicago.edu
eugenia.asci.uchicago.edu
eureka.alcf.anl.gov
fen.bluegene.bn1.gov
fenp.bluegene.bn1.gov
flash.uchicago.edu
flashviz.uchicago.edu
fleetwood.astro.sunysb.edu
fomalhaut
fornax.uchicago.edu
franklin.nersc.gov
fusion.lcrc.anl.gov
fusion.lcrc.uchicago.edu
gin.asci.uchicago.edu
gnu-ins
gnu-mpi2
gnu-ompi
hawkmoon.uchicago.edu
hera.llnl.gov
hopper.nersc.gov
hpc.msu.edu

hyades.ucsc.edu
hydra.si.edu
icc-9.0_fornax.uchicago.edu
ignition
intel-hedp
intel-mpi2
intel-ompi
intrepid.alcf.anl.gov
jacquard.nersc.gov
jaguar.ccs.ornl.gov
jaguar.nccs.gov
jub1.zam.kfa-juelich.de
karloff.lbl.gov
khorba.uchicago.edu
klaus-laptop
kraken.nics.tennessee.edu
lahey-mpi2
lenovolaptop
lituchi.uchicago.edu
login1.pads.ci.uchicago.edu
login2.pads.ci.uchicago.edu
lupin.uchicago.edu
macbro.uchicago.edu
maitreyi

mhd2.ascii.uchicago.edu
midway2.rcc.uchicago.edu
mira.alcf.anl.gov
mongchi.soe.ucsc.edu
mongchi.uchicago.edu
myristica.ascii.uchicago.edu
nag-mpi1
nag-mpi2
nagini.uchicago.edu
oakley.osc.edu
optix.cs.uoregon.edu
osel.uchicago.edu
p655-4.nic.uoregon.edu
pg-mpi2-32
purple.llnl.gov
pyramid.uchicago.edu
qsc.lanl.gov
r1.oit.ua.edu
ramsusii.mps.ohio-state.edu
ranger.tacc.utexas.edu
rc2.ua.edu
redstorm.sandia.gov
saguaro.fulton.asu.edu
scarf.rl.ac.uk

scooter.ascii.uchicago.edu
seaborg.nerc.gov
skeeter.ascii.uchicago.edu
sphere.uchicago.edu
splash.seas.gwu.edu
splash.seas.gwu.edu_KPD
sunspot.uchicago.edu
surveyor.alcf.anl.gov
tp-login2
tsoodzil.astro.uiuc.edu
tuxedo.uchicago.edu
uffda.ascii.uchicago.edu
variable.as.arizona.edu
variable.ph.ua.edu
vega.astr.nthu.edu.tw
vestalac1
watanlsn.watson.ibm.com
zeus.llnl.gov
zingiber.ascii.uchicago.edu
zingiber.uchicago.edu
```



Setup the Simulation

- Run a setup script

A simple example

```
./setup Template -2d -auto
```

```
./setup Template -2d -auto -maxblocks=2000 -nxb=8 -nyb=8
```

A more complicated example

```
~/codes/FLASH(sim_school*) » ./setup CoreCollapse/IDSA -auto -2d +cylindrical -objdir ccsn2dIdsa -nxb=16 -nyb=16 -max  
blocks=1400 threadBlockList=False +pm4dev threadWithinBlock=False +newMpole --with-unit=source/physics/RadTrans/RadTr  
ansMain/IDSA/rk2
```



Setup the Simulation

```
~/codes/FLASH(sim_school*) » ./setup
Processing Shortcut file: /Users/pan/codes/FLASH/bin/setup_shortcuts.txt
usage: setup <problem-name> [options] [VAR=VALUE]...
```

problem-name: see source/Simulation/SimulationMain directory
options:

(Science Options)

```
-auto -[123]d
-maxblocks=<#> -nxb=<#> -nyb=<#> -nzb=<#>
-with-unit=<unit> -with-library=<libname>[,args]
-without-unit=<unit> -without-library=<libname>
```

(Setup and Make Options)

