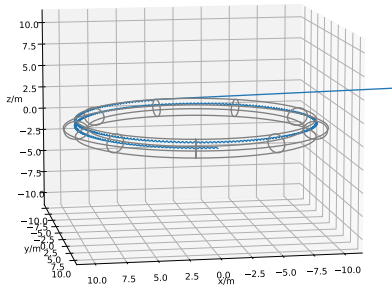


Exploring electric and magnetic forces using computer simulations

Nikolaj Roager Christensen

Student Colloquium in Physics and Astronomy, Aarhus University

March 2021



Wellcome

- ▶ Today's topic: particles in electric and magnetic fields

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- ▶ Todays topic: particles in electric and magnetic fields
- ▶ Explored using computer-simulations

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- ▶ Explored using computer-simulations
- ▶ Todays plan

Introduction (3 minutes)

Theory and background (13 minutes)

Steering particles with \vec{B} fields (14 min)

Introducing Electric fields (10 min)

Conclusion and Questions(5 min)

Introduction, what and why

- ▶ (Classical) Charged particles
in Electric and Magnetic
fields

Introduction, what and why

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- ▶ How can magnetic fields steer and collect particles

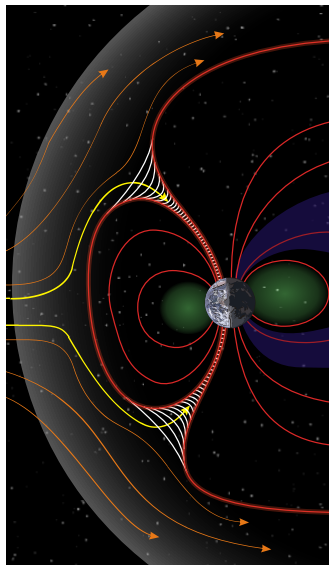
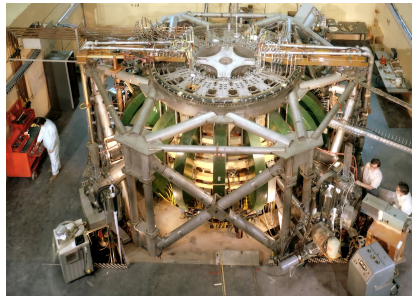


Illustration
originally from Nasa. Published
on wikipedia, in Public Domain

Introduction, what and why

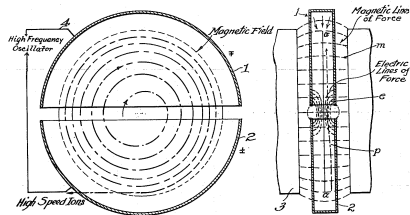
- ▶ (Classical) Charged particles in Electric and Magnetic fields
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- ▶ Real world examples:
 - ▶ Magnetic traps:
“Tokamak” style fusion reactors.



Princeton Large Torus in 1975,
image in Public Domain

Introduction, what and why

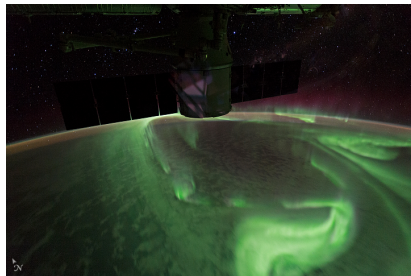
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Ernest O. Lawrence, 1934, U.S. Patent 1,948,384; image in Public Domain

Introduction, what and why

- ▶ (Classical) Charged particles in Electric and Magnetic fields
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 - ▶ The Aurora.



Nasa: Aurora Australis, Indian Ocean from ISS in 2017; image in Public domain

How: Numeric-simulation

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 - ▶ Runge Kutta Algorithm
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- ▶ Not the main focus.
- ▶ Simulations are not experiments!

Theory and background (13 minutes)

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Steering particles with \vec{B} fields (14 min)

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Theory: Electric and magnetic fields

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- ▶ Some repetition from Electrodynamics
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- ▶ Only 1 particle! so pre-programmed depending on the setup.

Theory: Alternative approach

► I used:

$$m\ddot{\vec{r}} = q(\dot{\vec{r}} \times \vec{B} + \vec{E}).$$

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$$\vec{E}(\vec{r}, t) = -\nabla\phi(\vec{r}, t)$$

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$$\vec{E}(\vec{r}, t) = -\nabla\phi(\vec{r}, t)$$

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- Hamiltonian equation of motion

$$\dot{r}_i = \frac{\partial \mathcal{H}}{\partial p_i},$$

$$\mathcal{H} = \frac{\vec{p}^2}{2m} + V(\vec{r}) \rightarrow \frac{(\vec{p}^2 - q\vec{A}(\vec{r}, t))^2}{2m} + q\phi(\vec{r}, t) + V(\vec{r}).$$

The ODE implementation

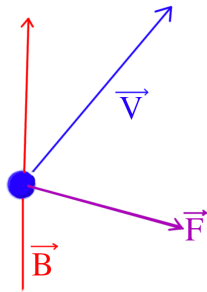
```
auto ODE = [...](const state_type Data, state_type &
    dDatadt, const double t)
{
    //Extract position and velocity from data
    ...

    //Get current force
    vec F = Charge*(Fields.get_Efield(pos,t)+
        cross(velocity,Fields.get_Bfield(pos,t)));

    vec dVdt = F*Inv_mass;    //get acceleration

    //Save derivative of data
    ...
};
```

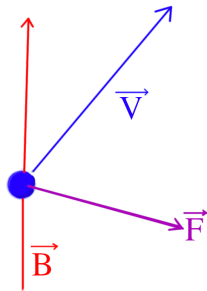
Known results, cyclotron motion \vec{B} fields



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- Magnetic forces do no work:

$$dW_{\vec{B}} = \vec{F}_B \cdot d\vec{r} \propto (\vec{v} \times \vec{B}) \cdot \vec{v} = 0.$$



