# Problem Set 5

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March 22, 2021

## Problem 1:

Solution:

Here we consider the condition,

$$\int_{x_1}^{x_2} dx \sqrt{2m(E - V(x))} = \hbar \pi (n + \frac{1}{2})$$

using the potential  $V(x) = \frac{1}{2}\omega x^2$ .

$$\int_{x_1}^{x_2} dx \sqrt{2m(E - \frac{1}{2}m\omega x^2)} =$$

$$\frac{1}{2}\sqrt{m(2E - m\omega^2 x^2)} \left[x + \frac{2E}{\omega\sqrt{2E - m\omega^2 x^2}} \arctan(\frac{\sqrt{m\omega}x}{\sqrt{2E - m\omega^2 x^2}})\right]|_{x=x_1}^{x=x_2} = \pi \frac{E}{\omega}$$

For the Harmonic oscillator, we have turning points at  $x_1 = -\sqrt{\frac{2E}{m\omega^2}}$  and  $x_2 = \sqrt{\frac{2E}{m\omega^2}}$ . Hence, we get

$$\pi\hbar(n+\frac{1}{2}) = \pi\frac{E_n}{\omega}$$

This gives us the result,

$$E_n = \hbar\omega(n + \frac{1}{2})$$

## Problem 2:

Solution:

Here, we can imagine a 3d particle that's bound to a spherically symmetrical potential. The wavefunctions have the form,  $R_l(r) = u_l(r)/r$ . We can start with the Schrodinger equatiton,

$$\left(-\frac{\hbar^2}{2m}\frac{\partial^2}{\partial r^2}r + \frac{\hbar^2 l(l+1)}{2mr^2} + V(r)\right)R_l(r) = ER_l(r)$$

We only care about the s states here, so we take l = 0. The equation will reduce to,

$$\left(-\frac{\hbar^2}{2m}\frac{\partial^2}{\partial r^2} + V(r)\right)u(r) = Eu(r)$$

Now, near the classical turning point, we can approximate the potential as a linear potential

$$V(r) = E - \alpha(r - a)$$

with  $\alpha$  as a constant. Hence,

$$\left(-\frac{\hbar^2}{2m}\frac{\partial^2}{\partial r^2} + E - \alpha(r-a)\right)u(r) = Eu(r)$$

$$\left(-\frac{\hbar^2}{2m}\frac{\partial^2}{\partial r^2} - \alpha(r-a)\right)u(r) = 0$$

we can let  $r_0 = (\frac{\hbar^2}{2m\alpha})^{1/3}$  and  $x = \frac{r-a}{r_0}$ . Hence, the equation reduces to

$$\left(\frac{\partial^2}{\partial x^2} + x\right)u(x) = 0$$

Allowing us to solve this equation using the boundary condition that  $u(x = -\frac{a}{r_0}) = 0$ . We are applying this boundary condition since we cannot allow the wave function R(r) to explode at r = 0. Hence, u = 0 at this point to avoid divergence. The solution to this equation is given by,

$$u(x) = c_1 Ai((-1)^3 x) + c_2 Bi((-1)^3 x)$$

with Ai and Bi as Airy functions and  $c_1$  and  $c_2$  as coefficients determined from the boundary conditions of the problem. However, we want this in

analytical form. Hence, we can transform the solution u to a Bessel function,  $u(x) = x^{1/2} f(\frac{2}{3} x^{3/2}) = x^{1/2} f(z)$  with f. This allows us to transform the expression to the n = 1/3 Bessel equation,

$$\frac{\partial^2 f}{\partial z^2} - \frac{1}{z} \frac{\partial f}{\partial z} + \left(1 - \frac{(1/3)^2}{z^2}\right) f = 0$$

This has the solution,

$$f(z) = c_1 J_0(z)$$

with,

$$J_0(z) = \sqrt{\frac{2}{\pi z}}\cos(z - \frac{\pi}{4})$$

Hence, the solution will be

$$u(x) = x^{1/2} \sqrt{\frac{3}{\pi x^{3/2}}} \cos(\frac{2}{3}x^{3/2} - \frac{\pi}{4})$$
$$= \sqrt{\frac{3}{\pi}} x^{-3/2} \cos(\frac{2}{3}x^{3/2} - \frac{\pi}{4})$$

