

15.2

$$\begin{aligned} M_A &= (\text{mass in amu}) \times \left(\frac{m_{\text{amu}} \text{ kg}}{1 \text{ a.m.u.}} \right) \\ &= (1 \text{ a.m.u.}) \left(\frac{m_{\text{amu}} \text{ kg}}{1 \text{ a.m.u.}} \right) \\ &= 1 \times 1.66 \times 10^{-27} \text{ kg.} \end{aligned}$$

$$M_{\text{Jeans}} = \dots$$

$$F = m a$$

$$\frac{GM(r)m}{r^2} = m a$$

$$r = r_0 + v_0 t + \frac{1}{2} a t^2$$

$$r = \frac{1}{2} a t^2$$

$$t = \sqrt{\frac{2r}{a}}$$

$$= \sqrt{\frac{2r \cdot r^2}{GM(r)}}$$

$$= \sqrt{\frac{2r^3}{GM(r)}}$$

$$, M(r) = \frac{4\pi r^3}{3} \rho$$

15.5 a)

$$\frac{dM}{dt} = \rho v \pi R^2 \underline{F_g}$$

velocity dispersion

$$= \rho v \pi R^2 \left(1 + \left(\frac{v_e}{v_\infty} \right)^2 \right)$$

$$= \rho v \pi R^2 \left(1 + \left(\frac{\left(\frac{GM}{R} \right)^{1/2}}{\left(\frac{2g_m}{C_D \rho_g A} \right)^{1/2}} \right) \right)$$

$\frac{3M}{4\pi R^3}$

Weight

effective cross-sectional area of body

local density of atmosphere

drag coefficient ≈ 1 for sphere

$$\frac{dR}{dt} = \left(\frac{3}{\pi} \right)^{1/2} \frac{\sigma_p n}{\rho} F_g$$

surface density: 100 kg m^{-2}

$$\left(\frac{GM_*}{a^3} \right)^{1/2} = v$$

Mass of star $3M_\odot$ in this question

mean motion

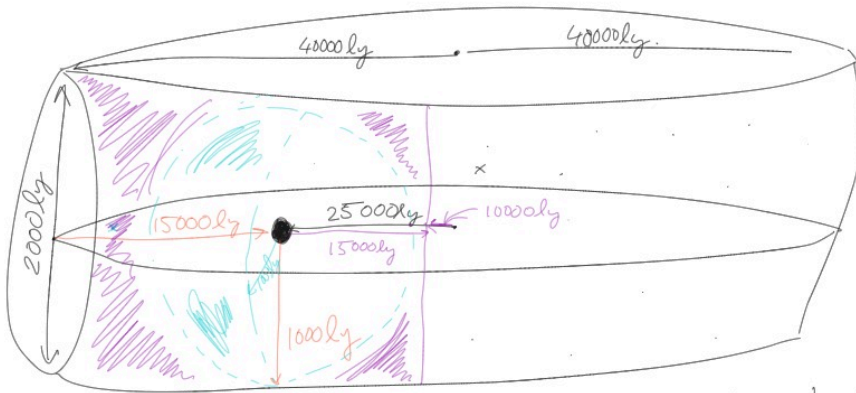
= distance

= 2 AU in this question

b) Isolation Mass:

$$M_{iso} \approx 5 \times 10^{20} (r_{AU}^2 \sigma_p)^{3/2} \left(\frac{M_\odot}{M_*} \right)^{1/2}$$

16)



$$\rho_{cc} = \frac{N_{cc}}{V}$$

$$N_{cc} = \rho_{cc} V$$

If N_{cc} is only a few (i.e. doesn't fit in Purple or blue regions)
then there's 1 cc / 30000 ly. \Rightarrow Only 2 civilizations?

$$\begin{aligned} d(N_{cc}) &= N_{cc} 4\pi r^2 \Rightarrow \\ &= \frac{4}{3} N_{cc}^3 = ? \end{aligned}$$