

# Geodynamic controls on melting in TERRA 3D mantle convection models

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## I. MOTIVATION

• He and Ar provide major constraints on Earth thermal evolution and outgassing that need to be satisfied within our geophysical understanding of the mantle.

• 40-50% of produced <sup>40</sup>Ar is currently in the atmosphere.

• High <sup>3</sup>He/<sup>4</sup>He and low <sup>40</sup>Ar/<sup>36</sup>Ar ratios in OIB samples suggest preservation of primordial or unprocessed material compared to MORB samples (figure 1).

• Melting creates compositional heterogeneity, segregates elements and allows transfer to the atmosphere. It is therefore a crucial process to understand for He and Ar signatures.

We investigate the links between geodynamics and melting, i.e. the impact of different amounts of Heat Producing Elements (HPEs), various viscosity settings and densities of enriched material (buoyancy number) on melting in terms of production and type of material processed.

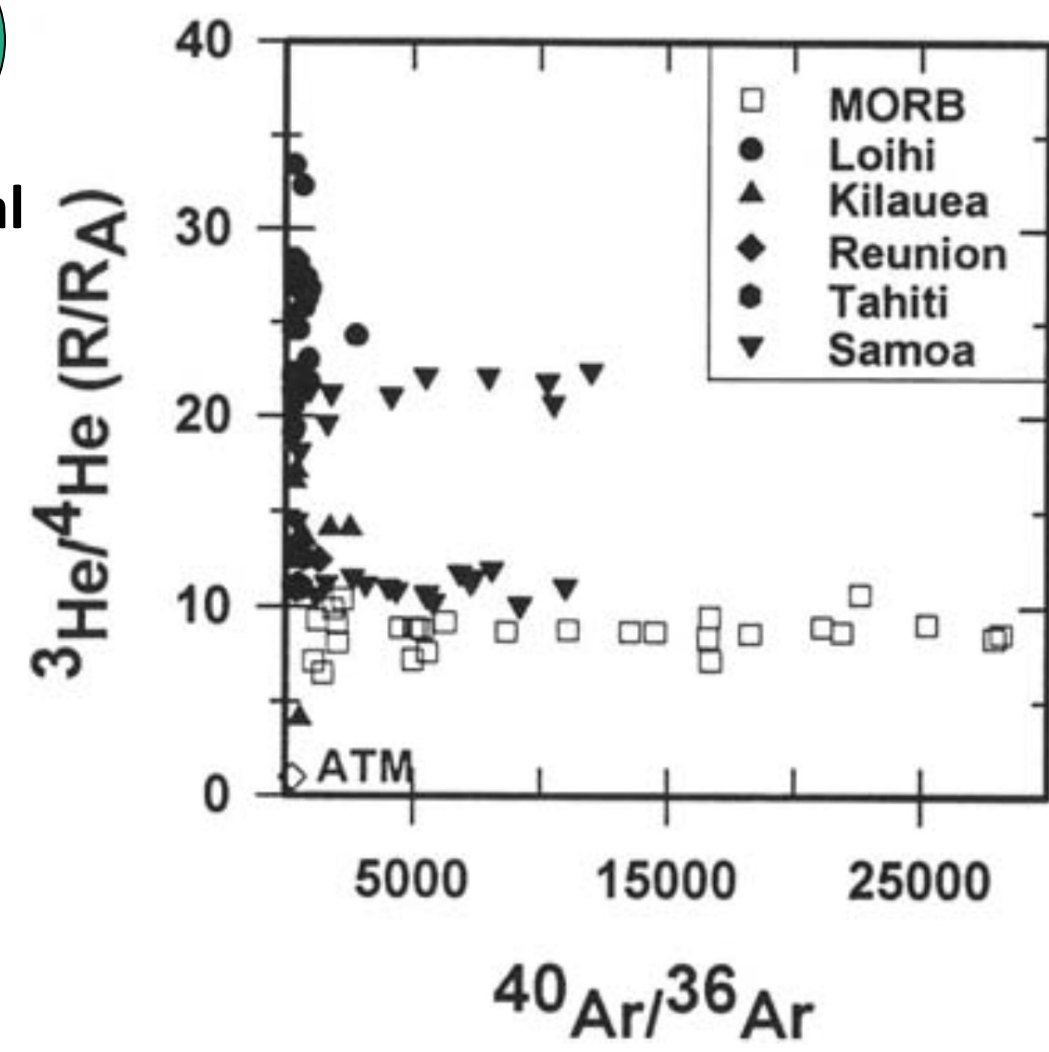


Figure 1 : Helium and argon ratios from OIBs and MORB samples. G.F. Davies, 2010

