

Astro dynamics Homework 5

a. The orbit is sufficient

$$r_J = 778.5 \times 10^6 \text{ Km} \quad r_E = 149.6 \times 10^6 \text{ Km}$$

$$r_p = r_E \quad e = 0.7 \quad r_a = \frac{149.6 \times 10^6 \text{ km}}{0.3} \cdot 1.7 = 847.7 \times 10^6 \text{ km}$$

$$a = \frac{r_p}{1-e} = \frac{r_a}{1+e} \quad r_a = \frac{r_p}{1-e} (1+e)$$

$$r_a > r_J$$

b.

$$r_{p,t} = r_E \quad e_t = 0.7 \quad r_{a,t} = 847.7 \times 10^6 \text{ Km}$$

$$V_{1,t} = \sqrt{\frac{2\mu_s}{r_{p,t}} - \frac{\mu_s}{\left(\frac{r_{p,t}}{1-e}\right)}} = 38832 \frac{\text{km}}{\text{s}}$$

$$V_E = \sqrt{\frac{\mu_s}{r_E}} = 29783 \quad V_1 = 9049$$

$$V_{2,t} = \sqrt{\frac{2\mu_s}{r_J} - \frac{\mu_s}{\left(\frac{r_{p,t}}{1-e}\right)}} = 8648.8 \quad V_J = \sqrt{\frac{\mu_s}{r_J}} = 13055$$

$$V_{2,\infty} = \sqrt{V_{2,t}^2 + V_J^2 - 2V_{2,t}V_J \cos(\phi)} \quad \sin \phi = e \quad \phi = 0.7754 \text{ rad}$$

$$V_{1,\infty} = 8639.5$$

$$V_{1,\infty} = V_{1,t} - V_E = 9049$$

$$1) V_{ph1} = \sqrt{V_{\infty 1}^2 + \frac{2 M_e}{r_e + 200}} = 14.25$$

$$V_{ce} = \sqrt{\frac{M_e}{r_e + 200}} = 7.7843$$

$$V_{ph2} = 60.15$$

$$V_{Gj} = 42.095$$

$$\Delta V_1 = 6.4661 \quad \Delta V_2 = 18.0557$$

$$\Delta V_{total} = 24.5219$$

$$2) \quad r_{ep} = r_e \quad r_{ta} = r_j$$

$$V_1 = \sqrt{\frac{2M_s}{r_e}} - \frac{2M_s}{r_e + r_j} =$$

$$V_j = \sqrt{\frac{M_s}{r_j}} = 13.0566$$

$$V_2 = \sqrt{\frac{2M_s}{r_j}} - \frac{2M_s}{r_e + r_j} = 7.4133 \quad V_2 = V_{2,a}$$

$$V_{200a,d} = V_j - V_{2,a} = 5.6433$$

all velocity for escape

$$V_{\text{escape}} = \sqrt{\frac{2M_s}{r}} = 18.4649$$

$$\delta_{max} = 2 \sin^{-1} \left(\frac{1}{1 + \frac{R_j V_\infty^2}{M_j}} \right) = 158.43^\circ$$

It can escape the solar system

$$V_{max} = V_{200a,d} + V_j = 18.7 > 18.46$$

3) $\vec{r}_1 = [-0.707, -0.707, 0] \text{ Au}$ TOF = 210 days

$\vec{v}_1 = [1.0423, -0.4124, 0] \text{ Au/TU}$

$|TU| = 58,13282 \text{ days}$

TOF

a) $\vec{r}_2 = [0, 1.6654, 0] \text{ Au}$

$\vec{v}_2 = [-0.6175, 0.2751, 0] \text{ AU/TU}$

b) $a = 1.3944 \text{ AU}$

$e = 0.4617$

$i = 0^\circ$

$\nu = \theta = 142.2^\circ$

c) $V_{\infty 1} = |\vec{v}_1| - \sqrt{\frac{M_s}{r_e}} = 0.1209 \frac{\text{AU}}{\text{TU}} = 3,6016 \frac{\text{km}}{\text{s}}$

