



Project II report
WENO schemes for the shallow water equations

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1 Introduction

In this project, Weighted Essentially Non-Oscillatory (WENO) schemes are implemented to solve the shallow water equations:

$$\partial_t \begin{bmatrix} h \\ m \end{bmatrix} + \partial_x \begin{bmatrix} m \\ \frac{m^2}{h} + \frac{1}{2}gh^2 \end{bmatrix} = \mathbf{S}(x, t) \quad (1)$$

where $h(x, t)$ is the depth of the water, $m(x, t)$ is the discharge, g is the gravitational constant and $\mathbf{S}(x, t)^1$ is a source term (vector). This system of hyperbolic PDEs is solved by considering appropriate boundary conditions (BCs) and initial conditions. Note that Eq. 1 can be rewritten in a more compact form as follows:

$$\partial_t \mathbf{q} + \partial_x \mathbf{f}(\mathbf{q}) = \mathbf{S}(x, t) \quad (2)$$

where $\mathbf{q}(x, t) = [h \ m]^T$. In what follows, this differential problem was solved in the spatial domain $\Omega = [0, 2]$ and over the temporal domain $\Omega_t = [0, 1]$. $g = 1$ is assumed throughout the whole project. All the implemented code can be found here on [GitHub](#).

2 Part I - Theoretical framework and motivation

A brief theoretical overview and motivation behind the ENO and WENO schemes are proposed². First of all, the motivation behind ENO schemes is to achieve a higher-order approximation of the solution to a conservation law. Indeed, due to Godunov's barrier theorem ([1]), monotone schemes and linear schemes have the downfall of only first-order convergence. ENO schemes rely on a higher-order reconstruction of the cell values at the interface thanks to polynomial interpolation. Each value at the interface is obtained by selecting the smoothest among all possible stencils that can be used for the polynomial interpolation, where the smoothness is quantified using divided differences. The possibility of choosing adaptively the most suitable stencil makes ENO schemes highly non-linear. Indeed, if the stencil could not be varied, we would simply be using a linear scheme, known for only allowing first-order convergence. Such a method achieves $\mathcal{O}(\Delta x^K)$ accuracy, with $K - 1$ the order of the polynomial reconstruction employed, while the dimension of the stencils used for reconstruction is K .

The motivation behind WENO schemes arises from the inefficient use of the available nodes from ENO schemes. Indeed, ENO schemes only allow for $\mathcal{O}(\Delta x^K)$ approximation in space, while evaluating $2K - 1$ nodes, i.e. we are not making efficient use of the nodes. Besides, as reported in [1], "The nonlinear nature of the stencil selection suggests that even small changes in the solution can result in a different stencil choice. Hence, even for smooth problems, the error behavior can be noisy and uneven". These are the motivations for the extension of ENO schemes to obtain WENO. In this case, all the stencils available for the polynomial interpolation of the interface values are weighed and used for the reconstruction. The weights are chosen adaptively, taking into account the smoothness of the solution, so as to achieve the desired non-linear behavior. In particular, the weights are such that in smooth regions we have $\mathcal{O}(\Delta x^{2K-1})$ convergence ensured, while in non-smooth we have an ENO scheme-like behavior, i.e. the smoothest stencils are preferred for the reconstruction, thus passing to a convergence of $\mathcal{O}(\Delta x^K)$.

3 Part II - About the implementation

The WENO scheme is implemented by exploiting some functions made available in [1]. The main functions are `part2_3.m` and `part2_4.m`, which implement the problem cases described in the project

¹In this context, a bold symbol is used for vectorial quantities.

²For a more rigorous discussion, the reader shall refer to [1].

description. The solution is evolved from time $t = 0$ to $t = T$ by the script `solver.m`. The initial condition is integrated over the cells using the `integral` command in Matlab. Then, at each time step, the solution is integrated in time using the Strong Stability Preserving (SSP) RK3 scheme, that is:

$$\begin{aligned} \mathbf{q}_1 &= \mathbf{q}^n + k \text{RHS}(\mathbf{q}^n, t^n) \\ \mathbf{q}_2 &= \frac{3}{4}\mathbf{q}^n + \frac{1}{4}(\mathbf{q}_1 + k \text{RHS}(\mathbf{q}_1, t^n + k)) \\ \mathbf{q}_3 &= \frac{1}{3}\mathbf{q}^n + \frac{2}{3}(\mathbf{q}_2 + k \text{RHS}(\mathbf{q}_2, t^n + \frac{k}{2})) \\ \mathbf{q}^{n+1} &= \mathbf{q}_3 \end{aligned}$$

Note that this scheme has an order of accuracy equal to 3. The timestep k is computed dynamically at each time according to a CFL condition:

$$k = \text{CFL} \frac{\Delta x}{\max_i(|u_i| + \sqrt{gh_i})} \quad (3)$$

with $\Delta x = \frac{2}{N}$, where N is the number of cells and 2 is the width of the spatial domain Ω . If a source term is present, this is integrated exactly over the cells, to speed up the computations. The RHS that appears in the RK3 scheme is updated using the function `evalRHS`. This function in turn calls the script `WENO.m` to perform the actual reconstruction. In this project, only $K = 2$ is considered, but the implemented code can deal with any value of K . Once the reconstructed values are computed, the chosen numerical flux (Lax-Friedrichs in the present work) is exploited to compute the flux across the cells.

Concerning the boundary conditions, two different scenarios are implemented. The first one is the case of periodic boundary conditions, that is, at each time step, the numerical solution at the first ($j = 1$) and last ($j = N + 1$) node of the spatial domain is updated by introducing fictitious “external nodes” (or ghost cells) to update the solution in the same manner as for the interior nodes. The number of fictitious nodes depends on the dimension of the stencil used to perform the reconstruction. This computation is performed by `apply_bc.m`. A slightly different approach is followed when specifying open boundary conditions: this time, at each time step, the numerical solution at the first ($j = 1$) and last ($j = N + 1$) node of the spatial domain is repeated to the left and right of the domain over fictitious nodes to be able to perform the reconstruction over the first and last cells. The problem is then solved, and the matrices `h` and `m` (containing the solution) are returned as outputs, as well as the discretized space and time vectors (`xc` and `tvec`)³.

4 Part III - Source term

The previously implemented method is tested against a problem for which an exact solution is available. The following smooth functions give the initial conditions for the problem:

$$h(x, 0) = h_0(x) = 1 + 0.5 \sin(\pi x), \quad m(x, 0) = m_0(x) = uh_0(x) \quad (4)$$

where u is the horizontal velocity, constant and equal to 0.25. The source term reads:

$$\mathbf{S}(x, t) = \left[\begin{array}{c} \frac{\pi}{2}(u - 1) \cos(\pi(x - t)) \\ \frac{\pi}{2}(-u + u^2 + gh_0(x - t)) \cos(\pi(x - t)) \end{array} \right] \quad (5)$$

³The reader may refer to Appendix A for insights into the code.

Using the method of characteristics, the exact solution to this problem can be readily computed:

$$h(x, t) = h_0(x - t), \quad m(x, t) = uh(x, t) \quad (6)$$

Regarding the numerical computations, periodic boundary conditions are selected (`bc = 'peri'`), and the time step k is chosen dynamically as described above with $\text{CFL} = 0.5$. The exact and the numerical solutions at the last time $T = 1$ are plotted in Fig. 1. Note that almost no difference between the exact and numerical solution for both $h(x, 1)$ and $m(x, 1)$ is noticeable. Indeed, despite using only $N = 500$ cells for discretizing the spatial domain, the two solutions overlap perfectly. Such an achievement is possible thanks to the high-order reconstruction of the WENO scheme, coupled with the order 3 accuracy of SSP-RK3.

