

# CONFINING HIDDEN SECTORS AT THE LHC AND BEYOND



**Carleton**  
**University**

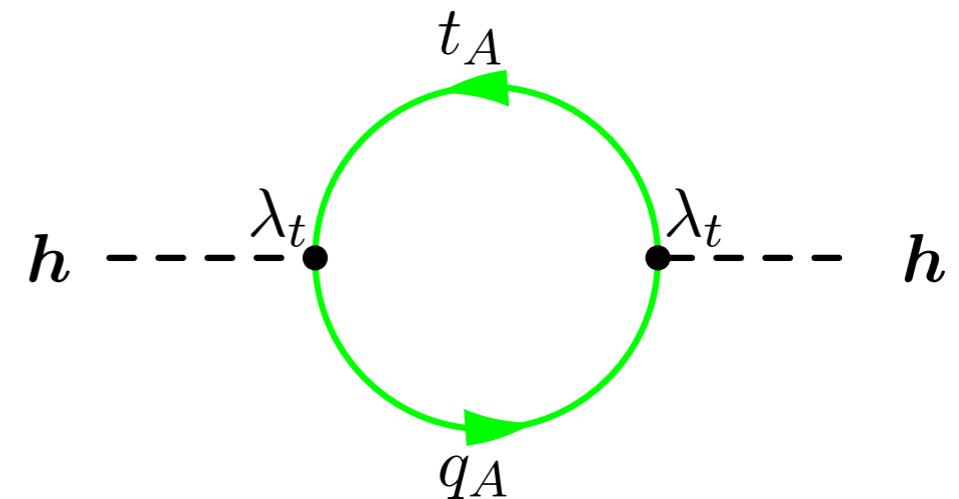
Department of Physics

In collaboration with Pedro  
Schwaller, Andreas Weiler, Dylan  
Linthorne, Wafia Bensalem, and  
Paul Archer-Smith.

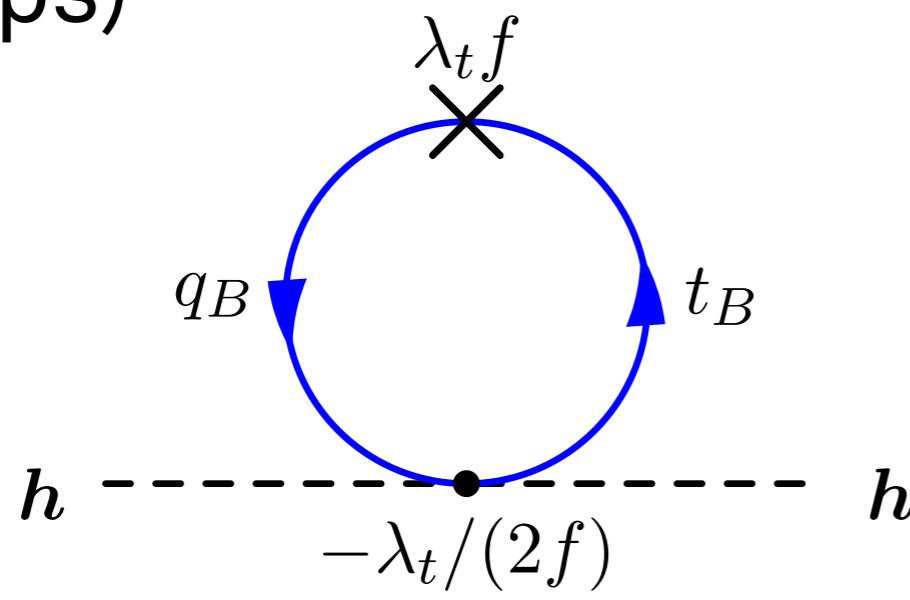
DANIEL STOLARSKI

# MOTIVATION 1

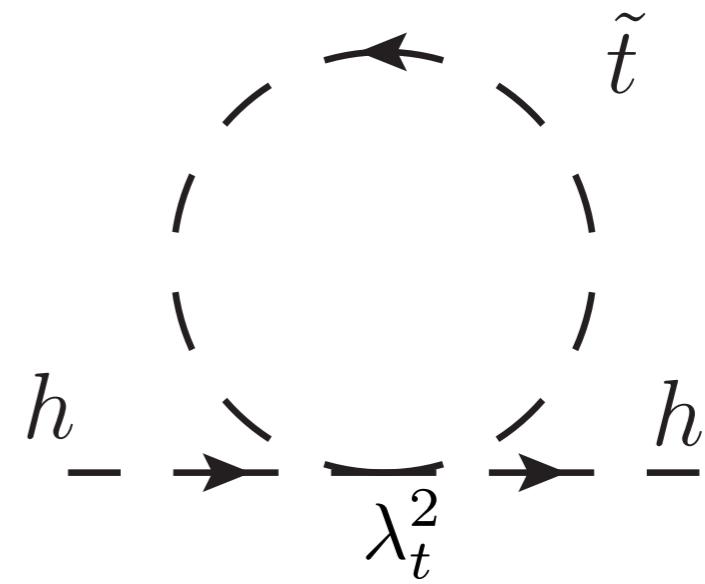
Gauge hierarchy problem:



Solved in composite Higgs (SUSY) with top-partners  
(stops)

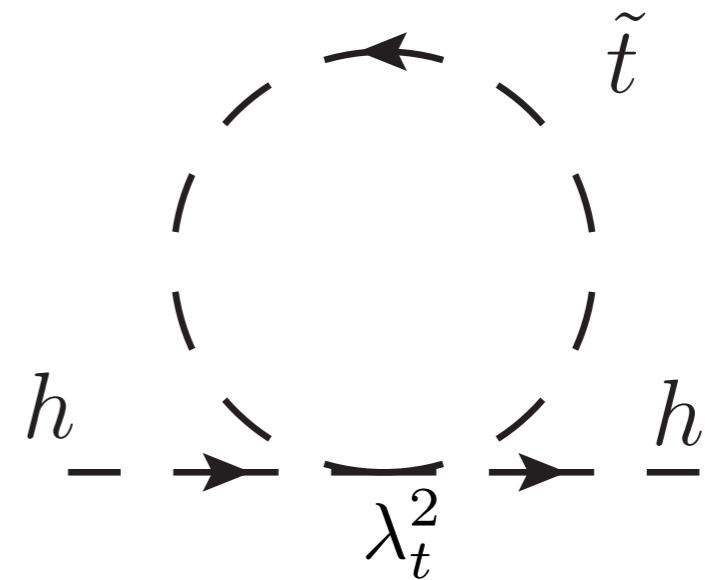
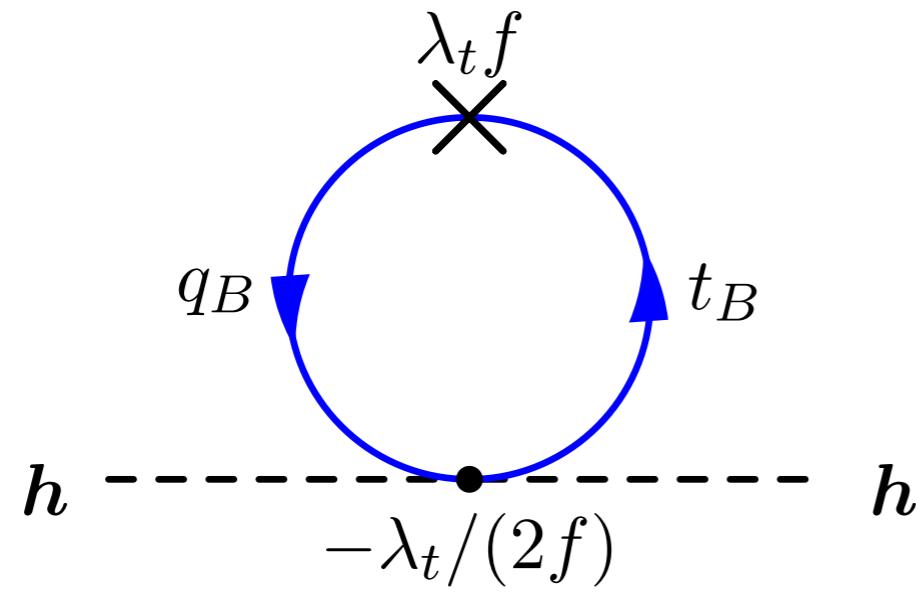


or



# TWIN HIGGS/FOLDED SUSY

No! But still need factor of 3.