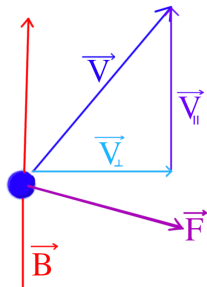
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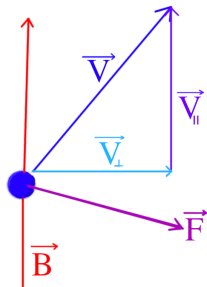
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Cyclotron motion



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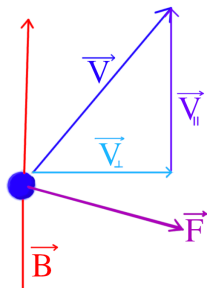
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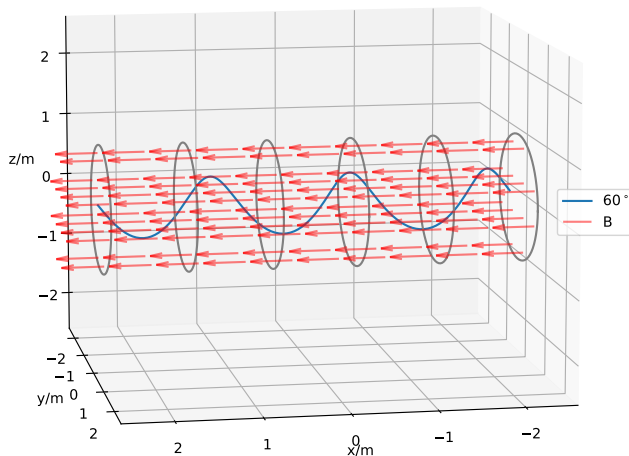
$$|\vec{F}_B| = |q(\vec{v} \times \vec{B})| = |qv_\perp B|.$$

- Same as Centripetal force:
Cyclotron motion
- Cyclotron radius and
frequency:

$$R = \frac{v_\perp m}{|q|B} \quad \omega_c = \frac{|q|B}{m}.$$

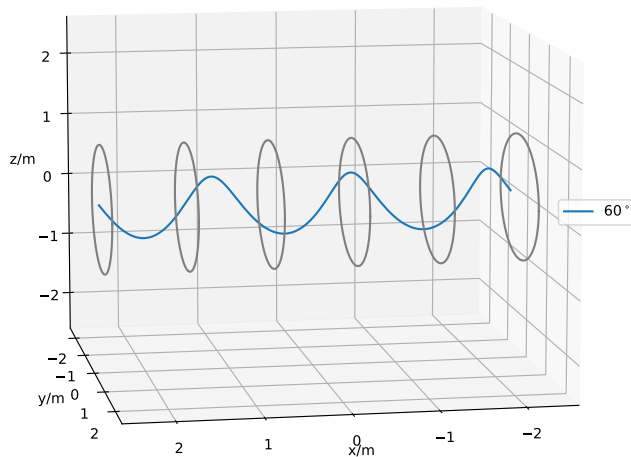


Protons in a Solenoid, 3D plots

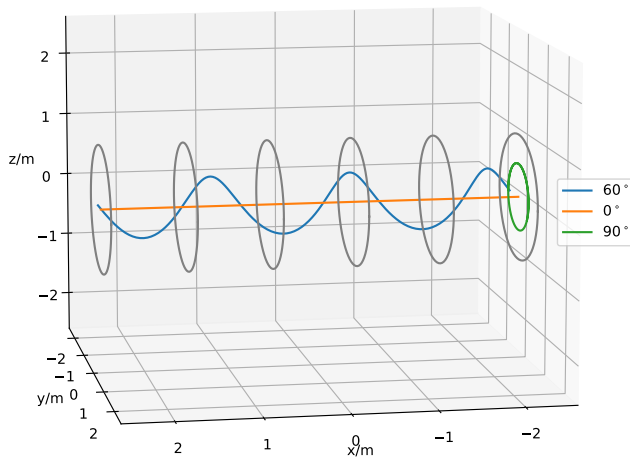


- ▶ Solenoid with $N = 1000$ turns per m , $I = 5$ A, $r = 1$ m, $|\vec{B}| \approx 6$ mT.
- ▶ Proton with $E_{kin} = 1$ MeV/ c^2 ($|v| \approx 3,195 \times 10^5$ m/s)

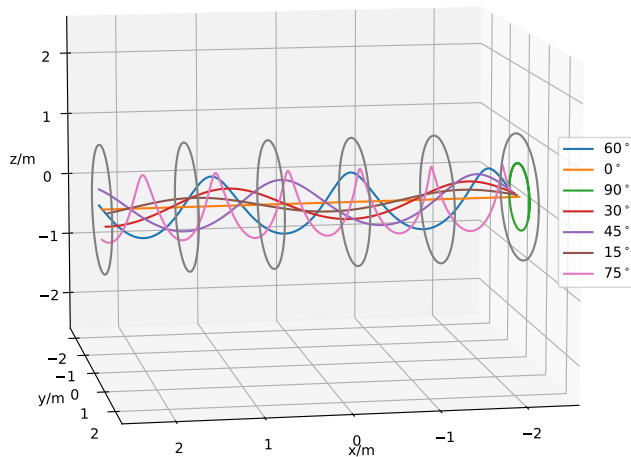
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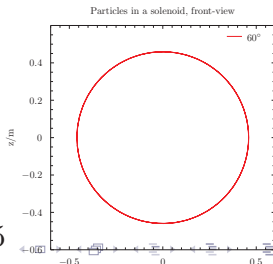
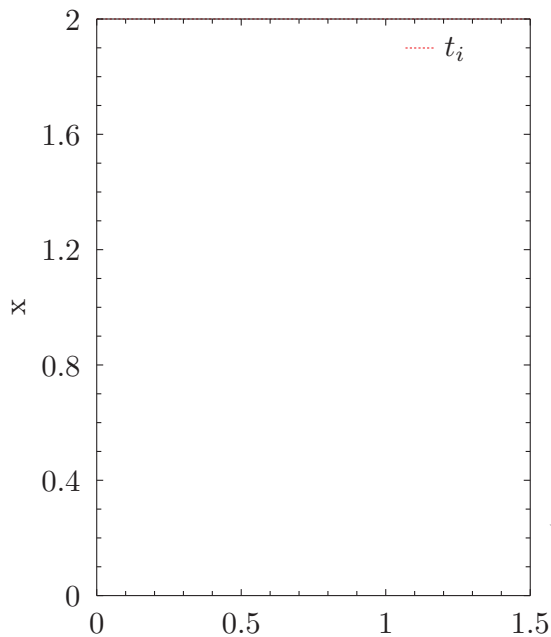
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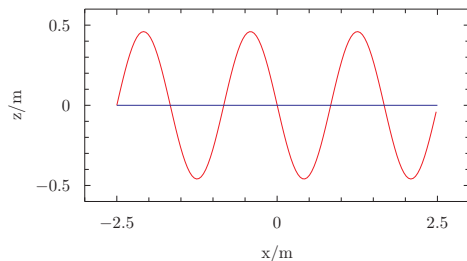


