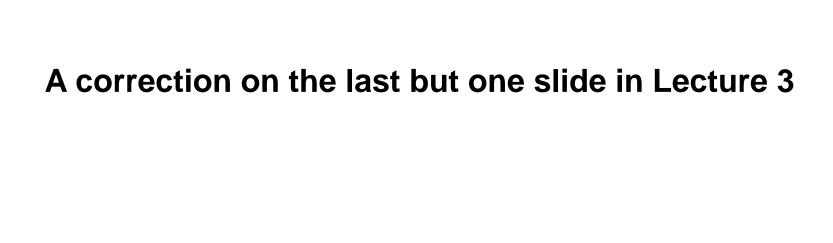
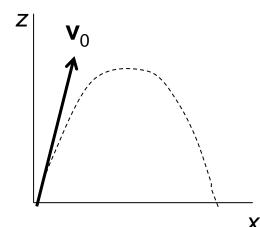
# PH101 Lecture 4

08.08.14



### **Example on motion in a uniform gravitational field:**

A <u>freely moving ball</u> under the influence of earth's gravity experiences a constant downward acceleration of *g*. Assuming that it is confined to *xz* plane, obtain its trajectory by assuming *z*-axis to be vertically upwards.



$$\vec{a} = -g \hat{k}$$

Assuming it to be released at t = 0 with an initial velocity  $\mathbf{v}_0$  and assuming position vector  $\mathbf{r}_0 = 0$ 

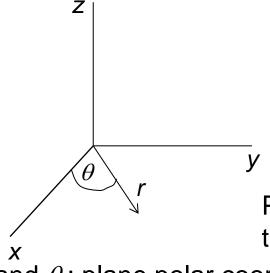
$$x = v_{0x}t$$

$$\vec{r}(t) = \vec{r}_0 + \vec{v}_0 t + \frac{1}{2} \vec{a} t^2$$
  $z = v_{0z} t - \frac{1}{2} g t^2$ 

$$\Rightarrow$$
 trajectory:  $z = v_{0z} \frac{x}{v_{0x}} - \frac{1}{2} g \frac{x^2}{v_{0x}^2} = \frac{v_{0z}}{v_{ox}} x - \frac{g}{2v_{0x}^2} x^2$ 

# Motion in plane polar coordinates

Cartesian coordinates are **not** suitable to model/discuss circular motion

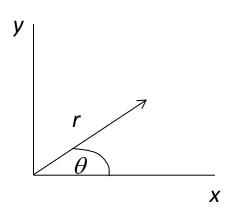


#### Polar coordinates

The cylindrical coordinate system, in which z axis is identical to cartesian system

Position in the xy-plane is given by the distance r from the z axis and its angle  $\theta$  with the x axis

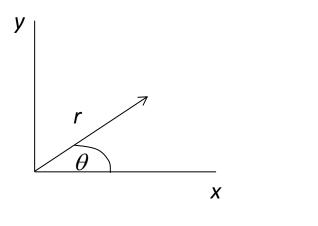
r and  $\theta$ : plane polar coordinates

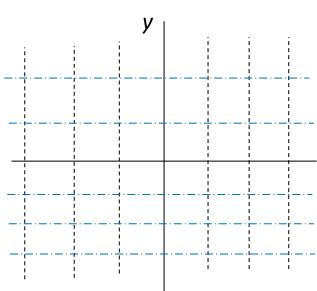


$$r = \sqrt{x^2 + y^2}$$

$$\frac{r\cos\theta}{r\sin\theta} = \frac{y}{x} \Rightarrow \theta = \tan^{-1}\frac{y}{x}$$