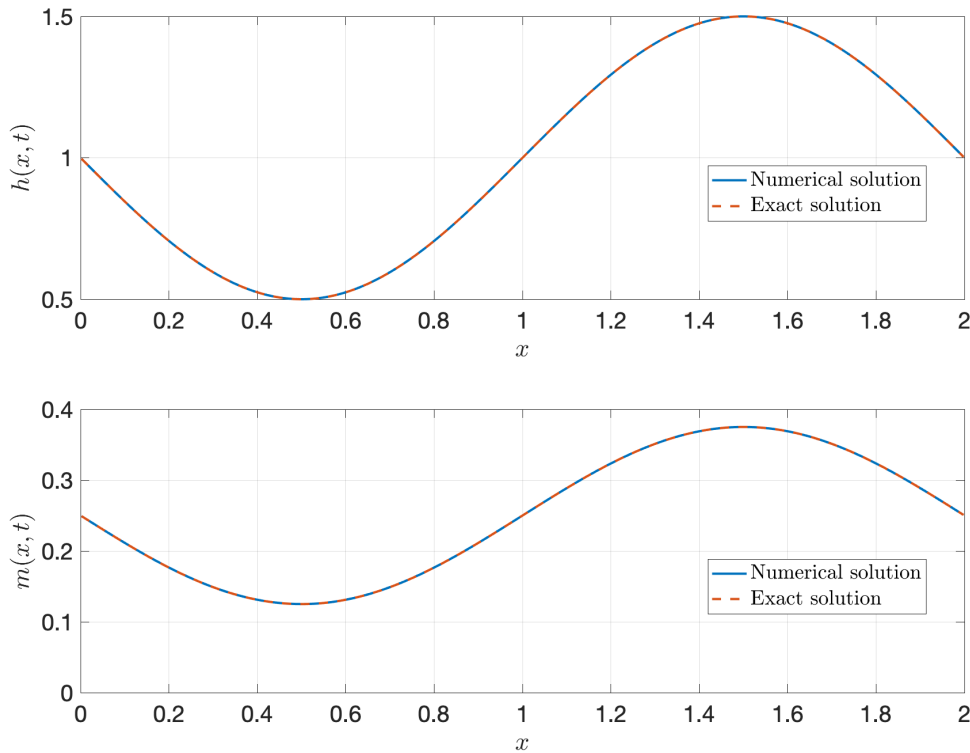


Figure 1: Comparison between exact and numerical solution for $h(x, 1)$ (top) and $m(x, 1)$ (bottom). Any difference between the two curves is barely discernible.

4.1 Error analysis

The order of convergence of the scheme is assessed by solving the problem multiple times varying the value of Δx , such that $\Delta x = 2^i$, $i = -6, \dots, -9$. Then, for each Δx the ℓ_2 norm of the error is computed and stored in the vectors `err_h_vec` and `err_m_vec`. The log-log plot of the error is shown in Fig. 2. In both cases it is possible to note that the solution converges with order 3 in space: this comes as no surprise since the adopted WENO scheme makes use of order-2 polynomial reconstruction of the interface values, i.e. order of convergence $2K - 1 = 3$, and the time-stepping method SSP-RK3 is also of order 3.

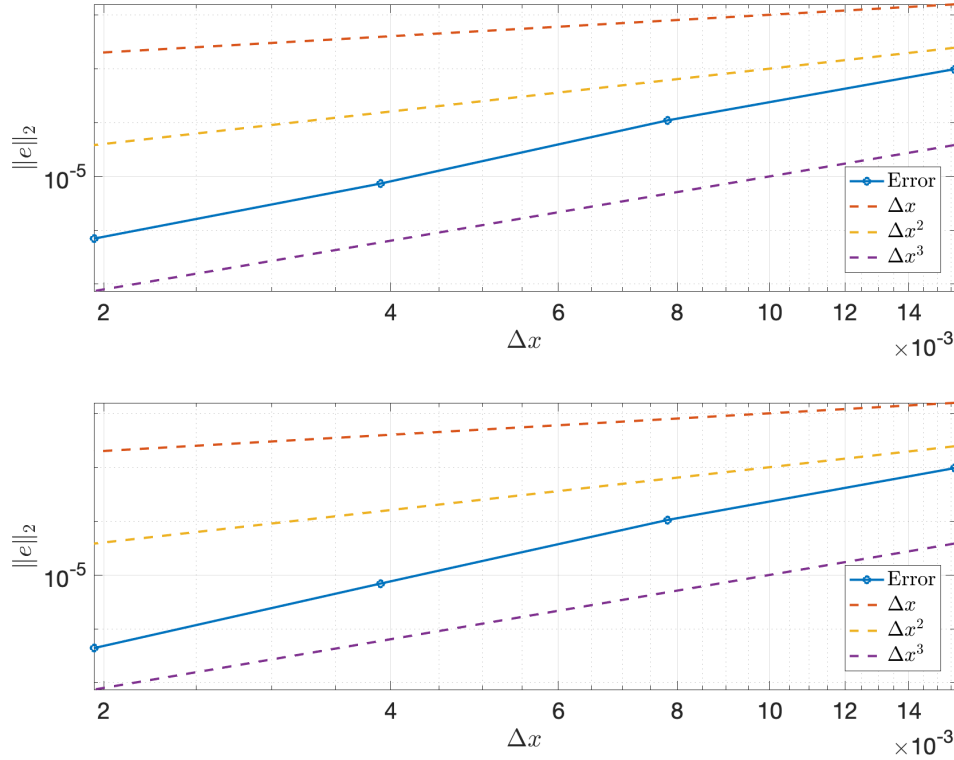


Figure 2: Log-log plot of the ℓ_2 norm of the error between exact and numerical solution at time $T = 1$ for $h(x, 1)$ (top) and $m(x, 1)$ (bottom). Order 1, 2, and 3 convergence rates are also plotted, to allow for comparison.

5 Part IV - Zero source term

The implemented numerical scheme is tested once again, this time with different parameters and no exact solution available. In particular, two sets of initial conditions are considered. The first set reads:

$$h(x, 0) = h_0(x) = 1 - 0.1 \sin(\pi x), \quad m(x, 0) = m_0(x) = 0 \quad (7)$$

while the second set of initial conditions is:

$$h(x, 0) = h_0(x) = 1 - 0.2 \sin(2\pi x), \quad m(x, 0) = m_0(x) = 0.5 \quad (8)$$

In both cases, periodic boundary conditions are considered and the source term is set to zero, i.e. $\mathbf{S} = \mathbf{0}$.

5.1 Numerical results

The results for the first and second set of initial conditions at the last time step $T = 1$ are plotted in Fig. 3 and Fig. 4, respectively.

Again, the numerical solution computed seems to be correctly converging also with a reasonably coarse spatial discretization if compared to the correct numerical solutions available from the previous project. The WENO scheme performs extremely well, even with not very regular solutions, as in the case shown in Fig. 4 for the second set of initial conditions. Note that the second set of initial conditions leads to a discontinuous solution at $T = 1$: this will affect the order of accuracy of the scheme.

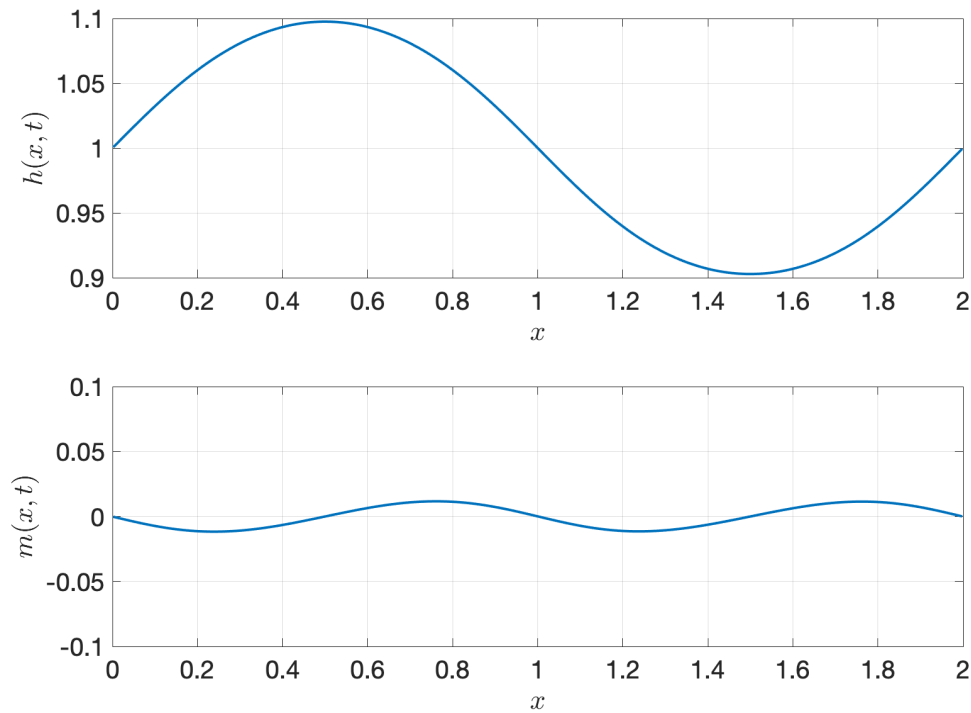


Figure 3: Numerical solution for $h(x, 1)$ (top) and $m(x, 1)$ (bottom) for the first set of initial conditions shown in Eq. 7.

