

# PHYS UN1602 Recitation Week 11 Worksheet

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## Problem 1

Consider the two oppositely traveling electric-field waves:

$$\vec{E}_1 = E_0 \cos(kz - \omega t)\hat{x} \text{ and } \vec{E}_2 = E_0 \cos(kz + \omega t)\hat{x}$$

The sum of these two waves is the standing wave  $2E_0 \cos(kz) \cos(\omega t)\hat{x}$ .

- Find the magnetic field associated with this standing electric wave by finding the  $\vec{B}$  fields associated with each of the above traveling  $\vec{E}$  fields, and then adding them.
- Find the magnetic field by instead using Maxwell's equations to find the  $\vec{B}$  field associated with the standing electric wave  $2E_0 \cos(kz) \cos(\omega t)\hat{x}$ .

*Solution.*

- The traveling  $\vec{B}$  fields must point in the  $\pm\hat{y}$  directions because they must be perpendicular to both the associated  $\vec{E}$  field and the direction of propagation, which is  $\pm\hat{z}$ . The magnitudes of the  $\vec{B}$  fields are  $E_0/c$ . The signs are determined by the fact that  $\vec{E} \times \vec{B}$  points in the direction of propagation. The two magnetic waves are therefore:

$$\vec{B}_1 = \frac{E_0}{c} \cos(kz - \omega t)\hat{y}$$

$$\vec{B}_2 = -\frac{E_0}{c} \cos(kz + \omega t)\hat{y}$$

The sum of these waves is  $\vec{B} = \frac{2E_0}{c} \sin(kz) \sin(\omega t)\hat{y}$ .

- We use the Maxwell equation  $\nabla \times \vec{E} = -\frac{\partial \vec{B}}{\partial t}$  to find  $\vec{B}$ . The curl of  $\vec{E} = 2E_0 \cos(kz) \cos(\omega t)\hat{x}$  is:

$$\nabla \times \vec{E} = \begin{vmatrix} \hat{x} & \hat{y} & \hat{z} \\ \frac{\partial}{\partial x} & \frac{\partial}{\partial y} & \frac{\partial}{\partial z} \\ 2E_0 \cos(kz) \cos(\omega t) & 0 & 0 \end{vmatrix} = -2kE_0 \sin(kz) \cos(\omega t)\hat{y}$$

Setting this equal to  $-\frac{\partial \vec{B}}{\partial t}$  gives (up to a constant):

$$\vec{B} = \frac{2kE_0}{\omega} \sin(kz) \sin(\omega t)\hat{y}$$

But we know that  $\omega/k = c$ , so we have the same formula for  $\vec{B}$  as derived in part a).

## Problem 2

We are familiar with the Lorentz boost in the following form (in the  $x$  direction):

$$\begin{aligned} ct' &= \gamma(ct - \beta x) \\ x' &= \gamma(x - \beta ct) \\ y' &= y \\ z' &= z \end{aligned}$$

where  $\beta = v/c$ . In this problem, we will write the Lorentz boost in a different form.

- a) Define a spacetime vector for an event as:

$$\mathbf{X} = \begin{pmatrix} ct \\ x \end{pmatrix}$$

Write the Lorentz transformation in matrix form using this 2-vector, i.e., express the transformation as:

$$\mathbf{X}' = \mathbf{\Lambda} \mathbf{X}$$

Give the  $2 \times 2$  matrix  $\mathbf{\Lambda}$  explicitly in terms of  $\gamma$  and  $\beta$ .

- b) Recall that  $\beta = v/c$ . Taylor expand  $\gamma = \frac{1}{\sqrt{1-\beta^2}}$  for small  $\beta$  (i.e.  $v \ll c$ ) and rewrite  $\mathbf{\Lambda}$  for this case. Show that this recovers the Galilean transformation, except for a small correction to  $t$ .
- c) Let's define a new variable called *rapidity*,  $\zeta$ , such that:

$$\beta = \tanh \zeta \text{ and } \gamma = \cosh \zeta$$

Show that  $\gamma\beta = \sinh \zeta$ .