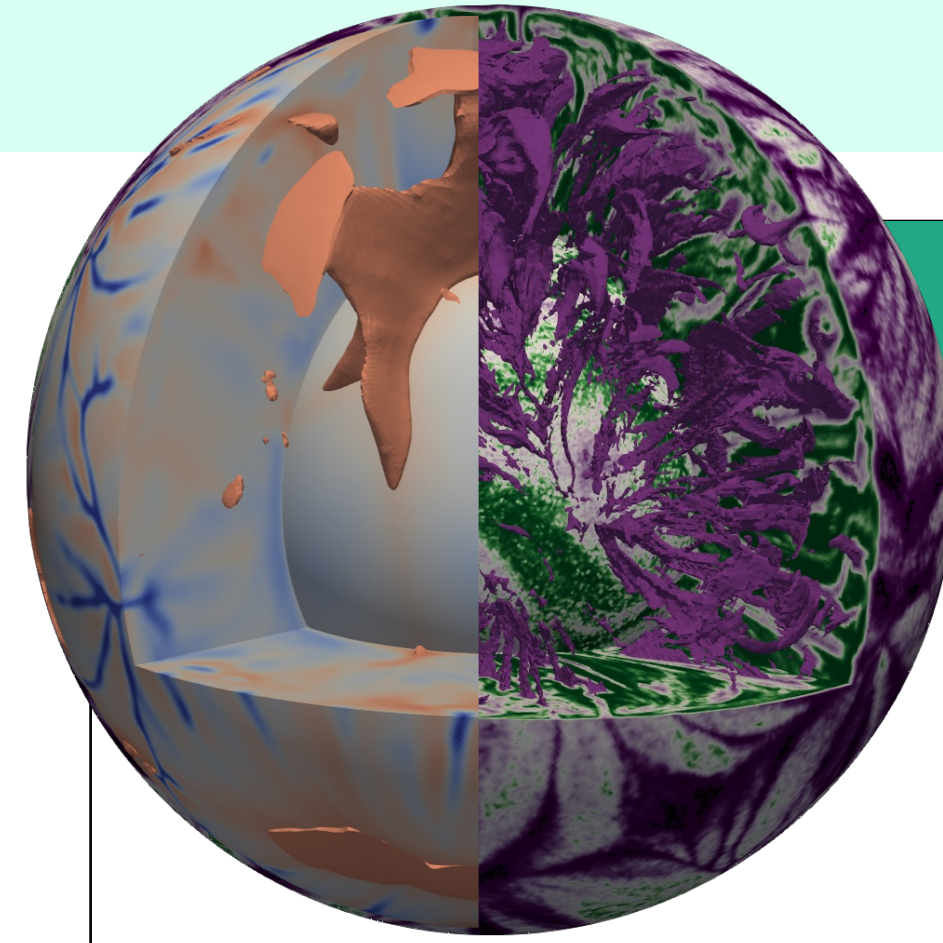


Figure 2: Visualization of a TERRA simulation with the temperature field (left side) and composition field (right side)

## II. METHODS: MELTING IN TERRA

• TERRA is a 3D finite element code solving mass, momentum and energy conservation equations for heat transfer in spherical shell (J. Baumgardner, 1983).

• Particles are advected by the flow and carry bulk composition and isotope information that impact the dynamics, as buoyancy is composition dependent and radiogenic heating is function of local U, Th and K concentrations (van Heck *et al*, 2016).

• Melting is differentiating the mantle from initial homogeneous composition C=0.2. It allows distribution of elements according to their geochemical behaviour.

• Solidus is a function of depth, temperature and composition (figure 3).

• Melting occurs when a particle surpasses its solidus. As a particle melts, it is depleted until the composition gives a melting temperature that is the particle's temperature (figure 3). The amount of depletion is redistributed to a receiving particle at the surface.