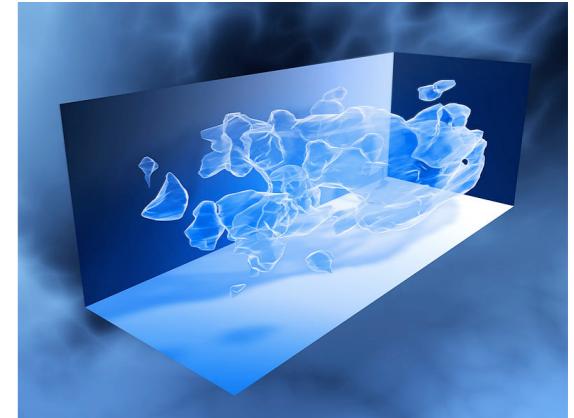
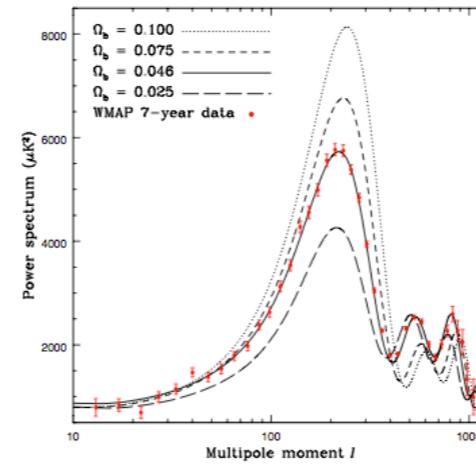
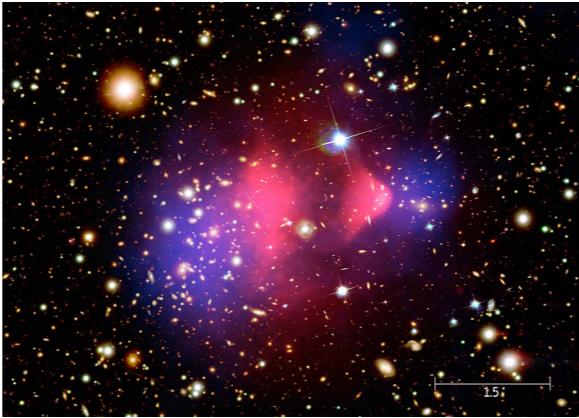
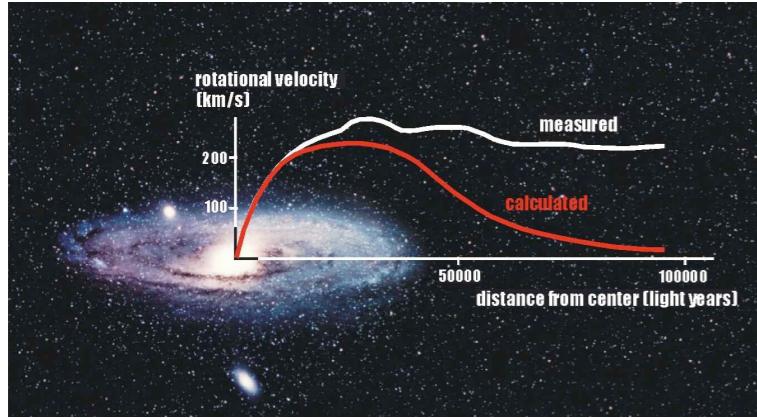
Chacko, Goh, Harnik, hep-ph/0506256.

Burdman, Chacko, Goh, Harnik, hep-ph/0609152.

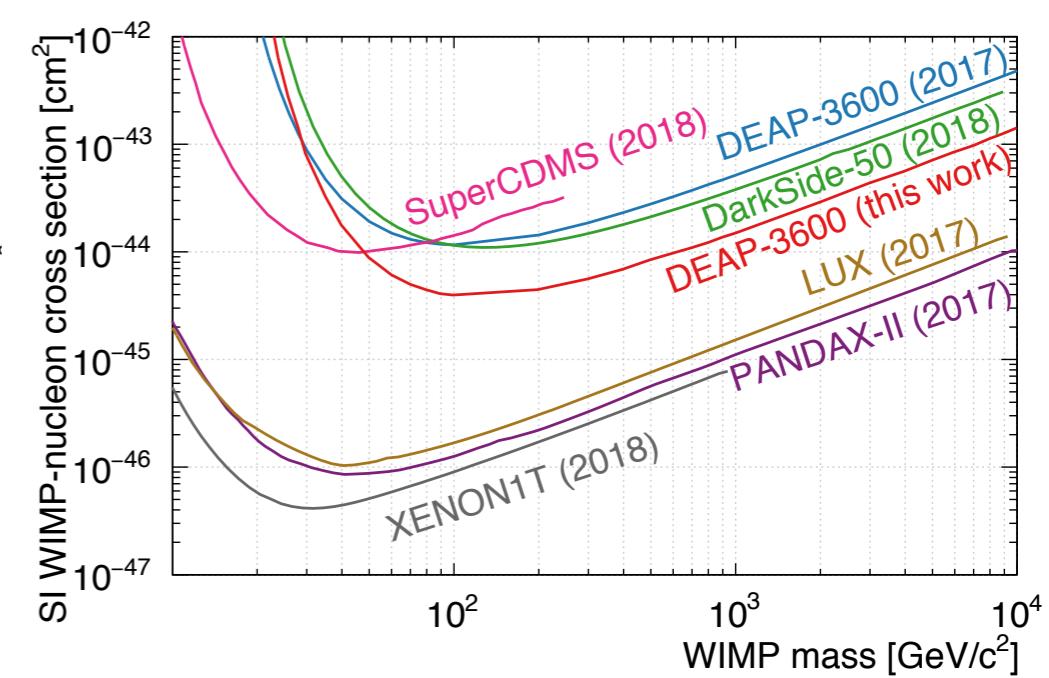
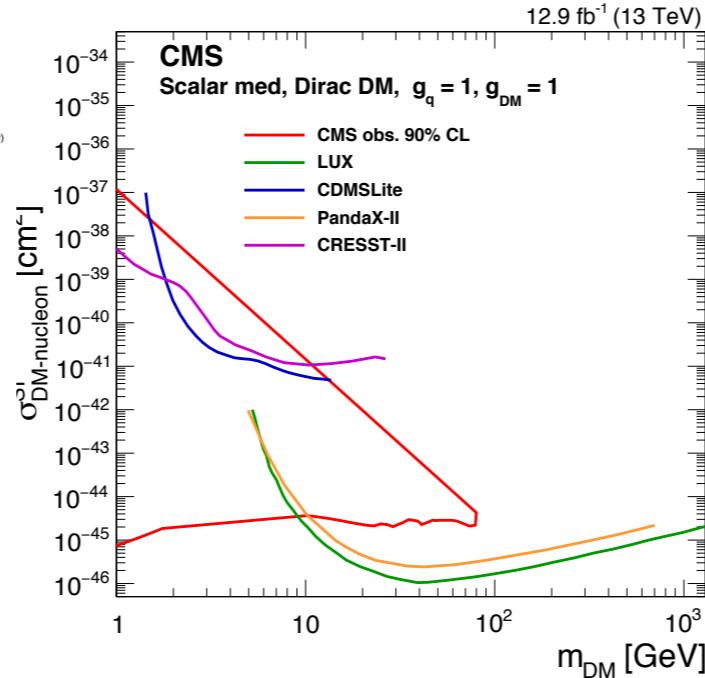
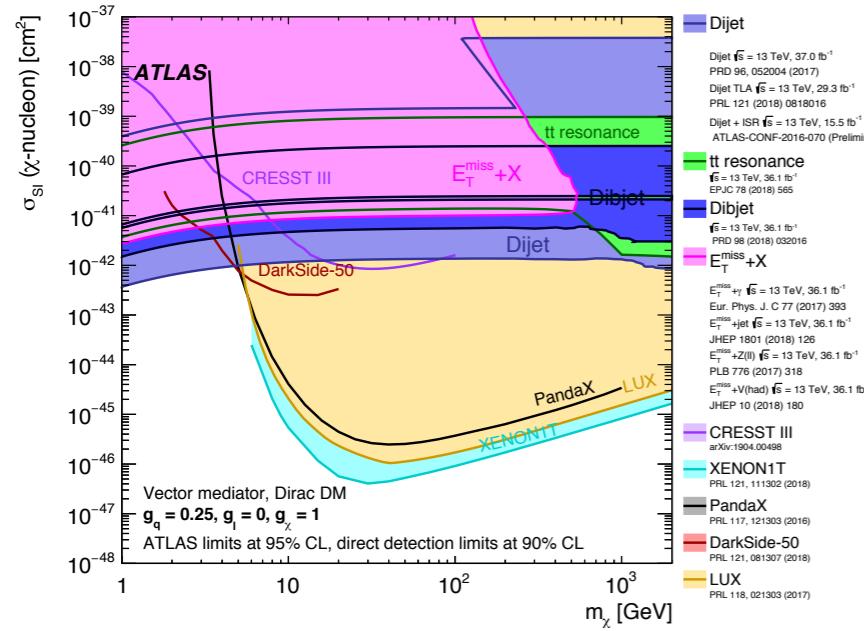
Most models have twin colour which confines around GeV scale (or slightly higher).

# MOTIVATION 2

We have seen dark matter in the sky.



But not in the lab.



# ASYMMETRIC DARK MATTER

$$\Omega_{DM} \simeq 5\Omega_B$$

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$$\Omega_{DM} = m_{DM} n_{DM}$$

$$\Omega_B = m_p n_B$$

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Controlled by complicated  
(known) QCD dynamics

