

Notes on Current Method Applied to New Grainy Image

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1 Introduction

Using a new test image, we see how robust our method is applied to a different (and most likely more realistic) image. The new image measures 1296 x 966 in size, compared to our old image size of 449 x 321. The result is that our image runs slower, but not linearly so.



Figure 1 I have used a different color table here (Stern Special) because the black and white image was too dim. The fiducials are staggered and can be seen extending off the solar limbs.

2 How Slow is Slow

Table 1 lays out where our code takes the most time. Part of the process of making the code faster will be looking at which routines are called sparsely but still consume a lot of computing time, like `sort()`, for example.

3 Code Flow Chart

Table 1. Time (Total elapsed: 0.23658586 s)

Routine	Times Called	Time Taken	Time Taken	A Number
SORT	64	0.02678	0.02678	1
CONVOL	1	0.02617	0.02617	1
SMOOTH	2	0.01668	0.01668	1
ROTATE	5	0.01565	0.01565	1
SHIFT	8	0.00713	0.00713	1
HISTOGRAM	2	0.00706	0.00706	1
LABEL_REGION	2	0.00641	0.00641	1
ERODE	2	0.00385	0.00385	1
TOTAL	141	0.00318	0.00318	1
DILATE	2	0.00262	0.00262	1
FLOAT	121	0.00200	0.00200	1
WHERE	38	0.00161	0.00161	1
MAX	10	0.00087	0.00087	1
RESOLVE_ROUTINE	1	0.00026	0.00026	1
BYTARR	10	0.00018	0.00018	1
REPLICATE	14	0.00006	0.00006	1
STRTRIM	50	0.00005	0.00005	1
FIX	21	0.00005	0.00005	1
FINDGEN	8	0.00004	0.00004	1
READF	10	0.00003	0.00003	1
PROFILER	1	0.00003	0.00003	1
CREATE_STRUCT	11	0.00003	0.00003	1
ON_ERROR	73	0.00003	0.00003	1
STRMID	34	0.00002	0.00002	1
FILE_LINES	1	0.00002	0.00002	1
GETTOK	2	0.00001	0.00002	0
STRTOK	10	0.00002	0.00002	1
REFORM	39	0.00002	0.00002	1
FLTARR	19	0.00002	0.00002	1
SQRT	62	0.00002	0.00002	1
SCOPE_VARFETCH	12	0.00002	0.00002	1
STRCOMPRESS	12	0.00002	0.00002	1
DOUBLE	10	0.00001	0.00001	1
OPENR	1	0.00001	0.00001	1
N_PARAMS	36	0.00001	0.00001	1
PRINT	1	0.00001	0.00001	1
INDGEN	3	0.00001	0.00001	1
MIN	8	0.00001	0.00001	1
FINITE	12	0.00001	0.00001	1
MESSAGE	1	0.00001	0.00001	1
DEFSYSV	2	0.00001	0.00001	1
STRING	12	0.00001	0.00001	1
STRLEN	30	0.00001	0.00001	1
FREE_LUN	1	0.00001	0.00001	1
CATCH	10	0.00001	0.00001	1
PRODUCT	5	0.00000	0.00000	1
ARRAY_EQUAL	5	0.00000	0.00000	1
BYTE	4	0.00000	0.00000	1
MAKE_ARRAY	2	0.00000	0.00000	1
SYSTIME	2	0.00000	0.00000	1
STRPOS	3	0.00000	0.00000	1
ABS	13	0.00000	0.00000	1
TAG_NAMES	1	0.00000	0.00000	1
ISA	9	0.00000	0.00000	1
SKIP_LUN	1	0.00000	0.00000	1
PTR_FREE	1	0.00000	0.00000	1
PTR_NEW	2	0.00000	0.00000	1
PTRARR	1	0.00000	0.00000	1
STRUPCASE	1	0.00000	0.00000	1
INTARR	1	0.00000	0.00000	1