Protons in a Solenoid, 2D plots

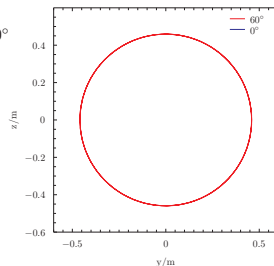


Protons in a Solenoid, 2D plots

Particles in a solenoid, side-view

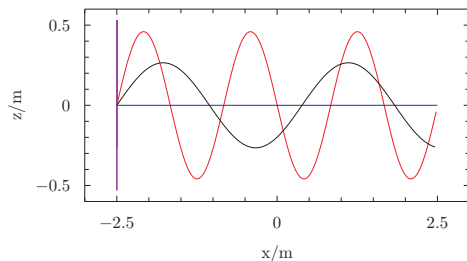


Particles in a solenoid, front-view

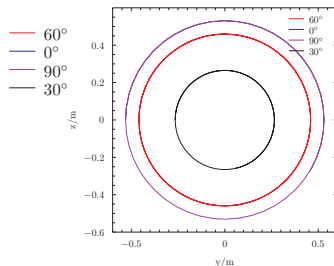


Protons in a Solenoid, 2D plots

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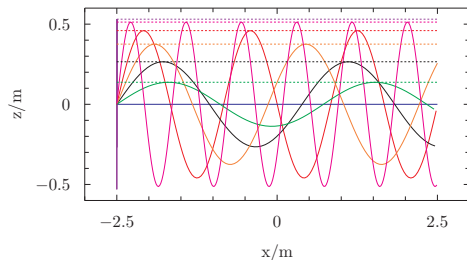


Particles in a solenoid, front-view

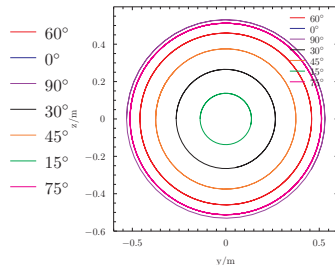


Protons in a Solenoid, 2D plots

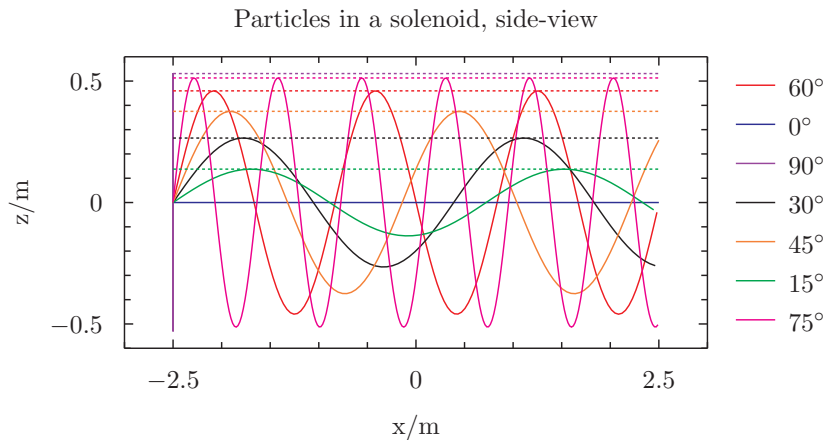
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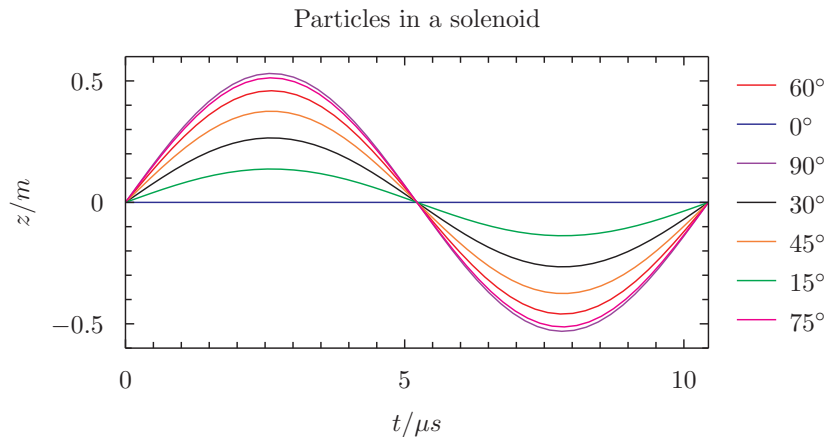


Protons in a Solenoid, 2D plots



$$R \approx 0,5 \text{ m} \sin(\theta) \quad T = \frac{2\pi}{\omega_c} \approx 10 \mu\text{s}$$