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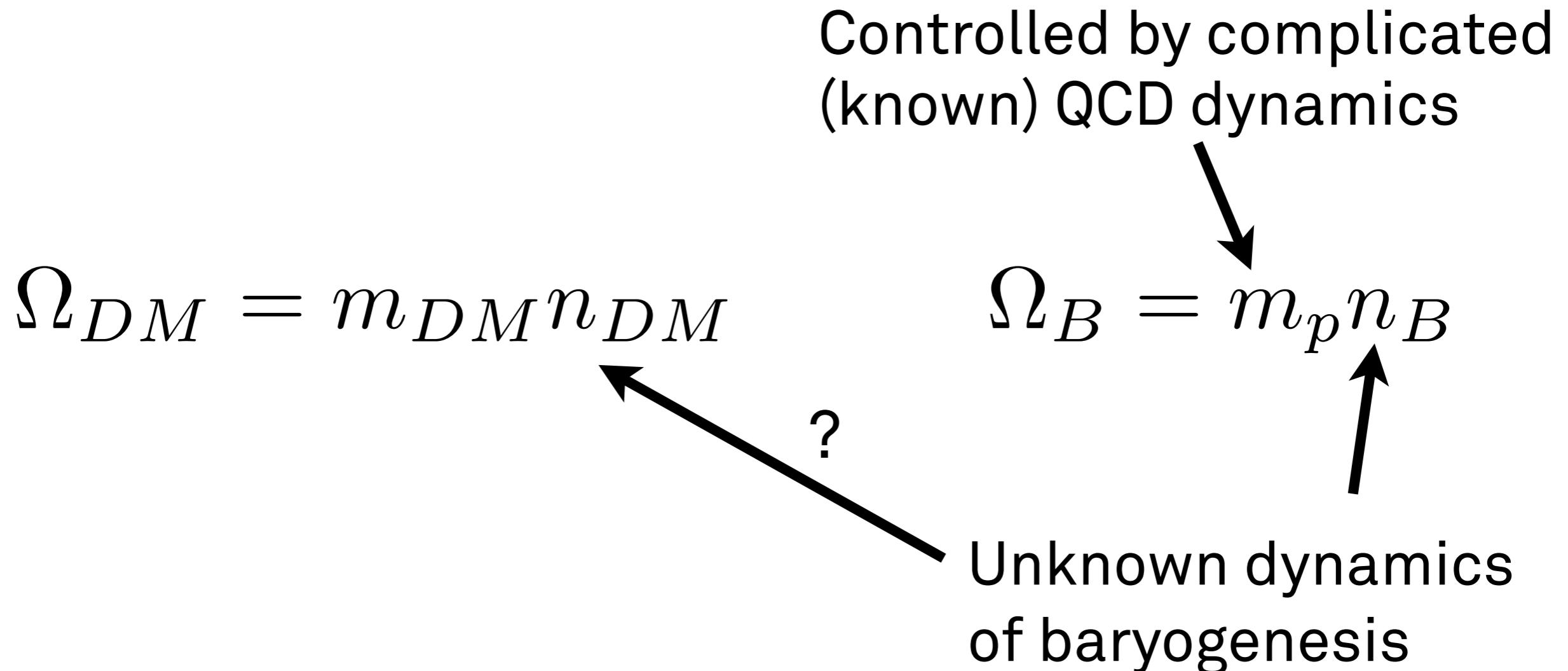
Controlled by complicated  
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$$\Omega_B = m_p n_B$$

Unknown dynamics  
of baryogenesis

# ASYMMETRIC DARK MATTER

$$\Omega_{DM} \simeq 5\Omega_B$$



# MANY PAPERS

S. Nussinov, Phys.Lett.B.165 (1985) 55.

D. B. Kaplan, Phys.Rev.Lett.B.68 (1992) 741-3.

...

D. E. Kaplan, M. A. Luty, K. M. Zurek, Phys.Rev.D 79 115016 (2009)  
[arXiv:0901.4117 [hep-ph]].

K. K. Boddy, J. L. Feng, M. Kaplinghat, and T. M. P. Tait, Phys. Rev. D. 89 11, 115017 (2014) [arXiv:1402.3629 [hep-ph]].

Agrawal, Kilic, Swaminathan, Trendafilova, Phys. Rev. D. 95 1 015031,  
(2017) [arXiv:1608.04745 [hep-ph]].

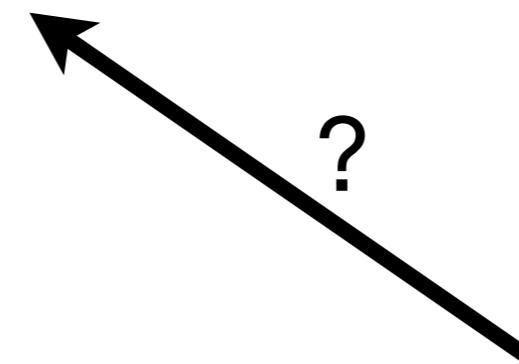
For a review see K. Petraki and R. R. Volkas, Int.J.Mod.Phys.A 28,

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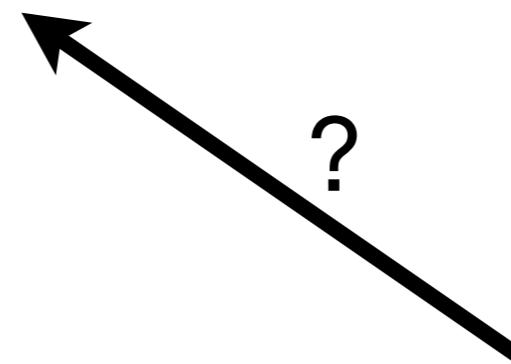
Can get  $n_{DM} \sim n_B$ , usually have to assume  $m_{DM} \sim m_B$ .

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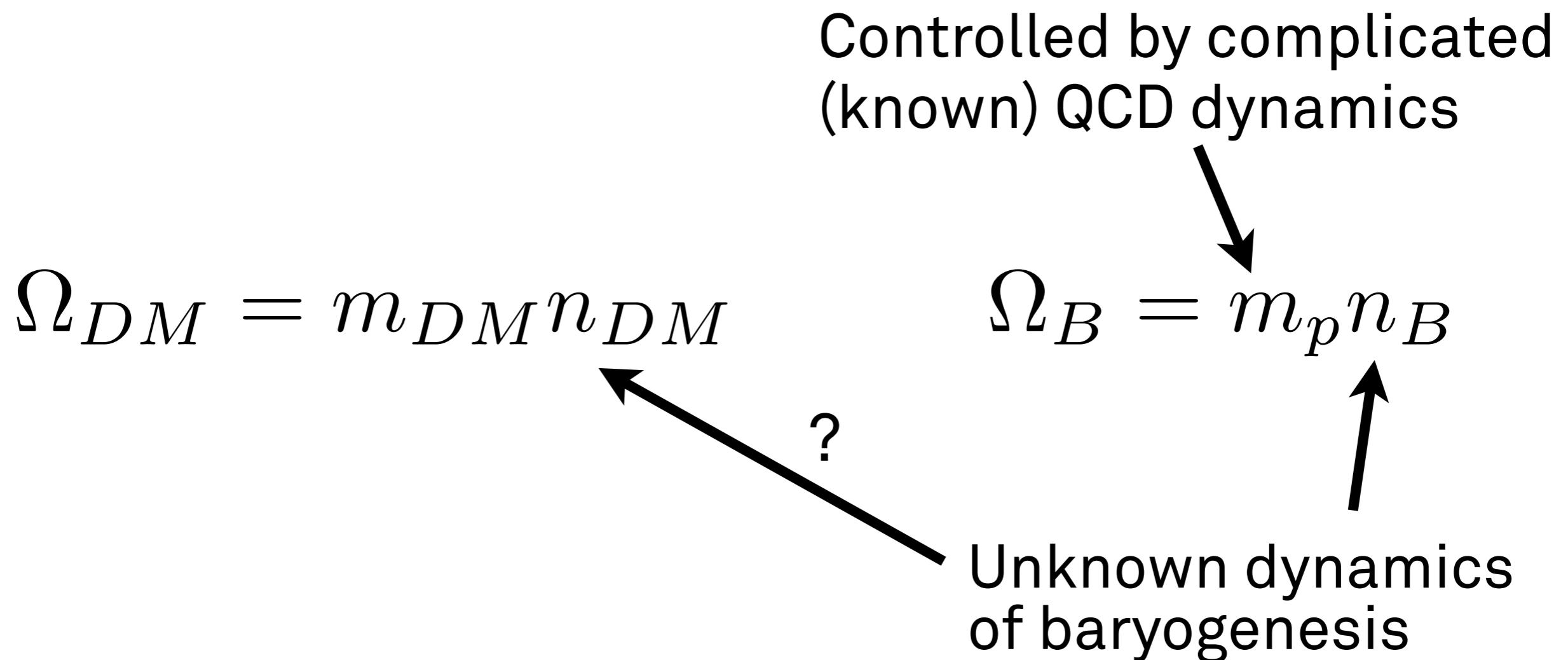
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Can get  $n_{DM} \sim n_B$ , usually have to assume  $m_{DM} \sim m_B$ .

Can we get **both**?

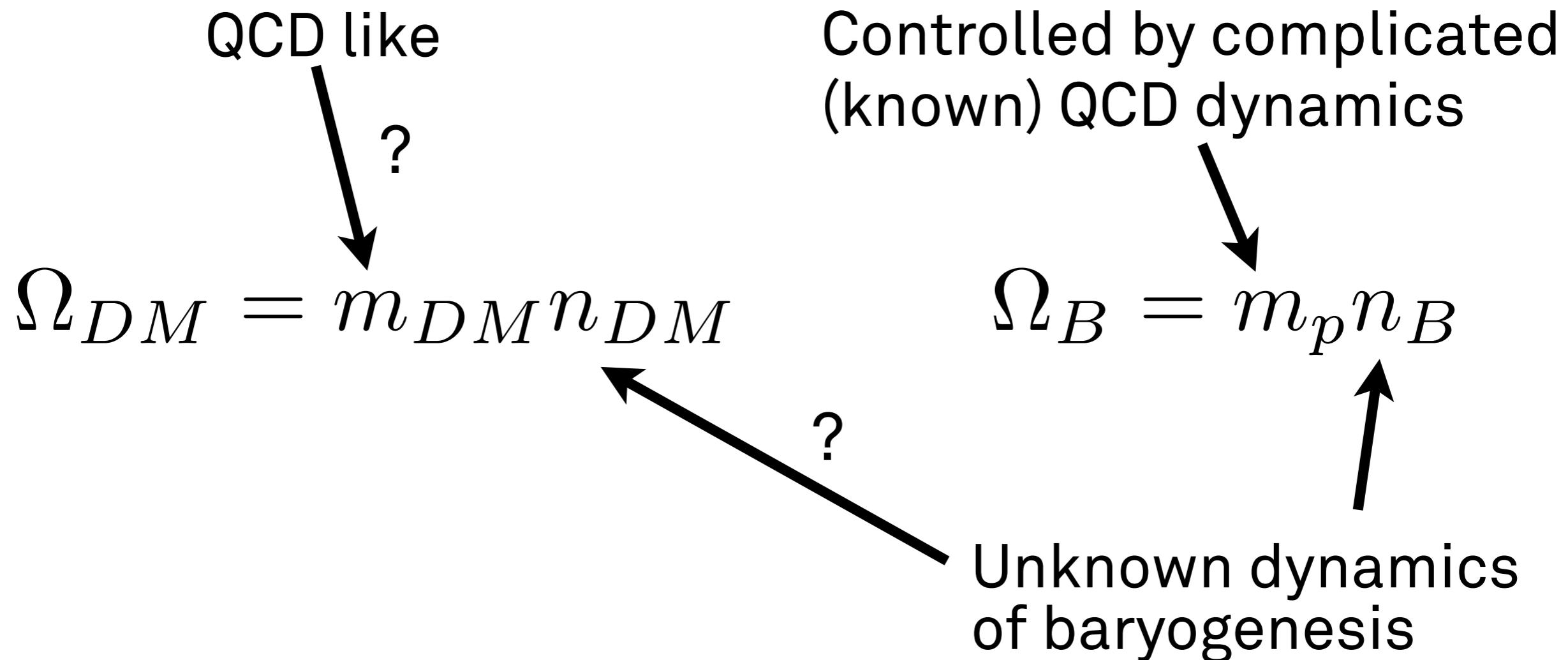
# GETTING THE MASS

$$\Omega_{DM} \simeq 5\Omega_B$$



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$$\Omega_{DM} \simeq 5\Omega_B$$