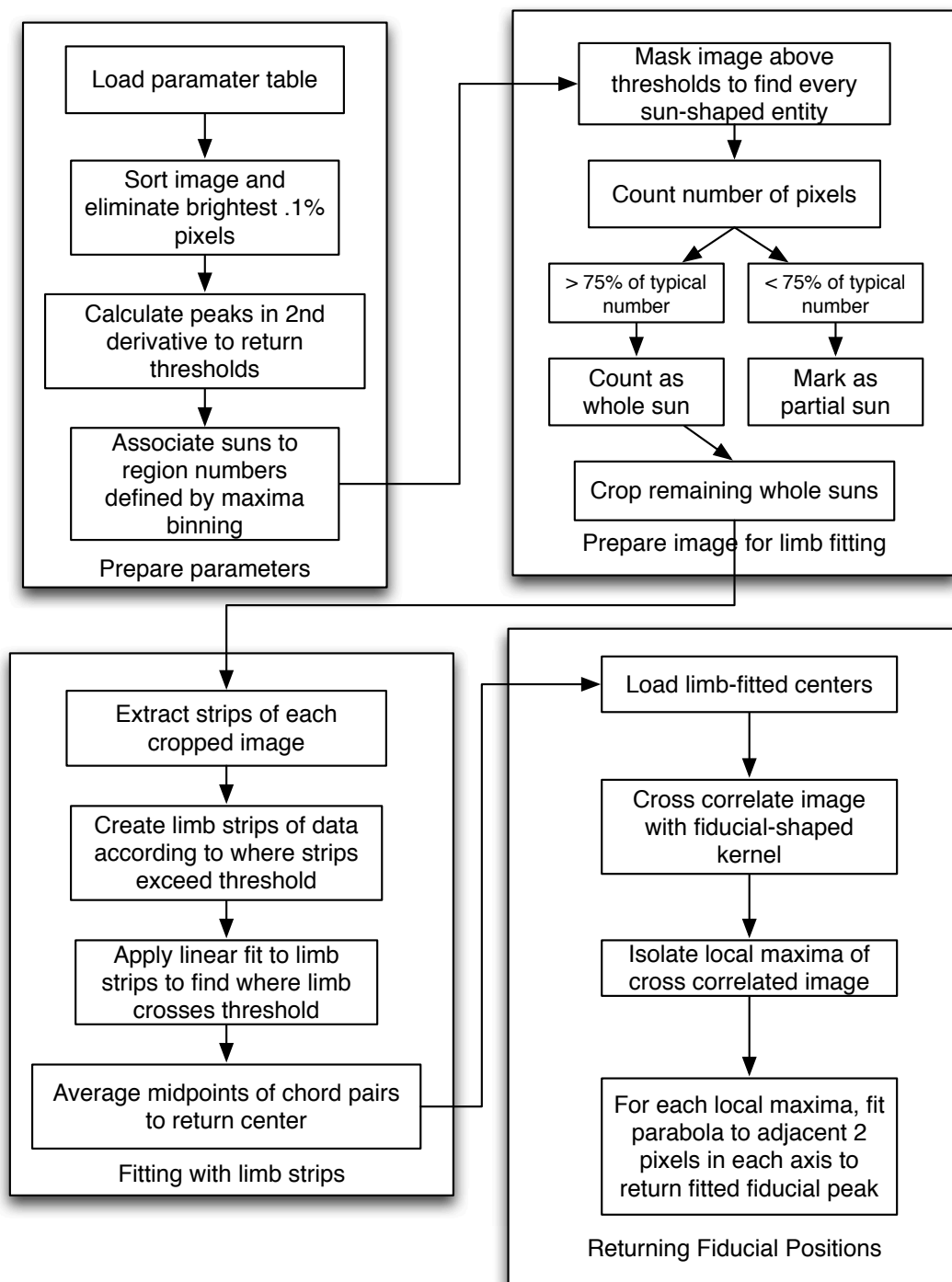


Figure 2

4 The Dreaded Nested For Loop

This is taken straight from Albert's C++ code:

```
1 for (int m = 1; m < correlation.rows-1; m++)
2 {
3     for (int n = 1; n < correlation.cols-1; n++)
4     {
5         thisValue = correlation.at<float>(m,n);
6         if(thisValue > threshold)
7         {
8             //Checks if cross correlated pixel is higher than adjacent pixels
9             if((thisValue > correlation.at<float>(m, n + 1)) &
10                (thisValue > correlation.at<float>(m, n - 1)) &
11                (thisValue > correlation.at<float>(m + 1, n)) &
12                (thisValue > correlation.at<float>(m - 1, n)))
13             {
14                 redundant = false;
15                 for (unsigned int k = 0; k < pixelFiducials.size(); k++)
16                 {
17                     // Checks if previous fiducial correlation values are within 2 fiducial lengths of each other. If so, use
18                     // the one with a higher correlation value
19                     if (abs(pixelFiducials[k].y - m) < fiducialLength*2 &&
20                        abs(pixelFiducials[k].x - n) < fiducialLength*2)
21                     {
22                         redundant = true;
23                         thatValue = correlation.at<float>((int) pixelFiducials[k].y,(int) pixelFiducials[k].x);
24                         Choose the "fiducial" with a higher correlation value
25                         if ( thisValue > thatValue)
26                         {
27                             pixelFiducials[k] = cv::Point2f(n,m);
28                         }
29                         // Break out of this because there should only be one instance of this per run
30                         break;
31                     }
32                 }
33                 // Regardless of whether or not the fiducial was replaced, break out of the loop
34                 if (redundant == true)
35                     continue;
36
37                 // If we're short a few entries for fiducials, extend the array
38                 if ( (int) pixelFiducials.size() < numFiducials)
39                 {
40                     pixelFiducials.add(n, m);
41                 }
42                 else
43                 {
44                     // Dealing with too many fiducials