```
-verbose=[DEBUG|INFO|WARN|IMPINFO|ERROR]
[-site=<site> | -ostype=<ostype>]
-makefile=<extension>
[-opt| -debug | -test ]
-objdir=<relative obj directory>
-defines=<defines> -unitsfile=<filename>
-datafiles=<wildcard> -parfile=<filename>
-fbs -nofbs -tau=<makefile>
```

(Misc Options)

```
-makehide -noclobber -portable -help
```

- * For GNU compatibility, options may be prefixed by -- instead of - as well
- * -unit and -library are considered equivalent to
-with-unit and -with-library respectively.
- * For information regarding the [VAR=VALUE] options and using 'setup variables'
refer to User's Guide.
- * To read how shortcuts work see README.shortcuts in your bin directory



Setup shortcuts

To use a shortcut add '+shortcut' to your setup line.

For example ./setup Sod -auto +ug

```
3t           ThreeT=1 --defines=FLASH_3T
3tr          +3t RadFlahThreeT=True
8wave         --with-unit=physics/Hydro/HydroMain/split/MHD_8Wave +grid --gridinterpolation=native
asynclaser    +laser useAsyncLaser=True --without-unit=physics/sourceTerms/EnergyDeposition/EnergyDeposition
Main/Laser/LaserIO
cartesian     --geometry=cartesian
chombo_amr   --unit=Grid/GridMain/Chombo/AMR --index-reorder Grid=Chombo --nofbs --makefile=chombo chomboCo
mpatibleHydro=True
chombo_ug    --unit=Grid/GridMain/Chombo/UG --index-reorder Grid=Chombo --maxblocks=1 --nofbs --makefile=ch
ombo chomboCompatibleHydro=True
cube16        --nxb=16 --nyb=16 --nzb=16
cube32        --nxb=32 --nyb=32 --nzb=32
cube64        --nxb=64 --nyb=64 --nzb=64
curv-pm2      +pm2 --unit=Grid/GridMain/paramesh/Paramesh2 --with-unit=Grid/GridMain/paramesh/Paramesh2/mono
tonic
curvilinear   --curvilinear
cylindrical   --geometry=cylindrical
ddt           use_ddt=True --defines=DDT --with-unit=source/monitors/Debugger/DebuggerMain/dmalloc
default        --with-library=mpi +io +grid --gridinterpolation=monotonic FlashAvoid0rrorry=True
dmalloc        use_dmalloc=True --defines=DMALLOC --with-unit=source/monitors/Debugger/DebuggerMain/dmalloc
dynamicmem-pm3
dynamicmem-pm40
gravmgdrid
gravmpole
gravpfftnofbs
grid
hdf5
hdf5typeio
io
laser         --unit=physics/sourceTerms/EnergyDeposition/EnergyDepositionMain/Laser --without-unit=Particle
```

s



Setup shortcuts (continues)