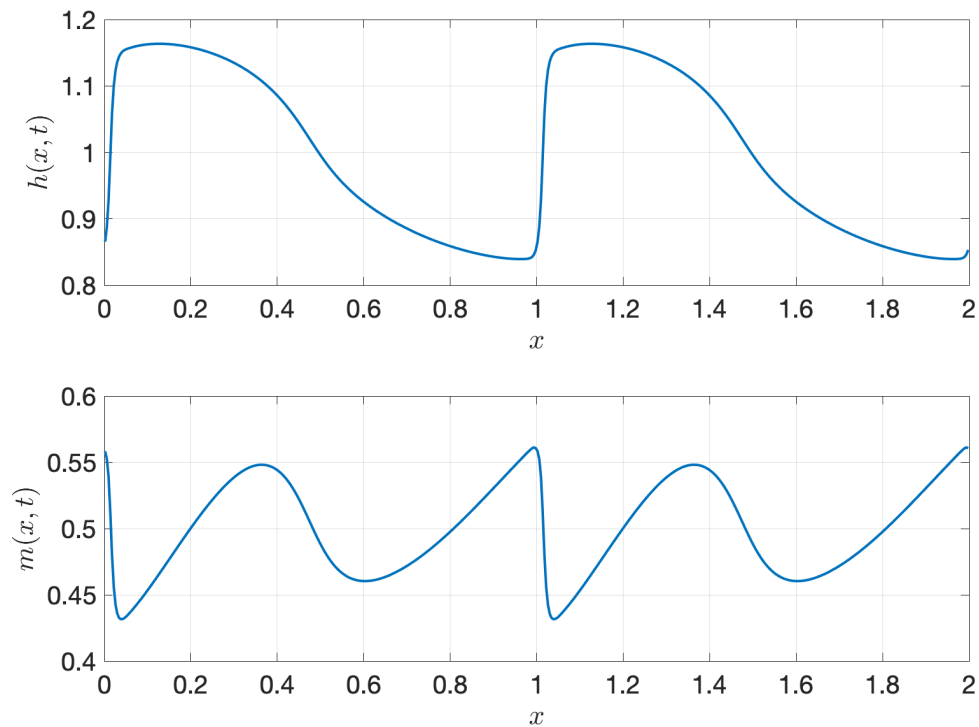


Figure 4: Numerical solution for $h(x, 1)$ (top) and $m(x, 1)$ (bottom) for the second set of initial conditions shown in Eq. 8.

5.2 Error analysis

As previously mentioned, since no exact solution is available for the case at hand, a slightly different approach for the error analysis is needed. In this case, the order of convergence of the scheme is assessed by solving the given problem multiple times varying the value of Δx , such that $\Delta x = 2^i$, $i = -6, \dots, -9$ again. Then, for each Δx , the ℓ_2 norm of the error is computed, using two numerical solutions: the one obtained with the chosen Δx , and, as reference solution, the numerical solution computed on a distinctly finer mesh with $N = 1500$. Similarly to the previous case, once the error is computed, it is stored in the vectors `err_h_vec1`, `err_m_vec1` for the first set of initial conditions, and `err_h_vec2`, `err_m_vec2` for the second one. The log-log plot of such a result is shown in Fig. 5 and Fig. 6, for the first and second set of initial conditions, respectively.

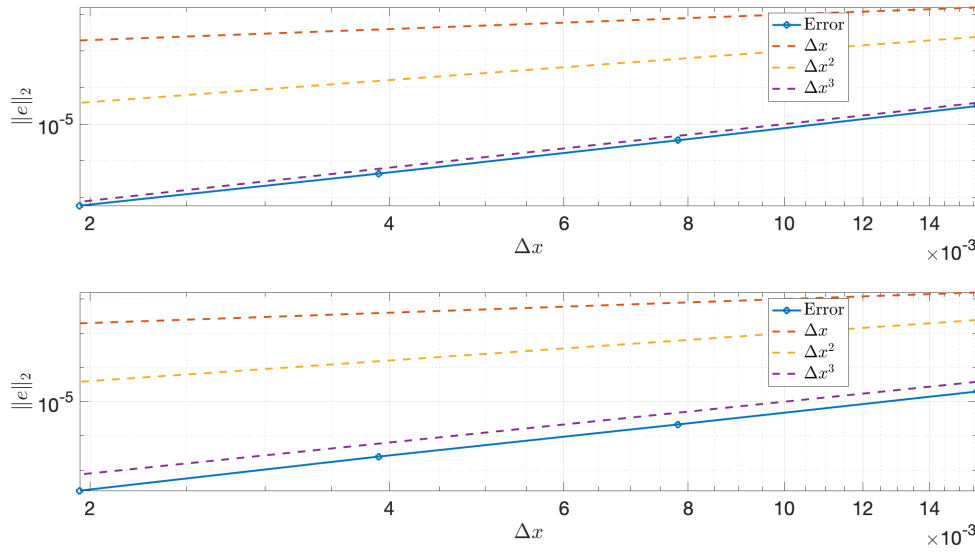


Figure 5: Log-log plot of the ℓ_2 norm of the error between reference and numerical solution at time $t = 1$ for $h(x, 1)$ (top) and $m(x, 1)$ (bottom), for the first set of initial conditions, shown in Eq. 7. Order 1, 2, and 3 convergence rates are also plotted, to ease the comparison.

In the first case, it can be noted that the order of convergence of the numerical solution is third order in space. On the other hand, the order of convergence for the second set of initial conditions, while being initially parallel to Δx , becomes then parallel to Δx^2 for smaller values of Δx . This can be explained because, in the presence of a discontinuity, WENO schemes reduce to ENO schemes, that is the accuracy of the scheme changes abruptly to $\mathcal{O}(\Delta x^K)$. We hence expect to reach second-order accuracy. Further work should focus on further reducing the experimented Δx to see a slope parallel to Δx^2 and confirm the theoretical results. Fig. 7 shows the reference solution that was employed for the error computation.

6 Conclusion

In the present project, WENO schemes were implemented for the solution of a system of hyperbolic PDEs. It was possible to show that, given the order of the polynomial for the reconstruction $K - 1$, WENO schemes reach an accuracy of $2K - 1$. This is true as long as no discontinuities appear in the solution: in such a case, WENO schemes reduce to ENO schemes, yielding an accuracy of order K , as it was proved empirically.

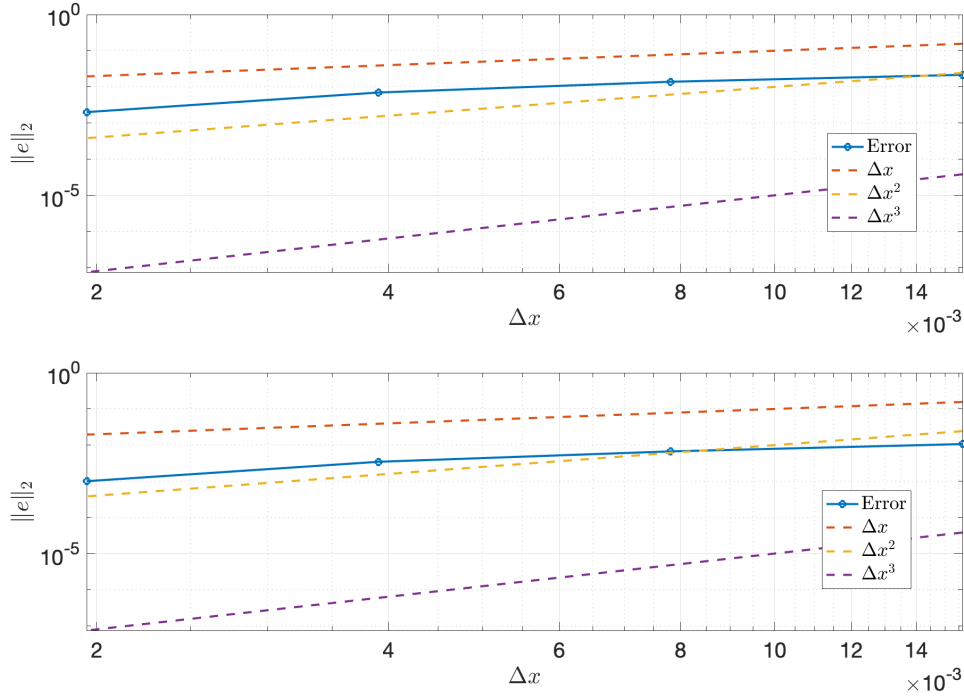


Figure 6: Log-log plot of the ℓ_2 norm of the error between reference and numerical solution at time $t = 1$ for $h(x, 1)$ (top) and $m(x, 1)$ (bottom), for the second set of initial conditions, shown in Eq. 8. Order 1, 2, and 3 convergence rates are also plotted, to ease the comparison.