# DARK QCD

Bai, Schwaller, PRD 13.

$$\Lambda_{\text{dQCD}}$$

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# DARK QCD

Propose new  $SU(N_d)$  “dark QCD,” dark quarks.  
Bai, Schwaller, PRD 13.

Dark matter is dark sector baryons with mass  $\sim \Lambda_{dQCD}$ .

Massive bifundamental fields decouple at mass  $M \gg \Lambda_{dQCD}$ .

# DARK QCD

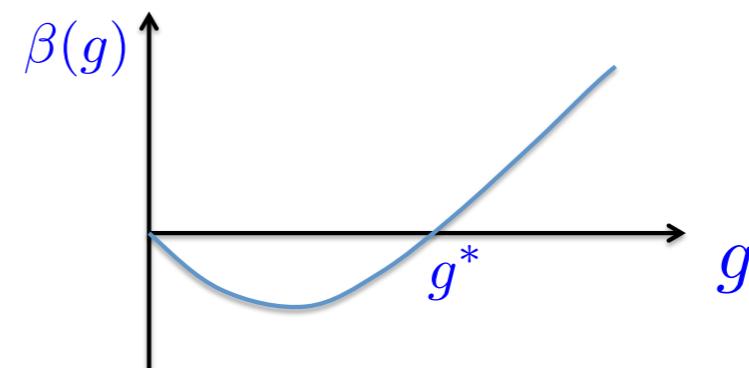
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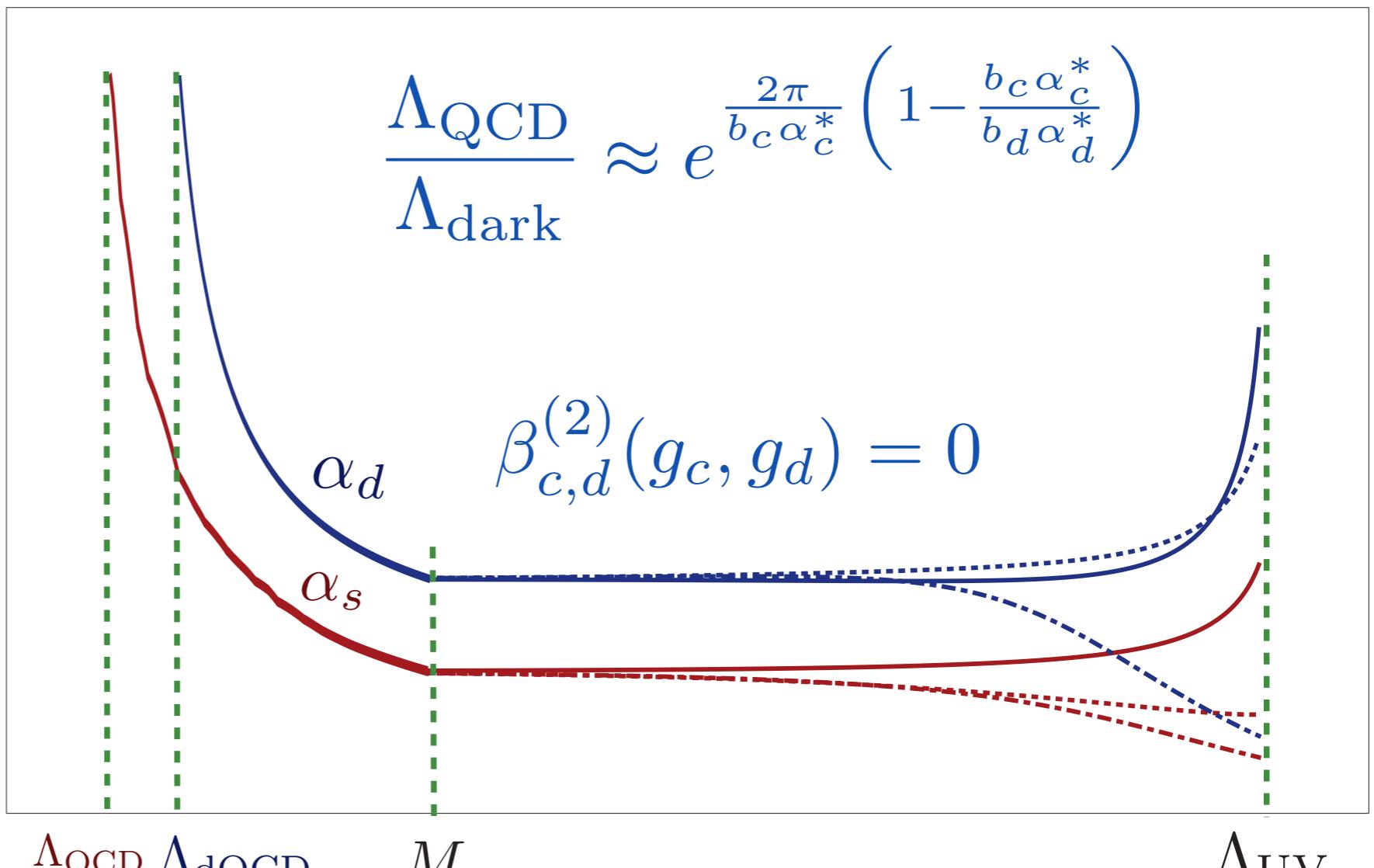
Massive bifundamental fields decouple at mass  $M \gg \Lambda_{dQCD}$ .

Search for model with perturbative fixed point.

$$\frac{dg}{dt} = \beta(g) = 0 \text{ for } g = g^*$$



# SCALES ARE RELATED



Bai, Schwaller, PRD 13.

# SCALES ARE RELATED

## Example

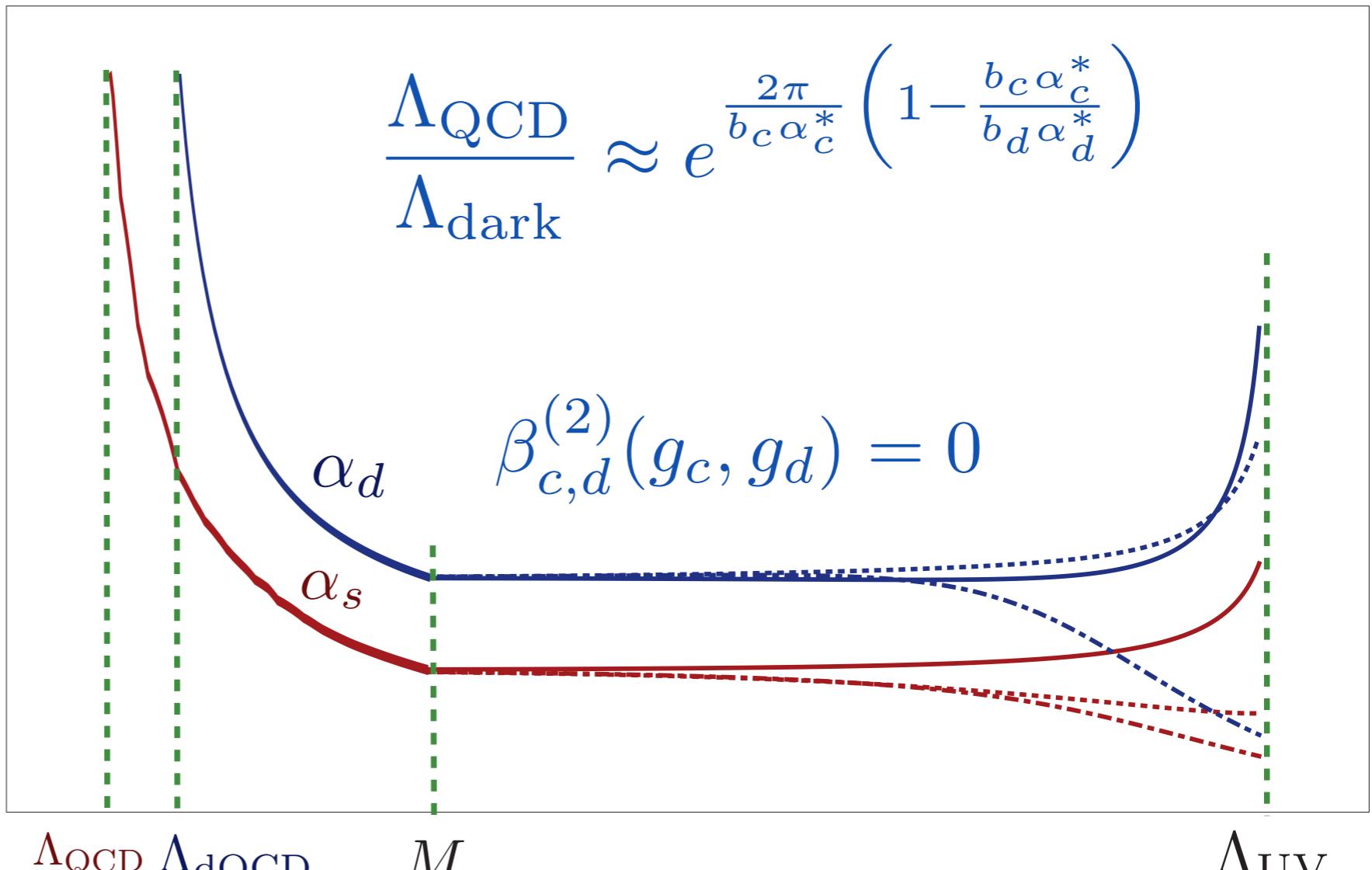
Fixed points:

$$\alpha_c^* = 0.090 \quad \alpha_d^* = 0.168$$

$$M = 870 \text{ GeV}$$

DM mass:

$$M_{DM} \approx 3.5 \text{ GeV}$$



Bai, Schwaller, PRD 13.