$V_{\text{perih}} = \sqrt{V_{\infty 1}^2 + \frac{2M_s}{r_e + 200}} = 11,5828$

$V_{\text{ce}} = \sqrt{\frac{M_e}{r_e + 200}} = 7.7843$ $\Delta V = 3.7985$

d) $V_{\infty 2} = |\vec{v}_2| - \sqrt{\frac{2M_s}{r_a} - \frac{2M_s}{r_a + r_p}} = 1.8321$

$\Delta \Delta = \Delta(3522.2) - \Delta(3380) = 230.95$

3)

e.

$$V_{pmh} = \sqrt{V_{\infty 2}^2 + \frac{2M_m}{R_n + 400}} = 5.1122$$

$$V_{mc} = \sqrt{\frac{M_m}{R_n + 400}} = 3.3747$$

$$\Delta V_m = V_{pmh} - V_{mc} = 1.7374$$

$$4) \quad r_{m_d} = [1.3386 \quad 0.4525 \quad -0.0234] \text{ AU}$$

$$v_{m_d} = [-0.2292 \quad 0.8403 \quad 0.0232] \text{ AU/TU}$$

$$r_{e_a} = [-0.9591 \quad 0.2526 \quad 0] \text{ AU}$$

$$v_{e_a} = [-0.2711 \quad -0.9711 \quad 0] \text{ AU/TU}$$

$$b) \quad a = 1$$

$$e = 0.0167$$

$$i = 5 \times 10^{-5}^\circ$$

$$\Omega = 348.75^\circ$$

$$\omega = 114.53^\circ$$

$$\nu_e = 61.95^\circ$$

$$\nu_m = 42.51^\circ$$

$$c) \quad V_{\infty} = 0.1274 \text{ AU/TU} = 3.7945 \frac{\text{km}}{\text{s}}$$

$$V_{\infty} = 0.1951 \text{ AU/TU} = 4.3211 \frac{\text{km}}{\text{s}} \quad \Delta V_1 = V_p - V_{\infty}$$

$$V_p = \sqrt{V_{\infty}^2 + \frac{2\mu_m}{r_m + 500 \text{ km}}} = 6.049 \quad \Delta V_2 = V_{p2} - V_{\infty}$$

$$V_{cm} = \sqrt{\frac{\mu_m}{r_m + 500 \text{ km}}} = 3.331 \quad \boxed{\Delta V_{total} = 6.7415 \frac{\text{km}}{\text{s}}}$$

$$V_p = \sqrt{V_{\infty}^2 + \frac{2\mu_e}{r_e + 300 \text{ km}}} = 11.749$$

$$V_{ce} = \sqrt{\frac{\mu_e}{r_c + 300 \text{ km}}} = 7.7258$$

```

clear;
clc;
close all;
mag = @(vector) sqrt(vector(1)^2 + vector(2)^2 + vector(3)^2); % magnitude of a 3
element vector
% constants for the sun
duS = 1.4959965e8; % conversion from du sun to km
au = duS; % a sun du is sometimes called an au
tuS = 5.0226757e6; % conversion from tu sun to seconds
muS = 1.3271544e11; % mu for sun km^3/s^2
vuS = duS/tuS;
% constants for earth
duE = 6378.136; % conversion from du earth to km
tuE = 806.8118744; % conversion from tu earth to seconds
muE = 3.986004418e5; % mu for earth
% constants for mars
muM = 4.305e4

```

muM = 43050

```

duM = 3380;
rj = 778.5e6;
muJ = 126.687e6;