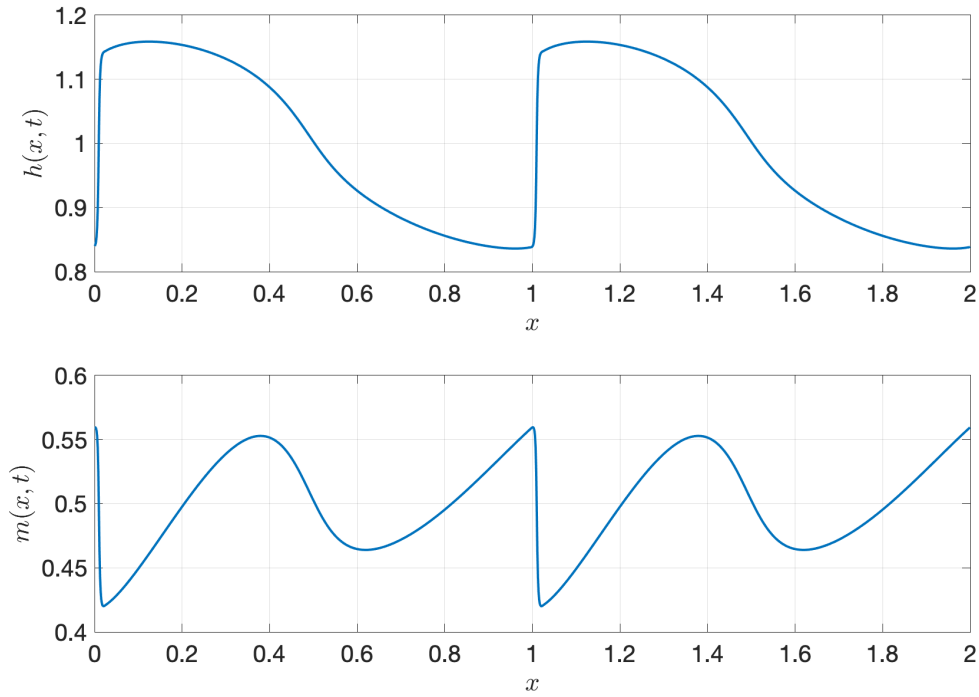


Figure 7: Numerical solution for $h(x, 1)$ (top) and $m(x, 1)$ (bottom) for the second set of initial conditions shown in Eq. 8, using a highly refined mesh featuring $N = 1500$ intervals.

References

- [1] Jan S. Hesthaven. *Numerical Methods for Conservation Laws*. Society for Industrial and Applied Mathematics, Philadelphia, PA, 2018. doi: 10.1137/1.9781611975109. URL <https://epubs.siam.org/doi/abs/10.1137/1.9781611975109>.

A MATLAB code

A.1 Part2_3

```
1 clear
2 close all
3 clc
4
5 %%% Code by Francesco Sala and Nicolo' Viscusi %%%
6
7 % Set to true if you want to see the animation of the solutions over
   time
8 animation = "True";
9
10 %% Resolution of the problem
11
12 % First problem (for integration of source term in space)
13 PROBLEM = 1;
14
15 % Definition of parameters
16 g = 1;
17 u = 0.25;
18
19 % Spatial domain
20 xspan = [0, 2];
21
22 % Temporal domain
23 tspan = [0, 1];
24
25 % Initial conditions
26 h0 = @(x) 1 + 0.5 * sin(pi * x);
27 m0 = @(x) u * h0(x);
28
29 % Number of grid points
30 N = 500;
31
32 % CFL condition
33 CFL = 0.5;
34
35 % Here we use periodic boundary condition as the option ('peri')
36 bc = 'peri';
37
38 % Choose order for WENO reconstruction
39 k = 2;
40
41 % Solve the problem
42 [h, m, xc, tvec] = solver(xspan, tspan, N, ...
43     CFL, g, h0, m0, @LaxFriedrichs, @flux_phys, bc, k, PROBLEM);
44
```

```

45 % We visualize the solution
46 if animation == "True"
47     figure(1)
48     for i = 1 : 1 : length(tvec)
49
50         subplot(2, 1, 1)
51         plot(xc, h(:, i), 'LineWidth', 2)
52         hold on
53         plot(xc, h0(xc - tvec(i)), '--', 'Linewidth', 2)
54         title(['$h(x, t)$ at $t = $ ', num2str(tvec(i))], ...
55             'Interpreter', 'latex')
56         xlabel('$x$', 'Interpreter', 'latex')
57         ylabel('$h(x, t)$', 'Interpreter', 'latex')
58         grid on
59         xlim([0 2]);
60         ylim([0.5 1.5]);
61         hold off
62         legend('Numerical solution', 'Exact solution', ...
63             'Interpreter', 'latex', 'Location', 'best')
64         set(gca, 'FontSize', 20)
65         drawnow
66
67         subplot(2, 1, 2)
68         plot(xc, m(:, i), 'LineWidth', 2)
69         hold on
70         plot(xc, u * h0(xc - tvec(i)), '--', 'Linewidth', 2)
71         title(['$m(x, t)$ at $t = $ ', num2str(tvec(i))], ...
72             'Interpreter', 'latex')
73         xlabel('$x$', 'Interpreter', 'latex')
74         ylabel('$m(x, t)$', 'Interpreter', 'latex')
75         grid on
76         xlim([0 2]);
77         ylim([0 0.4]);
78         hold off
79         legend('Numerical solution', 'Exact solution', ...
80             'Interpreter', 'latex', 'Location', 'best')
81         set(gca, 'FontSize', 20)
82         drawnow
83
84     end
85 end
86
87
88
89 %% Error analysis
90
91 % We solve the same problem for different values of \Delta x
92 delta_x_vec = 2.^-(6:9);
93

```

```

94 N_vec = (xspan(2) - xspan(1)) ./ delta_x_vec ;
95 err_h_vec = zeros(size(N_vec));
96 err_m_vec = zeros(size(N_vec));
97
98 for i=1:length(N_vec)
99     N = N_vec(i);
100
101     [h, m, xc, tvec] = solver(xspan, tspan, N, ...
102         CFL, g, h0, m0, @LaxFriedrichs, @flux_phys, bc, k, PROBLEM);
103
104     err_h_vec(i) = 1/sqrt(N) * norm(h(:, end) - h0(xc-tspan(2))');
105     err_m_vec(i) = 1/sqrt(N) * norm(m(:, end) - u*h0(xc-tspan(2))');
106 end
107
108
109 % Plot the error in loglog
110 figure(2)
111
112 subplot(2,1,1)
113 loglog(delta_x_vec, err_h_vec, "o-", "Linewidth", 2)
114 hold on
115 loglog(delta_x_vec, delta_x_vec, "--", "Linewidth", 2)
116 loglog(delta_x_vec, 10 * delta_x_vec.^2, "--", "Linewidth", 2)
117 loglog(delta_x_vec, 10 * delta_x_vec.^3, "--", "Linewidth", 2)
118 xlabel('$\Delta x$', 'Interpreter', 'latex')
119 ylabel("$|e|_2$", "Interpreter", "latex")
120 title("Error on \$(h(x,t))\$ at \$(t=1)\$", "Interpreter", "latex")
121 legend("Error", "\$(\Delta x)\$", "\$(\Delta x^2)\$", "\$(\Delta x^3)\$", ...
122     "interpreter", "latex", "location", "best")
123 set(gca, 'FontSize', 20)
124 grid on
125
126
127 subplot(2,1,2)
128 loglog(delta_x_vec, err_m_vec, "o-", "Linewidth", 2)
129 hold on
130 loglog(delta_x_vec, delta_x_vec, "--", "Linewidth", 2)
131 loglog(delta_x_vec, 10 * delta_x_vec.^2, "--", "Linewidth", 2)
132 loglog(delta_x_vec, 10 * delta_x_vec.^3, "--", "Linewidth", 2)
133 xlabel('$\Delta x$', 'Interpreter', 'latex')
134 ylabel("$|e|_2$", "Interpreter", "latex")
135 title("Error on \$(m(x,t))\$ at \$(t=1)\$", "Interpreter", "latex")
136 legend("Error", "\$(\Delta x)\$", "\$(\Delta x^2)\$", "\$(\Delta x^3)\$", ...
137     "interpreter", "latex", "location", "best")
138 grid on
139 set(gca, 'FontSize', 20)