```
newmpole +noMgrid +noDefaultMpole +gravMpole --with-unit=Grid/GridSolvers/Multipole_new
nodefaultmpole --without-unit=Grid/GridSolvers/Multipole
nofbs --nofbs +ug parallelIO=True
noio --without-unit=physics/sourceTerms/EnergyDeposition/EnergyDepositionMain/Laser/LaserIO --witho
ut-unit=IO
nolog --without-unit=monitors/LogFile
nomgrid --without-unit=physics/Gravity/GravityMain/Poisson/Multigrid
npg npg=True
parallelio parallelIO=True
pic +ug --unit=Grid/GridParticles/GridParticlesMove --without-unit=Grid/GridParticles/GridParticle
sMove/UG --without-unit=Grid/GridParticles/GridParticlesMove/UG/Directional
pipeline --unit=multiprocessorTools/Pipeline
pm2 +grid Grid=PM2
pm3 +pm40
pm40 +grid Grid=PM40
pm4dev pm4dev_basic FlashAvoid0rrery=True
pm4dev_clean +pm4dev_basic FlashAvoid0rrery=False
pm4dev_fixed +pm4dev ParameshLibraryMode=False
pnetcdf I0=pnetcdf
pnettypeio +io +parallelIO +pnetcdf typeI0=True
polar --geometry=polar
protonemission --unit=diagnostics/ProtonEmission --without-unit=Particles
protonimaging --unit=diagnostics/ProtonImaging --without-unit=Particles
ptdens --without-unit=Particles/ParticlesInitialization/Lattice --without-unit=Particles/ParticlesIni
tialization/WithDensity/CellMassBins --unit=Particles/ParticlesMain --unit=Particles/ParticlesInitialization/WithDensity --pa
rticlemethods=TYPE=passive,INIT=With_Density
ptio +ug --with-unit=Particles
purehydro physicsMode=hydro
rnf --3d --nxb=8 --nyb=16 --nzb=32 --nofbs +ug
serialio parallelIO=False
spherical --geometry=spherical
spherical-pm2 +pm2 +spherical
splithydro --with-unit=physics/Hydro/HydroMain/split --without-unit=physics/Hydro/HydroMain/unsplit Split
Driver=True
supportppmupwind SupportPpmUpwind=True
```

Compile source codes

- An object/ folder will be generated

cd object

make

```
mpi_unpack_edges.o mpi_unpack_fluxes.o mpi_unpack_tree_info.o mpi_wrapper_dble.o mpi_wrapper_int
mpi_wrapper_real.o nameSyntaxError.o nameValueLL_bcast.o nameValueLL_data.o nameValueLL_get.o na
alueLL_getNum.o nameValueLL_logRules.o nameValueLL_rules.o nameValueLL_set.o paramesh_comm_data.o
aramesh_interfaces.o paramesh_mpi_interfaces.o pc_interface.o pc_utilities.o physicaldata.o poiss
ist.o prolong_arrays.o rationalize_fetch_list.o removeNullChar.o rp_getArgument.o rp_getOpt.o rp_
s.o rp_storeIgnoredParams.o send_block_data.o setup_buildstamp.o setup_buildstats.o setup_flashRe
its.o timings.o tmr_buildSummary.o tmr\_create.o tmr_etime.o tmr_findTimerIndex.o tmr_getMaxCallSt
erParents.o tmr_init.o tmr_lookupIndex.o tmr_stackLib.o tree.o tree_search_for_surrblk.o umap.o
constRealZeroFn1.o ut_conversionInterface.o ut_convertToArrayIndices.o ut_convertToMemoryOffset.
Interface.o ut_getFreeFileUnit.o ut_hunt.o ut_interpolationInterface.o ut_parabolicInterpol.o ut_
o ut_qsort.o ut_qsortC.o ut_qsortInterface.o ut_quadraticInterpol.o ut_sortInterface.o ut_sortOnF
t_sysMemInterface.o ut_sysMemStats.o ut_sysMemSummaryStats.o ut_sys_mem_usage.o workspace.o -L/U
1-gnu/lib -lhdf5 -lhdf5_fortran -lz -DH5_USE_16_API
```

SUCCESS



Run Flash

- Requires
 - the executable file `flash4` and
 - the runtime parameter file `flash.par`



Runtime Parameters (flash.par)

```
1 # Runtime parameters for the Kelvin-Helmholtz instability simulation
2
3 # Parameters for initial model
4
5 sim_rho0 = 1.0
6
7
8 # Grid dimensionality and geometry
9
10 geometry = cartesian
11
12 # Size of computational volume
13
14 xmin      = -0.5
15 xmax      = 0.5
16 ymin      = -0.5
17 ymax      = 0.5
18
```



Runtime Parameters