```

vj = sqrt(muS/rj)

vj = 13.0566

%problem1

phi = asin(0.7)

phi = 0.7754

v2infin = sqrt(8.6488^2+vj^2-2*8.6488*vj*cos(phi))

v2infin = 9.1646

vph1 = sqrt(9.049^2+2*muE/(duE+200))

vph1 = 14.2504

vce = sqrt(muE/(duE+200))

vce = 7.7843

Vph2 = sqrt(v2infin^2 + 2*muJ/(71492+200))

Vph2 = 60.1514

vcj = sqrt(muJ/71492)

vcj = 42.0957

```
dv1 = vph1 - vce
```

```
dv1 = 6.4661
```

```
dv2 = Vph2 - vcj
```

```
dv2 = 18.0557
```

```
dvtotal = dv1 + dv2
```

```
dvtotal = 24.5219
```

```
%problem2
```

```
v2prob2 = sqrt(2*muS/rj - 2*muS/(au + rj))
```

```
v2prob2 = 7.4133
```

```
vinfprob2 = vj - v2prob2
```

```
vinfprob2 = 5.6433
```

```
vescape = sqrt(2*muS/rj)
```

```
vescape = 18.4649
```

```
sdelta = 2*asind(1/(1+71492*(vinfprob2^2)/muJ))
```

```
sdelta = 158.4357
```

```
vmax = vinfprob2 + vj
```

```
vmax = 18.6999
```

```
%problem3
```

```
tof3 = 210*24*60*60/tuS;
```

```
r1 = [-0.707 -0.707 0]; % AU
```

```
v1 = [1.0423 -0.4124 0]; % AU/TU
```

```
[r2,v2] = timeOfFlight(r1,v1,tof3)
```

```
check = 1.0000
```

```
r2 = 1x3
```

```
0.0000 1.6654 0
```

```
v2 = 1x3
```

```
-0.6175 0.2751 0
```

```
elementsproblem3b = rv2elements(r2,v2,1)
```

```
elementsproblem3b = struct with fields:
```

```

a: 1.3444
e: 0.4617
i: 0
bigOmega: NaN
omega: NaN
nu: 142.2054
u: NaN

vinf1 = (mag(v1) - 1)*vuS

vinf1 = 3.6016

vpeh = sqrt(vinf1^2+2*muE/(duE+200))

vpeh = 11.5828

vce = sqrt(muE/(duE+200))

vce = 7.7843

deltav1 = vpeh - vce

deltav1 = 3.7985

vinf2 = abs(mag(v2)*vuS - sqrt((2*muS)/249.3e6 - (2*muS)/(249.3e6+206.7e6)))

vinf2 = 1.8321

delta1 = 3522.2/vinf2*sqrt(vinf2^2 + 2*muM/3522.2)

delta1 = 1.0137e+04

delta2 = 3380/vinf2*sqrt(vinf2^2 + 2*muM/3380)

delta2 = 9.9056e+03

deltadelta = delta1 - delta2

deltadelta = 230.9589

vpmh = sqrt(vinf2^2 + 2*muM/(3380+400))

vpmh = 5.1122

vcm = sqrt(muM/(3380+400))

vcm = 3.3747

deltavm = vpmh - vcm

deltavm = 1.7374

%problem4

re0 = [0.9422 -0.3614 -1.49e-7]; % AU
ve0 = [0.3419 0.93 8.543e-7]; % AU/TU

```

```

rm0 = [1.3871 0.1765 -0.0304]; % AU
vm0 = [-0.0714 0.8766 0.0201]; % AU/TU

% departure time in tu
ti = 706*24*60*60/tuS;

% time of flight in tu
tof = 210*24*60*60/tuS;

% departure position and velocity of mars in heliocentric frame
[rd,vd] = timeOffFlight(rm0,vm0,ti)

```

```

check = 1.0000
rd = 1x3
    1.3386    0.4525   -0.0234
vd = 1x3
   -0.2292    0.8403    0.0232

```

```

% arrival time from start
ta = ti + tof;

% departure position and velocity of earth in heliocentric frame
[ra,va] = timeOffFlight(re0,ve0,ta)