```

A.2 Part2_4

```

1 clear
2 close all
3 clc
4
5 %%% Code by Francesco Sala and Nicolò Viscusi %%%
6
7 % Set to true if you want to see the animation of the solutions over
   time
8 animation = "True";
9
10 %% Resolution of the problem (first set of initial conditions)
11
12 % Second problem
13 PROBLEM = 2;
14
15 % Definition of parameters
16 g = 1;
17 u = 0.25;
18
19 % Spatial domain
20 xspan = [0, 2];
21
22 % Temporal domain
23 tspan = [0, 1];
24
25 % Initial conditions
26 h01 = @(x) 1 - 0.1 * sin(pi * x);
27 m01 = @(x) 0 * x;
28
29 % Number of grid points
30 N = 500;
31
32 % Number of time steps
33 CFL = 0.5;
34
35 % Here we use periodic boundary condition as the option ('peri')
36 bc = 'peri';
37
38 % Choose order for WENO reconstruction
39 k = 2;
40
41 % Solve the problem
42 [h1, m1, xc1, tvec1] = solver(xspan, tspan, N, ...
43     CFL, g, h01, m01, @LaxFriedrichs, @flux_phys, bc, k, PROBLEM);
44
45 % We visualize the solution
46 if animation == "True"
47

```

```

48     figure(1)
49
50     for i = 1 : 1 : length(tvec1)
51
52         subplot(2, 1, 1)
53         plot(xc1, h1(:, i), 'LineWidth', 2)
54         title(['$h(x, t)$ at $t = $ ', num2str(tvec1(i))], ...
55             'Interpreter', 'latex')
56         xlabel('$x$', 'Interpreter', 'latex')
57         ylabel('$h(x, t)$', 'Interpreter', 'latex')
58         grid on
59         xlim([0 2]);
60         ylim([0.9 1.1]);
61         set(gca, 'FontSize', 20)
62         drawnow
63
64         subplot(2, 1, 2)
65         plot(xc1, m1(:, i), 'LineWidth', 2)
66         title(['$m(x, t)$ at $t = $ ', num2str(tvec1(i))], ...
67             'Interpreter', 'latex')
68         xlabel('$x$', 'Interpreter', 'latex')
69         ylabel('$m(x, t)$', 'Interpreter', 'latex')
70         grid on
71         xlim([0 2]);
72         ylim([-0.1 0.1])
73         set(gca, 'FontSize', 20)
74         drawnow
75
76     end
77
78 end
79
80
81
82 %% Error analysis (first set of initial conditions)
83
84 % Generate a reference solution using sufficiently fine mesh
85 [h1_ex, m1_ex, xvec1_ex, ~] = solver(xspan, tspan, 1500, ...
86     CFL, g, h01, m01, @LaxFriedrichs, @flux_phys, bc, k, PROBLEM);
87
88 % We solve the same problem for different values of \Delta x
89 delta_x_vec = 2.^-(6:9);
90
91 N_vec = (xspan(2) - xspan(1)) ./ delta_x_vec;
92 err_h_vec1 = zeros(size(N_vec));
93 err_m_vec1 = zeros(size(N_vec));
94
95 for i = 1 : length(N_vec)
96

```

```

97     N = N_vec(i);
98
99     [h1, m1, xc1, ~] = solver(xspan, tspan, N, ...
100         CFL, g, h01, m01, @LaxFriedrichs, @flux_phys, bc, k, PROBLEM);
101
102     % We now want to compare h1(:, end) with h1_ex(:, end),
103     % but this second vector is defined on a different grid xvec1_ex
104     % We interpolate h1_ex(:, end) on the grid xc1
105     h1ex_interp = interp1(xvec1_ex, h1_ex(:, end), xc1);
106     m1ex_interp = interp1(xvec1_ex, m1_ex(:, end), xc1);
107
108     % Compute norm 2 of the error
109     err_h_vec1(i) = 1/sqrt(N) * norm(h1(:, end)' - h1ex_interp);
110     err_m_vec1(i) = 1/sqrt(N) * norm(m1(:, end)' - m1ex_interp);
111
112 end
113
114
115 % Plot the error in loglog
116 figure(2)
117
118 subplot(2,1,1)
119 loglog(delta_x_vec, err_h_vec1, "o-", "Linewidth", 2)
120 hold on
121 loglog(delta_x_vec, delta_x_vec, "--", "Linewidth", 2)
122 loglog(delta_x_vec, 10 * delta_x_vec.^2, "--", "Linewidth", 2)
123 loglog(delta_x_vec, 10 * delta_x_vec.^3, "--", "Linewidth", 2)
124 xlabel('$\Delta x$', 'Interpreter', 'latex')
125 ylabel("$|e|_2$", "Interpreter", "latex")
126 title("Error on \h(x,t)\ at \t=1\)", "Interpreter", "latex")
127 legend("Error", "\(\Delta x\)", "\(\Delta x^2\)", "\(\Delta x^3\)", ...
128     "interpreter", "latex", "location", "best")
129 set(gca, 'FontSize', 20)
130 grid on
131
132
133 subplot(2,1,2)
134 loglog(delta_x_vec, err_m_vec1, "o-", "Linewidth", 2)
135 hold on
136 loglog(delta_x_vec, delta_x_vec, "--", "Linewidth", 2)
137 loglog(delta_x_vec, 10 * delta_x_vec.^2, "--", "Linewidth", 2)
138 loglog(delta_x_vec, 10 * delta_x_vec.^3, "--", "Linewidth", 2)
139 xlabel('$\Delta x$', 'Interpreter', 'latex')
140 ylabel("$|e|_2$", "Interpreter", "latex")
141 title("Error on \m(x,t)\ at \t=1\)", "Interpreter", "latex")
142 legend("Error", "\(\Delta x\)", "\(\Delta x^2\)", "\(\Delta x^3\)", ...
143     "interpreter", "latex", "location", "best")
144 grid on
145 set(gca, 'FontSize', 20)

```



```

146
147
148
149 %% Resolution of the problem (second set of initial conditions)
150
151 % Initial conditions
152 h02 = @(x) 1 - 0.2 * sin(2 * pi * x);
153 m02 = @(x) 0.5 + 0*x;
154
155 % Number of grid points
156 N = 500;
157
158 % Solve the problem
159 [h2, m2, xc2, tvec2] = solver(xspan, tspan, N, ...
160     CFL, g, h02, m02, @LaxFriedrichs, @flux_phys, bc, k, PROBLEM);
161
162 % We visualize the solution
163 if animation == "True"
164
165     figure(3)
166
167     for i = 1 : 1 : length(tvec2)
168
169         subplot(2, 1, 1)
170         plot(xc2, h2(:, i), 'LineWidth', 2)
171         title(['$h(x, t)$ at $t = $ ', num2str(tvec2(i))], ...
172             'Interpreter', 'latex')
173         xlabel('$x$', 'Interpreter', 'latex')
174         ylabel('$h(x, t)$', 'Interpreter', 'latex')
175         grid on
176         xlim([0 2]);
177         ylim([0.8 1.2]);
178         set(gca, 'FontSize', 20)
179         drawnow
180
181         subplot(2, 1, 2)
182         plot(xc2, m2(:, i), 'LineWidth', 2)
183         title(['$m(x, t)$ at $t = $ ', num2str(tvec2(i))], ...
184             'Interpreter', 'latex')
185         xlabel('$x$', 'Interpreter', 'latex')
186         ylabel('$m(x, t)$', 'Interpreter', 'latex')
187         grid on
188         xlim([0 2]);
189         ylim([0.4 0.6])
190         set(gca, 'FontSize', 20)
191         drawnow
192
193     end
194

