Problem Set 10

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Problem 1:

Solution:

Here we consider a 1-dimensional scattering problem where a particle scatters with a reflection and transmission coefficient r and t. These coefficients satisfy the condition that $|t|^2 + |r|^2 = 1$. Hence, we can can derive a 1D version of the optical theorem.

(a) Now given $f(\theta = \pi) = r$ and $f(\theta = 0) = t - 1$. These are given by,

$$f(\theta = 0) = \langle \theta = 0 | T | \theta = 0 \rangle = 1 - \langle \theta = 0 | S | \theta = 0 \rangle 1 - t$$
$$f(\theta = \pi) = \langle \theta = \pi | T | \theta = 0 \rangle = -r$$

Hence,

$$|f(\theta = \pi)|^2 = |r|^2$$
$$|f(\theta = 0)|^2 = (1 - t)(1 - t^*) = 1 - t^* - t - t^* + |t|^2$$

Now if we add these probabilities $|f(\theta=\pi)|^2+|f(\theta=0)|^2$ we get

$$|f(\theta=0)|^2 + |f(\theta=\pi)|^2 = 2 - t - t^* = 2Re[f(\theta=0)]$$

This gives us the outcomes for 1-t and $1-t^*$,

$$1 - t = Re[f(\theta = 0)] + iIm[f(\theta = 0)]$$

$$1 - t^* = Re[f(\theta = 0)] - iIm[f(\theta = 0)]$$

which we obtained from the above expression. These above expressions relate $|f(\theta)|^2$ to $Im(f(\theta = 0))$ and $Re(f(\theta = 0))$ giving a relationship for the 1D optical theorem.

(b) We can now verify the optical theorem for a spherically symmetric potential in terms of δ_l . We start with our expression from the partial waves expansion,

$$f(\theta) = \frac{1}{k} \sum_{l} (2l+1)e^{i\delta_l} \sin \delta_l P_l(\cos \theta)$$

Now evaluating this at $\theta = 0$,

$$f(\theta = 0) = \frac{1}{k} \sum_{l} (2l+1)e^{i\delta_{l}} \sin \delta_{l} P_{l}(\cos(\theta = 0))$$
$$= \frac{1}{k} \sum_{l} (2l+1)e^{i\delta_{l}} \sin \delta_{l}$$

We are only interested in the imaginary part of this scattering amplitude,

$$Im f(\theta = 0) = \frac{1}{k} \sum_{l} (2l+1) \sin^2 \delta_l$$

This will be the left hand side of the expression. Evaluating the right hand side requires finding the total cross section,

$$\int d\Omega |f(\theta)|^2 = \int_0^{2\pi} d\phi \int_0^{\pi} d\theta |f(\theta)|^2$$
$$= \frac{2\pi}{k^2} \int_{-1}^1 d(\cos\theta) \sum_{l,l'} (2l+1)(2l'+1)e^{i(\delta_l - \delta_{l'})} P_l(\cos\theta) P_l'(\cos\theta) \sin\delta_l \sin\delta_{l'}$$

Recall the orthonormality condition that $\int_{-1}^{1} P_l(x) P_{l'}(x) = \frac{2\delta_{l,l'}}{2l+1}$. Reducing to,

$$\frac{2\pi}{k^2} \sum_{l,l'} (2l+1)(2l'+1)e^{i(\delta_l - \delta_{l'})} \frac{2\delta_{l,l'}}{2l+1} \sin \delta_l \sin \delta_{l'} = \frac{4\pi}{k^2} \sum_{l} (2l+1) \sin^2 \delta_l$$

This gives us,

$$\frac{4\pi}{k}Imf(\theta=0) = \int d\Omega |f(\Omega)|^2$$

Problem 2:

Solution:

Here we compute the lifetime of a 2p state of Hydrogen which undergoes spontaneous emission of a photon. For instance, we can consider the transition from $2p_z$ to the 1s state. This requires employing Fermi's golden rule, given by

$$\frac{1}{\tau} = \frac{2\pi}{\hbar} \sum_{\epsilon,\lambda} |\langle \phi_{2p_z} n_{\lambda,\epsilon} + 1 = 1 | V | \phi_{2p_z}, n_{\lambda,\epsilon} = 0 \rangle|^2 \delta(E_{2p} - E_{1s} - \hbar \omega_{\epsilon})$$