```
19 # Boundary conditions
20
21 xl_boundary_type = "reflect"
22 xr_boundary_type = "reflect"
23
24 yl_boundary_type = "reflect"
25 yr_boundary_type = "reflect"
26
27
28 # Simulation (grid, time, I/O) parameters
29
30 cfl = 0.8
31 basenm = "my_sim_"
32 restart = .false.
33
```



Runtime Parameters

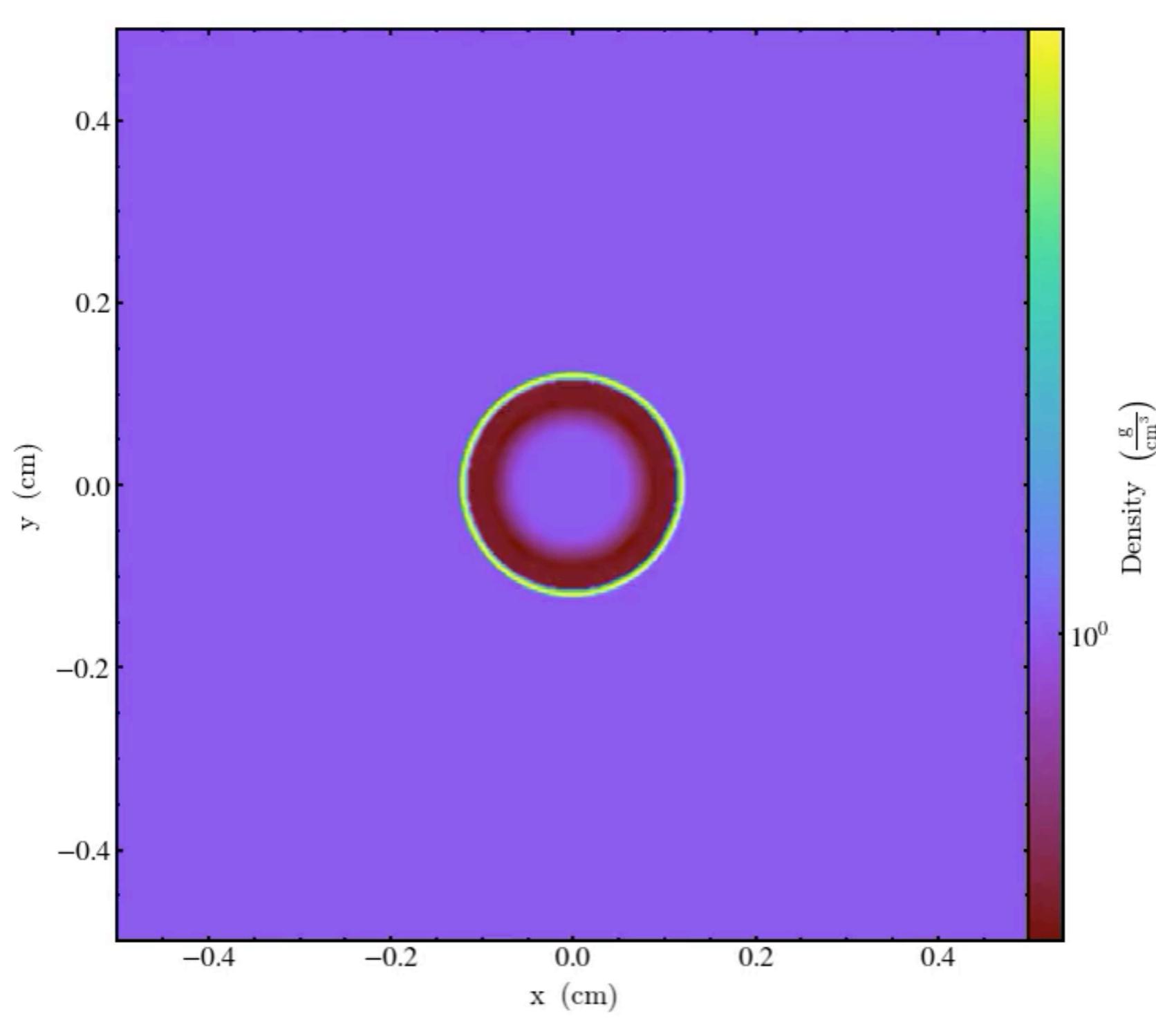
```
34 # checkpoint file output parameters
35 checkpointFileIntervalTime = 0.5
36 checkpointFileIntervalStep = 0
37 checkpointFileNameNumber = 0
38
39 # plotfile output parameters
40 plotfileIntervalTime = 0.1
41 plotfileIntervalStep = 0
42 plotfileNameNumber = 0
43
44 nend = 1000000
45 tmax = 10.0
46
47 run_comment = "my first flash simulation"
48 log_file = "my_sim.log"
49 eintSwitch = 1.e-4
50
51
52 plot_var_1 = "dens"
53 plot_var_2 = "pres"
54 plot_var_3 = "temp"
55 plot_var_4 = "velx"
56 plot_var_5 = "vely"
--
```

