



Harvard-Chile Data Science School

An Innovative Learning Exchange

Visualization of Supernovae Classification

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Outline

1. Introduction
2. Getting the information
 - a. Tension as an uncertainty measure
 - b. Making the csv file
3. Visualizing changes through time
4. Demo
5. Future Work

Introduction: Why Visualization?

- Vision is the most important channel for receiving information from our environment



Figure 1: Scientific visualization, using data from a computer simulation. It shows a Supernova 1987A. *Credits: NASA, ESA, and F. Summers and G. Bacon (STScI); Simulation Credit: S. Orlando (INAF-Osservatorio Astronomico di Palermo)*

Introduction: Why Visualization?

- It's hard for humans detect patterns from numbers

	id_sn	time	RA	DECL	MWEBV	REDSHIFT FINAL	g	r	i	z	g_e	r_e	i_e	z_e	SIM REDSHIFT	class
0	393133	1.10850	34.5	-5.5	0.0227	0.695	36.7300	8.8670	-16.9800	-1.8180	30.750	3.599	7.060	2.926	0.7001	2
1	393133	2.09000	34.5	-5.5	0.0227	0.695	-5.3990	0.4444	6.1790	3.0800	9.315	3.863	6.318	2.922	0.7001	2
2	393133	3.10750	34.5	-5.5	0.0227	0.695	0.6672	4.0030	-0.3458	-0.5002	7.694	2.655	4.790	3.017	0.7001	2
3	393133	7.13700	34.5	-5.5	0.0227	0.695	6.9330	-3.9910	1.7490	-2.4030	3.789	1.758	3.200	2.936	0.7001	2
4	393133	8.15075	34.5	-5.5	0.0227	0.695	0.4684	-2.3700	0.3952	2.4730	2.000	1.875	2.783	3.190	0.7001	2

- We need to translate these measures in order to show it clearly

Presentation and Exploratory Graphics

- Presentation graphics are static. They're useful for presenting summaries of information

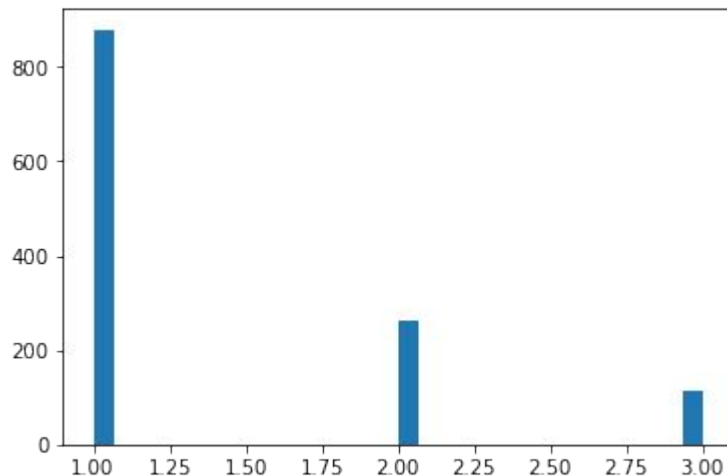
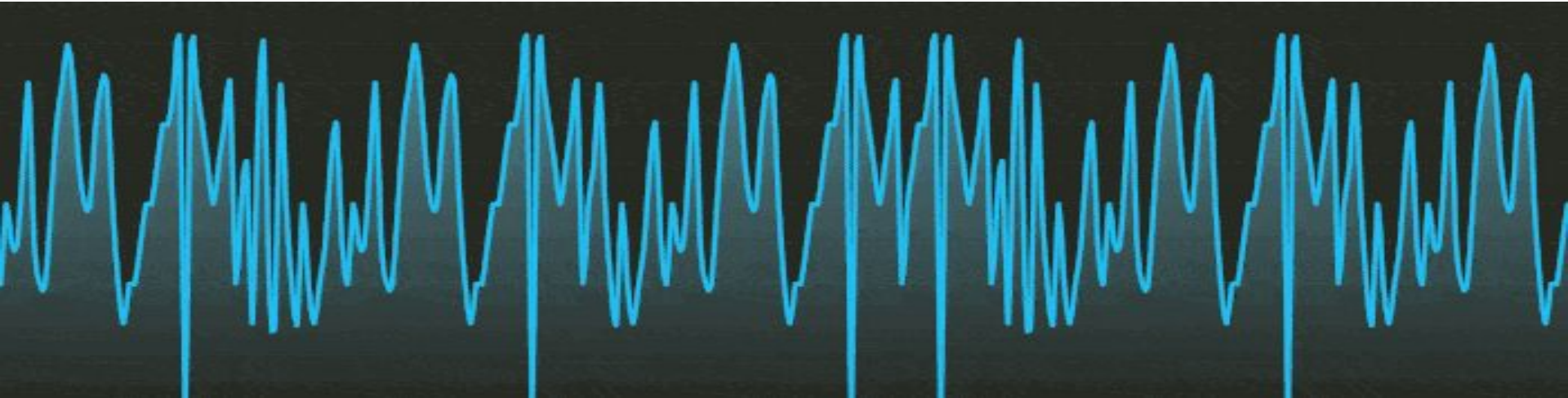