```

```

195 end
196
197
198
199 %% Error analysis (second set of initial conditions)
200
201 % Generate a reference solution
202 [h2_ex, m2_ex, xvec2_ex, ~] = solver(xspan, tspan, 1500, ...
203     CFL, g, h02, m02, @LaxFriedrichs, @flux_phys, bc, k, PROBLEM);
204
205 % We solve the same problem for different values of \Delta x
206 delta_x_vec = 2.^(6:9);
207
208 N_vec = (xspan(2) - xspan(1)) ./ delta_x_vec;
209 err_h_vec2 = zeros(size(N_vec));
210 err_m_vec2 = zeros(size(N_vec));
211
212 for i = 1 : length(N_vec)
213
214     N = N_vec(i);
215
216     [h2, m2, xc2, ~] = solver(xspan, tspan, N, ...
217         CFL, g, h02, m02, @LaxFriedrichs, @flux_phys, bc, k, PROBLEM);
218
219     % We now want to compare h1(:, end) with h1_ex(:, end),
220     % but this second vector is defined on a different grid xvec1_ex
221     % We interpolate h1_ex(:, end) on the grid xc2
222     h2ex_interp = interp1(xvec2_ex, h2_ex(:, end), xc2);
223     m2ex_interp = interp1(xvec2_ex, m2_ex(:, end), xc2);
224
225     err_h_vec2(i) = 1/sqrt(N) * norm(h2(:, end)' - h2ex_interp);
226     err_m_vec2(i) = 1/sqrt(N) * norm(m2(:, end)' - m2ex_interp);
227
228 end
229
230
231 % Plot the error
232 figure(4)
233
234 subplot(2,1,1)
235 loglog(delta_x_vec, err_h_vec2, "o-", "Linewidth", 2)
236 hold on
237 loglog(delta_x_vec, 10 * delta_x_vec, "--", "Linewidth", 2)
238 loglog(delta_x_vec, 100 * delta_x_vec.^2, "--", "Linewidth", 2)
239 loglog(delta_x_vec, 10 * delta_x_vec.^3, "--", "Linewidth", 2)
240 xlabel('$\Delta x$', 'Interpreter', 'latex')
241 ylabel('$||e||_2$', "Interpreter", "latex")
242 title("Error on \((h(x,t))\) at \((t=1)\)", "Interpreter", "latex")
243 legend("Error", "\(\Delta x\)", "\(\Delta x^2\)", "\(\Delta x^3\)", ...

```

```

244     "interpreter", "latex", "location", "best")
245 set(gca, 'Fontsize', 20)
246 grid on
247
248
249 subplot(2,1,2)
250 loglog(delta_x_vec, err_m_vec2, "o-", "Linewidth", 2)
251 hold on
252 loglog(delta_x_vec, 10 * delta_x_vec, "--", "Linewidth", 2)
253 loglog(delta_x_vec, 100 * delta_x_vec.^2, "--", "Linewidth", 2)
254 loglog(delta_x_vec, 10 * delta_x_vec.^3, "--", "Linewidth", 2)
255 xlabel('$\Delta x$', 'Interpreter', 'latex')
256 ylabel("$|e|_{2}$", "Interpreter", "latex")
257 title("Error on \$(m(x,t))\$ at \$(t=1)\$", "Interpreter", "latex")
258 legend("Error", "\$(\Delta x)\$", "\$(\Delta x^2)\$", "\$(\Delta x^3)\$", ...
259     "interpreter", "latex", "location", "best")
260 grid on
261 set(gca, 'Fontsize', 20)

```

A.3 Integrate the problem in time

```

1  function [h, m, xc, tvec] = solver(xspan, tspan, N, CFL, g, h0, m0, ...
2      flux, flux_phys, bc, k, PROBLEM)
3  % SOLVER - Solve a hyperbolic conservation law using the SSP-RK3 scheme
4  %
5  % [h, m, xc, tvec] = SOLVER(xspan, tspan, N, CFL, g, h0, m0, ...
6  %                          flux, flux_phys, bc, k, PROBLEM)
7  % solves a hyperbolic conservation law using the SSP-RK3 scheme
8  % and returns the solutions for water height 'h', discharge 'm',
9  % spatial grid locations 'xc' and the time vector 'tvec'.
10 %
11 % Input:
12 %     xspan - Spatial domain [x_start, x_end].
13 %     tspan - Time domain [t_start, t_end].
14 %     N     - Number of spatial grid points.
15 %     CFL   - Courant-Friedrichs-Lewy (CFL) number for adaptive
16 %            time-stepping.
17 %     g     - Gravitational constant (g=1).
18 %     h0    - Function handle for the initial water height
19 %            condition.
20 %     m0    - Function handle for the initial discharge condition
21 %
22 %     flux   - Flux function for the conservative variables.
23 %     flux_phys - Flux function in physical variables.
24 %     bc     - String specifying the type of boundary condition.
25 %            Supported values: 'peri' (periodic), 'open'
26 %            (open boundary).
27 %     k     - Order of accuracy for WENO reconstruction.

```

```

27 %          PROBLEM      - Integer specifying the problem type.
28 %                          Supported values: 1 or 2.
29 %
30 %      Output:
31 %          h            - Matrix of water height solutions over time.
32 %          m            - Matrix of discharge solutions over time.
33 %          xc           - Spatial grid locations.
34 %          tvec         - Time vector.
35 %
36 % Authors: [Francesco Sala, Nicolo' Viscusi]
37 % January 2024
38
39
40 % Define preliminary variables
41 dx      = (xspan(2) - xspan(1)) / N;
42 xc      = (xspan(1) + 0.5 * dx) : dx : (xspan(2) - 0.5 * dx);
43 xf      = linspace(xspan(1), xspan(2), N + 1);
44 h       = zeros(N,1);
45 m       = zeros(N,1);
46 tvec    = 0;
47
48 % We compute cell-averages of the initial condition
49 for j = 1 : N
50     h(j) = integral(h0, xf(j), xf(j+1), 'AbsTol', 1e-14) / dx;
51     m(j) = integral(m0, xf(j), xf(j+1), 'AbsTol', 1e-14) / dx;
52 end
53
54 % Initialize polynomial coefficients (of degree k-1)
55 Crec = zeros(k + 1, k);
56 for r=-1:k-1
57     Crec(r+2,:) = ReconstructWeights(k,r);
58 end
59
60 % Initialize linear weights
61 dw = LinearWeights(k,0);
62
63 % Compute smoothness indicator matrices
64 beta = zeros(k,k,k);
65 for r=0:k-1
66     x1 = -1/2 + [-r:1:k-r];
67     beta(:,:,r+1) = betarcalc(x1,k);
68 end
69
70 % Store the initial condtion in q, to have size(q) = [2, Nspacenodes]
71 q = [h';m'];
72
73 % We can now start solving the problem
74 t = 0;
75

```

```

76 while (t < tspan(2))
77
78     vel = q(2, :) ./ q(1,:);
79
80     % Adaptive time-step
81     dt = dx * CFL / max(abs(vel) + sqrt(g * q(1,:)));
82
83     % DEBUGGING
84     if dt < 1e-5
85         disp(h)
86         disp(m)
87         error("The solution is exploding")
88     end
89
90     if(t + dt >= tspan(2))
91         dt = tspan(2) - t;
92     end
93
94     qold = q;
95
96     % Define variables that will store the source terms
97     Source1 = zeros(2, length(xf)-1);
98     Source2 = zeros(2, length(xf)-1);
99     Source3 = zeros(2, length(xf)-1);
100
101     % Integrate once for all the source term and evaluate it at
102     % different
103     % time steps for later use from SSP-RK3
104     if PROBLEM == 1
105
106         Source1 = integrate_source(xf(1:end-1), xf(2:end), ...
107             t, PROBLEM);
108         Source2 = integrate_source(xf(1:end-1), xf(2:end), ...
109             t + dt, PROBLEM);
110         Source3 = integrate_source(xf(1:end-1), xf(2:end), ...
111             t + 0.5 * dt, PROBLEM);
112     end
113
114
115     % We use Runge-Kutta Strong Stability Preserving scheme to
116     % integrate
117     % in time (SSP-RK3) - see exercise sheet for algorithm
118
119     % SSP-RK3 Stage1
120     RHS = evalRHS(q, N, dt, dx, flux, flux_phys, bc, Crec, ...
121         k, xf, dw, beta);
122     q1 = qold + dt/dx * (RHS + Source1);

```

```

123 % SSP-RK3 Stage2
124 RHS = evalRHS(q1, N, dt, dx, flux, flux_phys, bc, Crec, ...
125     k, xf, dw, beta);
126 q2 = 3*qold/4.0 + (q1 + dt/dx * (RHS + Source2))/4.0;
127
128 % SSP-RK3 Stage3
129 RHS = evalRHS(q2, N, dt, dx, flux, flux_phys, bc, Crec, ...
130     k, xf, dw, beta);
131 q3 = qold/3.0 + 2.0*(q2 + dt/dx * (RHS + Source3))/3.0;
132
133 q = q3;
134
135 t = t + dt;
136
137 % Store the new solution
138 h = [h, q(1, :)'];
139 m = [m, q(2, :)'];
140 tvec = [tvec, t];
141
142 end