45                     minValue = std::numeric_limits<float>::infinity();
46                     minIndex = -1;
47                     for (int k = 0; k < numFiducials; k++)
48                     {
49                         if (correlation.at<float>((int) pixelFiducials[k].y,(int) pixelFiducials[k].x)
50                            < minValue)
51                         {
52                             minIndex = k;
53                             minValue = correlation.at<float>((int) pixelFiducials[k].y,(int) pixelFiducials[k].x);
54                         }
55                     }
56                     if (thisValue > minValue)
57                     {
58                         pixelFiducials[minIndex] = cv::Point2f(n, m);
59                     }
60                 }
61             }
62         }
63     }
```

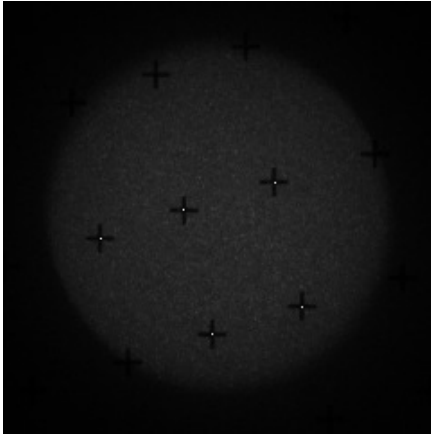
Table 2. Comparison of Fiducial Positions

Fiducial Number	Albert's X	My X	Albert's Y	My Y
0	674.6796	N/A	151.0038	N/A
1	796.3074	N/A	195.0324	N/A
2	740.4443	741.185	210.6342	211.289
3	690.2598	690.985	226.1973	226.961
4	643.4235	644.227	241.8869	242.636
5	755.8672	756.764	279.6622	280.443
6	706.0065	706.809	295.3022	295.957

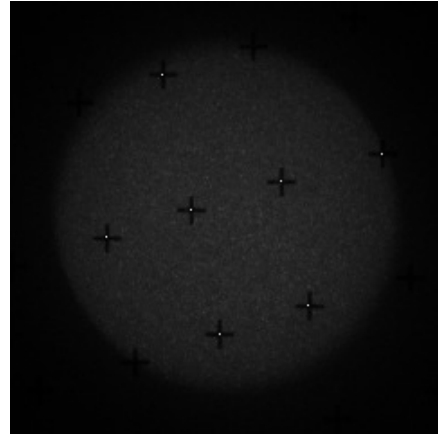
Table 3. Side Crop Test

Amount Cropped from Limb (pixels)	$x_{\text{True}} - x_{\text{Cropped}}$	$y_{\text{True}} - y_{\text{Cropped}}$
10	-1.17771	-0.0108643
20	-4.07970	-0.0376663
30	-7.63260	-0.0522766
40	-11.7287	-0.0585175
50	-16.2043	-0.0185776
60	-20.9117	-0.0872879
70	-25.7588	-0.277687
80	-30.8586	-0.321724
90	-36.1318	-0.318489

5 Comparison to Albert's Code



(a) The fiducials I find



(b) The fiducials Albert finds

Figure 3 My code can't pick up two fiducials due to one or many of the following factors: different kernel, different convolution method, different threshold.

In Table 2 The fiducial positions are typically within 1 pixel of Albert's calculated positions, which is pretty good.

6 Partial Sun Checking

We're motivated to keep some center data regardless of how cut off a sun may be. To do this, we must quantify the poorness of the fit as more sun is cut off. Figures 4 and 5 aim to quantify the worsenings of the evaluated centers. We start by lining up the edge of the image to the solar limb then cropping in 10 columns.