### Constant coordinate lines in cartesian & plane polar coordinates





---- : *x* = constant

----: *y* = constant

Lines of constant x & constant y are st lines  $\perp$  to each other

---- : *r* = constant

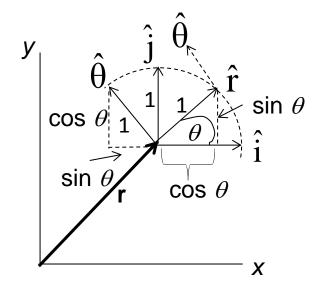
----:  $\theta$  = constant

Lines of constant  $\theta$  are also st lines

Lines of constant *r* are concentric circles

### Base vectors in plane polar coordinates are

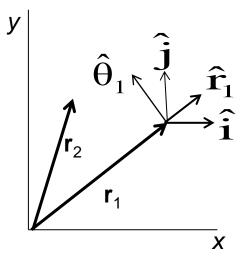
# $\hat{r} \& \hat{\theta}$ and are perpendicular to each other



$$\hat{\mathbf{r}} = \hat{\mathbf{i}}\cos\theta + \hat{\mathbf{j}}\sin\theta$$

$$\hat{\theta} = -\hat{i}\sin\theta + \hat{j}\cos\theta$$

They point in the direction of increasing r and increasing  $\theta$  and they both vary with position.



Naturally 
$$\hat{\mathbf{r}} \cdot \hat{\boldsymbol{\theta}} = 0$$

Check for  $\theta = 0$  and  $\pi/2$ 

$$\vec{r} = x\hat{i} + y\hat{j}$$

In plane polar coordinates:  $\vec{r} = r \hat{r}$ 

$$\Rightarrow$$

$$x\hat{i} + y\hat{j} = r(\hat{i}\cos\theta + \hat{j}\sin\theta)$$

$$\Rightarrow$$

$$x = r \cos \theta$$

$$y = r \sin \theta$$
As one would expect

 $\vec{r} = r \hat{r}$ 

x and y are needed in cartesian coordinates to specify position vector r!

is not a fixed vector

Orientation  $\theta$  of  $\widehat{\mathbf{r}}$  is needed as well as r to find out how far one is from the origin

$$\Rightarrow \vec{r} = r \hat{\mathbf{r}}(\theta)$$

 $\hat{r}$  and  $\hat{\theta} \perp$  to each other and also are unit vectors

$$\Rightarrow |\hat{\mathbf{r}}| = 1 \text{ and } |\hat{\boldsymbol{\theta}}| = 1$$

Thus if

$$\vec{A} = A_r \hat{\mathbf{r}} + A_\theta \hat{\mathbf{\theta}}$$

$$\vec{B} = B_r \hat{\mathbf{r}} + B_\theta \hat{\boldsymbol{\theta}}$$

$$\vec{A}.\vec{B} = A_r B_r + A_\theta B_\theta$$

Is this true everywhere in space?

 $\hat{\mathbf{r}} \ \& \ \hat{\boldsymbol{\theta}}$  must refer to the same point in space!

# **Velocity in polar coordinates**

Position **r** of a particle in plane polar coordinates is

$$\vec{r} = r \hat{r}$$

**r** changes as the particle moves

$$\vec{v} = \frac{d\vec{r}}{dt} = \frac{dr}{dt}\hat{r}$$

 $\hat{\mathbf{r}}$  depends on heta

As the particle moves  $\theta$  changes with time and hence  $\widehat{\mathbf{1}}$  also changes

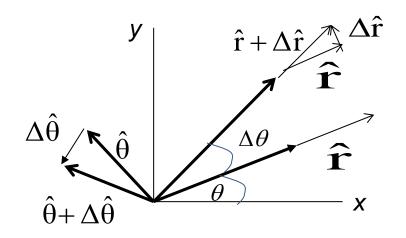
Correct v is

$$\vec{v} = \left(\frac{dr}{dt}\right)\hat{\mathbf{r}} + r\left(\frac{d\,\hat{\mathbf{r}}}{dt}\right)$$

Compare with **v** in cartesian coordinates

$$\vec{v} = \frac{d}{dt} (x \hat{i} + y \hat{j}) = \dot{x} \hat{i} + \dot{y} \hat{j}$$