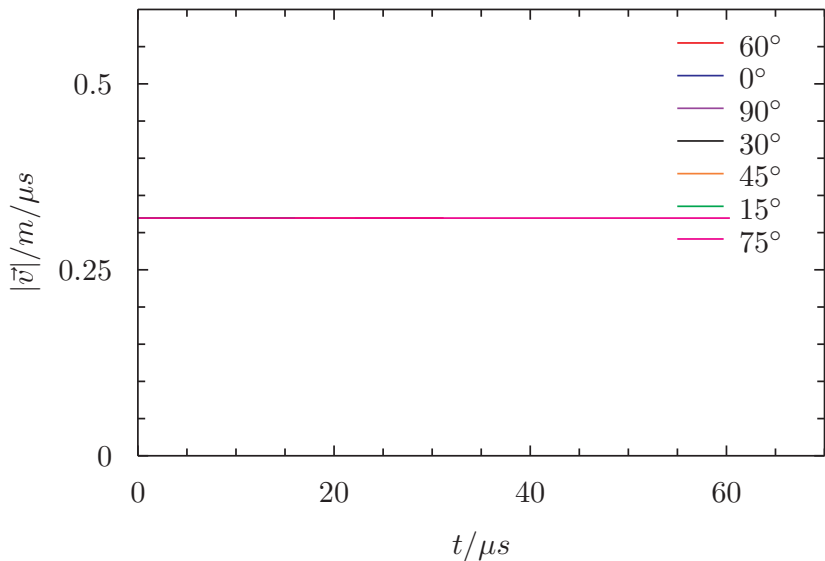
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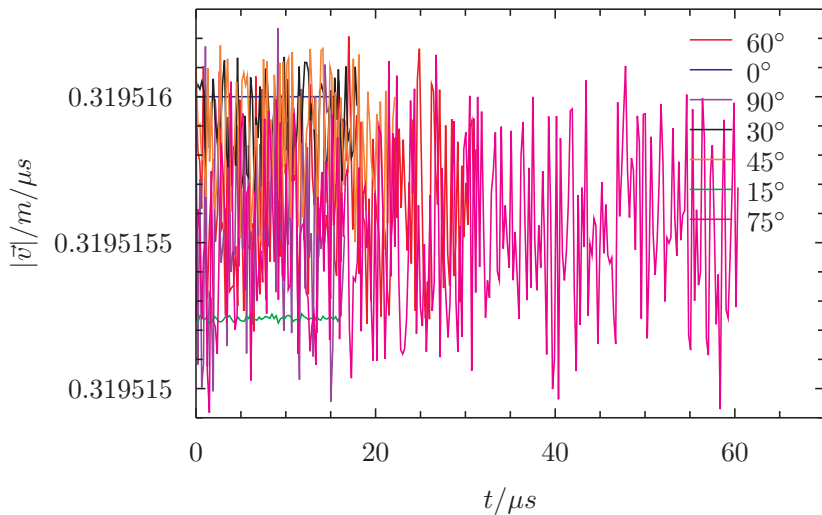
Sanity check, no work

speed over time should be constant



Sanity check, no work

speed over time (closer look)



Steering particles with \vec{B} fields (14 min)

Introduction (3 minutes)

Theory and background (13 minutes)

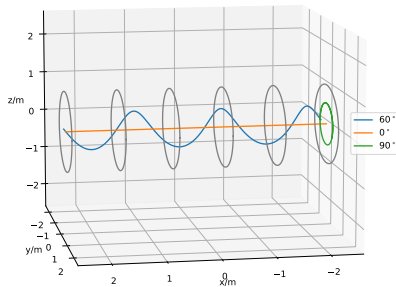
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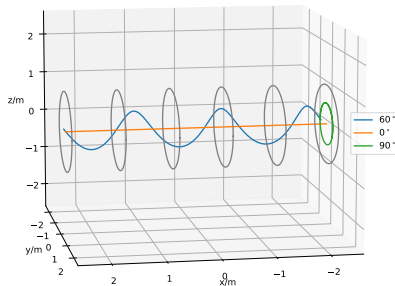
how can \vec{B} field steer particles

- The parallel velocity is untouched.



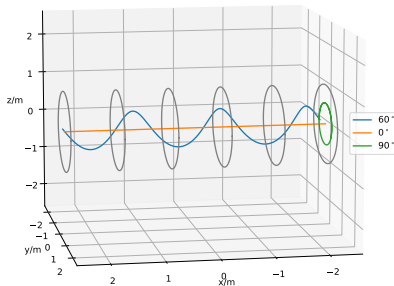
how can \vec{B} field steer particles

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how can \vec{B} field steer particles

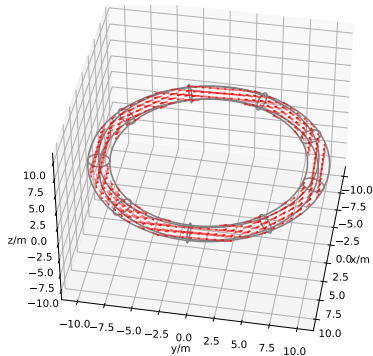
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- ▶ What if instead the field curves.



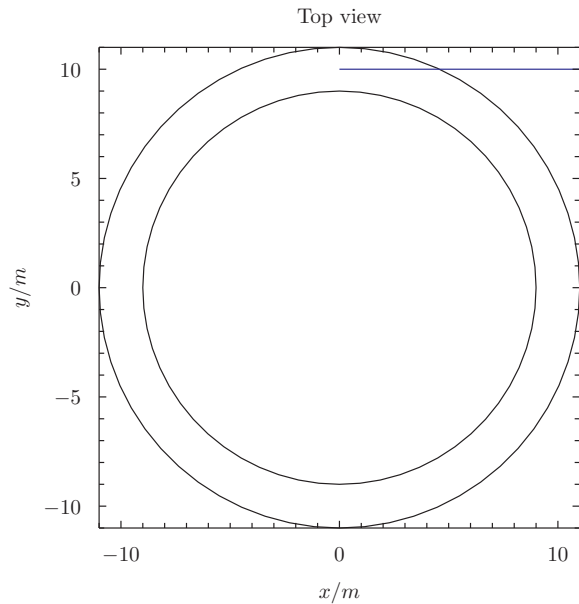
how can \vec{B} field steer particles

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- ▶ Example torus:

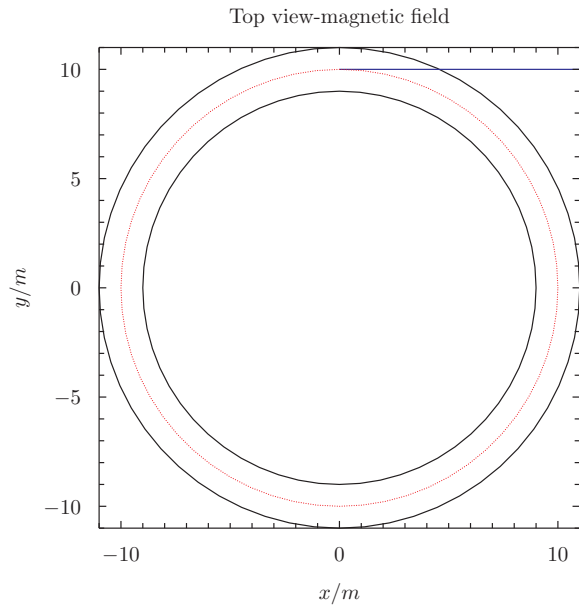
$$\vec{B}(z, r, \phi) = \hat{\phi} \mu_0 \frac{N_{tot} I}{2\pi r} = \hat{\phi} B(R) \frac{R}{r}.$$



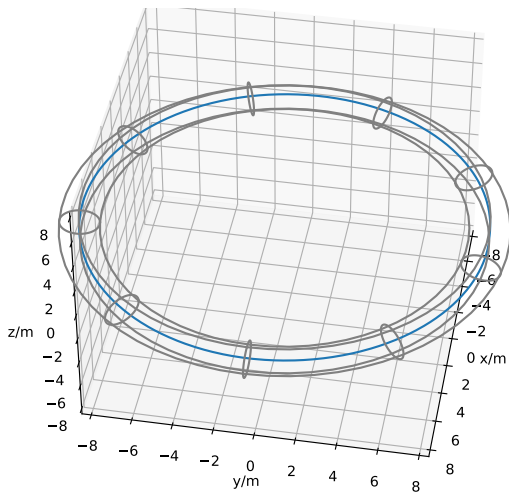
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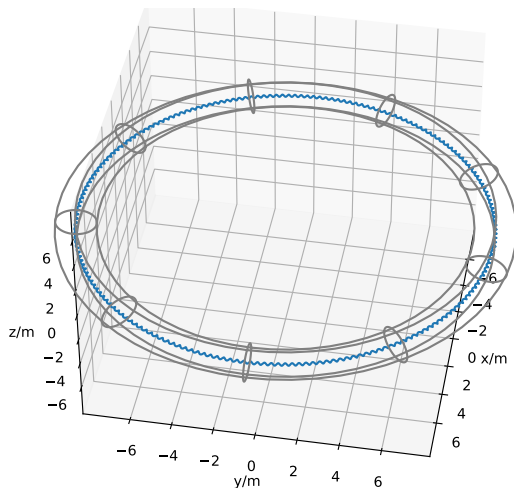
Containing protons in a Torus



Proton

starting at $(0, 10, 0)$, $\vec{v}_0 = (0.3, 0, 0)\text{m}/\mu\text{s}$. $|B| \approx 60\text{ mT}$

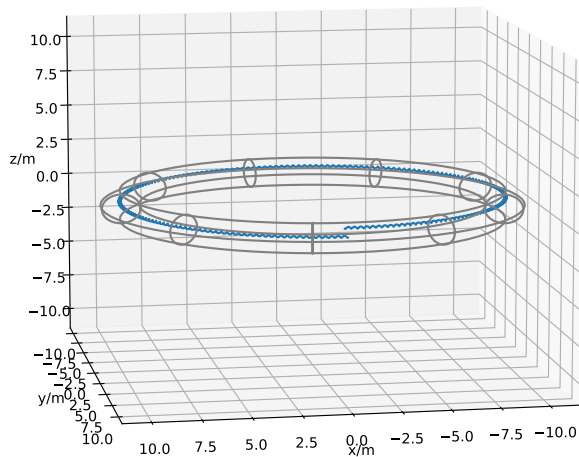
Containing protons in a Torus



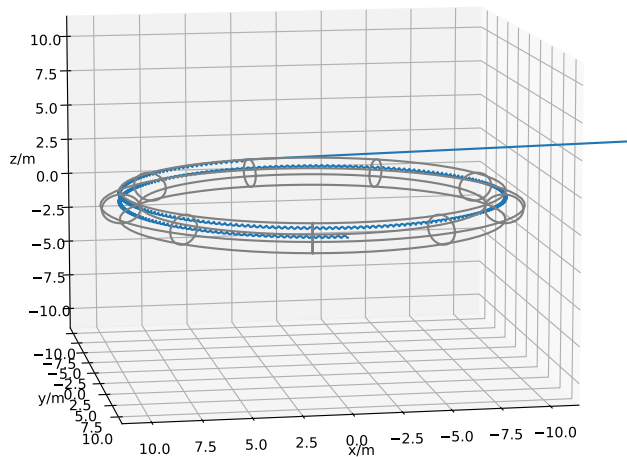
Proton

starting at $(0, 10, 0)$, $\vec{v}_0 = (0.3, 0.4, 0)m/\mu s$.

Instability, Error?

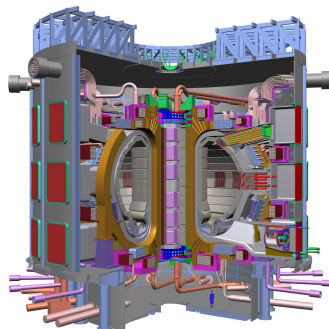


Instability, Error?



Not a torus

- ▶ “Tokamak” style reactors are NOT toruses.