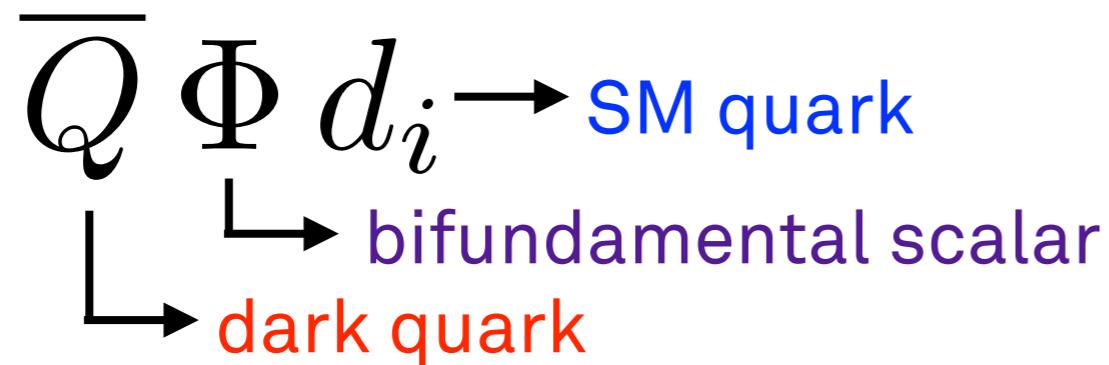
# DARK MATTER

Can co-generate DM and baryon asymmetry.

$$\overline{Q} \Phi d_i$$

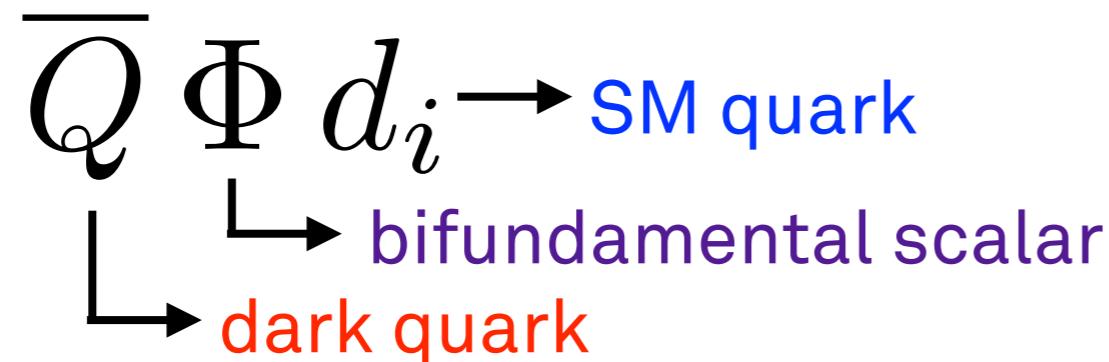
# DARK MATTER

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# DARK MATTER

Can co-generate DM and baryon asymmetry.



Dark matter is strongly self interacting — potentially solves various problems of cold dark matter.

- Cusp vs core
- Missing satellites

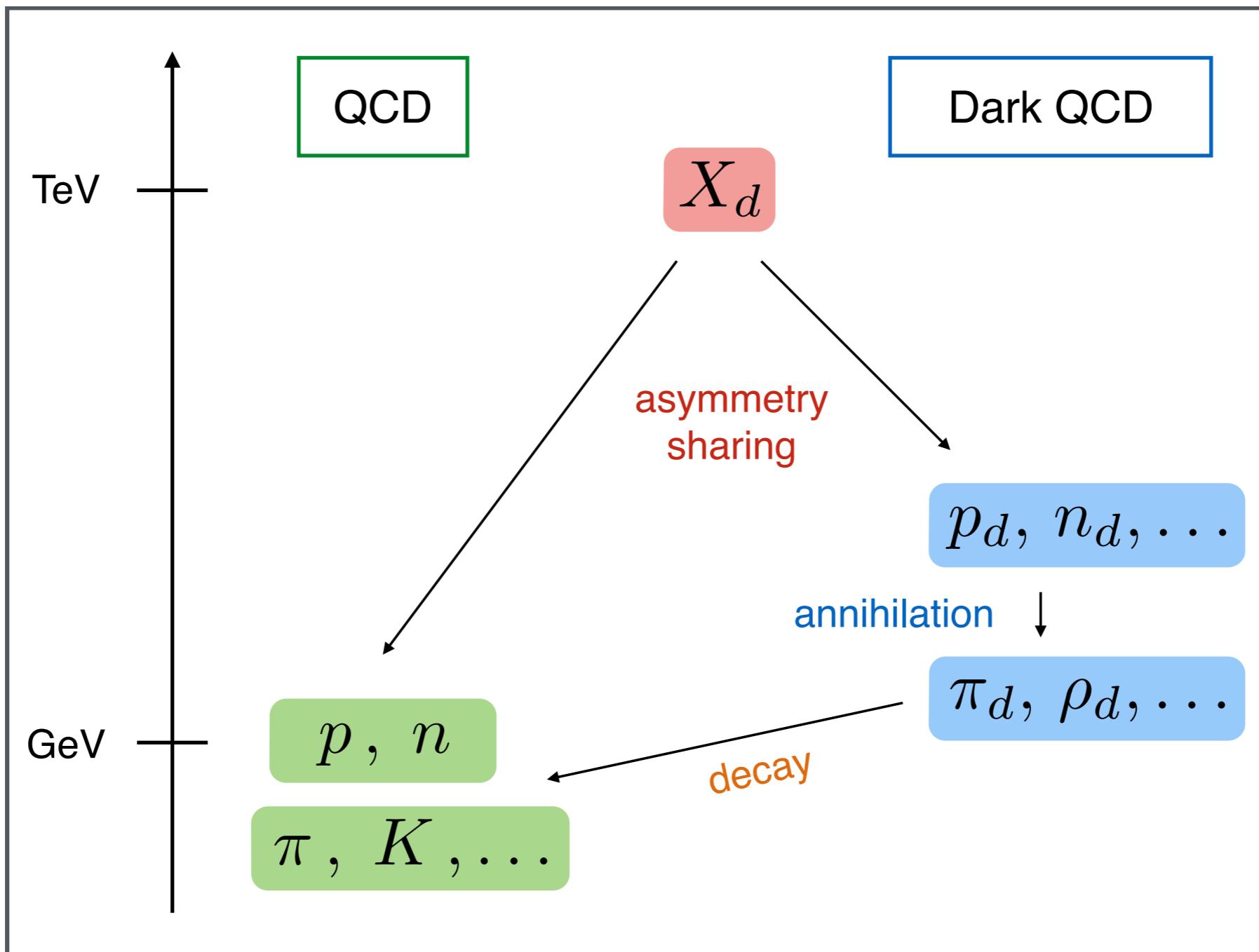
Rocha et. al. '12. Peter et. al. '12.

Vogelsberger, Zavala, Loeb, '12.

Zavala, Vogelsberger, Walker '12.