```

```

check = 1.0000
ra = 1x3
   -0.9591    0.2526    0.0000
va = 1x3
   -0.2711   -0.9711   -0.0000

```

```

% earth nu
earthElements = rv2elements(ra,va,1)

```

```

earthElements = struct with fields:
    a: 0.9999
    e: 0.0167
    i: 5.0010e-05
    bigOmega: 348.7544
    omega: 114.5398
    nu: 61.9497
    u: 176.4894

```

```
earthElements.nu
```

```
ans = 61.9497
```

```

% mars nu
marsElements = rv2elements(rd,vd,1);
marsElements.nu

```

```
ans = 42.5160
```

```
% phase 2: we have to r vectors find v vectors with gauss's  
[v1,v2] = gaussSolution(rd,ra,tof)
```

```
check = 1.0000  
check = 1.0000  
v1 = 1x3  
-0.2313    0.7202   -0.0192  
v2 = 1x3  
-0.1696   -1.0696    0.0324
```

```
transferElements = rv2elements(rd,v1,1)
```

```
transferElements = struct with fields:  
    a: 1.1866  
    e: 0.1916  
    i: 1.7241  
    bigOmega: 165.2437  
    omega: 37.2303  
    nu: 176.2133  
    u: 213.4436
```

```
transferElements = rv2elements(ra,v2,1)
```

```
transferElements = struct with fields:  
    a: 1.1866  
    e: 0.1916  
    i: 1.7241  
    bigOmega: 165.2437  
    omega: 37.2303  
    nu: 322.7698  
    u: 1.0177e-04
```

```
vinf1 = mag(v1 - vd)*vuS;  
  
vinf2 = mag(v2 - va)*vuS;  
  
vph1 = sqrt(vinf1^2+2*muM/(duM+500));  
  
vph2 = sqrt(vinf2^2+2*muE/(duE+300));  
vcm = sqrt(muM/(duM+500));  
vce = sqrt(muE/(duE+300));  
deltav1 = vph1 - vcm;  
deltav2 = vph2 - vce;  
totaldeltav = deltav1 + deltav2
```

```
totaldeltav = 6.7415
```

```
function [V1s,V2s,V1l,V2l] = gaussSolution(r1,r2,tof)  
% this function finds the solution to gausses problem when given values  
% for position in DU and a TOF in TU. The s are the short way and  
% the l is the long way.
```

```
zguess = 10;
```

```

mag = @(vector) sqrt(vector(1)^2 + vector(2)^2 + vector(3)^2); % magnitude of a
3 element vector

dtheta = acos(dot(r1,r2)/(mag(r1)*mag(r2))); % the angle between the position
vectors

A = (sqrt(mag(r1)*mag(r2))*sin(dtheta))/sqrt(1-cos(dtheta));

error = 1;
znplus1 = zguess;

while error > 1*10^(-5)
    [s,c] = fofZ(znplus1);

    y = mag(r1) + mag(r2) - A*(1-znplus1*s)/sqrt(c);

    x = sqrt(y/c);

    t = x^3*s + A*sqrt(y);

    error = abs(tof - t);

    dsdz = (c - 3*s)/2*znplus1;

    dcdz = (1 - znplus1*s - 2*c)/2*znplus1;

    dtdz = x^3 * (dsdz - 3*s*dcdz/(2*c)) + A/8 * (3*s*sqrt(y)/c + A/x);

    znplus1 = znplus1 + (tof-t)/dtdz;

end

f = 1 - y/mag(r1);
fdot = -x/(mag(r1)*mag(r2))*(1-znplus1*s);
g = A*sqrt(y);
gdot = 1 - y/mag(r2);

V1s = (r2 - f*r1)/g;
V2s = (gdot*r2 - r1)/g;

check = f*gdot-fdot*g

dthetal = 2*pi-dtheta; % the angle between the position vectors

A1 = (sqrt(mag(r1)*mag(r2))*sin(dthetal))/sqrt(1-cos(dthetal));