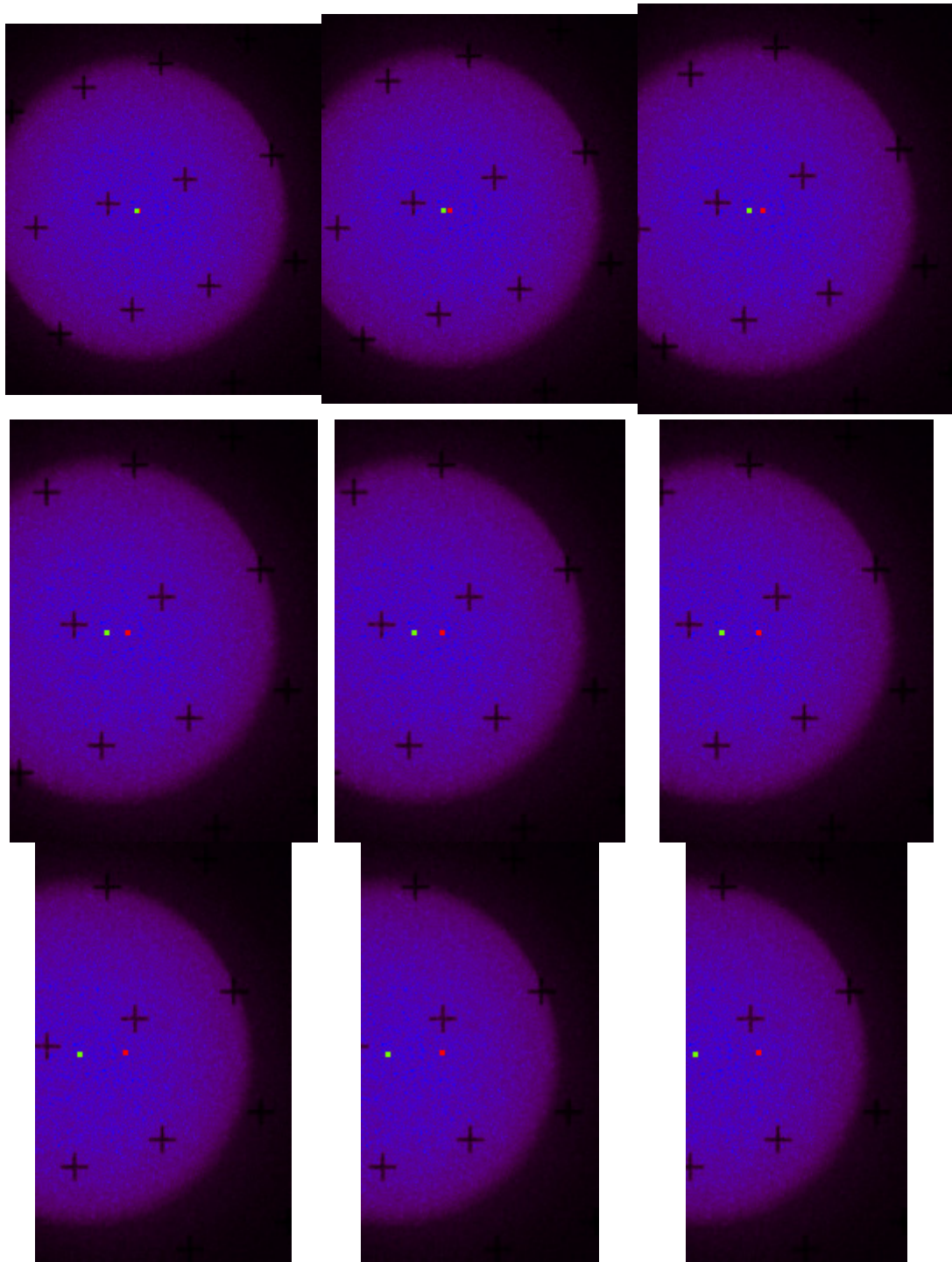


Figure 4 The green pixel is the image's true center and the red pixel is the center of the cropped image. The images are cropped 10 columns at a time.