- d) Using these identities, rewrite the Lorentz transformation matrix  $\mathbf{\Lambda}$  in terms of hyperbolic functions  $\cosh \zeta$  and  $\sinh \zeta$ .
- e) Compare the matrix you obtained to a standard 2D rotation matrix:

$$R(\theta) = \begin{pmatrix} \cos \theta & -\sin \theta \\ \sin \theta & \cos \theta \end{pmatrix}$$

How are they similar? How are they different?

- f) Show that under this transformation, the spacetime interval

$$s^2 = -(ct)^2 + x^2$$

is preserved, i.e.  $s'^2 = s^2$ . Compare this to the way distances are preserved under circular rotations in Euclidean space. What's the analogy?

*Solution.*

- a) From the Lorentz transformation:

$$\begin{aligned} ct' &= \gamma(ct - \beta x) \\ x' &= \gamma(x - \beta ct) \end{aligned}$$

Writing this in matrix form gives:

$$\mathbf{\Lambda} = \begin{pmatrix} \gamma & -\gamma\beta \\ -\gamma\beta & \gamma \end{pmatrix}$$

b)

$$\gamma = \frac{1}{\sqrt{1-\beta^2}} \approx 1 + \frac{1}{2}\beta^2 \dots$$

$$\Rightarrow \mathbf{\Lambda} = \begin{pmatrix} 1 + \frac{1}{2}\beta^2 & -\beta - \frac{1}{2}\beta^3 \\ -\beta - \frac{1}{2}\beta^3 & 1 + \frac{1}{2}\beta^2 \end{pmatrix} \approx \begin{pmatrix} 1 & -\beta \\ -\beta & 1 \end{pmatrix}$$

$$\begin{pmatrix} ct' \\ x' \end{pmatrix} = \mathbf{\Lambda} \begin{pmatrix} ct \\ x \end{pmatrix} \approx \begin{pmatrix} ct - \frac{v}{c}x \\ x - vt \end{pmatrix}$$

This is the same as a Galilean transformation except  $t' = t - \frac{v}{c^2}x$  instead of  $t$ , but for  $v \ll c$  this correction is virtually negligible.

c)

$$\gamma\beta = \cosh \zeta \tanh \zeta = \frac{\cosh \zeta \cdot \sinh \zeta}{\cosh \zeta} = \sinh \zeta$$

d)

$$\mathbf{\Lambda} = \begin{pmatrix} \cosh \zeta & -\sinh \zeta \\ -\sinh \zeta & \cosh \zeta \end{pmatrix}$$

e) They are very similar except for two differences: replacing trigonometric functions with their hyperbolic equivalents, and the bottom entry having a minus sign in the Lorentz transform case (as opposed to the rotation matrix having it be positive), making the Lorentz transform matrix fully symmetric.

f)

$$\begin{aligned} s'^2 &= -(\cosh \zeta ct - \sinh \zeta x)^2 + (-\sinh \zeta ct + \cosh \zeta x)^2 \\ &= -(\cosh^2 \zeta c^2 t^2 - 2 \cosh \zeta \sinh \zeta ct x + \sinh^2 \zeta x^2) \\ &\quad + (\sinh^2 \zeta c^2 t^2 - 2 \cosh \zeta \sinh \zeta ct x + \cosh^2 \zeta x^2) \\ &= -\cosh^2 \zeta c^2 t^2 + 2 \cosh \zeta \sinh \zeta ct x - \sinh^2 \zeta x^2 \\ &\quad + \sinh^2 \zeta c^2 t^2 - 2 \cosh \zeta \sinh \zeta ct x + \cosh^2 \zeta x^2 \\ &= (-\cosh^2 \zeta + \sinh^2 \zeta) c^2 t^2 + (\cosh^2 \zeta - \sinh^2 \zeta) x^2 \\ &= -c^2 t^2 + x^2 = s^2 \end{aligned}$$

In Euclidean space, circular rotations preserve:

$$r^2 = x^2 + y^2$$

This is preserved under rotation matrices using cos and sin. Similarly, Lorentz boosts preserve:

$$s^2 = -(ct)^2 + x^2$$

but using cosh and sinh. We can hence think of Lorentz boosts as "rotations" in spacetime, preserving a different (hyperbolic) type of distance — the spacetime interval.