Figure 2: Histogram of supernovae sorted by type

Presentation and Exploratory Graphics

- Exploratory graphs are used to look for results. In the context of supernovae classification, they have to show us changes through time

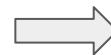
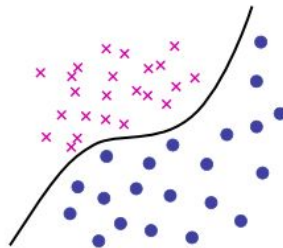
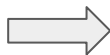
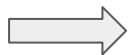
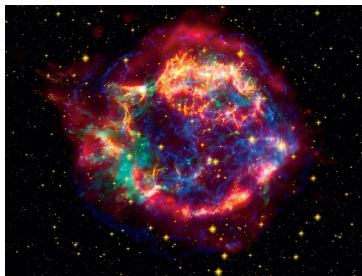


The problem

We want to explore the results from an online classification of supernovae. In order to do this, we have to visualize:

1. **Probability of type classification** for each supernova
2. **Tension** between classes
3. Global information

Online Classification

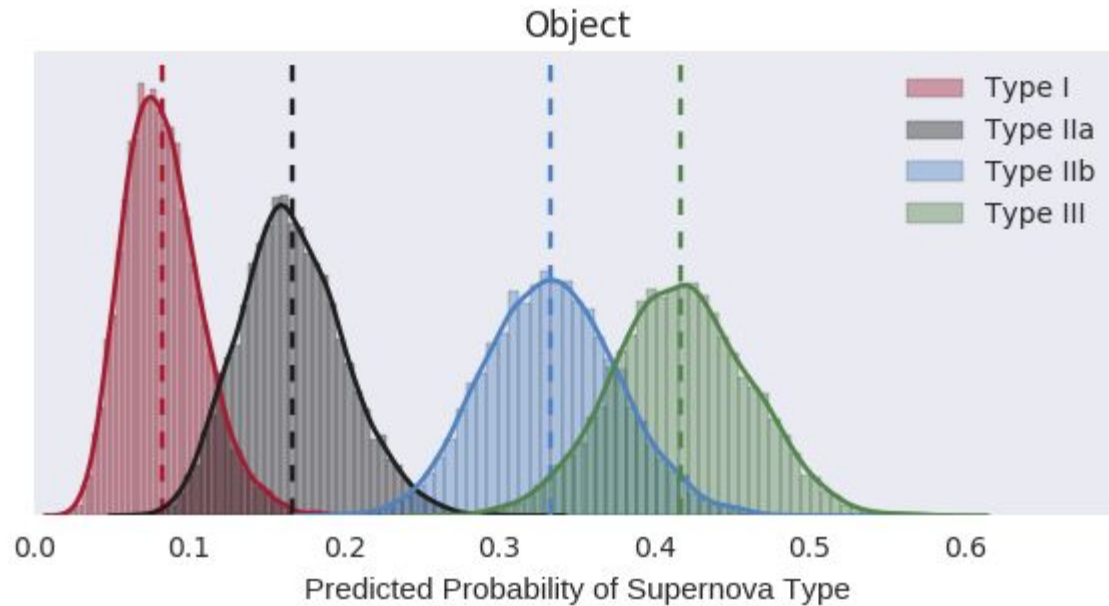


We receive
information from the
universe via
observations at each
time step

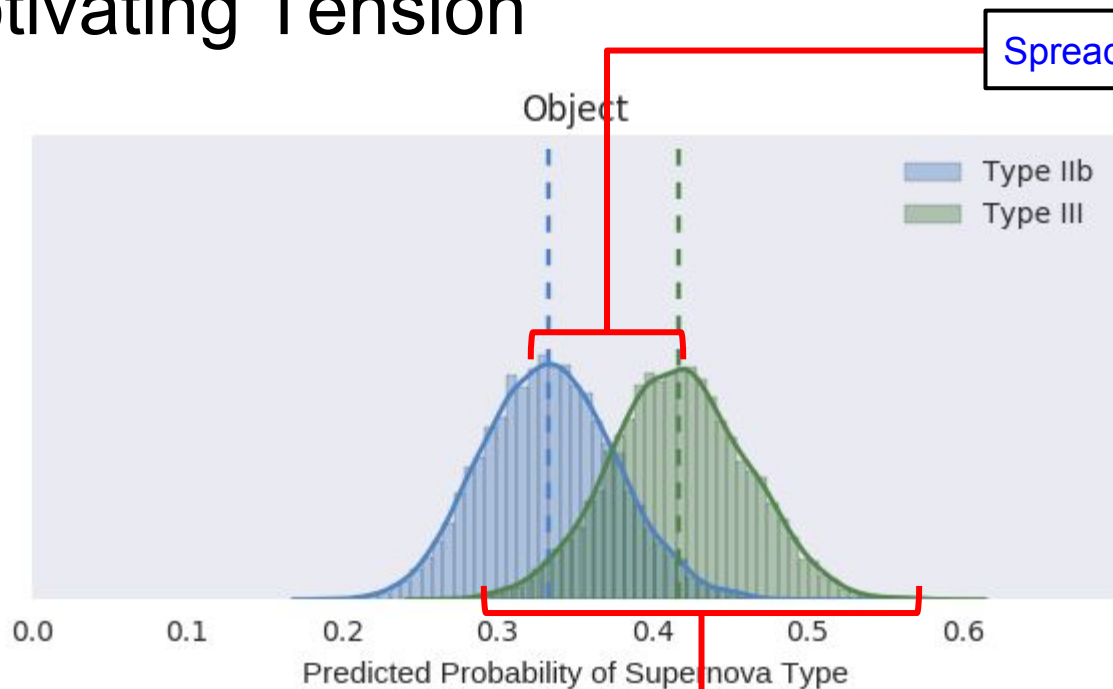
We train a model &
make
classifications

We want to
visualize the
process & results

Motivating Tension



Motivating Tension

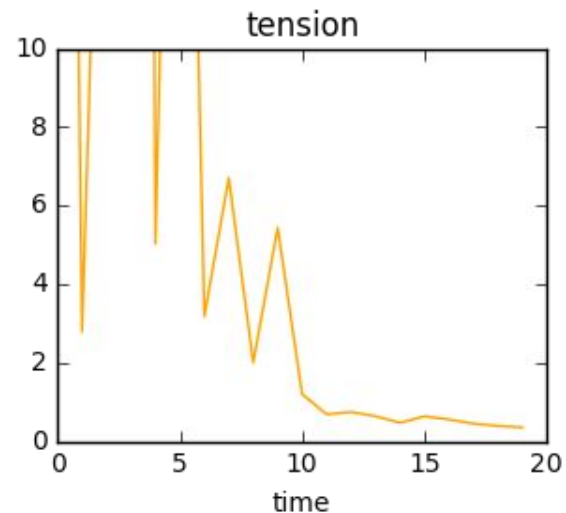
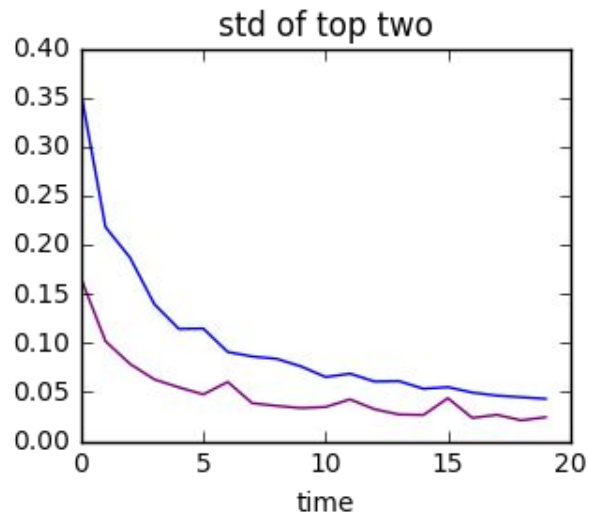
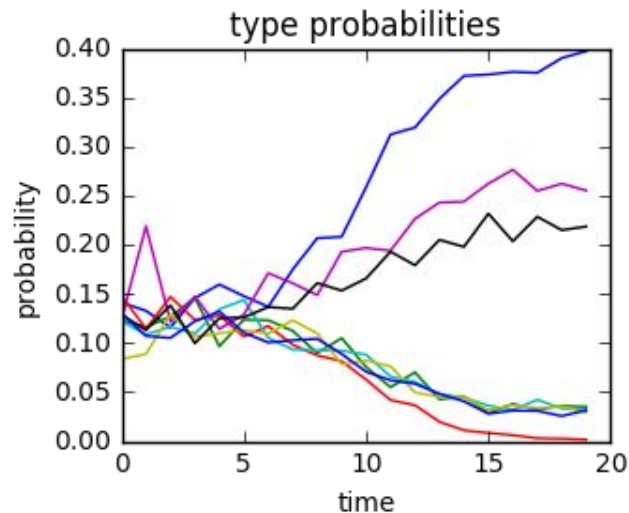