 $\left(rac{d \, \hat{\mathbf{r}}}{dt}
ight)$ :  $\hat{\mathbf{r}}$ does not change as one moves radially in or out  $\hat{\mathbf{r}}$  changes only if heta changes



$$\begin{aligned} \left| \Delta \hat{\mathbf{r}} \right| &\approx \left| \hat{\mathbf{r}} \right| \Delta \theta = \Delta \theta \\ \frac{\left| \Delta \hat{\mathbf{r}} \right|}{\Delta t} &\approx \frac{\Delta \theta}{\Delta t} \qquad \text{In the limit } \Delta t \to 0 \qquad \left| \frac{d \hat{\mathbf{r}}}{d t} \right| = \frac{d \theta}{d t} \end{aligned}$$

As heta increases,  $\hat{\mathbf{r}}$  swings in the direction  $\hat{\mathbf{\Theta}}$ 

$$\frac{d\,\hat{\mathbf{r}}}{dt} = \dot{\theta}\,\hat{\boldsymbol{\theta}}$$

Alternatively

$$\frac{d\hat{\mathbf{r}}}{dt} = \frac{d}{dt} \left( \cos \theta \, \hat{\mathbf{i}} + \sin \theta \, \hat{\mathbf{j}} \right) : \left( \frac{d}{d\theta} \cos \theta \right) \frac{d\theta}{dt} = -\sin \theta \, \dot{\theta}$$

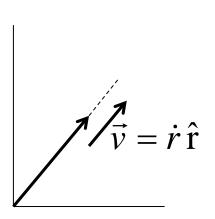
$$\left( \frac{d}{d\theta} \sin \theta \right) \frac{d\theta}{dt} = \cos \theta \, \dot{\theta}$$

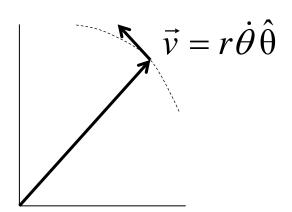
$$\Rightarrow \frac{d\hat{\mathbf{r}}}{dt} = -\hat{\mathbf{i}} \sin \theta \, \dot{\theta} + \hat{\mathbf{j}} \cos \theta \, \dot{\theta} = (-\hat{\mathbf{i}} \sin \theta + \hat{\mathbf{j}} \cos \theta) \, \dot{\theta}$$

$$\Rightarrow \frac{d\hat{\mathbf{r}}}{dt} = \dot{\theta} \, \hat{\theta}$$

Thus, velocity of the particle will be

$$\vec{v} = \dot{r}\hat{r} + r\dot{\theta}\hat{\theta}$$





Acceleration

$$\vec{a} = \frac{d\vec{v}}{dt} = \frac{d}{dt} \left( \dot{r} \, \hat{\mathbf{r}} + r \dot{\theta} \, \hat{\theta} \, \right) = \ddot{r} \, \hat{\mathbf{r}} + \dot{r} \, \frac{d \, \hat{\mathbf{r}}}{dt} + \dot{r} \dot{\theta} \, \hat{\theta} + r \ddot{\theta} \, \hat{\theta} + r \dot{\theta} \, \frac{d \, \hat{\theta}}{dt}$$

$$d \, \hat{\mathbf{r}} \qquad \dot{\partial} \, \hat{\mathbf{o}}$$

$$\frac{d\,\hat{\mathbf{r}}}{dt} = \dot{\theta}\,\hat{\mathbf{\theta}}$$

$$\frac{d\theta}{dt} = -\dot{\theta}\,\hat{\mathbf{r}}$$

Centripetal acceleration

Coriolis acceleration

$$\vec{a} = (\ddot{r} - r\dot{\theta}^2)\hat{\mathbf{r}} + (r\ddot{\theta} + 2\dot{r}\dot{\theta})\hat{\mathbf{\theta}}$$

Linear acceleration in the radial direction

Linear acceleration in the tangential direction