5500 Problem Set 6 Nikko Cleri October 12, 2020

Question I.

Starting form HW5 P3: Calculate the grand partition function, and use it to derive Boyle's law pV = NkT.

From HW5 we can start by saying:

$$\mu = kT \ln n\lambda^{3}$$

$$= kT \ln \frac{N\lambda^{3}}{V}$$

$$Z_{N} = \frac{1}{N!} \left(\frac{V}{\lambda^{3}}\right)^{N}$$

$$\mathcal{Z} = \sum_{N} z^{N} Z_{N}$$

$$= \sum_{N} \frac{z^{N}}{N!} \left(\frac{V}{\lambda^{3}}\right)^{N}$$

$$= e^{\frac{zV}{\lambda^{3}}}$$

We will now do the following to find Boyle's Law:

$$\Omega - -kT \ln \mathcal{Z}$$

$$= -kT \left(\frac{zV}{\lambda^3}\right)$$

$$= -pV$$

$$pV = kT \left(\frac{zV}{\lambda^3}\right)$$

$$kT \frac{\partial}{\partial \mu} \log \mathcal{Z} = kT\beta z \frac{\partial}{\partial z} \ln e^{zV/\lambda^3}$$

$$N = \frac{zV}{\lambda^3}$$

$$\implies pV = NkT$$

gives us Boyle's Law.

Question II.

Consider an ultrarelativistic Fermi gas, where the dispersion relation for the particles is $\epsilon = cp~p$ being the momentum. Find the Fermi energy.

In the ultrarelativistic limit the relation $\epsilon_F = \frac{\hbar^2 k^2}{2m}$ no longer holds, and we have

$$\begin{split} \epsilon_F &= p_F c \\ p_F &= \hbar k \\ &= \hbar \left(\frac{6\pi^2 n}{g}\right)^{1/3} \\ \epsilon_F &= \hbar \left(\frac{6\pi^2 n}{g}\right)^{1/3} c \end{split}$$

Question III.

In a magnetic field the spin degeneracy of the electron gas is lifted: the magnetic energy for the z component of the spin m_S is $E = -g\mu_B B m_S$. Hence, in equilibrium the number of electrons in the spin-up and spin-down states is no longer the same, and the electron gas acquires a macroscopic magnetization M (magnetic moment/unit volume). Show that the resulting "Pauli paramagnetic susceptibility" $\chi = \frac{\partial M}{\partial B}$ at zero temperature and zero magnetic field is

$$\chi = \frac{g^2 \mu_B^2}{4V} \mathcal{D}(\epsilon_F) = \frac{3g^2 \mu_B^2 n}{8\epsilon_F}$$

Here $g \approx 2$ and \mathcal{D} is the density of one particle energy eigenstates

$$\mathcal{D}(\epsilon) = \frac{\partial \mathcal{N}}{\partial \epsilon}$$

 $\mathcal{N}(\epsilon)$, in turn, is the number of one-particle energy eigenstates with kinetic energy $\leq \epsilon$, as if there were no magnetic field at all.

Things to consider: This is obviously a case in which the different magnetic states are not degenerate. Moreover, since you are finding the susceptibility at zero magnetic field, an analysis up to the lowest nontrivial order in B suffices.

The energy of a particle in a magnetic field is given by

$$\epsilon = \frac{p^2}{2m} - \mu^* \cdot \mathbf{B}$$

The numbers of occupied energy levels in the two groups will be

$$N^{+} = \frac{4\pi V}{3h^{3}} (2m(\epsilon_{F} + \mu^{*}B))^{3/2}$$
$$N^{-} = \frac{4\pi V}{3h^{3}} (2m(\epsilon_{F} - \mu^{*}B))^{3/2}$$

and the magnetic oment is given by:

$$\begin{split} M &= \mu_B (N^+ - N) \\ &= \frac{4\pi \mu_B V (2m)^{3/2}}{3h^3} ((\epsilon_F + \mu_B B)^{3/2} - (\epsilon_F - \mu_B B)^{3/2}) \\ \chi &= \lim_{B \to 0} \left(\frac{M}{VB}\right) \\ &= \frac{4\pi \mu_B^2 (2m)^{3/2} \epsilon_F^{1/2}}{h^3} \end{split}$$

Applying the known expression for the Fermi energy gives the desired answer.

$$\chi = \frac{3g^2 \mu_B^2 n}{8\epsilon_F}$$