Spread

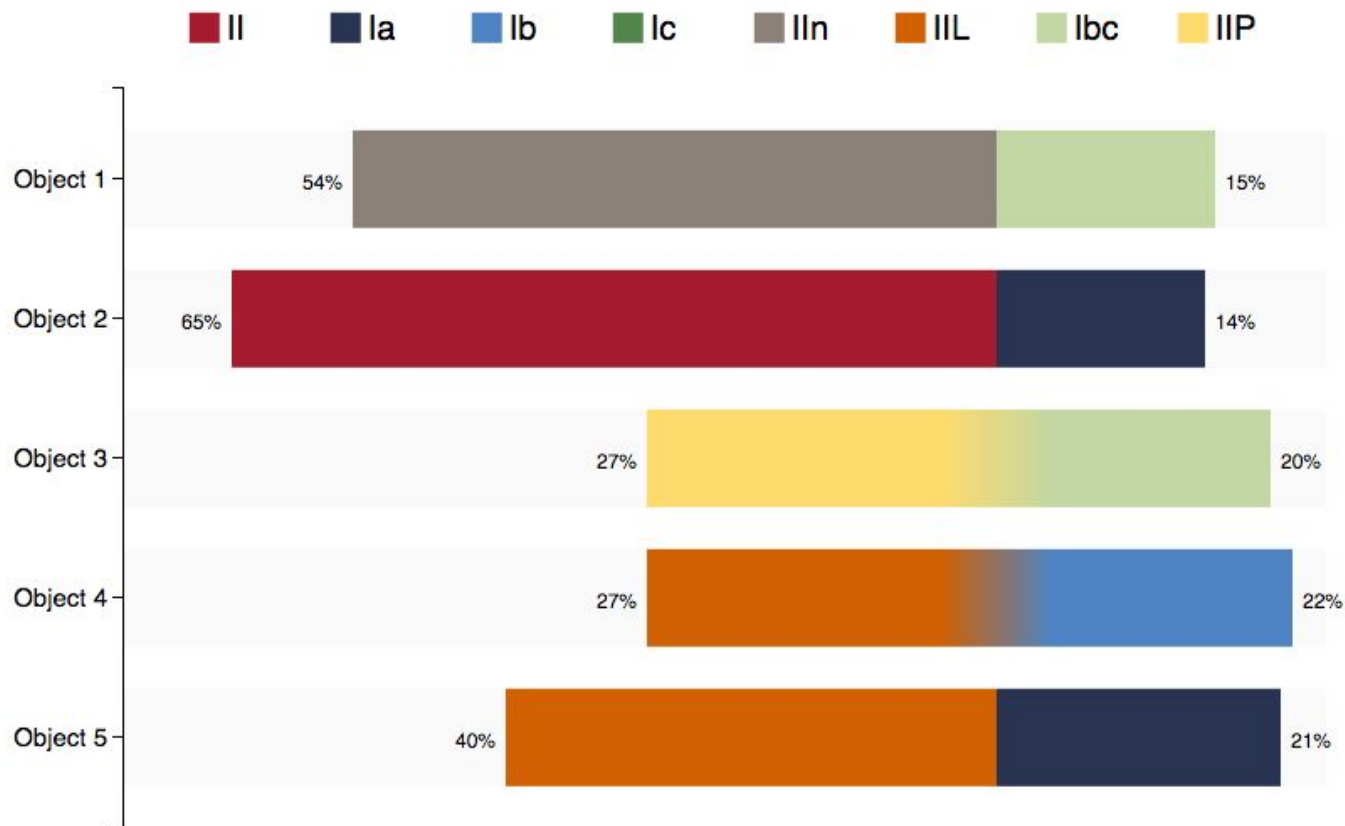
$$\frac{|P1 - P2|}{\sqrt{\sigma_1^2 + \sigma_2^2}}$$

Distribution "fatness"

Simulating Online Predictions



Our Visualization



Future Work

- Time Elapse
- Online Feedback - Have we achieved classification yet? If not, should we keep observing?
- Visualize thousands of objects at different levels of granularity
- Accessibility for color blindness

