NTU DIP 2020 Spring HW1 Report

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Problem 0: WARM-UP

(a)

• Read the raw file into a 1D numpy array, then reshape it to the given image size. Write the reshaped array into the jpg format file. The result image result1.jpg is same as the given image in spec.



result1.jpg (sample1.raw)

(b)

• Read the color image in 3-channel, with each represents R, G, B. The convert function is refer to this <u>link (https://bit.ly/33EZyNZ)</u>. The result image result2.jpg is shown as below.





sample2.jpg

result2.jpg

(c)

• By changing the coordinate of every pixel, we can generate the 90 degrees counterclockwise image <code>result3.jpg</code> and diagonally mirrored image <code>result4.jpg</code>. The transfer function from <code>result2.jpg</code> to <code>result3.jpg</code> is I3(i,j) = I2(j,w-1-i) (w represents the image width), and the transfer function from <code>result2.jpg</code> to <code>result4.jpg</code> is I4(i,j) = I2(j,i). The <code>result images result3.jpg</code> and <code>result4.jpg</code> are shown as below.







result4.jpg

Problem 1: IMAGE ENHANCEMENT

In this problem, the function cv2.calcHist() is used for plotting histogram.

(a)

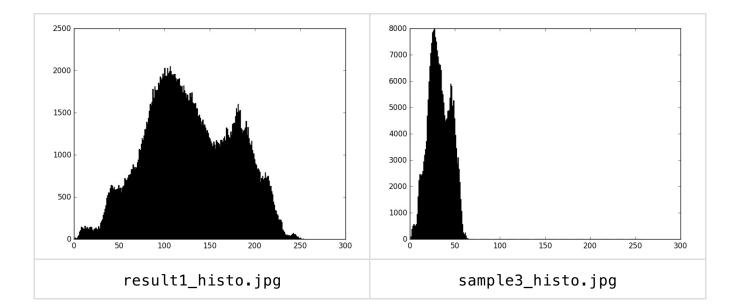
• The histograms of result1.jpg and sample3.jpg are shown as below. Oberve that the max gray-value in result1.jpg is 255, while the max gray-value in sample3.jpg is 63. Thus, we can use the transfer function S3'(i,j) = S3(i,j) to make sample3.jpg look like result1.jpg . (S3 represents the original sample3.jpg , while S3' represents the result image.)





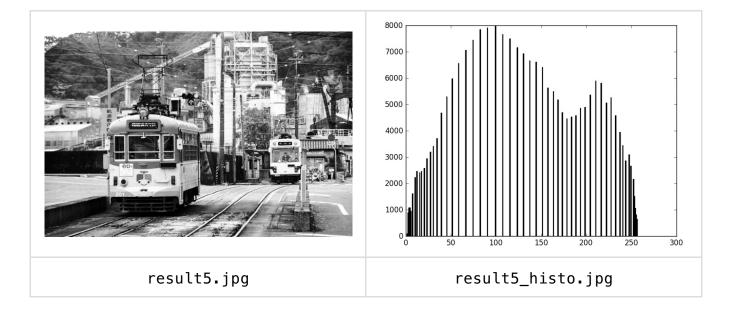
result1.jpg

sample3.jpg



(b)

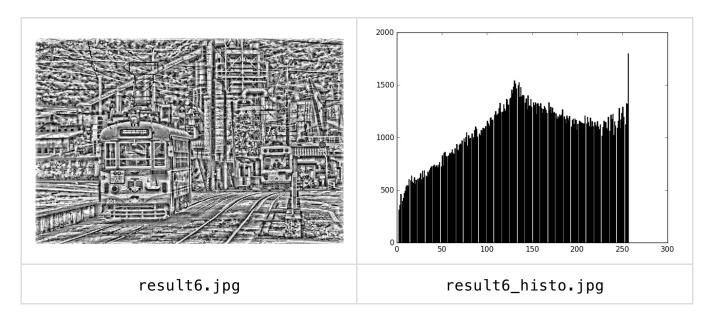
Apply global histogram equilization to sample3.jpg.
The result image result5.jpg and its histogram result5_histo.jpg is shown as below.



(c)

• Apply local histogram equilization to sample3.jpg with window size 15×15 .

The result image result6.jpg and its histogram result6_histo.jpg is shown as below.



(d)

• Comparing two images above, we can found that the global histogram equalized image enhances the contrast and the average brightness of the original dark image, while local histogram equalized image emphasizes more small details in every small region.

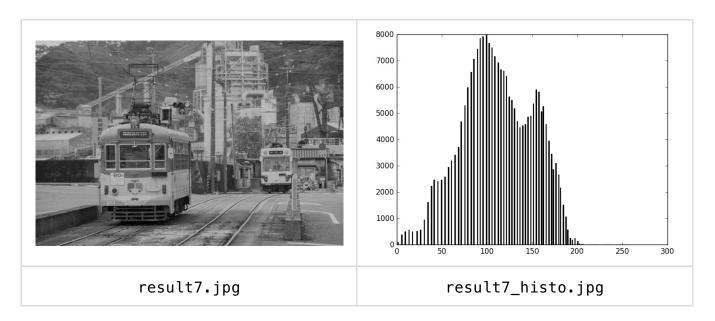
(e)

Log Transform:

Apply log transform to sample3.jpg with transfer function

$$R7(i,j) = rac{ln(1+aS3(i,j))}{ln2} \ (a=3).$$

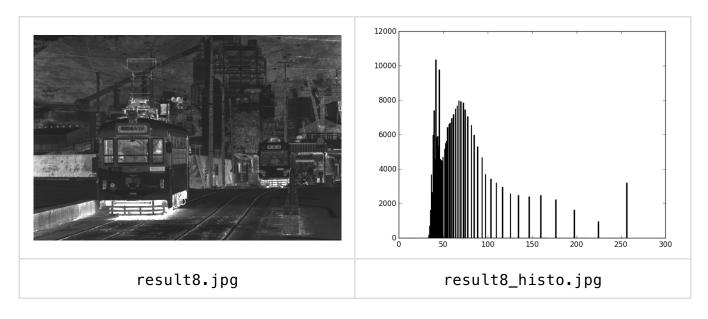
The result image result7.jpg and its histogram result7_histo.jpg are shown as below.