- Too big to fail

# GENERAL PICTURE



# PHENOMENOLOGY

# DARK QCD

Confining  $SU(N_c)$  gauge group with  $N_f$  flavours.

$$Q_i \quad \bar{Q}_j \quad G_d^{\mu\nu}$$

This sector is QCD like, and it confines at a scale.

$$\Lambda_d \sim 1 - 10 \text{ GeV}$$

At the confining scale we have all the usual states.

$$p_d \quad \pi_d \quad \text{Zoo}_d$$

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$p_d$

$\pi_d$

$Zoo_d$

Stable

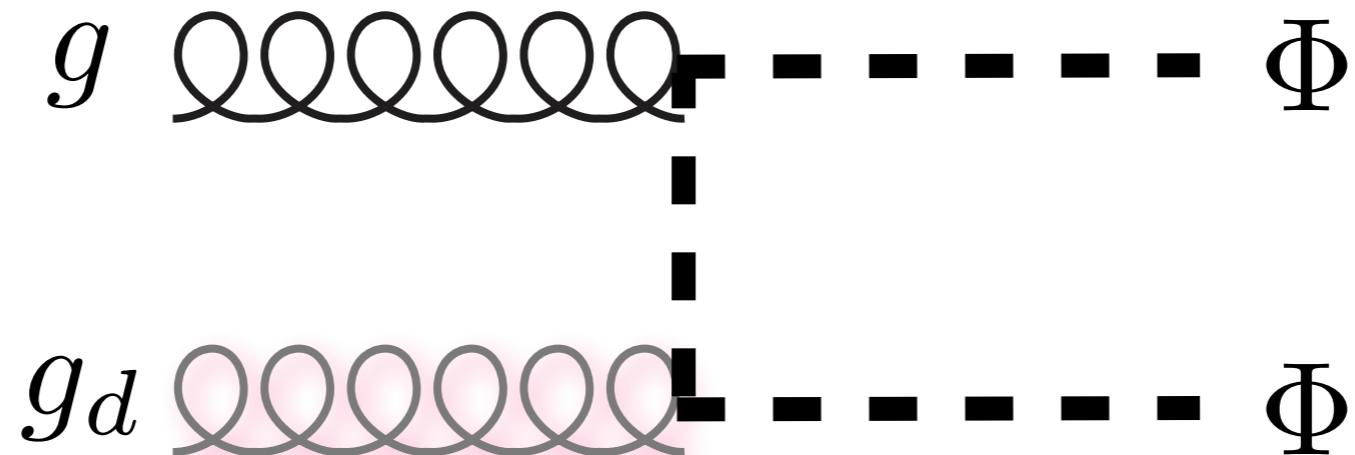
Decays on  
collider times

# MEDIATORS

Motivated by getting comparable asymmetries, put in heavy mediator which couples to SM and dark sector

$$M_\Phi \gg \Lambda_d$$

Example 1:  $\Phi$  is a scalar charged under both colour and dark colour.

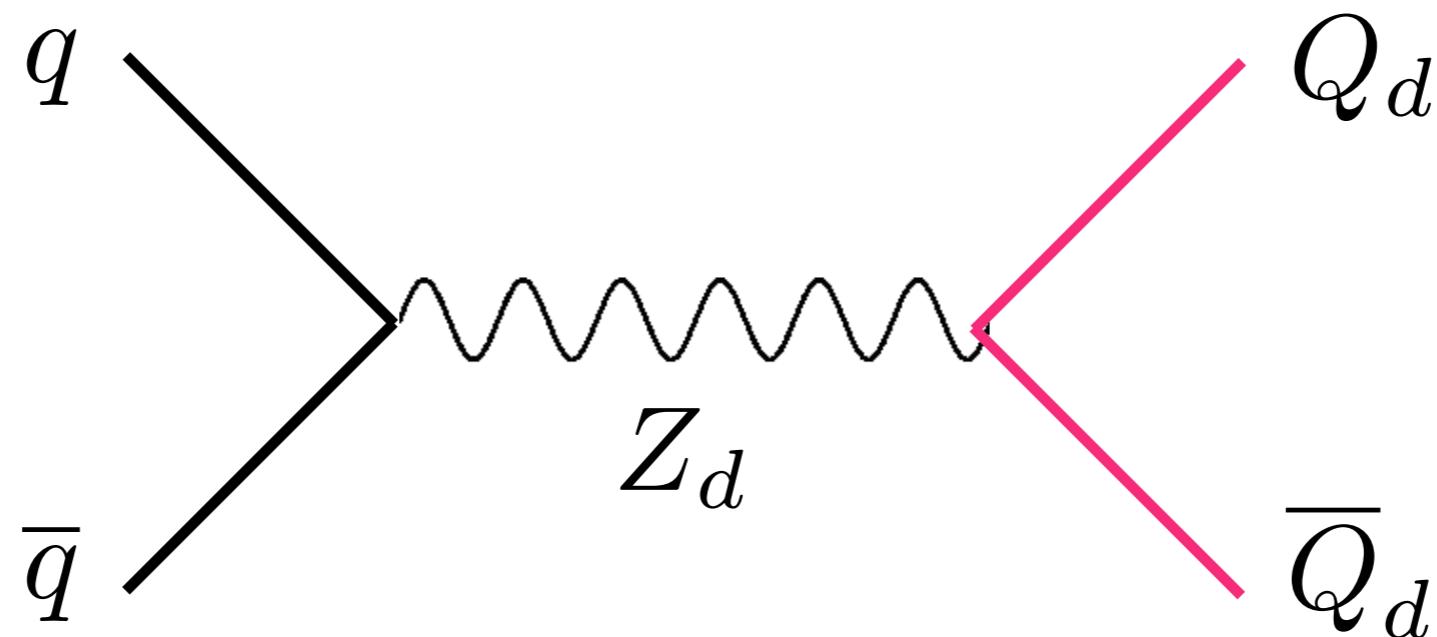


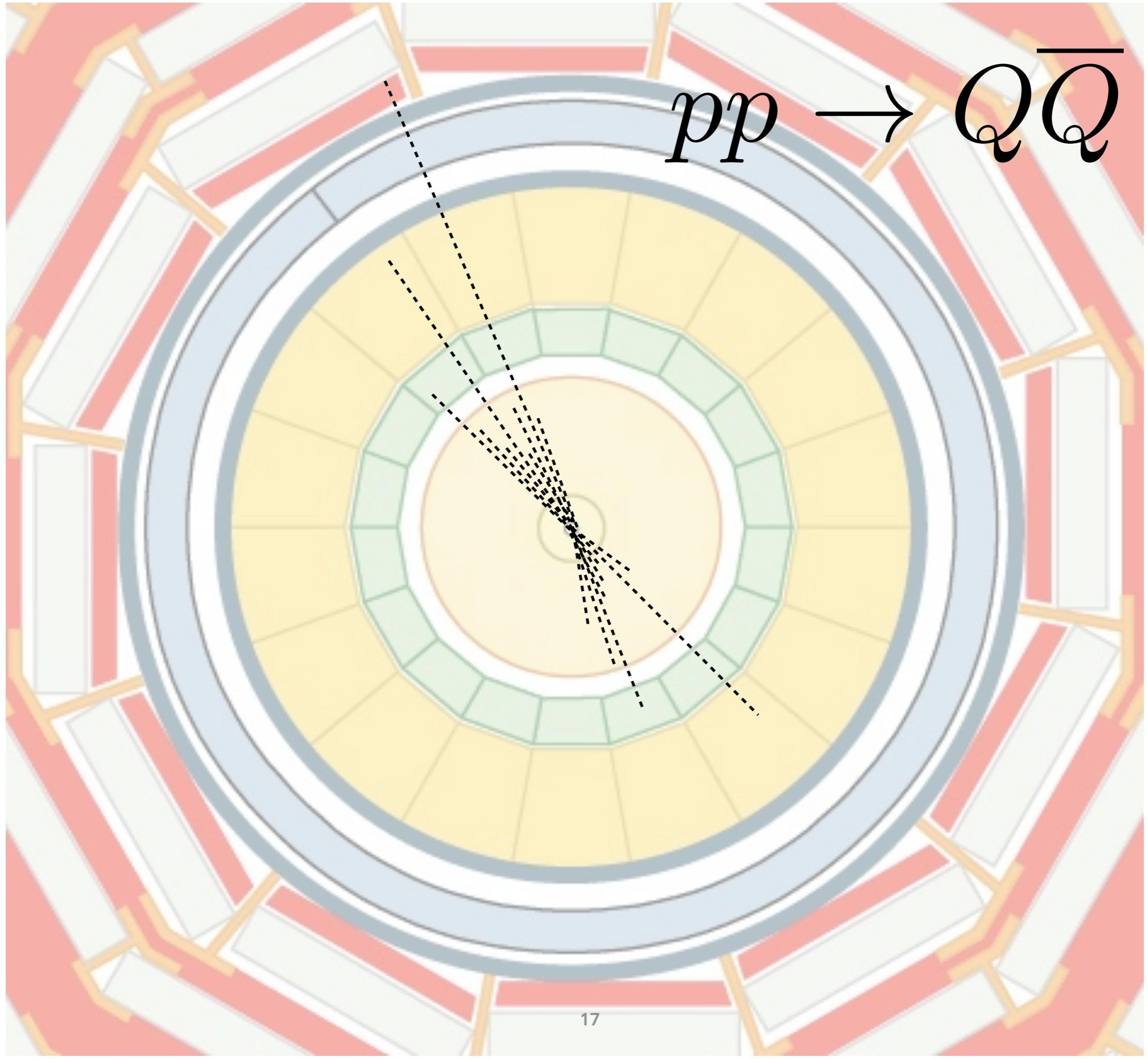
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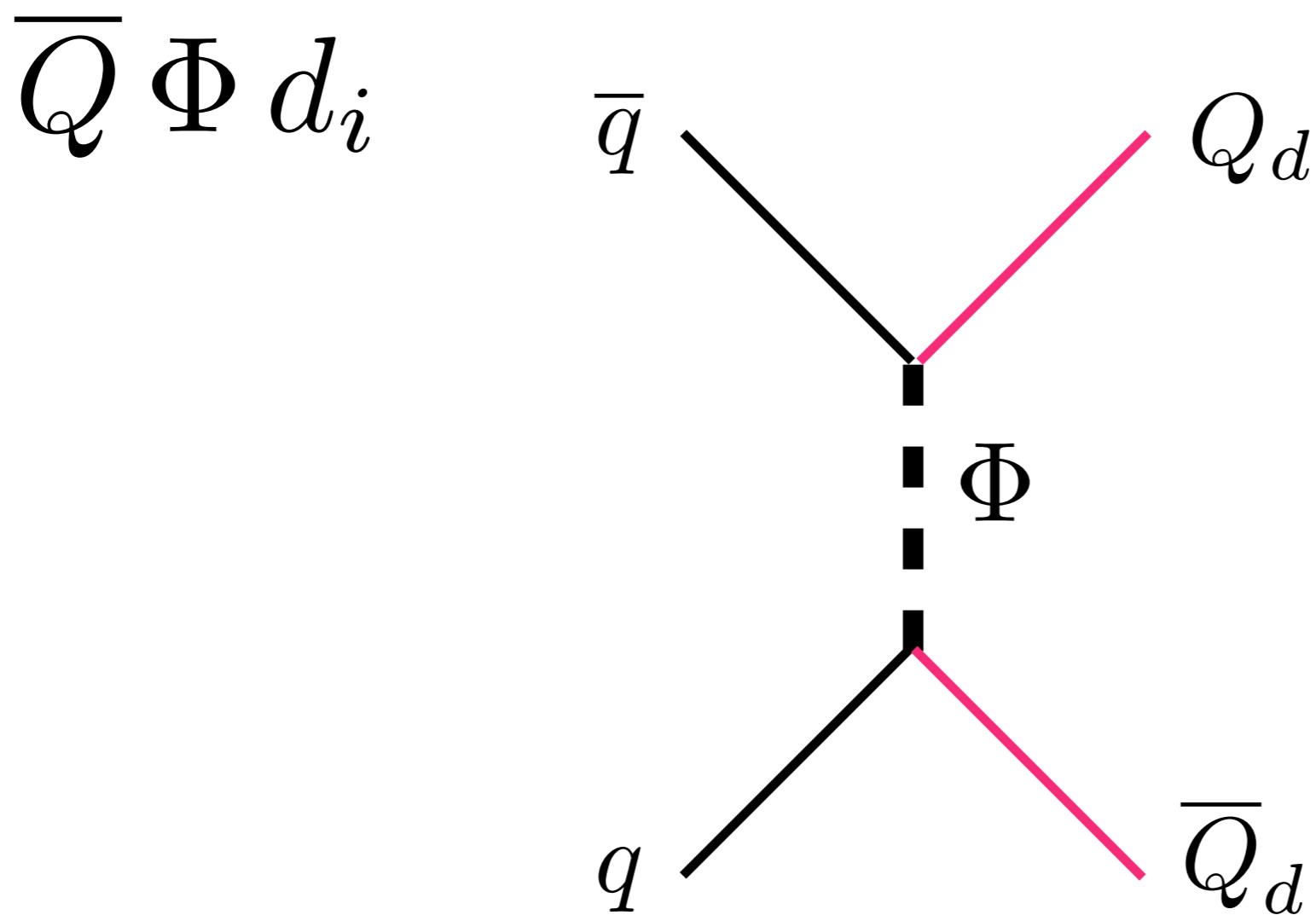
Example 2:  $Z_d$  is a vector that couples to quarks and dark quarks. **Strassler, Zurek, PLB 07.**