```

```

error = 1;
znplus1 = zguess;

while error > 1*10^(-5)
    [s,c] = fofZ(znplus1);

    y = mag(r1) + mag(r2) - A1*(1-znplus1*s)/sqrt(c);

    x = sqrt(y/c);

    t = x^3*s + A1*sqrt(y);

    error = abs(tof - t);

    dsdz = (c - 3*s)/2*znplus1;

    dc当地 = (1 - znplus1*s - 2*c)/2*znplus1;

    dtdz = x^3 * (dsdz - 3*s*dc当地/(2*c)) + A1/8 * (3*s*sqrt(y)/c + A1/x);

    znplus1 = znplus1 + (tof-t)/dtdz;

end

f = 1 - y/mag(r1);
fdot = -x/(mag(r1)*mag(r2))*(1-znplus1*s);
g = A1*sqrt(y);
gdot = 1 - y/mag(r2);

V11 = (r2 - f*r1)/g;
V21 = (gdot*r2 - r1)/g;

check = f*gdot - fdot*g

function [s,c] = fofZ(z) % this function finds the s and c values for a given z

if z > 0.1
    s = (sqrt(z)-sin(sqrt(z)))/z^(3/2);
    c = (1-cos(sqrt(z)))/z;
elseif z < -0.1
    s = (sinh(-z)-sqrt(-z))/(-z)^(3/2);
    c = (1 - cosh(-z))/z;
else
    s = 1/6 - z/factorial(5) + z^2/factorial(7) - z^3/factorial(9);
    c = 1/2 - z/factorial(4) + z^2/factorial(6) - z^3/factorial(8);
end
end

```

```

end

function [r1,v1] = timeOfFlight(r0,v0,tof)
% this function finds some new r and v vectors from an initial vector and a
% time of flight everything must be in canonical units

xguess = 1000;

mag = @(vector) sqrt(vector(1)^2 + vector(2)^2 + vector(3)^2); % magnitude of a
3 element vector

elements = rv2elements(r0,v0,1);

error = 1;
x = xguess;

while error > 1*10^(-5)

znplus1 = x^2/elements.a;

[s,c] = fofZ(znplus1);

t = x^3*s + dot(r0,v0)*x^2*c+mag(r0)*x*(1-znplus1*s);

rn = x^2*c+dot(r0,v0)*x*(1-znplus1*s)+mag(r0)*(1-znplus1*c);

error = abs(tof - t);

dtdx = rn;

x = x + (tof-t)/dtdx;

end

f = 1 - elements.a/mag(r0)*(1 - cos(x/sqrt(elements.a)));

fdot = -sqrt(elements.a)/(mag(r0)*rn)*sin(x/sqrt(elements.a));

g = (elements.a)^2/sqrt(elements.a)*(dot(r0,v0)/sqrt(elements.a)*(1-cos(x/
sqrt(elements.a)))+mag(r0)/elements.a*sin(x/sqrt(elements.a)));

gdot = 1 - elements.a/rn + elements.a/rn*cos(x/sqrt(elements.a));

check = f*gdot-fdot*g

r1 = f*r0 + g*v0;
v1 = fdot*r0 + gdot*v0;

```

```
function [s,c] = fofZ(z) % this function finds the s and c values for a given z

if z > 0.1
    s = (sqrt(z)-sin(sqrt(z)))/z^(3/2);
    c = (1-cos(sqrt(z)))/z;
elseif z < -0.1
    s = (sinh(-z)-sqrt(-z))/(-z)^(3/2);
    c = (1 - cosh(-z))/z;
else
    s = 1/6 - z/factorial(5) + z^2/factorial(7) - z^3/factorial(9);
    c = 1/2 - z/factorial(4) + z^2/factorial(6) - z^3/factorial(8);

end
end

end
```