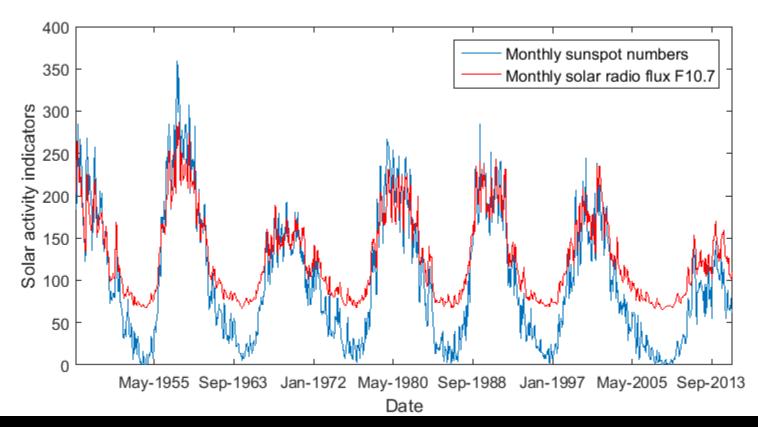


"Experimental Data Processing"

Laboratory work 1
Relationship between solar radio flux F10.7
and sunspot number

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Main indicator of solar activity



Sunspot number

$$R = k(n + 10g)$$

n – number of observed sunspots

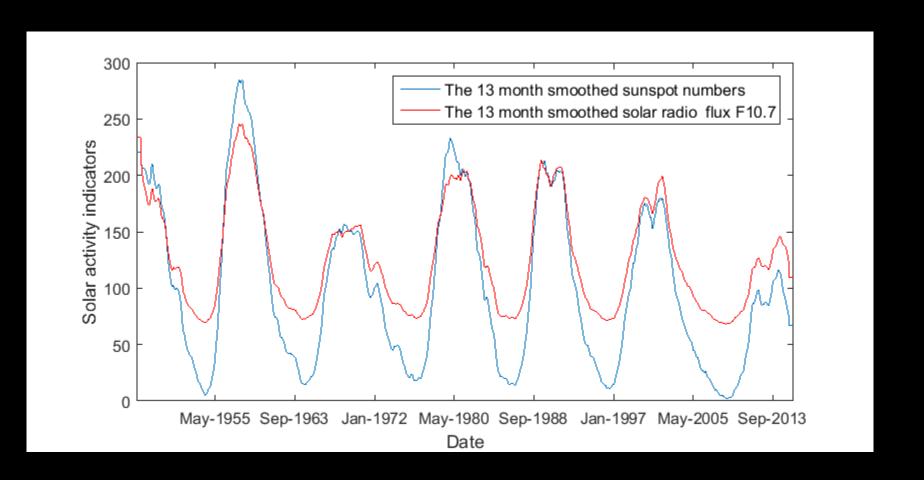
g – number of observed sunspot groups

k - coefficient of a telescope

Solar radio Flux at 10.7 cm

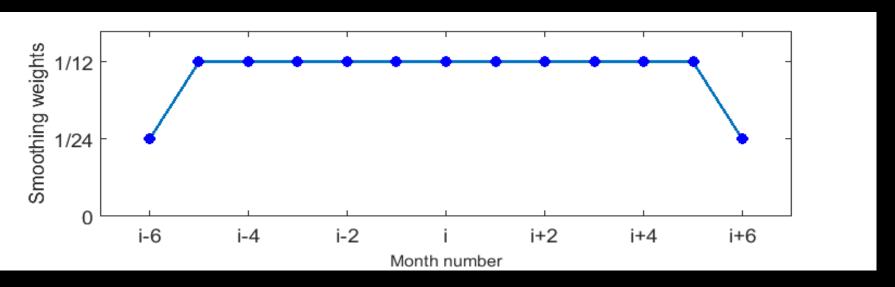
A measurement of radio emission at a wavelength of 10.7 cm (2800 MHz) from all sources present on the solar disk

Smoothing: 13-month running mean



13-month sequent monthly mean sunspot numbers





13-month running mean \overline{R}

$$\frac{1}{24}R_{i-6} + \frac{1}{12}(R_{i-5} + R_{i-4} + \dots + R_{i-1} + R_i + R_{i+1} + \dots + R_{i+5}) + \frac{1}{24}R_{i+6}$$

Multi-dimensional linear regression

$$F_i = \beta_0 + \beta_1 R_i + \beta_2 R_i^2 + \beta_3 R_i^3 + \varepsilon_i$$

 $i=1,\ldots N$

*F_i*Dependent
variable
Regressand

 $oldsymbol{eta_j}$ Coefficients of regression

 R_i Independent variable Regressor

ε_i
Unbiased
uncorrelated
Gaussian noise
with constant
variance

Coefficients
$$\beta_j$$
 are determined by LSM

$$\sum_{i=1}^{N} \varepsilon_i^2 \rightarrow min$$

Multi-dimensional linear regression

$$F = \begin{vmatrix} f_1 \\ f_2 \\ \vdots \\ f_N \end{vmatrix} \Rightarrow \begin{cases} \text{Vector of dependent variables} \end{cases}$$

$$\beta = \begin{vmatrix} \beta_0 \\ \beta_1 \\ \beta_2 \\ \beta_3 \end{vmatrix} \Rightarrow \begin{cases} \text{Vector of coefficients} \end{cases}$$

$$R = \begin{vmatrix} 1 & r_1 & r_{1^2} & r_1^3 \\ 1 & r_2 & r_2^2 & r_3^3 \\ \dots & \dots & \dots & \dots \\ 1 & r_N & r_N^2 & r_N^3 \end{vmatrix} \Rightarrow \text{inc}$$

$$F = R \cdot \beta + \varepsilon$$

$$\widehat{\boldsymbol{\beta}} = \left(R^T R\right)^{-1} R^T F$$

Linear Regression Analysis, G.A.F. Seber and J. Lee, Wiley, N.Y., 2003

Estimation error of solar radio flux F10.7

Covariance matrix of estimation error
$$\sigma^2 = \frac{1}{N-1} \sum_{i=1}^{N} (f_i - \widehat{f}_i)^2$$