Expanding out gives,

$$u_{\pm}(x) = \frac{2}{\sqrt{\kappa(x)}} \left( \exp\left(\frac{i}{\hbar} \int_{-\pi}^{x} \kappa(x') dx' - \frac{\pi}{4}\right) + \exp\left(\frac{-i}{\hbar} \int_{-\pi}^{x} \kappa(x') dx' - \frac{\pi}{4}\right) \right)$$

Allowing us to substitute in  $x = \frac{r-a}{r_0}$ .

$$u_I(r) = \sqrt{\frac{3}{\pi}} \sqrt{\frac{r_0^3}{(r-a)^3}} \left( \exp\left(\frac{i}{\hbar} \int_r^a \kappa(r') dr' - \frac{\pi}{4}\right) + \exp\left(-\frac{i}{\hbar} \int_r^a \kappa(r') dr' - \frac{\pi}{4}\right) \right)$$

which is our solution for the side that's left of the turning point. To the right of the turning point, we have

$$u_{II} \sim \frac{1}{\sqrt{\kappa(r)}} \exp(-\int_a^r dr' \kappa(r'))$$

We have that u(r=0)=0. To meet our condition that u(0)=0, we must have that

$$\frac{2\pi}{h} \int_0^a \kappa(r')dr' - \frac{\pi}{4} = n + \frac{\pi}{2}$$

with n as an integer. This expression simplifies to,

$$2\int_0^a \kappa(r')dr' = 2\int_0^a \sqrt{2m(E - V(r))}dr' = (n + \frac{3}{4})h$$

which is the quantization condition for bound s states. Now for a potential that goes as  $r^{-p}$ , we have

$$2\lim_{r\to\infty} \int_0^a \sqrt{2m(E-r^{-p})} dr = 2\int_0^a \lim_{r\to\infty} \sqrt{2m(E-r^{-p})} dr = 2\int_0^a \sqrt{2mE} dr = 2\sqrt{2mE} = (n+\frac{3}{4})h$$

so we can see that there are an infinite number of states here. Hence, the quantization condition is satisfied here. Now, if p > 2 we get that near r = 0, the potential will blow up rapidly. Hence, the integral will diverge at 0. Hence, the quantization condition will not be satisfied.

#### Problem 3:

Solution:

Here, we can write the Hamiltonian of the system as

$$\mathcal{H} = \frac{p^2}{2m} - \lambda \delta(x)$$

At t=0, the potential suddenly changes to  $\mu$ . We can now determine the probability of transition from the old ground state to the new ground state. We know that the initial ground state has the form,

$$\psi(x, t = -\epsilon) = \begin{cases} \sqrt{\frac{-m\lambda}{\hbar^2}} e^{\frac{-m\lambda}{\hbar^2}x} & x < 0\\ \sqrt{\frac{-m\lambda}{\hbar^2}} e^{\frac{m\lambda}{\hbar^2}x} & x > 0 \end{cases}$$

the new ground state is given by,

$$\psi(x,t=+\epsilon) = \begin{cases} \sqrt{\frac{-m\mu}{\hbar^2}} e^{\frac{-m\mu}{\hbar^2}x} & x < 0\\ \sqrt{\frac{-m\mu}{\hbar^2}} e^{\frac{m\mu}{\hbar^2}x} & x > 0 \end{cases}$$

Hence, the probability of transition from the old to new state is given by  $\langle \psi(x, t = -\epsilon) | \psi(x, t = +\epsilon) \rangle$ . We can evaluate this for the left and right hand side. For x < 0, we have

$$\langle \psi_0(x, t = -\epsilon) | \psi_0(x, t = +\epsilon) \rangle = \sqrt{\frac{m\lambda}{\hbar^2}} \sqrt{\frac{m\mu}{\hbar^2}} \int_{-\infty}^0 e^{\frac{-m(\lambda + \mu)}{\hbar^2} x} dx = -\frac{\sqrt{\mu\lambda}}{\lambda + \mu}$$

Then for x > 0,

$$\langle \psi_0(x, t = -\epsilon) | \psi_0(x, t = +\epsilon) \rangle = -\frac{\sqrt{\mu \lambda}}{\lambda + \mu}$$

Hence, for combining both sides we have the probability amplitude:  $\langle \psi_0(x, t = -\epsilon) | \psi_0(x, t = +\epsilon) \rangle = -\frac{2\sqrt{\mu\lambda}}{\lambda+\mu}$ . Giving us probability,

$$|\langle \psi_0(x, t = -\epsilon)|\psi_0(x, t = +\epsilon)\rangle|^2 = \frac{4\mu\lambda}{(\lambda + \mu)^2}$$