# PION DECAY

Operator used to generate asymmetry mediates decay:

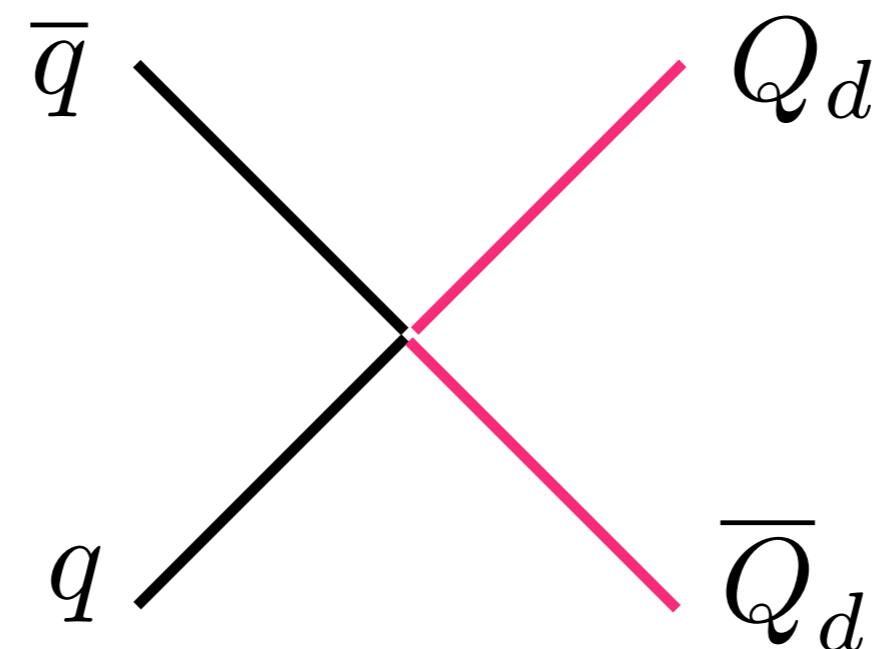


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Operator used to generate asymmetry mediates decay:

$$\overline{Q} \Phi d_i$$

Integrate out  $\Phi$

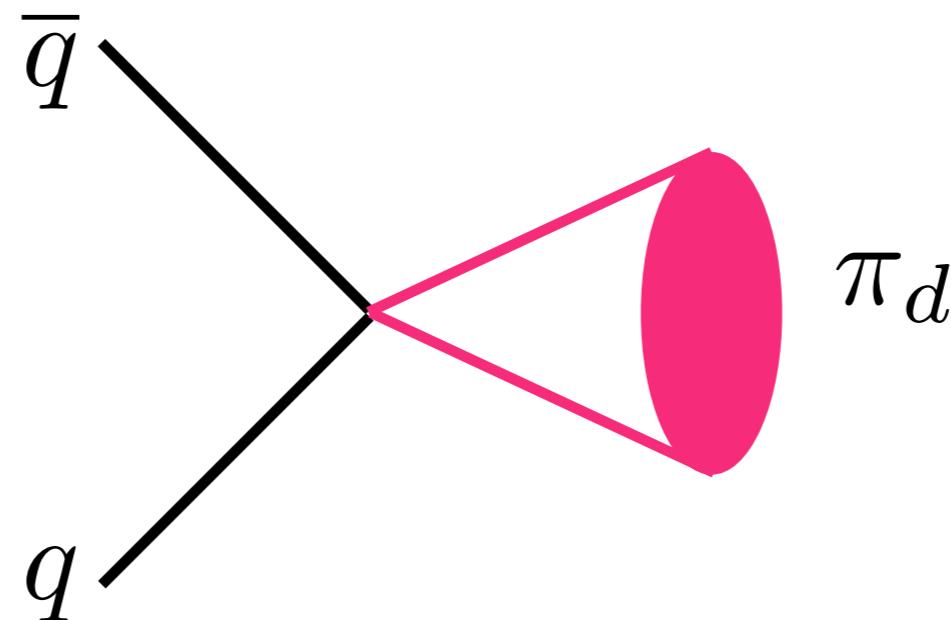


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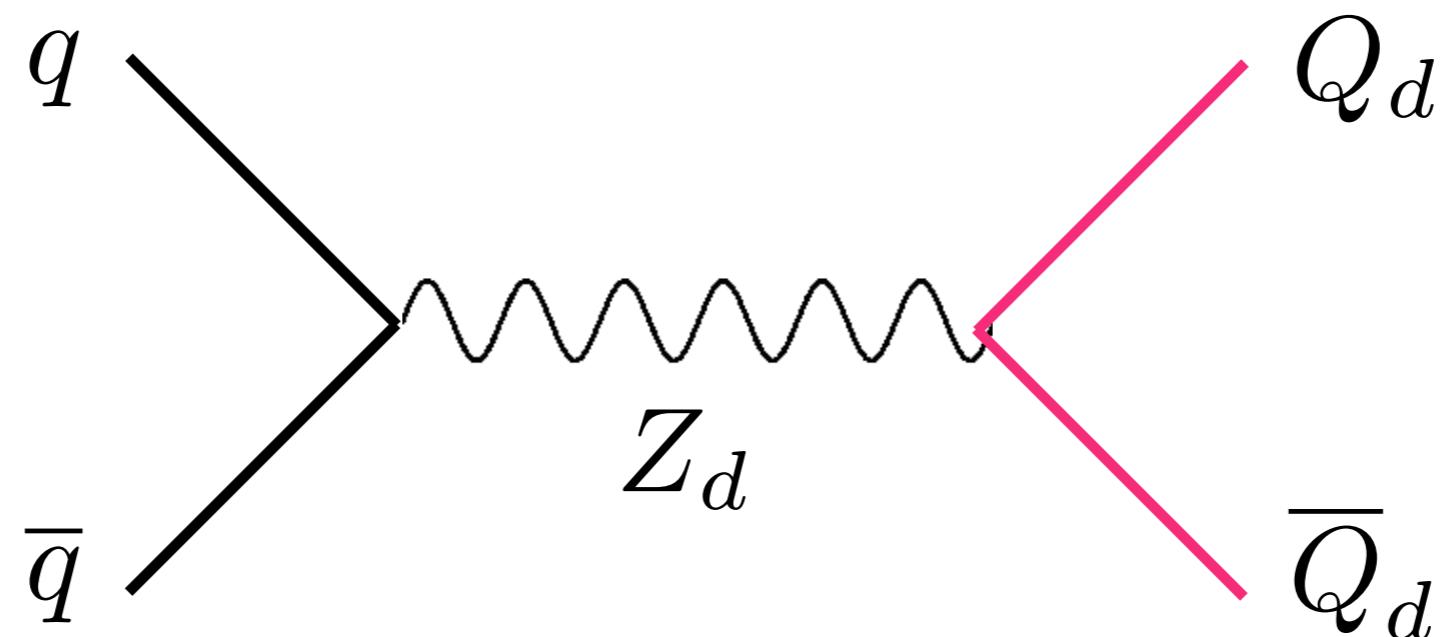
Integrate out  $\Phi$