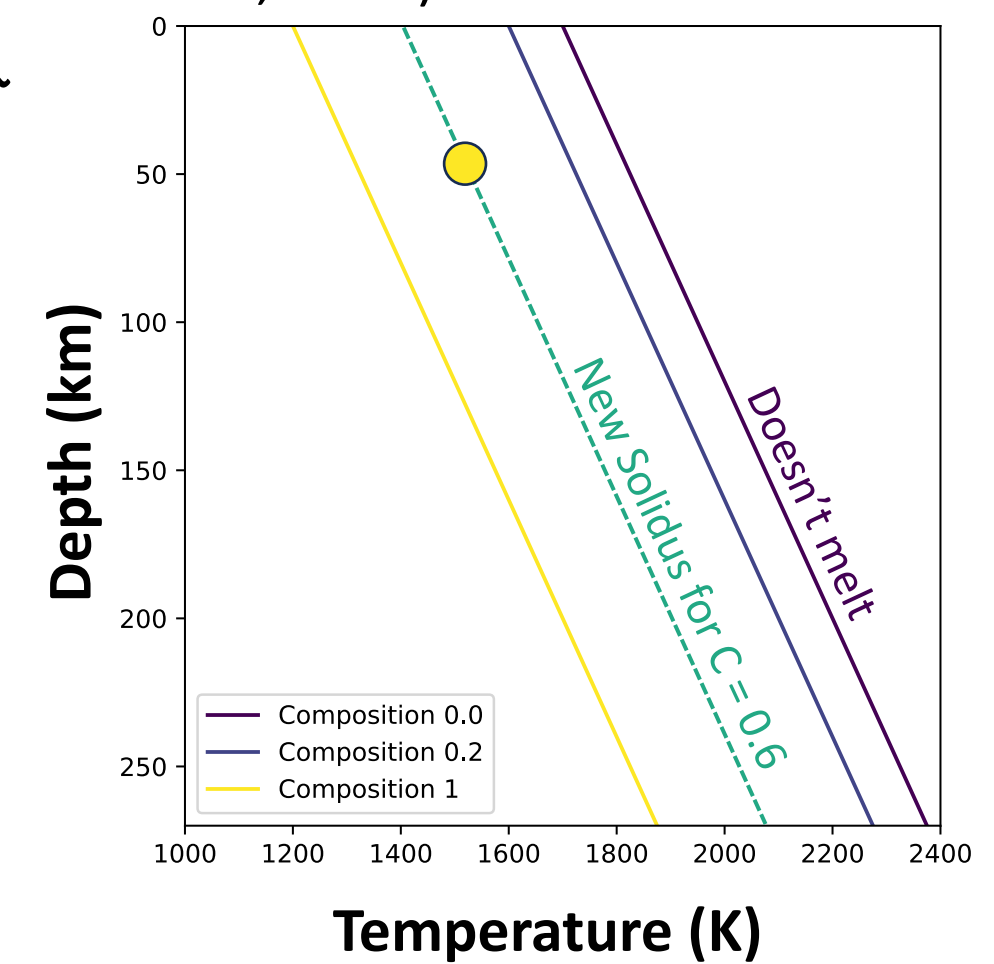
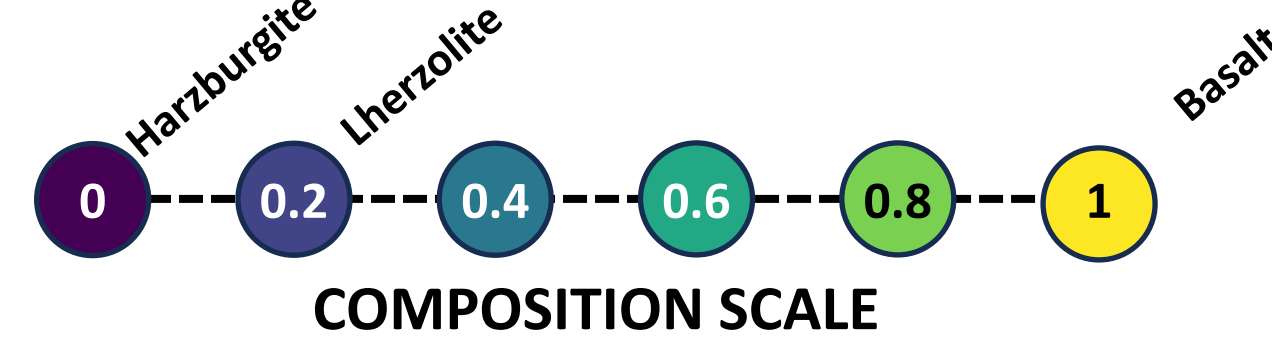