```
59 # AMR refinement parameters
60 lrefine_min = 1
61 lrefine_max = 6
62 refine_var_1 = "dens"
63 refine_var_2 = "pres"
64
```



Outputs

- my_sim.log
- my_sim.dat
- my_sim_hdf5_chk_*
- my_sim_hdf5_plt_cnt_*



See the afternoon lectures for data analysis and visualization



Runtime parameters documentation

- <http://flash.uchicago.edu/site/flashcode/>
[user support/rpDoc 4p61.py](#)

RUNTIME PARAMETERS DOCUMENTATION FOR FLASH RELEASE 4.6.1

All Chapters

Burn
Conductivity
Cool
Cosmology
Debugger
Deleptonize
Diffuse
Driver
EnergyDeposition
Eos
Flame
Gravity
Grid
Heat
Heatexchange
Hydro
IO
IncompNS
Ionize
LogFile
MagneticResistivity
MassDiffusivity
Opacity
Particles
PhysicalConstants
PlasmaState
Polytropes
PrimordialChemistry
Profiler
ProtonEmission
ProtonImaging
RadTrans
RayTrace
RungeKutta
Simulation
Stir
ThomsonScattering
Timers
TreeRay
Turb
Viscosity
XrayImaging



FLASH API

- <http://flash.uchicago.edu/site/flashcode/>

user support/robodoc-FLASH4 4p61/

FLASH4.6.1 API

ROBODoc 4.99 manual

Generated from /asc/asci2/site/flashcode/secure/release_4p6/ with ROBODoc v4.99.8 on Mon Sep 02 01:17:10 2019

[Sourcefiles] [Functions] [Modules]

FUNCTIONS