For  $\mu = \lambda/2$ , we have

$$|\langle \psi_0(x, t = -\epsilon) | \psi_0(x, t = +\epsilon) \rangle|^2 = \frac{2\lambda^2}{(3/2\lambda)^2} = \frac{8}{9}$$

Let  $\gamma = \mu/\lambda$ . Hence, our total probability expression becomes

$$|\langle \psi_0(x, t = -\epsilon)|\psi_0(x, t = +\epsilon)\rangle|^2 = \frac{4\gamma}{(1+\gamma)^2} \sim \frac{1}{\gamma}$$

Hence, for  $\gamma = \mu/\lambda >> 1$  the transition probability will be very small. Conversely, for  $\gamma = \mu/\lambda << 1$  the transition probability will be very large.

# Problem 4:

Solution:

Considering now a spin 1/2 particle with magnetic moment  $\mu$  in a magnetic field,

$$\mathbf{B}(t) = B_0(\cos\theta(t), \sin\theta(t), 0)$$

such that,

$$\theta(t) = \frac{\pi}{2} \frac{1}{1 + e^{t/t_0}}$$

The Hamiltonian of the system is given by,  $\mathcal{H} = -\mu \cdot \mathbf{B}$ . This comes out to (general case),

$$\mathcal{H} = -\frac{\mu}{2}(\cos\theta(t)\sigma_x + \sin\theta(t)\sigma_y)$$

Early on or when  $t \to -\infty$ , we have that the dominant term of the Hamiltonian corresponds to  $\sigma_y$  (with  $\theta = \pi/2$  in this limit). For later times or when  $t \to +\infty$ , the dominant part of the Hamiltonian corresponds to  $\sigma_y$ . Hence,

$$\mathcal{H}(t=-\infty) = -\frac{1}{2}\mu B_0 \sigma_y$$

$$\mathcal{H}(t=+\infty) = -\frac{1}{2}\mu B_0 \sigma_x$$

Now for each case, we can look at the limit of large  $t_0$  and small  $t_0$  corresponding to the adiabatic and sudden approximations that will be used.

(i) We can now find the probability as a function of time for the particle to be in its ground state for  $t >> t_0$ . Here, the particle is in state  $|y; -\rangle$  but the ground state is  $|x; -\rangle$ . For large  $t_0$  (adiabatic), Hence, in the adiabatic limit we expect the particle to stay in the ground state with probability P = 1. In the sudden limit, we can the probability  $|\langle -; y|x; -\rangle|^2$  We find,

$$\langle -; y | x; - \rangle = \frac{1}{2} (1, -1)^T (1, -i) = \frac{1+i}{2}$$

Hence,  $\langle -; y | x; - \rangle|^2 = \frac{1}{2}$ .

(ii) We can now compute the expected value  $\langle \mathbf{S}(t) \rangle$  for  $t >> t_0$ . We start with the wave function of the system given by,

$$|\psi(t)\rangle = e^{-\frac{i}{\hbar}E_{+}t}|x;+\rangle\langle+;x|y;+\rangle + e^{-\frac{i}{\hbar}E_{-}t}|x;-\rangle\langle-;x|y;+\rangle$$
$$= \frac{1}{\sqrt{2}}(e^{i\pi/4}e^{i\pi/2\omega_{B}t}|x;+\rangle + e^{-i\pi/4}e^{-i\pi/2\omega_{B}t}|x;-\rangle)$$

with  $\omega_B = \frac{\mu B_0}{\hbar}$ .

$$=\frac{1}{\sqrt{2}}(ie^{\frac{i}{2}\omega_B t}|x;+\rangle+e^{-\frac{i}{2}\omega_B t}|x;-\rangle)$$