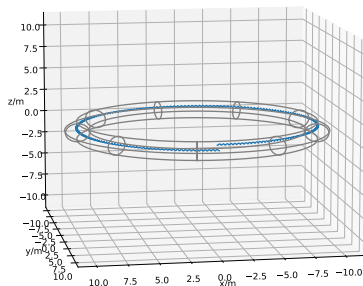


Iter: International Thermonuclear
Experimental Reactor, Image
published by U.S. Department of
Energy, placed in Public Domain

Not a torus

- ▶ “Tokamak” style reactors are NOT toruses.
- ▶ Remember:

$$\vec{B}(z, r, \phi) = \vec{\hat{\phi}} B(R) \frac{R}{r}.$$

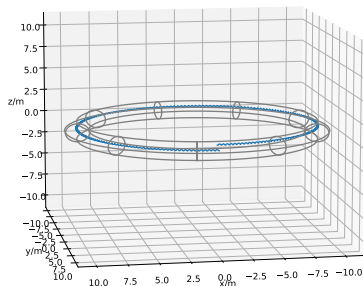


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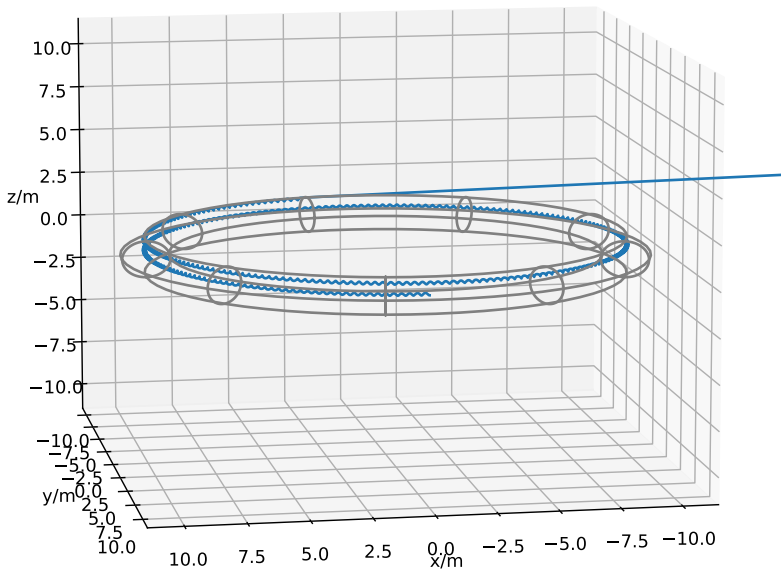
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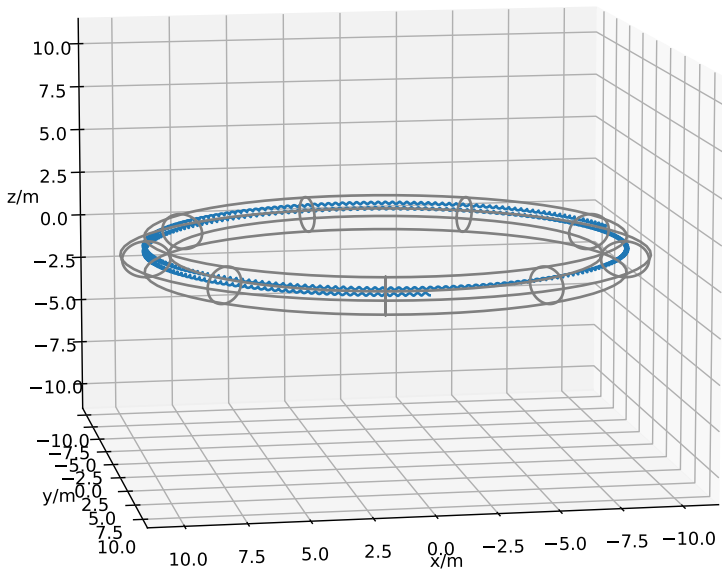
- ▶ Does lead to drift.



Solving the problem



Solving the problem





- ▶ Soviet Stamp from 1987 showing concept of Tokamak, Image in Public Domain
- ▶ Need to consider “current” of the plasma.



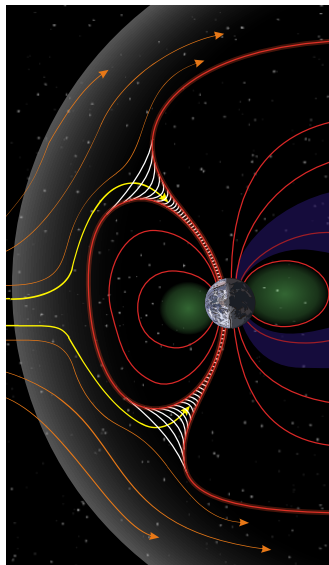
- ▶ Soviet Stamp from 1987 showing concept of Tokamak, Image in Public Domain
- ▶ Need to consider “current” of the plasma.
- ▶ More stabilization and control is needed.



- ▶ Soviet Stamp from 1987 showing concept of Tokamak, Image in Public Domain
- ▶ Need to consider “current” of the plasma.
- ▶ More stabilization and control is needed.
- ▶ Particles do follow magnetic fields!

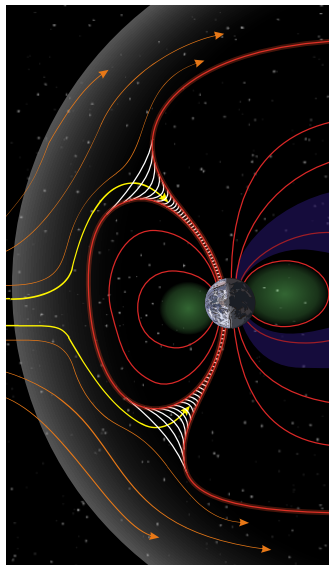
Another example, Aurora

- ▶ Solar wind (ions) vs. Earth Magnetic field.
- ▶ Illustration originally from Nasa. Published on wikipedia, in Public Domain (Cropped to fit page)