Figure 3: Solidi and melting in TERRA

## III. RESULTS

TABLE: Presented simulations

Presented simulations are incompressible and Boussinesq

Run name	Reference Viscosity (Pa.s)	Viscosity Structure	Proportion of HPEs	Density excess of enriched material (%)
ref	3.00E+22	2 layers	1	2
02_HPEs	3.00E+22	2 layers	0.2	2
05_HPEs	3.00E+22	2 layers	0.5	2
0_densdif	3.00E+22	2 layers	1	0
6_densdif	3.00E+22	2 layers	1	6
10_densdif	3.00E+22	2 layers	1	10
3.10 <sup>21</sup> _visc	3.00E+21	2 layers	1	2
1.10 <sup>22</sup> _visc	1.00E+22	2 layers	1	2
3.10 <sup>23</sup> _visc	3.00E+23	2 layers	1	2
3_layers	5.00E+21	3 layers	1	2

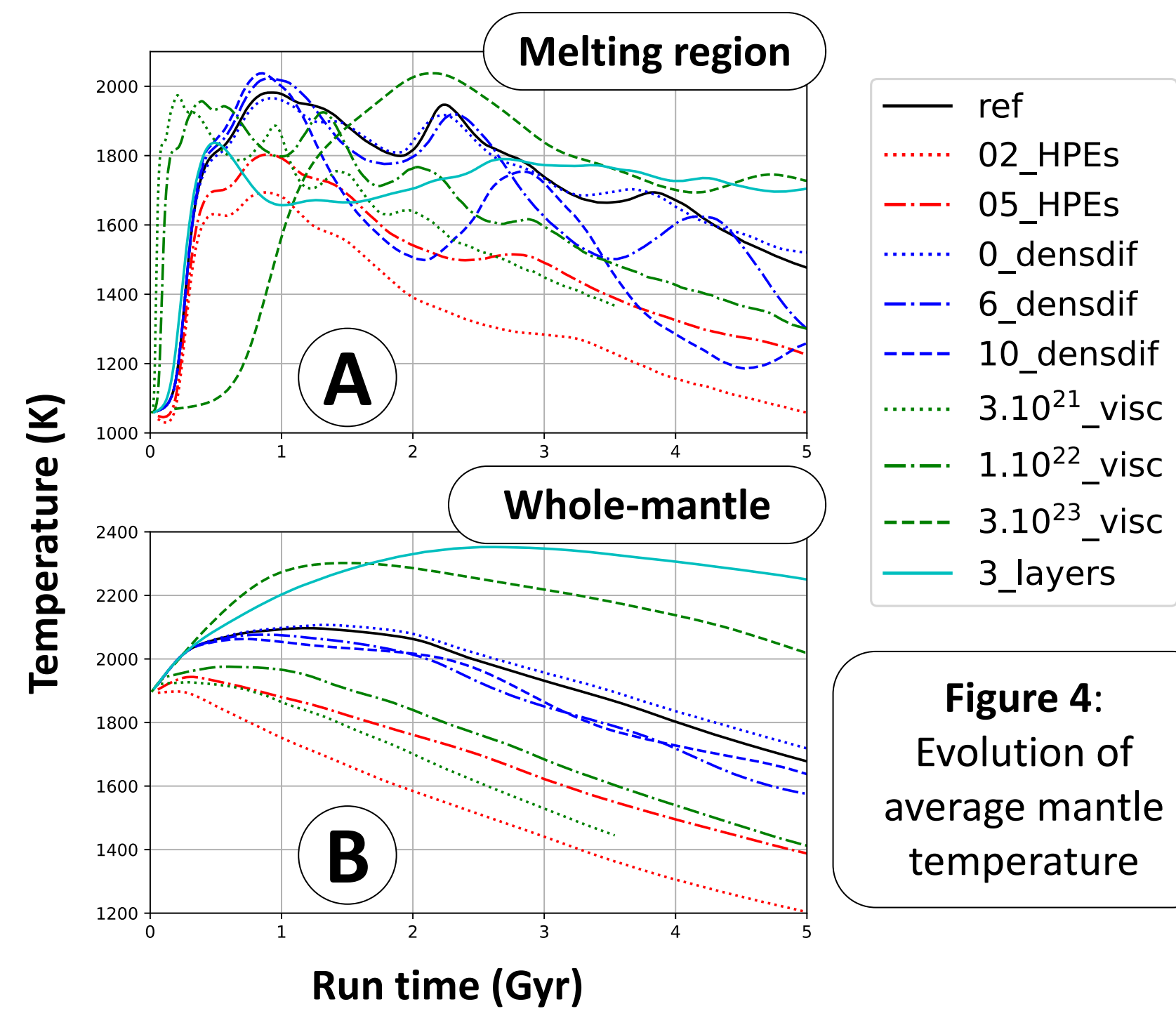


Figure 4: Evolution of average mantle temperature

## Viscosity structure and surface velocities control the type of material melting

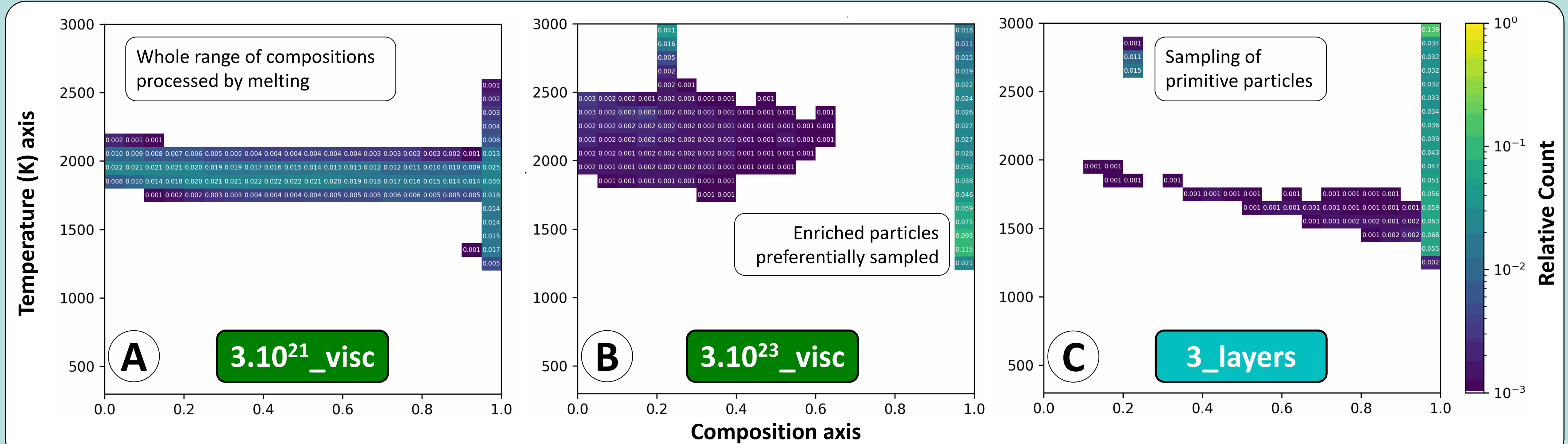


Figure 6: 2D histograms categorizing melting particles by their composition (x axis) and temperature (y axis) for named models at 2 Gyr (similar processing rates). Color scale and bin numbers represent relative count of particles (1 is the sum of all melting particles)

• Surface velocities dictate the renewal rate of material within the melting region. Higher surface velocities, encouraged by lower viscosity, favour higher processing rates (figure 5) and decompression melting (figure 6A).

• However, with a lower reference viscosity, the mantle can get rid of its heat efficiently, the average mantle temperature decreases (figure 4A, 4B), as a result the processing rate decreases accordingly (figure 5).

• By imposing a higher reference viscosity (figure 6B) or a highly viscous lid (3 layers viscosity structure, figure 6C), the mantle can't get rid of its heat as efficiently (figure 4A, 4B). Consequently, the melt production varies drastically and reaches higher values for model 3\_layers (figure 5).

## Less internal heating favours preservation of primitive material

• When the amount of HPEs is reduced (model 05\_HPEs and 02\_HPEs), less heat is provided and the mantle starts cooling earlier in the simulation (figure 4B). This results in less melt production (figure 5) because of a global lower temperature difference between the mantle and the solidus (figure 4A).

• Because of lower processing rates, some primitive particles are cooling so that they can't be processed by melting (figure 7).

• The extraction of the continental crust, thought to happen rather early in Earth history (Hawkesworth *et al*, 2019), could concentrate enough incompatibles and HPEs to promote preservation of primitive particles (figure 7), reproduce more realistic processing rates (figure 5) and increase longevity of piles in the deep mantle (figure 8).