We're interested in computing  $\langle \psi | \sigma_x | \psi \rangle$ ,  $\langle \psi | \sigma_y | \psi \rangle$ ,  $\langle \psi | \sigma_z | \psi \rangle$ . Hence,

$$\langle \psi | \sigma_z | \psi \rangle = \frac{\hbar}{4} (-ie^{-i/2\omega_B t} \langle x; + | + e^{i/2\omega_B t} \langle x; - |)(ie^{i/2\omega_B t} | x; - \rangle + e^{-i/2\omega_B t} | x; + \rangle) \frac{\hbar i}{4} (e^{it/\omega_b} - e^{it/\omega_b}) =$$

$$= -\frac{\hbar}{2} \sin \omega_B t$$

For  $\langle \psi | \sigma_x | \psi \rangle$ , we get

$$\frac{1}{2}([ie^{-\frac{i}{2}\omega_B t}\langle x;+|+e^{\frac{i}{2}\omega_B t}\langle x;-|]\sigma_x[ie^{\frac{i}{2}\omega_B t}|x;+\rangle+e^{-\frac{i}{2}\omega_B t}|x;-\rangle])$$

$$=\frac{1}{2}([ie^{-\frac{i}{2}\omega_Bt}\langle x;+|+e^{\frac{i}{2}\omega_Bt}\langle x;-|][ie^{\frac{i}{2}\omega_Bt}|x;+\rangle-e^{-\frac{i}{2}\omega_Bt}|x;-\rangle])=\frac{1}{2}$$

Then for  $\langle \psi | \sigma_y | \psi \rangle = 0$ . Hence,

$$\langle \mathbf{S} \rangle = \hbar(1/2, 0, -\frac{1}{2} \sin \omega_B t)$$

This is the result for the sudden limit. In the adiabatic limit, we simply have  $\langle \mathbf{S} \rangle = (0, \frac{\hbar}{2}, 0)$ .

(ii) For  $t_0$  in the sudden limit, we should compare the timescale to  $t_0 << \frac{\hbar}{\Delta E}$  which corresponds to the time it takes for the wave function to evolve from the y to x and  $\Delta E$  is the corresponding energy scale. Like wise, in the adiabatic limit we must have  $t_0 >> \frac{\hbar}{\Delta E}$ .

### Problem 5:

Solution:

(a)

Here, we take the Hamiltonian  $H(\mathbf{R}) = (X\sigma_x + Y\sigma_y + Z\sigma_z)$ . This can be expanded as,

$$H(\mathbf{R}) = \begin{pmatrix} Z & X - iY \\ X + iY & -Z \end{pmatrix} = -R \begin{pmatrix} \cos \theta & \sin \theta e^{-i\phi} \\ \sin \theta e^{i\phi} & -d \cos \theta \end{pmatrix}$$

This Hamiltonian has eigenvalues  $\pm \sqrt{x^2 + y^2 + z^2} = \pm |\mathbf{R}|$ . Hence, the ground state energy of the system is  $-|\mathbf{R}|$ . Its corresponding eigenvector is,

$$|0\rangle = \begin{pmatrix} \cos\theta/2\\ \sin\theta/2e^{i\phi} \end{pmatrix} e^{i\alpha(R,\theta,\phi)}$$

with  $\alpha$  as our phase.

(b) Now we can compute Berry's potential given by,

$$\mathbf{A}_0(\mathbf{R}) = i \langle \Psi_0(\mathbf{R}) | \nabla_{\mathbf{R}} \Psi_0(\mathbf{R}) \rangle$$

Hence, we need  $\nabla_{\mathbf{R}}\Psi_0(\mathbf{R})$ . Here,

$$\begin{split} \frac{\partial \Psi_0}{\partial R} \hat{r} &= i \begin{pmatrix} \cos \theta/2 \\ \sin \theta/2 e^{i\phi} \end{pmatrix} e^{i\alpha(R,\theta,\phi)} \frac{\partial \alpha}{\partial R} \\ & \frac{\partial \Psi_0}{\partial \theta} \hat{\theta} = \frac{1}{R} \frac{\partial \Psi_0}{\partial \theta} \\ &= \frac{1}{R} \begin{pmatrix} (\cos \theta/2) \frac{\partial \alpha}{\partial \theta} - \frac{1}{2} \sin \theta/2) \\ e^{i\phi} (\frac{1}{2} (\cos(\theta/2) + i \frac{\partial}{\partial \theta})) \end{pmatrix} e^{i\alpha} \\ & \frac{\partial \Psi_0}{\partial \phi} \hat{\phi} = \frac{1}{R \sin \theta} \frac{\partial \Psi_0}{\partial \phi} = \frac{i}{R \sin \theta} \left( \begin{pmatrix} \cos \theta/2 \frac{\partial \alpha}{\partial \phi} \\ i \sin(\theta/2) e^{i\phi} (1 + \frac{\partial \alpha}{\partial \phi}) \end{pmatrix} e^{i\alpha(R,\theta,\phi)} \right) \end{split}$$