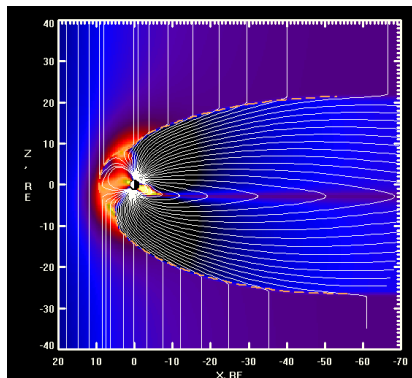
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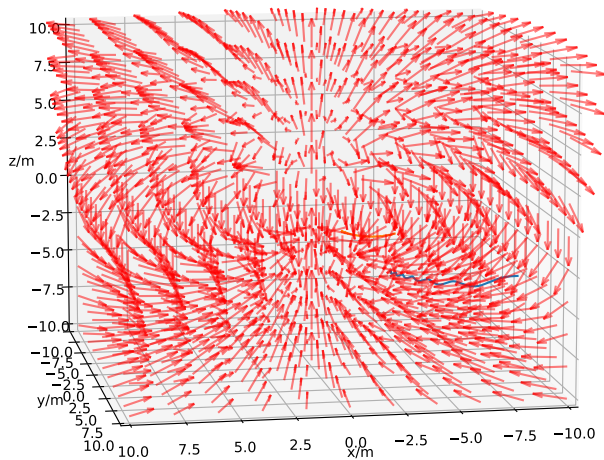
Another example, Aurora

- ▶ Solar wind (ions) vs. Earth Magnetic field.
- ▶ Why not my simulation here?
- ▶ Here is someone elses simulation
- ▶ Illustration originally from Nasa. Published on wikipedia, in Public Domain (Cropped to fit page)

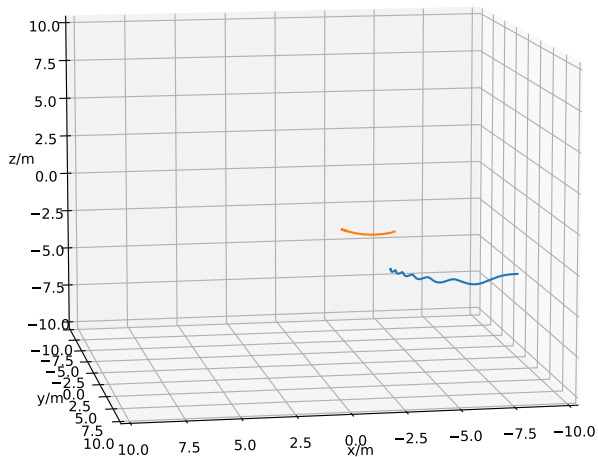


Frame from animation of the Earth magnetic field during increased activity, by N. A. Tsyganenko, Published under the GNU General Public License V. 3

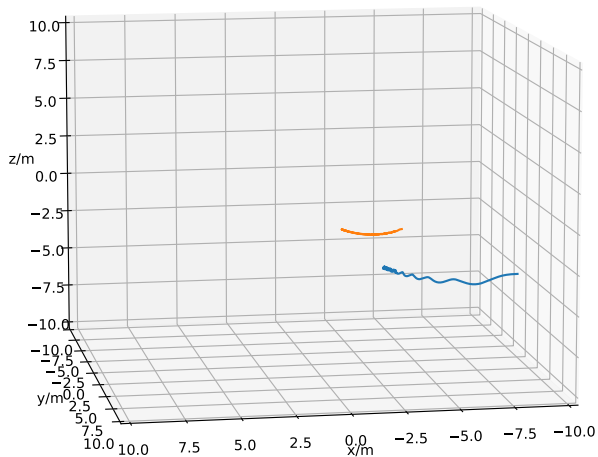
Aurora, bad simulation



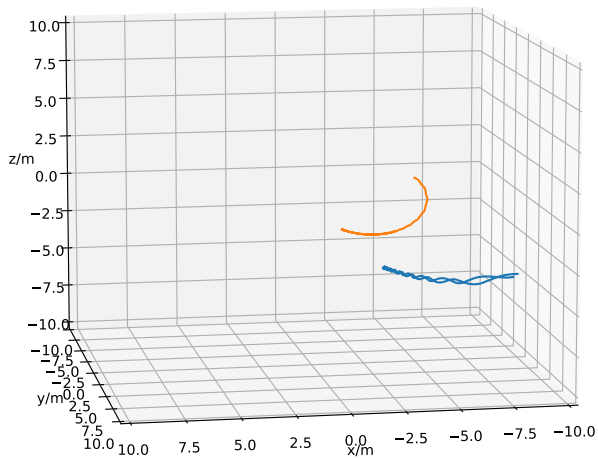
Aurora, bad simulation



Aurora, bad simulation



Aurora, bad simulation



Introducing Electric fields (10 min)

Introduction (3 minutes)

Theory and background (13 minutes)

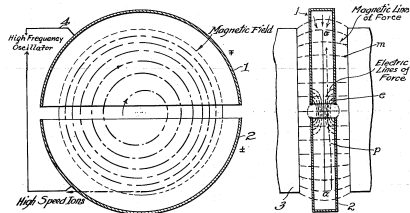
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Conclusion and Questions (5 min)

Electric fields

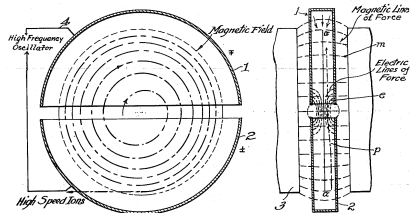
- Electric forces do work.



Ernest O. Lawrence, 1934, U.S. Patent 1,948,384; image in Public Domain.

Electric fields

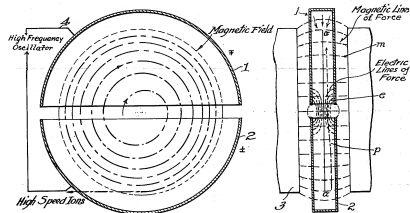
- ▶ Electric forces do work.
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Electric fields

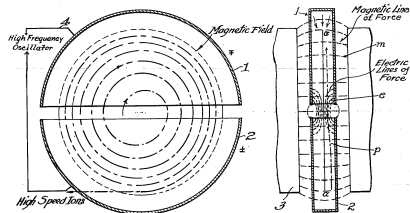
- ▶ Electric forces do work.
- ▶ Can be used in particle accelerators.
- ▶ Practical example, the Cyclotron.