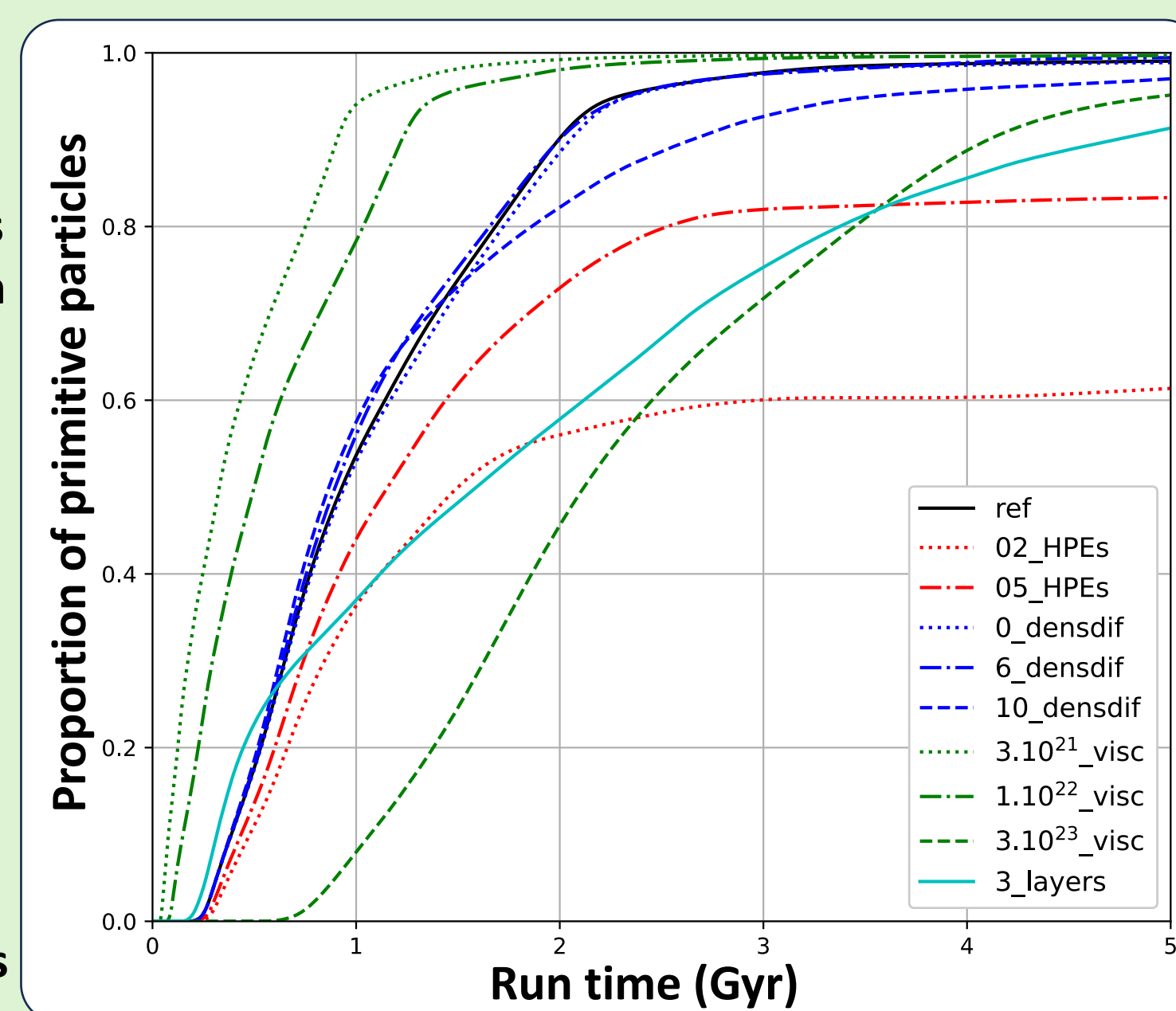


Figure 7: Proportion of primitive particles processed by melting

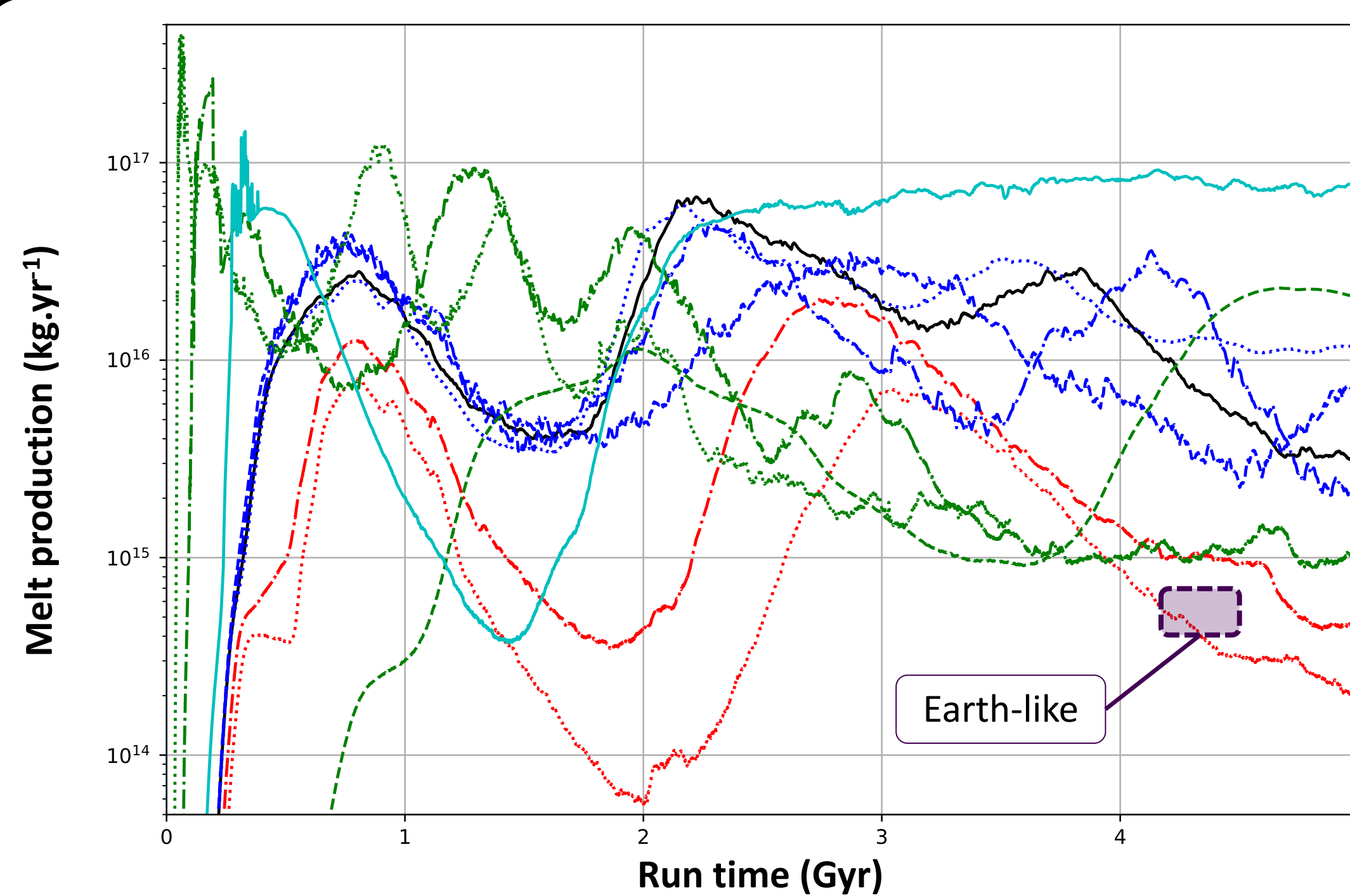


Figure 5: Processing rate in function of time. Current processing rate (box) from J.P. Morgan, 1998

## Intrinsically dense enriched material retain <sup>40</sup>Ar in thermo-chemical piles longer

• When the excess density of enriched material is increased, the residence time of slabs in the deep mantle is prolonged.

• The first generation of slabs is highly enriched in <sup>40</sup>K because they form early (half-life of <sup>40</sup>K is 1.2 Gyr).

• When this first generation of slabs is reprocessed by melting, <sup>40</sup>Ar product of <sup>40</sup>K is released to the atmosphere, leading to an increased outgassing rate (figure 8).

• The higher the density excess of enriched material, the later and the less efficient the outgassing of <sup>40</sup>Ar.