7 Glaring Problems

I was having trouble with proper thresholding but it was alleviated with increasing the smoothing parameter. *Go parameter block!*

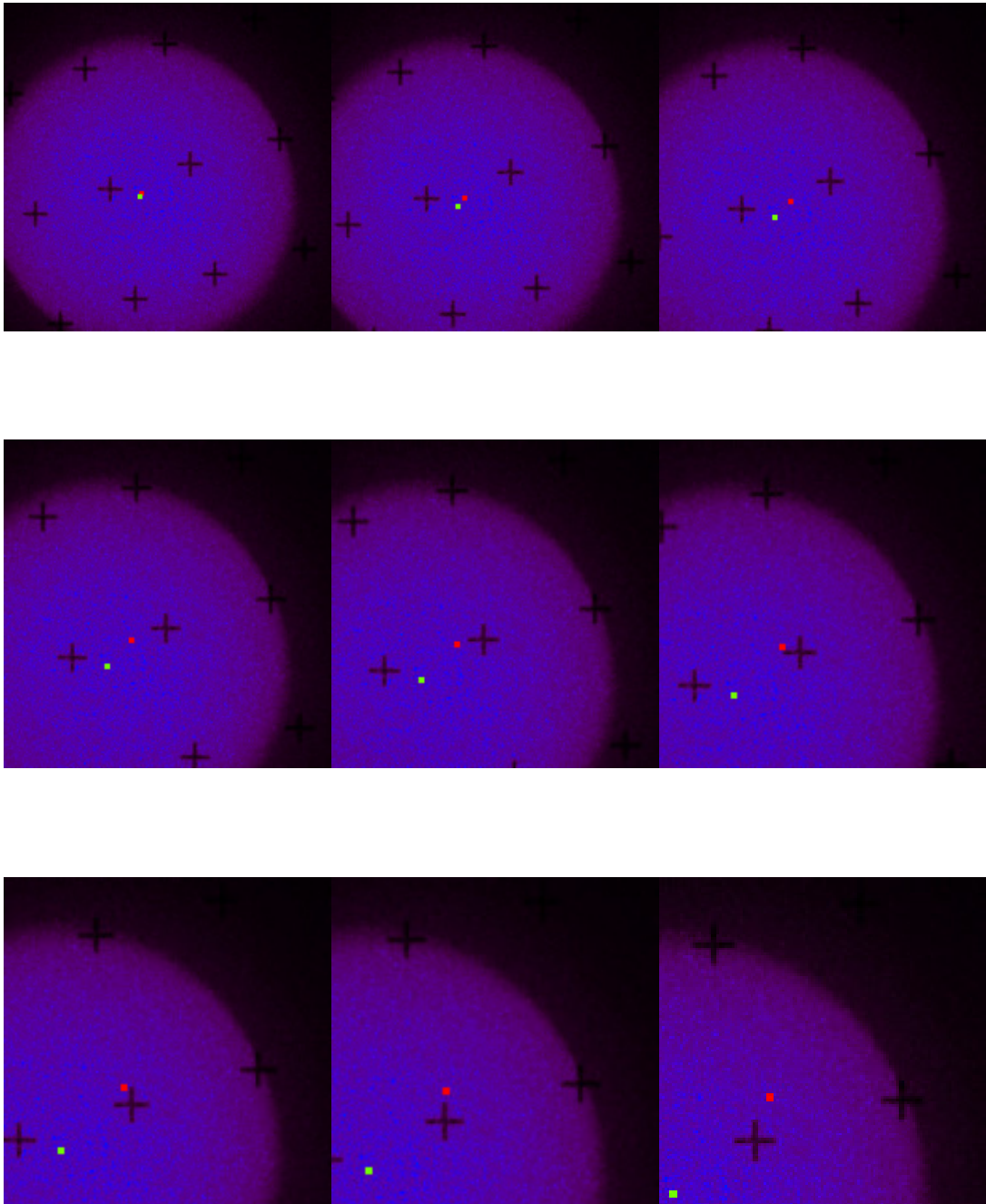


Figure 5 The green pixel is the image's true center and the red pixel is the center of the cropped image. The images are cropped 10 columns at a time.

Table 4. Corner Crop Test

Amount Cropped from Limb (pixels)	$x_{\text{True}} - x_{\text{Cropped}}$	$y_{\text{True}} - y_{\text{Cropped}}$
10	-1.17902	-1.22132
20	-4.23825	-4.28215
30	-8.41805	-8.49775
40	-13.2540	-13.3160
50	-18.2548	-18.0202
60	-23.0267	-22.9181
70	-27.5755	-27.8987
80	-32.1102	-32.4790
90	-36.6139	-37.0980

Table 5. Comparison of Center Positions

Method	X Position	Y Position
Mine	710.811	230.695
Albert's	709.7835	230.1023