```

A.4 Evaluate RHS

```

1 function RHS = evalRHS(U, N, dt, dx, flux, flux_phys, bc, Crec, ...
2     k, xf, dw, beta)
3
4 % EVALRHS - Evaluate the right-hand side (RHS) of a numerical scheme
5 %           for a
6 %           hyperbolic conservation law.
7 %   RHS = EVALRHS(U, N, dt, dx, flux, flux_phys, bc, Crec, k, xf, dw,
8 %   beta)
9 %   computes the RHS of a hyperbolic conservation law based on the
10 %   given
11 %   solution vector 'U', numerical parameters, and flux functions.
12 %
13 %   Input:
14 %       U           - Solution matrix with two components.
15 %       N           - Number of grid points in the spatial domain.
16 %       dt          - Time step size.
17 %       dx          - Spatial grid spacing.
18 %       flux        - Flux function for the conservative variables.
19 %       flux_phys   - Physical flux.
20 %       bc          - String specifying the type of boundary condition.
21 %                   Supported values: 'peri' (periodic), 'open'
22 %                   (open boundary).
23 %       Crec        - Reconstruction method coefficient
24 %                   (for WENO reconstruction).
25 %       k           - Order of accuracy for WENO reconstruction.

```

```

24 %           xf           - Spatial grid locations.
25 %           dw           - WENO weights.
26 %           beta        - Parameter for WENO reconstruction.
27 %
28 %   Output:
29 %       RHS             - Computed right-hand side vector for
30 %                         the conservation law.
31 %
32 %   Note:
33 %       This function applies appropriate boundary conditions,
34 %       performs WENO reconstruction, and computes the flux differences
35 %       to obtain the RHS.
36 %
37 % Authors: [Francesco Sala, Nicolo' Viscusi]
38 % January 2024
39
40 % Apply appropriate boundary conditions
41 U = apply_bc(U, bc, k);
42
43 % Obtain reconstructed states
44 hl = zeros(1, N+2);
45 hr = zeros(1, N+2);
46
47 ml = zeros(1, N+2);
48 mr = zeros(1, N+2);
49
50 for i = 1:N+2
51     [hl(i), hr(i)] = WENO(xf, U(1, i:(i+2*(k-1))))', k, Crec, dw, beta);
52     [ml(i), mr(i)] = WENO(xf, U(2, i:(i+2*(k-1))))', k, Crec, dw, beta);
53 end
54
55 qr = [hr(2:N+1); mr(2:N+1)]; ql = [hl(2:N+1); ml(2:N+1)];
56 qm = [hr(1:N); mr(1:N)]; qp = [hl(3:N+2); ml(3:N+2)];
57
58 fluxval1 = flux(flux_phys, qr, qp, dx, dt);
59 fluxval2 = flux(flux_phys, qm, ql, dx, dt);
60
61 RHS = - (fluxval1-fluxval2);
62
63 end

```

A.5 Physical flux

```

1 function f = flux_phys(q)
2
3 % FLUX_PHYS - Computes the physical flux function for the shallow
4 %             water equations.
5 %
6 %   f = flux_phys(q)

```

```

7 %
8 % INPUTS:
9 %   q   - Vector of state variables [h, m], where
10 %       h: Water depth
11 %       m: Discharge
12 %
13 % OUTPUT:
14 %   f   - Vector representing the physical flux corresponding to the
15 %       input
16 %       state q.
17 %
18 % DESCRIPTION:
19 %   This function calculates the physical flux for the shallow water
20 %   equations.
21 %
22 % Authors: [Francesco Sala, Nicolo' Viscusi]
23 % January 2024
24
25 g = 1;
26 h = q(1, :);
27 m = q(2, :);
28
29 f = [m;
30      m.^2./h + 1/2*g*h.^2];
31
32 return

```

A.6 Numerical flux

```

1 function fval = LaxFriedrichs(flux_phys, Ul, Ur, dx, dt)
2 % LAXFRIEDRICHS - Compute the Lax-Friedrichs flux at a cell interface.
3 %
4 %   fval = LAXFRIEDRICHS(flux_phys, Ul, Ur, dx, dt) computes the
5 %   Lax-Friedrichs flux at a cell interface based on the
6 %   physical flux function, left-state 'Ul', right-state 'Ur',
7 %   spatial grid spacing 'dx', and time step size 'dt'.
8 %
9 % Input:
10 %   flux_phys - Physical flux function.
11 %   Ul        - Left-state values.
12 %   Ur        - Right-state values.
13 %   dx        - Spatial grid spacing.
14 %   dt        - Time step size.
15 %
16 % Output:
17 %   fval      - Computed Lax-Friedrichs flux at the cell interface.
18 %
19 % Authors: [Francesco Sala, Nicolo' Viscusi]
20 % January 2024

```



```

21
22
23 fval = 0.5 * (flux_phys(Ul) + flux_phys(Ur)) - 0.5 * dx/ dt * (Ur - Ul)
24 ;
25 return;

```

A.7 Apply boundary conditions

```

1 function U = apply_bc(Ui, bc, m)
2 % APPLY_BC - Apply boundary conditions to a given solution matrix.
3 %
4 % U = APPLY_BC(Ui, bc, m) applies boundary conditions to the input
5 % matrix
6 % UI based on the specified boundary condition type 'BC' and the
7 % number of
8 % ghost points 'M'. The function returns the modified matrix 'U' with
9 % added ghost points.
10 %
11 % Input:
12 %     Ui    - Input matrix without boundary conditions.
13 %     bc    - String specifying the type of boundary condition.
14 %             Supported values: 'peri' (periodic), 'open' (open
15 %             boundary).
16 %     m     - Number of ghost points to be added on each side.
17 %
18 % Output:
19 %     U     - Matrix with applied boundary conditions.
20 %
21 % Function available on Moodle for MATH-459 course
22 % January 2024
23
24 switch bc
25
26     case 'peri'
27
28         U = [Ui(:, end-m+1 : end), Ui, Ui(:, 1:m)];
29
30     case 'open'
31
32         U = [repmat(Ui(:, 1), 1, m), Ui, repmat(Ui(:, end), 1, m)];
33
34 end

```

A.8 Integrate source term in space exactly

```

1 function intS = integrate_source(xa, xb, t, PROBLEM)
2
3 % INTEGRATE_SOURCE - Compute the spatial integral of the source term
4 % for
5 %             a given hyperbolic problem at a time t, using the
6 %             exact solution of the integral.
7 %
8 % intS = INTEGRATE_SOURCE(xa, xb, t, PROBLEM) computes the spatial
9 % integral of the source term for a specified problem at a given time
10 % 't'.
11 %
12 % Input:
13 %     xa      - Left spatial boundary.
14 %     xb      - Right spatial boundary.
15 %     t       - Time at which the source term is evaluated.
16 %     PROBLEM - Integer specifying the problem type.
17 %               Supported values: 1 or 2.
18 %
19 % Output:
20 %     intS     - Computed spatial integral of the source term.
21 %
22 % Parameters:
23 %     u        - Constant parameter (problem-dependent).
24 %     g        - Constant parameter (problem-dependent).
25 %
26 % Problem Descriptions:
27 %     PROBLEM = 1: Computes the integral for the source term.
28 %     PROBLEM = 2: Returns zero for the source term.
29 %
30 % Authors: [Francesco Sala, Nicolò Viscusi]
31 % January 2024
32
33 u = 0.25;
34 g = 1;
35
36 if PROBLEM == 1
37     int_x_S = @(xa,xb) [0.5*(u-1).*(sin(pi*(t-xa))-sin(pi*(t-xb)))+
38         1/8*sin(pi*(t-xb)).*(g*sin(pi*(t-xb)) -4*(g+(u-1)*u))- ...
39         1/8*sin(pi*(t-xa)).*(g*sin(pi*(t-xa)) -4*(g+(u-1)*u))];
40     intS = int_x_S(xa,xb);
41
42 elseif PROBLEM == 2
43
44     int_x_S = @(xa,xb) [0*xa; 0*xa];
45     intS = int_x_S(xa, xb);
46
47 end