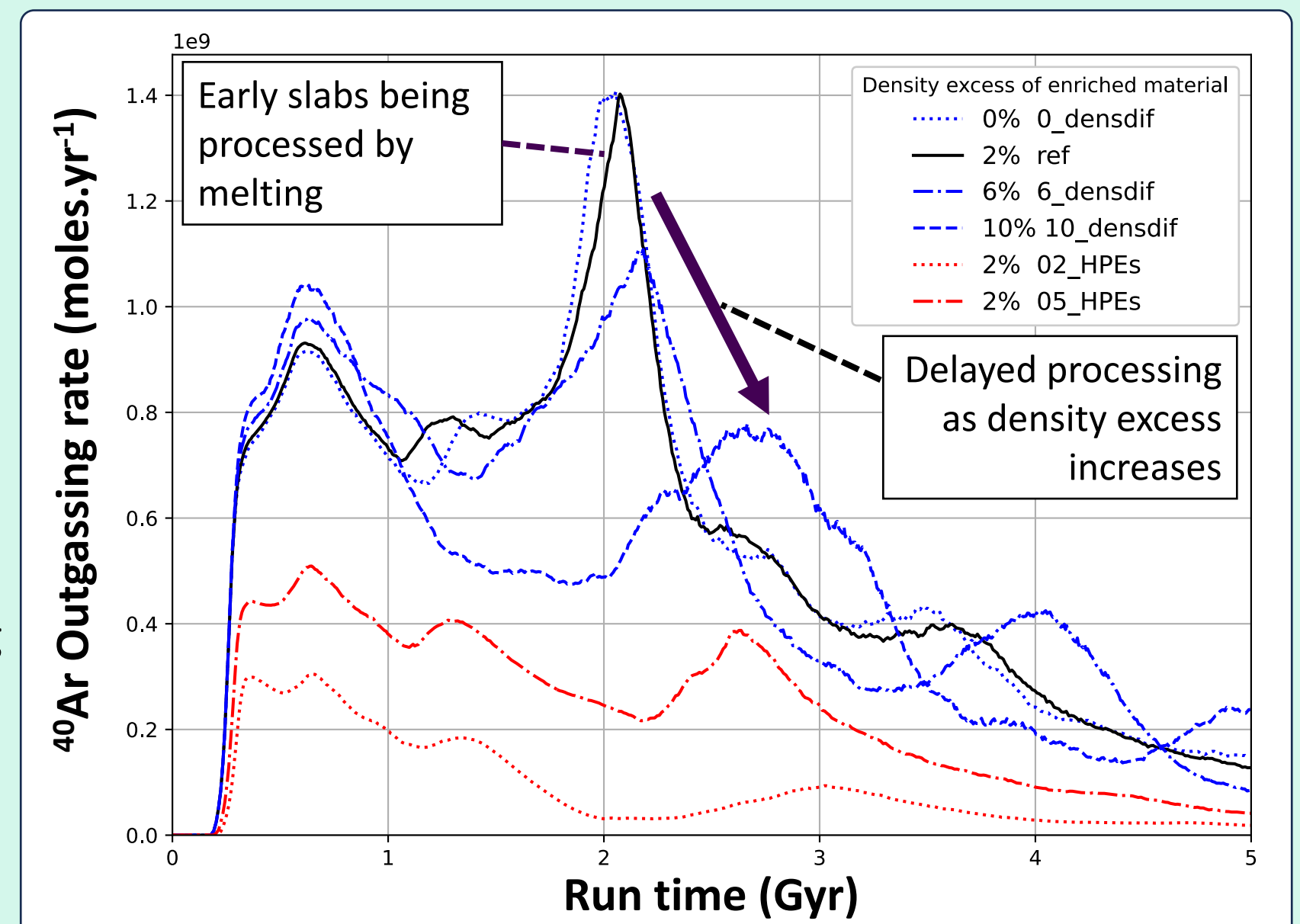


Figure 8: Outgassing rate of <sup>40</sup>Ar

## IV. FUTURE WORK

### Realistic processing rates

- In order to have more realistic processing rates (figure 5), implement a more sophisticated solidus.
- It has also been shown that taking into account the latent heat in melting can reduce the processing rate by half (Li *et al*, 2016)

### Earth-like models

- Circulation models (plate motion history)
- Influence of continental crust extraction on pile stability, <sup>40</sup>Ar outgassing and primitive material preservation

### Other things to look at

- Investigate what type of material thermochemical piles are made of (compositions, ages)

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