We can now obtain the inner product  $\langle \Psi_0 | \nabla_{\mathbf{R}} \Psi_0 \rangle$ . Computing this for each component gives,

$$\langle \Psi_0 | \nabla_R \Psi_0 \rangle = i (\sin^2 \theta / 2 + \cos^2 \theta / 2) \frac{\partial \alpha}{\partial R} = i \frac{\partial \alpha}{\partial R}$$

For the  $\theta$  component, we have

$$\langle \Psi_0 | \nabla_{\theta} \Psi_0 \rangle = \cos^2(\theta/2) \frac{\partial \alpha}{\partial \theta} - \frac{1}{2} \cos(\theta/2) \sin(\theta/2) + \frac{1}{2} \cos(\theta/2) \sin(\theta/2) + i \sin^2(\theta/2) \frac{\partial \alpha}{\partial \theta}$$
$$= \frac{1}{R} \frac{\partial \alpha}{\partial \theta} (\cos^2(\theta/2) + i \sin^2(\theta/2))$$

Lastly, we can compute this for the  $\phi$  component. We get,

$$\langle \Psi_0 | \nabla_{\phi} \Psi_0 \rangle = \frac{i}{R \sin \theta} (\cos^2 \theta \frac{\partial \alpha}{\partial \phi} + i \sin^2 \theta / 2(1 + \frac{\partial \alpha}{\partial \phi}))$$

Hence, these three components define  $\mathbf{A} = i(\langle \Psi_0 | \nabla_R \Psi_0 \rangle, \langle \Psi_0 | \nabla_\theta \Psi_0 \rangle, \langle \Psi_0 | \nabla_\phi \Psi_0 \rangle)$  in spherical coordinates. Now to obtain the Berry curvature, we can compute the curl of  $\mathbf{A}$ . This is defined as,

$$\mathbf{V} = \nabla \times \mathbf{A} = \frac{1}{R \sin \theta} \left( \frac{\partial}{\partial \theta} (A_{\phi} \sin \theta) - \frac{\partial A_{\theta}}{\partial \phi} \right) \hat{\mathbf{R}} + \frac{1}{r} \left( \frac{1}{\sin \theta} \frac{\partial A_{R}}{\partial \phi} - \frac{\partial}{\partial R} (RA_{\phi}) \right) \hat{\theta} + \frac{1}{R} \left( \frac{\partial}{\partial R} (RA_{\theta}) - \frac{\partial A_{R}}{\partial \theta} \right) \hat{\phi}$$

With,

$$A_R = -\frac{\partial \alpha}{\partial R}$$

$$A_{\theta} = \frac{1}{R} \frac{\partial \alpha}{\partial \theta} (\cos^{2}(\theta/2) + i \sin^{2}(\theta/2))$$

$$A_{\phi} = -\frac{1}{R \sin \theta} (\cos^{2}(\theta/2) \frac{\partial \alpha}{\partial \phi} + \sin^{2}\theta/2(1 + \frac{\partial \alpha}{\partial \phi}))$$

Letting Mathematica do the heavy lifting to compute the above expression, we arrive at

$$\mathbf{V} = -\frac{\hat{R}}{2R^2}$$

(c) We can now find a point where  $\mathbf{A}_0$  is singular which occurs in the vector potential of a magnetic monopole. We can find a guage  $\alpha$  where  $|\Psi_0|$  is continuous everywhere except at X = Y = 0, Z > 0 (south pole). We can also do this for the north pole. For the south pole, we want  $\alpha = 0$ . The corresponding components of  $\mathbf{A}_0$  here will be (corresponding to  $\theta = \pi$ ),

$$A_R = 0$$

$$A_\theta = 0$$

$$A_\phi = -\frac{1}{R}$$

(d) We can now compute the Berry phase for a given Guage. Let's take  $\alpha=0$ .

$$\gamma = \oint_C \mathbf{A} \cdot d\mathbf{R}$$

such that we integrate over  $\phi$ .

$$\gamma = \int_0^{2\pi} \left( -\frac{1}{R\sin\theta} (\sin^2\theta/2) \right) d\phi$$
$$= -\frac{2\pi \sin^2\theta/2}{R\sin\theta}$$

This relates to the solid angle subtended by C, as it relates to the phase difference over the closed loop C for any chosen point  $R, \theta$ . This expression is only true for the guage that we chose. Hence, this quantitty is not guage invariant.