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Public Domain.

Electric fields

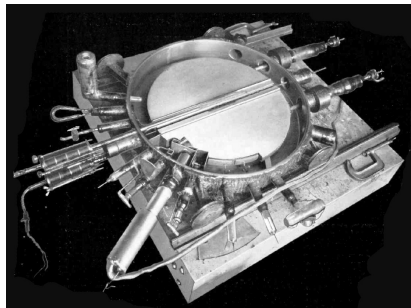
- ▶ Electric forces do work.
- ▶ Can be used in particle accelerators.
- ▶ Practical example, the Cyclotron.
- ▶ Single gap, oscillating field.
- ▶ Uses Cyclotron frequency



Ernest O. Lawrence, 1934, U.S. Patent 1,948,384; image in Public Domain.

Simulating the Cyclotron

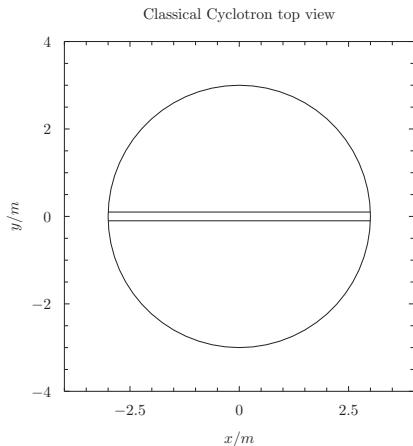
- I test same \vec{B} as before (around 6 mT), so $T \approx 10 \mu\text{s}$



Photography of Lawrence's 27-inch proof of concept Cyclotron, in Public Domain.

Simulating the Cyclotron

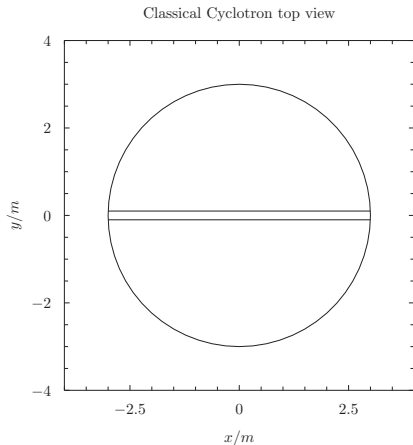
- ▶ I test same \vec{B} as before (around 6 mT), so $T \approx 10 \mu\text{s}$
- ▶ $E_{\text{max}} = 1 \text{ kV/m}$ (For display purposes), gap 20 cm, radius 3 m.



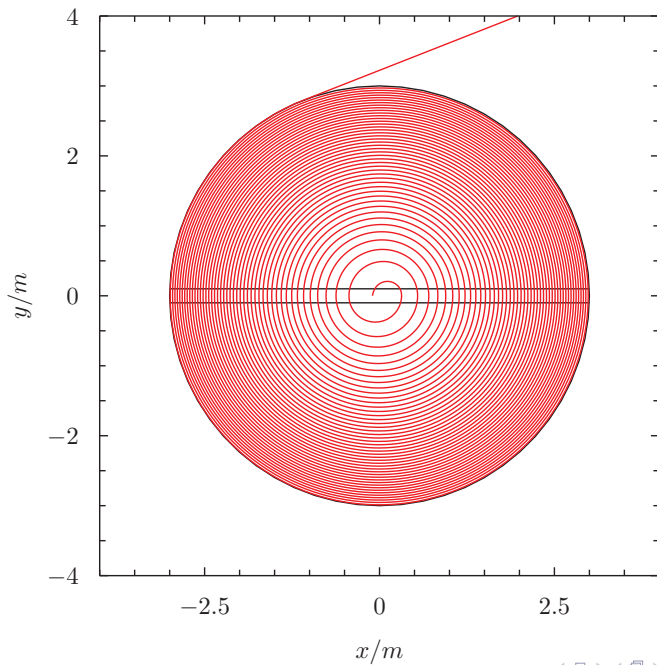
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- ▶ $E_{\text{max}} = 1 \text{ kV/m}$ (For display purposes), gap 20 cm, radius 3 m.
- ▶ Energy depends only on the radius and \vec{B} field:

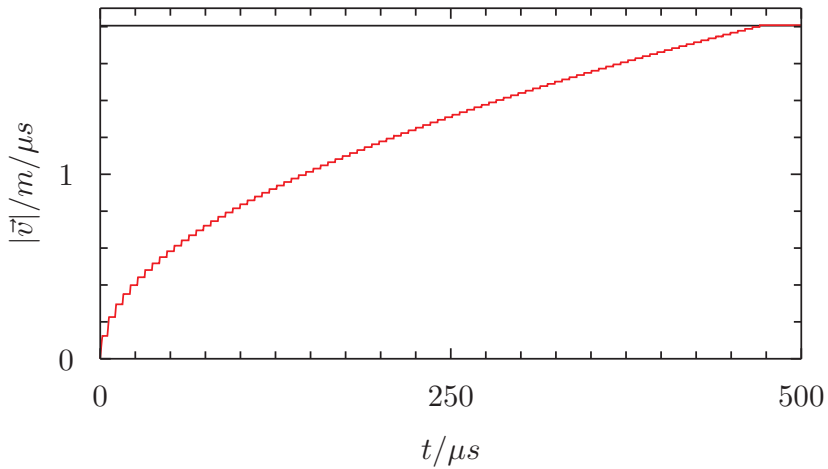
$$\frac{R|q|B}{m} = v_{\perp}$$



Classical Cyclotron top view



speed over time



— Theoretical speed limit — speed

Limitations of the Simulation

- Key points changed by relativity. $\vec{v}m \rightarrow \gamma\vec{v}m$, for $\gamma = 1/\sqrt{1 - v^2/c^2}$.

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- ▶ Cyclotron needs to be synchronized.
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- ▶ Simulating the Electric and Magnetic field would be better.

Questions