```

48 `return`

A.9 Functions by Jan H. Hesthaven

A.9.1 WENO reconstruction

```

1  function [um,up] = WENO(xloc,uloc,m,Crec,dw,beta);
2  % Purpose: Compute the left and right cell interface values using an
   WENO
3  % approach based on 2m-1 long vectors uloc with cell
4  % Set WENO parameters
5  % Function implemented by Jan S. Hesthaven
6
7  p = 1; q = m-1; vareps = 1e-6;
8
9  % Treat special case of m=1 - no stencil to select
10 if (m==1)
11     um = uloc(1); up = uloc(1);
12 else
13     alphas = zeros(m,1); alphap = zeros(m,1);
14     upl = zeros(m,1); uml = zeros(m,1); betar = zeros(m,1);
15
16     % Compute um and up based on different stencils and
17     % smoothness indicators and alpha
18     for r=0:m-1;
19         umh = uloc(m-r+[0:m-1]);
20         upl(r+1) = Crec(r+2,:)*umh; uml(r+1) = Crec(r+1,:)*umh;
21         betar(r+1) = umh'*beta(:, :, r+1)*umh;
22     end;
23
24     % Compute alpha weights - classic WENO
25     alphap = dw./(vareps+betar).^(2*p);
26     alphas = flipud(dw)./(vareps+betar).^(2*p);
27
28     % Compute alpha weights - WENO-Z
29     % tau = abs(betar(1) - betar(m));
30     % if mod(m,2)==0
31     % tau = abs(betar(1)-betar(2) - betar(m-1) + betar(m));
32     % end
33     % alphap = dw.*(1 + (tau./(vareps+betar)).^q);
34     % alphas = flipud(dw).*(1 + (tau./(vareps+betar)).^q);
35
36     % Compute nonlinear weights and cell interface values
37     um=alphas'*uml/sum(alphas); up=alphap'*upl/sum(alphap);
38 end
39 return

```

A.9.2 Reconstruction weights

```

1 function [c] = ReconstructWeights(m,r)
2 % Purpose: Compute weights c_ir for reconstruction
3 %            $v_{\{j+1/2\}} = \sum_{j=0}^{m-1} c_{\{ir\}} v_{\{i-r+j\}}$ 
4 %           with  $m=order$  and  $r=shift$  ( $-1 \leq r \leq m-1$ ).
5 % Function implemented by Jan S. Hesthaven
6
7 c = zeros(1,m); fh = @(s) (-1)^(s+m)*prod(1:s)*prod(1:(m-s));
8 for i=0:m-1
9     q = linspace(i+1,m,m-i);
10    for q=(i+1):m
11        if (q~=r+1)
12            c(i+1) = c(i+1) + fh(r+1)/fh(q)/(r+1-q);
13        else
14            c(i+1) = c(i+1) - (harmonic(m-r-1)-harmonic(r+1));
15        end
16    end
17 end
18 end

```

A.9.3 Qcalc

```

1 function [Qelem] = Qcalc(D,m,l);
2 % Purpose: Evaluate entries in the smoothness indicator for WENO
3 % Function implemented by Jan S. Hesthaven
4
5 [x,w] = LegendreGQ(m); xq = x./2; Qelem = 0;
6 for i=1:m+1
7     xvec = zeros(m-l+1,1);
8     for k=0:m-l xvec(k+1) = xq(i)^k./prod(1:k); end;
9     Qelem = Qelem + (xvec'*D*xvec)*w(i)/2;
10 end
11 return

```

A.9.4 betarcalc

```

1 function [errmat] = betarcalc(x,m)
2 % Purpose: Compute matrix to allow evaluation of smoothness indicator
3 %           in
4 %           WENO based on stencil [x] of length m+1 .
5 %           Returns sum of operators for l=1..m-1
6 % Function by Jan H. Hesthaven
7
8 % Evaluate Lagrange polynomials
9 [cw] = lagrangeweights(x);
10
11 % Compute error matrix for l=1..m-1
12 errmat = zeros(m,m);
13 for l=2:m

```

```

13 % Evaluate coefficients for derivative of Lagrange polynomial
14 dw = zeros(m,m-1+1);
15 for k=0:(m-1)
16     for q=0:m-1
17         dw(q+1,k+1) = sum(cw((q+2):m+1,k+1+1));
18     end
19 end
20
21 % Evaluate entries in matrix for order 'l'
22 Qmat = zeros(m,m);
23 for p=0:m-1
24     for q=0:m-1
25         D = dw(q+1,:)'*dw(p+1,:);
26         Qmat(p+1,q+1) = Qcalc(D,m,l);
27     end
28 end
29 errmat = errmat + Qmat;
30 end
31 return

```

A.9.5 betarcalc

```

1 function [errmat] = betarcalc(x,m)
2 % Purpose: Compute matrix to allow evaluation of smoothness indicator
   in
3 %           WENO based on stencil [x] of length m+1 .
4 %           Returns sum of operators for l=1..m-1
5 % Function by Jan H. Hesthaven
6
7 % Evaluate Lagrange polynomials
8 [cw] = lagrangeweights(x);
9
10 % Compute error matrix for l=1..m-1
11 errmat = zeros(m,m);
12 for l=2:m
13     % Evaluate coefficients for derivative of Lagrange polynomial
14     dw = zeros(m,m-1+1);
15     for k=0:(m-1)
16         for q=0:m-1
17             dw(q+1,k+1) = sum(cw((q+2):m+1,k+1+1));
18         end
19     end
20
21     % Evaluate entries in matrix for order 'l'
22     Qmat = zeros(m,m);
23     for p=0:m-1
24         for q=0:m-1
25             D = dw(q+1,:)'*dw(p+1,:);
26             Qmat(p+1,q+1) = Qcalc(D,m,l);

```

```

27     end
28 end
29 errmat = errmat + Qmat;
30 end
31 return

```

A.9.6 LegendreGQ

```

1 function [x,w] = LegendreGQ(m);
2 % function [x,w] = LegendreGQ(m)
3 % Purpose: Compute the m'th order Legendre Gauss quadrature points, x,
4 %           and weights, w
5 % Function implemented by Jan S. Hesthaven
6
7 if (m==0) x(1)=0; w(1) = 2; return; end;
8
9 % Form symmetric matrix from recurrence.
10 J = zeros(m+1); h1 = 2*(0:m);
11 J = diag(2./(h1(1:m)+2).*...
12         sqrt((1:m).*((1:m)).*((1:m)).*((1:m))./(h1(1:m)+1)./(h1(1:m)+3))
13         ,1);
14 J(1,1)=0; J = J + J';
15
16 %Compute quadrature by eigenvalue solve
17 [V,D] = eig(J); x = diag(D); w = 2*(V(1,:))'.^2;
18 return

```

A.9.7 LinearWeights

```

1 function [d]=LinearWeights(m,r0)
2 % Purpose: Compute linear weights for maximum accuracy 2m-1,
3 % using stencil shifted $r_0=-1,0$ points upwind.
4 % Function implemented by Jan S. Hesthaven
5
6 A = zeros(m,m); b = zeros(m,1);
7
8 % Setup linear system for coefficients
9 for i=1:m
10     col = ReconstructWeights(m,i-1+r0);
11     A(1:(m+1-i),i) = col(i:m)';
12 end
13
14 % Setup righthand side for maximum accuracy and solve
15 crhs = ReconstructWeights(2*m-1,m-1+r0);
16 b = crhs(m:(2*m-1))'; d = A\b;
17 return

```

A.9.8 Lagrange weights

```
1 function [cw] = lagrangeweights(x)
2 % Purpose: Compute weights for Taylor expansion of Lagrange polynomial
3 %          based on x and evaluated at 0.
4 %          Method due to Fornberg (SIAM Review, 1998, 685-691)
5 % Function implemented by Jan S. Hesthaven
6
7 np = length(x); cw=zeros(np,np);
8 cw(1,1)=1.0; c1 = 1.0; c4 = x(1);
9 for i=2:np
10     mn = min(i,np-1)+1;
11     c2 = 1.0; c5 = c4; c4 = x(i);
12     for j=1:i-1
13         c3 = x(i)-x(j); c2 = c2*c3;
14         if (j==i-1)
15             for k=mn:-1:2
16                 cw(i,k) = c1*((k-1)*cw(i-1,k-1)-c5*cw(i-1,k))/c2;
17             end
18             cw(i,1) = -c1*c5*cw(i-1,1)/c2;
19         end
20         for k=mn:-1:2
21             cw(j,k) = (c4*cw(j,k)-(k-1)*cw(j,k-1))/c3;
22         end
23         cw(j,1) = c4*cw(j,1)/c3;
24     end
25     c1=c2;
26 end
27 return
```