MULTIPLE CHOICE. Choose the one alternative that best completes the statement or answers the question.

| 1) According to our theory of solar system formation, which law best explains why the solar nebula | | | |
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| spun faster as it shrank in size? | | | |
| A) the law of universal gravitation | | | |
| B) the law of conservation of angular momentum | | | |
| C) Einstein's law $E = mc^2$ | | | |
| D) the law of conservation of energy | | | |
| 2) What is the primary basis upon which we divide the ingredients of the solar nebula into four | 2) | | |
| categories (hydrogen/helium; hydrogen compound; rock; metal)? | _, | | |
| A) the atomic mass numbers of various materials | | | |
| B) the amounts of energy required to ionize various materials | | | |
| C) the locations of various materials in the solar nebula | | | |
| D) the temperatures at which various materials will condense from gaseous form to solid | | | |
| form | | | |
| 2) According to our present theory of color system formation which of the following statements | 3) | | |
| 3) According to our present theory of solar system formation, which of the following statements about the growth of terrestrial and jovian planets is <i>not</i> true? | 3) | | |
| A) The jovian planets began from planetesimals made only of ice, while the terrestrial | | | |
| planets began from planetesimals made only of rock and metal. | | | |
| B) The terrestrial planets formed inside the frost line of the solar nebula and the jovian | | | |
| planets formed beyond it. | | | |
| C) Both types of planet begun with planetesimals growing through the process of accretion, | | | |
| but only the jovian planets were able to capture hydrogen and helium gas from the solar | | | |
| nebula. | | | |
| D) Swirling disks of gas, like the solar nebula in miniature, formed around the growing | | | |
| jovian planets but not around the growing terrestrial planets. | | | |
| 4) Why are terrestrial planets denser than jovian planets? | 4) | | |
| A) Actually, the jovian planets are denser than the terrestrial planets. | , | | |
| B) Gravity compresses terrestrial planets to a higher degree, making them denser. | | | |
| C) The Sun's gravity gathered dense materials into the inner solar system. | | | |
| D) Only dense materials could condense in the inner solar nebula. | | | |
| 5) Current evidence suggests that many massive jovian planets orbit at very close orbital distances | 5) | | |
| to their stars. How do we think these planets ended up on these close orbits? | 5) | | |
| A) These planets migrated inward after being born on orbits much farther from their stars. | | | |
| B) Despite their large masses, these planets are terrestrial in nature and therefore could form | | | |
| in their inner solar systems. | | | |
| C) These planets are jovian in nature and were able to form close to their stars because their | | | |
| solar nebulas were very cold in temperature. | | | |
| D) These planets were captured from other solar systems. | | | |
| · • • • • • • • • • • • • • • • • • • • | | | |
| 6) Suppose you are using the Doppler technique to look for planets around another star. What must you do? | 6) | | |
| A) Compare many spectra of the star taken over a period of many months or years. | | | |
| B) Compare many spectra of an orbiting planet taken over a period of many months or years. | | | |
| C) Carefully examine a single spectrum of an orbiting planet. | | | |

D) Compare the brightness of the star over a period of many months or years.

| | 7) In general, which type of planet would you expect to cause the largest Doppler shift in the | 7) |
|-------|--|-----|
| | spectrum of its star? | |
| | A) a massive planet that is close to its star | |
| | B) a low-mass planet that is far from its star C) a massive planet that is far from its star | |
| | • | |
| | D) a low-mass planet that is close to its star | |
| | 8) Suppose a planet is discovered by the Doppler technique and is then discovered to have transits. | 8) |
| | In that case, we can determine all the following about the planet <i>except</i> | |
| | A) its physical size (radius). | |
| | B) its rotation period. | |
| | C) its precise mass. | |
| | D) its density. | |
| | E) its orbital period. | |
| | O) Von de como et a como et a la como et a l | 0) |
| | 9) You observe a star very similar to our own Sun in size and mass. This star moves very slightly | 9) |
| | back and forth in the sky once every four months, and you attribute this motion to the effect of | |
| | an orbiting planet. What can you conclude about the orbiting planet? | |
| | A) The planet must be farther from the star than Neptune is from the Sun. | |
| | B) The planet must have a mass about the same as the mass of Jupiter. | |
| | C) The planet must be closer to the star than Earth is to the Sun. | |
| | D) You do not have enough information to say anything at all about the planet. | |
| | 10) All the following statements about known extrasolar planets are true. Which one came as a | 10) |
| | surprise to scientists who expected other solar systems to be like ours? | |
| | A) Most of the planets orbit stars that are quite nearby compared to the scale of the entire | |
| | Milky Way Galaxy. | |
| | B) In some cases, we've found more than one planet orbiting the same star. | |
| | C) Most of the planets are quite massive—much more like Jupiter than like Earth. | |
| | D) Some of the planets orbit their star more closely than Mercury orbits the Sun. | |
| D 11 | | |
| Probl | ems | |
| | 11) Assume a planet (mass = 1 $M_{jupiter}$ = 1.9 x 10 ³⁰ g) is in circular orbit at 5.2 11) | |
| | au around its 1M☉ - host star. | |
| | (a) Calculate the orbital period of the planet. | |
| | (b) Using the properties of the centre of mass, calculate the radius of the | |
| | star's orbitabout the centre of mass of the system. | |
| | (c) If this system is at 10 pc, what is the angular amplitude (in arcsec) of | |
| | the star's motion in response to the gravitational force exerted on it by the | |
| | exoplanet? Assumethe inclination of the orbit is 90°, such that sin(i)=1. | |
| | (d) By how much will a spectral line from the star be Doppler-shifted due | |
| | to the exoplanet's orbital motion? Hint: You might find it more useful to use | |
| | Kepler's thirdlaw in terms of velocities (as we derived in the lecture | |
| | "Binary star systems"). | |
| | (e) Recalculate item (d) now assuming that this exoplanet orbits at 0.05 | |
| | àu (a factor of100 smàller). | |
| | | |

| 12) | Calculate | the relative | e flux of the | e star that is | blocked b | by a transi | ting |
|-----|-----------|----------------|---------------|----------------|-----------|-------------|------|
| | exoplanet | t in the follo | wing syste | ems: | | | Ŭ |

| 12) | |
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| | |

- (a) an Earth-sized planet (radius = 6400 km) orbiting around an M dwarf star with radius of 0.2 R_{\odot} .(
- b) a Jupiter-sized planet (radius = $7.2x10^4$ km) orbiting around a solar twin (ie., star with same mass and radius as the Sun).