Here, V is simply a perturbative potential corresponding to the coupling of the quantized electromagnetic field to the atom. The coupling will have the form $V = -\frac{e}{mc} \mathbf{A} \cdot \mathbf{P}$. **A** is the gauge of the field and **P** is the momentum operator for the atom. Here **A** will take the following form,

$$\mathbf{A}(\mathbf{r}) = \sum_{\lambda, \mathbf{k}} \sqrt{\frac{2\pi\hbar c^2}{\Omega\omega_k}} \hat{\epsilon}_{k,\lambda} (a_{k,\lambda} + a_{-k,\lambda}^{\dagger}) e^{i\mathbf{k}\cdot\mathbf{r}}$$

which results in the following coupling potential (Ω here is the volume).

$$V = -\frac{e}{m} \sum_{k,\lambda} \sqrt{\frac{2\pi\hbar}{\Omega\omega_k}} \hat{\epsilon}_{k,\lambda} \cdot \mathbf{p} e^{i\mathbf{k}\cdot\mathbf{r}} (a_{k,\lambda} + a_{-k,\lambda}^{\dagger})$$

Employing the dipole approximation and the fact that we are only considering stimulated emission, this would give us,

$$V = -\frac{e}{m} \sum_{k,\lambda} \sqrt{\frac{2\pi\hbar}{\Omega\omega_k}} \hat{\epsilon}_{k,\lambda} \cdot \mathbf{p} a_{-k,\lambda}^{\dagger}$$

We are now ready to compute the corresponding matrix element.

$$\langle \phi_{1s} n_{\lambda,\epsilon} + 1 = 1 | V | \phi_{2p_z}, n_{\lambda,\epsilon} = 0 \rangle = -\frac{e}{m} \sum_{k,\lambda} \sqrt{\frac{2\pi\hbar}{\Omega\omega_k}} \langle \phi_{1s} | \hat{\epsilon}_{k,\lambda} \cdot \mathbf{p} | \phi_{2p_z} \rangle \langle n_{\lambda,\epsilon} + 1 = 1 | a_{-k,\lambda}^{\dagger} | n_{\lambda,\epsilon} = 0 \rangle$$
$$= -\frac{e}{m} \sum_{k,\lambda} \sqrt{\frac{2\pi\hbar}{\Omega\omega_k}} \langle \phi_{1s} | \hat{\epsilon}_{k,\lambda} \cdot \mathbf{p} | \phi_{2p_z} \rangle$$

Let's expand the matrix element involving $\hat{\epsilon}_{k,\lambda} \cdot \mathbf{p}$. We know from the Heisenberg equation of motion that, $\mathbf{p} = \frac{m}{i\hbar} [\mathbf{r}, \mathcal{H}_0]$. Hence,

$$\hat{\epsilon}_{k,\lambda} \cdot \mathbf{p} = \frac{m}{i\hbar} [(\hat{\epsilon}_{k,\lambda} \cdot \mathbf{r}) \mathcal{H}_0 - \mathcal{H}_0 (\hat{\epsilon}_{k,\lambda} \cdot \mathbf{r})]$$

Hence, the matrix element we evaluate is

$$\langle \phi_{1s} | \hat{\epsilon}_{k,\lambda} \cdot \mathbf{p} | \phi_{2p_z} \rangle = \frac{m}{i\hbar} (E_{2p} - E_{1s}) \langle \phi_{1s} | \hat{\epsilon}_{k,\lambda} \cdot \mathbf{r} | \phi_{2p_z} \rangle$$
$$= \frac{m}{i\hbar} (E_{2p} - E_{1s}) (\hat{\epsilon}_{q,\lambda} \cdot \hat{z}) \langle \phi_{1s} | z | \phi_{2p_z} \rangle$$