• Inverse Log Transform:

Apply inverse log transform to <code>sample3.jpg</code> with transfer function $R8(i,j)=rac{0.1}{R7(i,j)}~(0.1\leq R7(i,j)<1) \lor R8(i,j)=1~(R7(i,j)<0.1)$

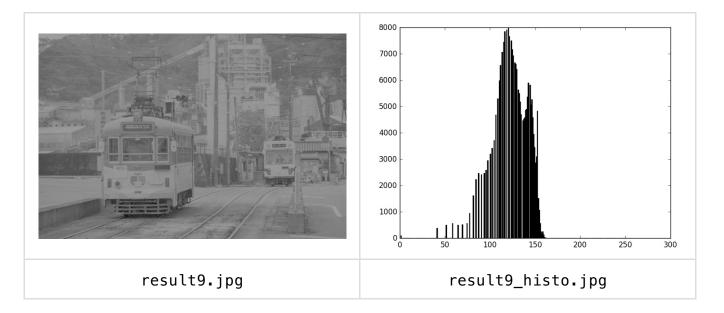
The result image result8.jpg and its histogram result8_histo.jpg are shown as below.



• Power-Law Transform:

Apply power-law transform to sample3.jpg with transfer function $R9(i,j)=[S3(i,j)]^p\ (p=1/3).$

The result image result9.jpg and its histogram result9_histo.jpg are shown as below.



Summary :

- Among these three enhanced images, result7.jpg seems to be the most similar to result1.jpg. The distribution of its histogram is also similar to result1.jpg.
- result8.jpg looks dark in most part of the entire image, but it seems to have a few bright regions. These bright regions are definitely the darker regions in result7.jpg.
- Though applying Power-law transform (result9.jpg) increases average brightness, the improvement of overall contrast is not apparent.

Problem 2: NOISE REMOVAL

Since noises are added randomly, execution result may differ.

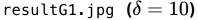
 \bullet Below is sample4.jpg , which is the original image and the standard in calculating PSNR value in (e) .



(a)

- The function $Gau(i,j)=I(i,j)+\delta\cdot N(\mu,\sigma)$ adds **Gaussian** Noise to the given image. (I: original image, Gau: image with gaussian noise, δ : amplitude, $N(\mu,\sigma)$: normal distribution with mean μ and variance σ , here $\mu=0$ and $\sigma=1$)
- The two images below are corrupted with Gaussian noise, with amplitude $\delta=10$ and 30.







resultG2.jpg ($\delta=20$)

(b)

- Define a threshold t ($0 \le t \le 1$), the function SNP(i,j) adds **salt and pepper** noise to the given image, determined by probability p with uniform distribution $U(\mu,\sigma)$. (Here $\mu=0$ and $\sigma=1$)
- For each pixel:

If p < t, then set SNP(i,j) to 0. If $t \le p \le 1-t$, then set SNP(i,j) to I(i,j). If p > 1-t, then set SNP(i,j) to 255. (I(i,j) represents the original image, while SNP(i,j) stands for image contaminated with salt and pepper noise.)

• The two images below are contaminated with salt and pepper noise, with threshold t=0.005 and 0.01.







resultS2.jpg (t=0.01)

(c)

• Since Gaussian noise is a kind of Uniform noise, performing low-pass filtering is a better choice to remove noise. We set the 3×3 mask as $\frac{1}{(b+2)^2}$

$$egin{bmatrix} 1 & b & 1 \ b & b^2 & b \ 1 & b & 1 \end{bmatrix}$$
 . Here I apply $b=3$.

• The two images below are results of images in (a) after low-pass filtering with mask above.



resultR1.jpg



resultR2.jpg

(d)

- Since Salt and pepper noise is a kind of Impulse noise, performing non-linear filtering is a better choice to remove noise. Here I choose outlier detection as the way to remove two noise images in **(b)**, and I set $\varepsilon=65$
- The two images below are results of images in (a) after low-pass filtering with mask above.





resultR3.jpg

resultR4.jpg

(e)

• The PSNR value of resultR1.jpg and resultR2.jpg are listed below:

Image	resultR1.jpg	resultR2.jpg
PSNR	33.79	26.48

Apparently, resultR1.jpg is better than resultR2.jpg, since resultR1.jpg has lower MSE, which means the mean of each pixel in resultR1.jpg has smaller error to the original image.

ullet The PSNR value of resultR3.jpg and resultR4.jpg are listed below:

Image	resultR3.jpg	resultR4.jpg
PSNR	37.56	34.75

Most of the **pepper** noise has been removed in both images. However, some **salt** noise (especially in resultR4.jpg) still exists in the background. I guess that the gray value of background is close to 255 (gray value of white pixel)

• According to <u>Wikipedia (https://bit.ly/39frNEI)</u>, typical PSNR value of 8-bit grayscale image is between $30db\sim50db$.

Bonus

• First, I apply median filtering with window size 7×5 , then I perform the low-pass filtering using the same mask in (c). The image below is the result image <code>result_bonus.jpg</code> after two-step process.