```
diagnostics/ProtonEmission/ProtonEmission
diagnostics/ProtonEmission/ProtonEmission_finalize
diagnostics/ProtonEmission/ProtonEmission_init
diagnostics/ProtonImaging/ProtonImaging
diagnostics/ProtonImaging/ProtonImaging_finalize
diagnostics/ProtonImaging/ProtonImaging_init
diagnostics/ThomsonScattering/ThomsonScattering
diagnostics/ThomsonScattering/ThomsonScattering_finalize
diagnostics/ThomsonScattering/ThomsonScattering_init
diagnostics/XrayImaging/XrayImaging
diagnostics/XrayImaging/XrayImaging_finalize
diagnostics/XrayImaging/XrayImaging_init
Driver/Driver_abortFlash
Driver/Driver_abortFlashC
Driver/Driver_checkMPIErrorCode
Driver/Driver_computeCelllocations
Driver/Driver_computeDt
Driver/Driver_diagnostics
Driver/Driver_driftBlock
Driver/Driver_driftSetSrcLoc
Driver/Driver_driftUnk
Driver/Driver_evolvewFlash
Driver/Driver_finalizeFlash
Driver/Driver_finalizeSourceTerms
Driver/Driver_getComm
Driver/Driver_getDt
Driver/Driver_getElapsedWCTime
Driver/Driver_getMype
Driver/Driver_getNstep
Driver/Driver_getNumProcs
Driver/Driver_getSimTime
Driver/Driver_getTimeStamp
Driver/Driver_init
Driver/Driver_initFlash
Driver/Driver_initMaterialProperties
Driver/Driver_initNumericalTools
Driver/Driver_initParallel
Driver/Driver_initSourceTerms
Driver/Driver_logMemoryUsage
Driver/Driver_mpiThreadSupport
Driver/Driver_putTimeStamp
Driver/Driver_sendOutputData
Driver/Driver_setupParallelEnv
Driver/Driver_sourceTerms
Driver/Driver_superTimeStep
Driver/Driver_verifyInitDt
flashUtilities/Pipeline/Pipeline_finalize
flashUtilities/Pipeline/Pipeline_globalCheckStatus
flashUtilities/Pipeline/Pipeline_globalCheckStructure
flashUtilities/Pipeline/Pipeline_init
flashUtilities/Pipeline/Pipeline_localActivate
flashUtilities/Pipeline/Pipeline_localCreate
flashUtilities/Pipeline/Pipeline_localDeactivate
Particles/Particles_finalize
Particles/Particles_getCountPerBlk
Particles/Particles_getGlobalNum
Particles/Particles_getLocalNum
Particles/Particles_init
Particles/Particles_initData
Particles/Particles_initForces
Particles/Particles_initPositions
Particles/Particles_longRangeForce
Particles/Particles_manageLost
Particles/Particles_mapFromMesh
Particles/Particles_putLocalNum
Particles/Particles_sendOutputData
Particles/Particles_shortRangeForce
Particles/Particles_sinkAccelGasOnSinksAndSinksOnGas
Particles/Particles_sinkAdvanceParticles
Particles/Particles_sinkComputeDt
Particles/Particles_sinkCreateAccrete
Particles/Particles_sinkInit
Particles/Particles_sinkMarkRefineDerefine
Particles/Particles_sinkMoveParticles
Particles/Particles_sinkSortParticles
Particles/Particles_sinkSumAttributes
Particles/Particles_sinkSyncWithParticles
Particles/Particles_specifyMethods
Particles/Particles_unitTest
Particles/Particles_updateAttributes
Particles/Particles_updateGridVar
Particles/Particles_updateRefinement
PhysicalConstants/PhysicalConstants_get
PhysicalConstants/PhysicalConstants_init
PhysicalConstants/PhysicalConstants_list
PhysicalConstants/PhysicalConstants_listUnits
PhysicalConstants/PhysicalConstants_unitTest
physics/Cosmology/Cosmology_cdmPowerSpectrum
physics/Cosmology/Cosmology_computeDeltaCrit
physics/Cosmology/Cosmology_computeDt
physics/Cosmology/Cosmology_computeVariance
physics/Cosmology/Cosmology_finalize
physics/Cosmology/Cosmology_getOldRedshift
physics/Cosmology/Cosmology_getParams
physics/Cosmology/Cosmology_getRedshift
physics/Cosmology/Cosmology_init
physics/Cosmology/Cosmology_massoLength
physics/Cosmology/Cosmology_redshiftHydro
physics/Cosmology/Cosmology_redshiftToTime
physics/Cosmology/Cosmology_sendOutputData
physics/Cosmology/Cosmology_solveFriedmannEqn
physics/Cosmology/Cosmology_unitTest
physics/Diffuse/Diffuse
physics/Diffuse/Diffuse_advanceID
physics/Diffuse/Diffuse_computeDt
physics/Diffuse/Diffuse_computeFluxLimiter
```



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Happy coding!