Now evaluating the matrix element,

$$\langle \phi_{1s}|z|\phi_{2p_z}\rangle = \frac{1}{\sqrt{\pi}a_0^{3/2}} \frac{1}{4\sqrt{2\pi}a_0} \int_0^\infty dr r^2 \int_0^{2\pi} d\phi \int_0^\pi d\theta \sin\theta e^{-r/a_0} (re^{-r/2a_0}) \cos\theta (r\cos\theta)$$

$$=\frac{2\pi}{4\sqrt{2}\pi a_0^{5/2}}\int_0^\infty dr r^4 e^{-3r/a_0}\int_0^\pi d\theta\cos^2\theta\sin\theta=\frac{2\pi}{4\sqrt{2}\pi a_0^{5/2}}(\frac{8a_0^5}{81})(\frac{2}{3})=\frac{\sqrt{2}128}{243}a_0$$

To evaluate the total rate, we can work in the continuous limit by converting the discrete sum over \mathbf{q} states to an integral

$$\frac{1}{\tau} = \sum_{\lambda} \int_{\Omega_{q}} d^{3}\mathbf{q} \left[\frac{e^{2}}{2\pi\hbar^{2}cq} (E_{2p} - E_{1s})^{2} (\hat{\epsilon} \cdot \hat{z})^{2} |\langle \phi_{1s}|z|\phi_{2p_{z}} \rangle|^{2} \delta(E_{2ps} - E_{1s} - \hbar qc) \right]$$

$$= \sum_{\lambda} \int_{0}^{2\pi} d\phi \int_{0}^{\pi} d\theta \sin\theta (\hat{\epsilon}_{k,\lambda} \cdot \hat{z})^{2} \left[\frac{e^{2} (E_{2p} - E_{1s})^{3}}{2\pi\hbar^{4}c^{3}} |\langle \phi_{1s}|z|\phi_{2p} \rangle|^{2} \right]$$

$$= 2\pi \int_{0}^{\pi} d\theta \sin^{3}\theta \left[\frac{e^{2} (E_{2p} - E_{1s})^{3}}{2\pi\hbar^{4}c^{3}} |\langle \phi_{1s}|z|\phi_{2p} \rangle|^{2} \right]$$

$$= \frac{8\pi}{3} \left[\frac{e^{2} (E_{2p} - E_{1s})^{3}}{2\pi\hbar^{4}c^{3}} |\langle \phi_{1s}|z|\phi_{2p} \rangle|^{2} - \frac{4}{3} \left[\frac{e^{2} (E_{2p} - E_{1s})^{3}}{\hbar^{4}c^{3}} \frac{2(128)^{2}}{243^{2}} \right] = \frac{8(128)^{2} (E_{2p} - E_{1s})^{3}}{3(243)^{2}\hbar^{4}c^{3}} = (\frac{2}{3})^{8} \frac{me^{10}}{\hbar^{6}c^{3}}$$

Here, $E_{1s} = -\frac{m_e e^4}{2(4\pi\epsilon_0)\hbar^2}$ and $E_{2p} = -\frac{m_e e^4}{8(4\pi\epsilon_0)\hbar^2}$. This gives us an expression now for the transition time τ . Evaluating this gives us the numerical result that $\tau \sim 1.6$ ns.

Problem 3:

Solution:

Now consider a gas of atoms in equilibrium where the population of atoms n is given by $n=e^{-E_n/kT}$. Given that the energy of photons we consider is $\hbar\omega=E_n-E_m$, we can determine a relation for the population of photons. We know that during emission, we have a transition from the Fock state $|n_q\rangle$ to $|n_q+1\rangle$. Hence, the transition rate will go as

$$|\langle n_q + 1|a + a^{\dagger}|n_q\rangle|^2 = n_q + 1$$

Giving an emission rate,

$$\frac{1}{\tau_{emiss}} = (n_q + 1)A$$

Then for absorption, we have the opposite process given by

$$|\langle n_q - 1|a + a^{\dagger}|n_q\rangle|^2 = n_q$$

Hence, our rates are

$$\frac{1}{\tau_{emiss}} = (n_q + 1)A$$
$$\frac{1}{\tau_{abs}} = n_q A$$

Giving us a relationship for the photon numbers in states n and m (should be equal rates),

$$\frac{N_n}{\tau_{emiss}} = \frac{N_m}{\tau_{abs}}$$

Where $N_n = e^{-E_n/kT}$ and $N_m = e^{-E_m/kT}$.