Dark pion  
decays to  
quarks

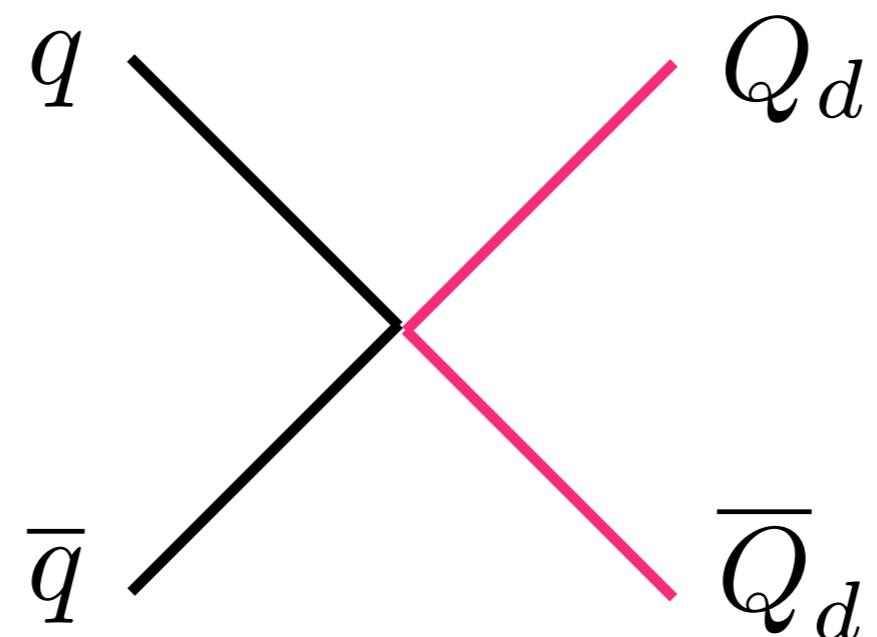
# PION DECAY

Same story for  $Z_d$  model:



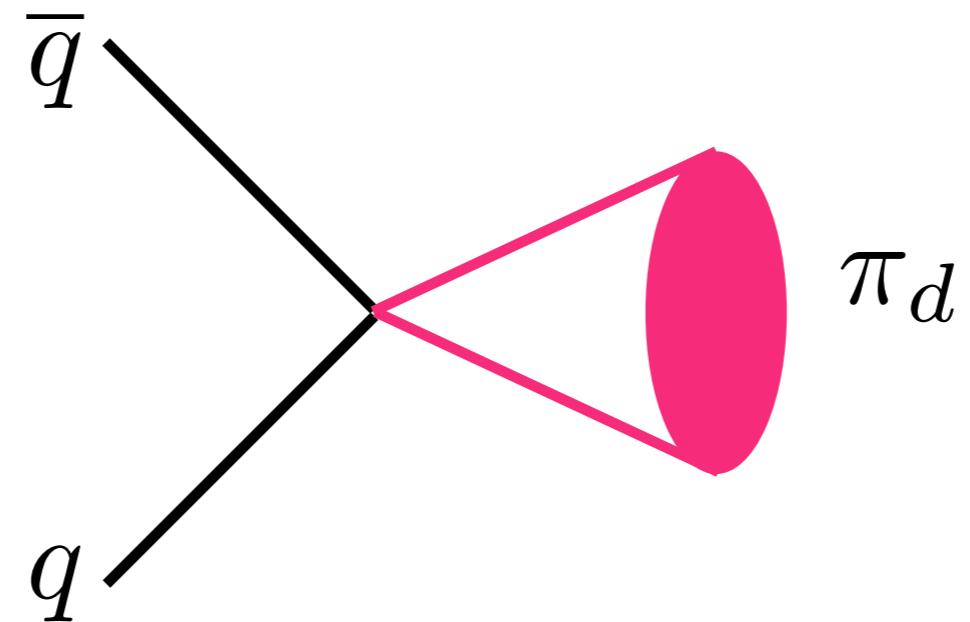
# PION DECAY

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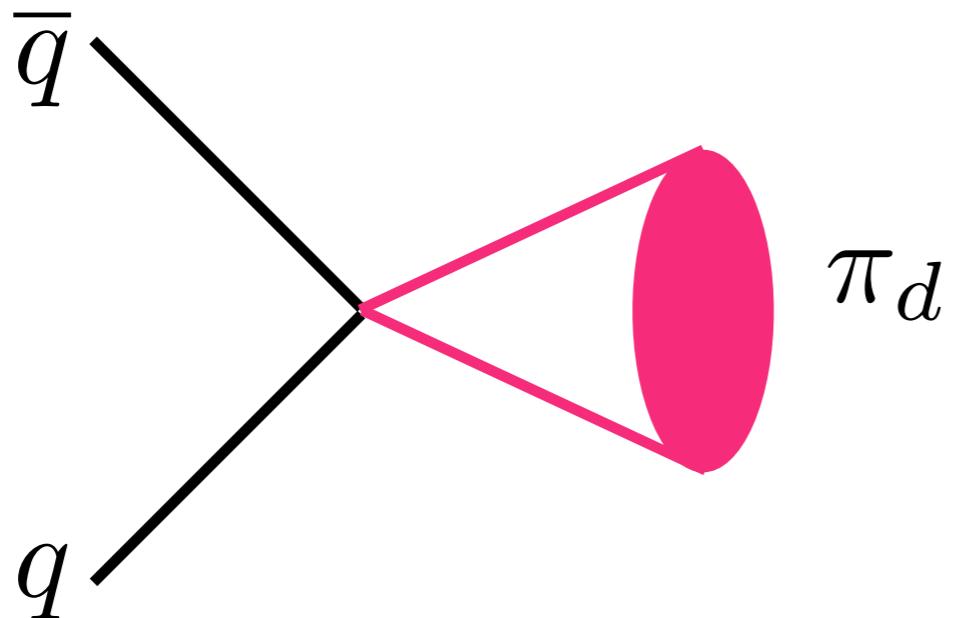


# PION DECAY

Same story for  $Z_d$  model:



# DECAY LENGTH

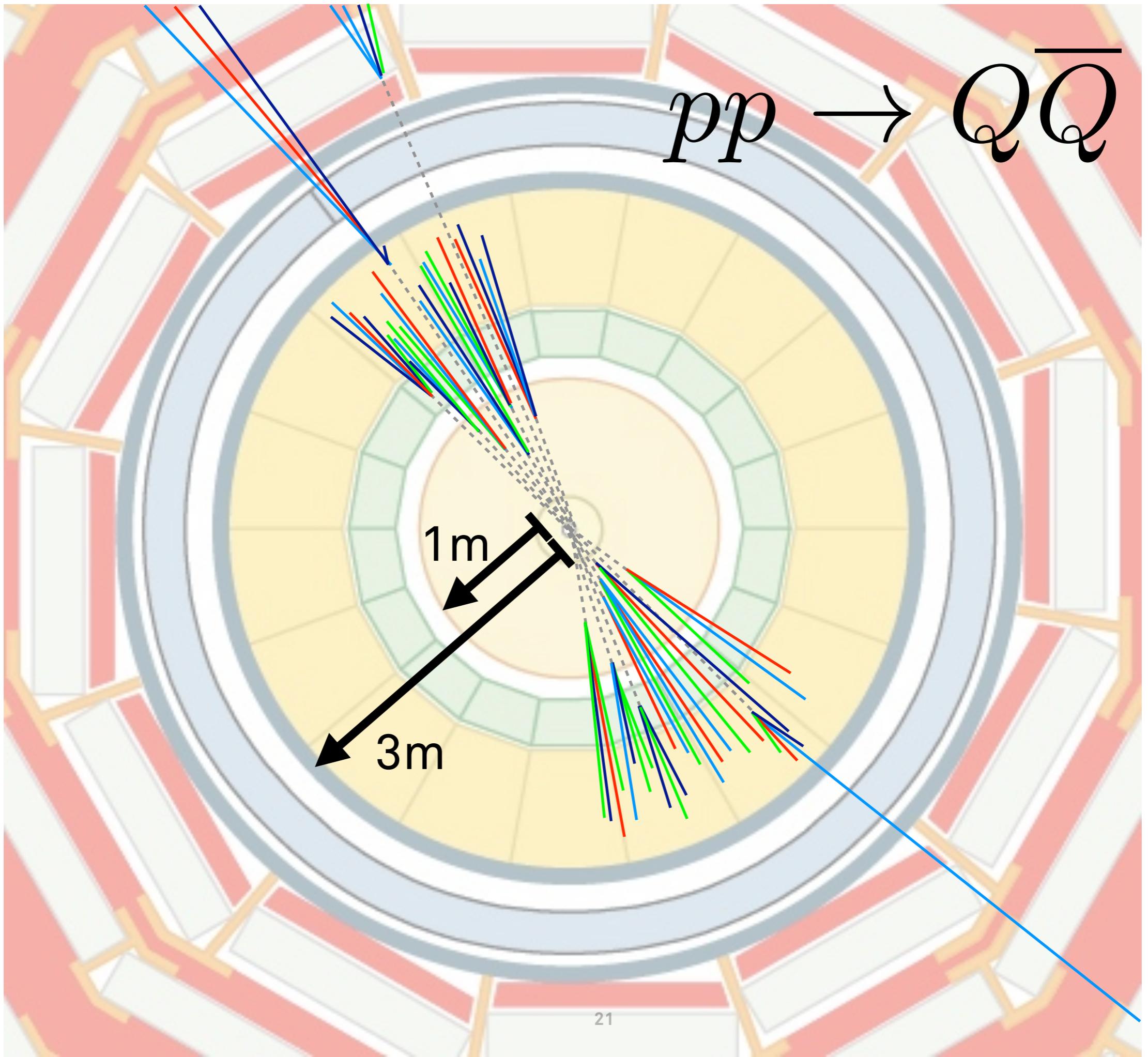


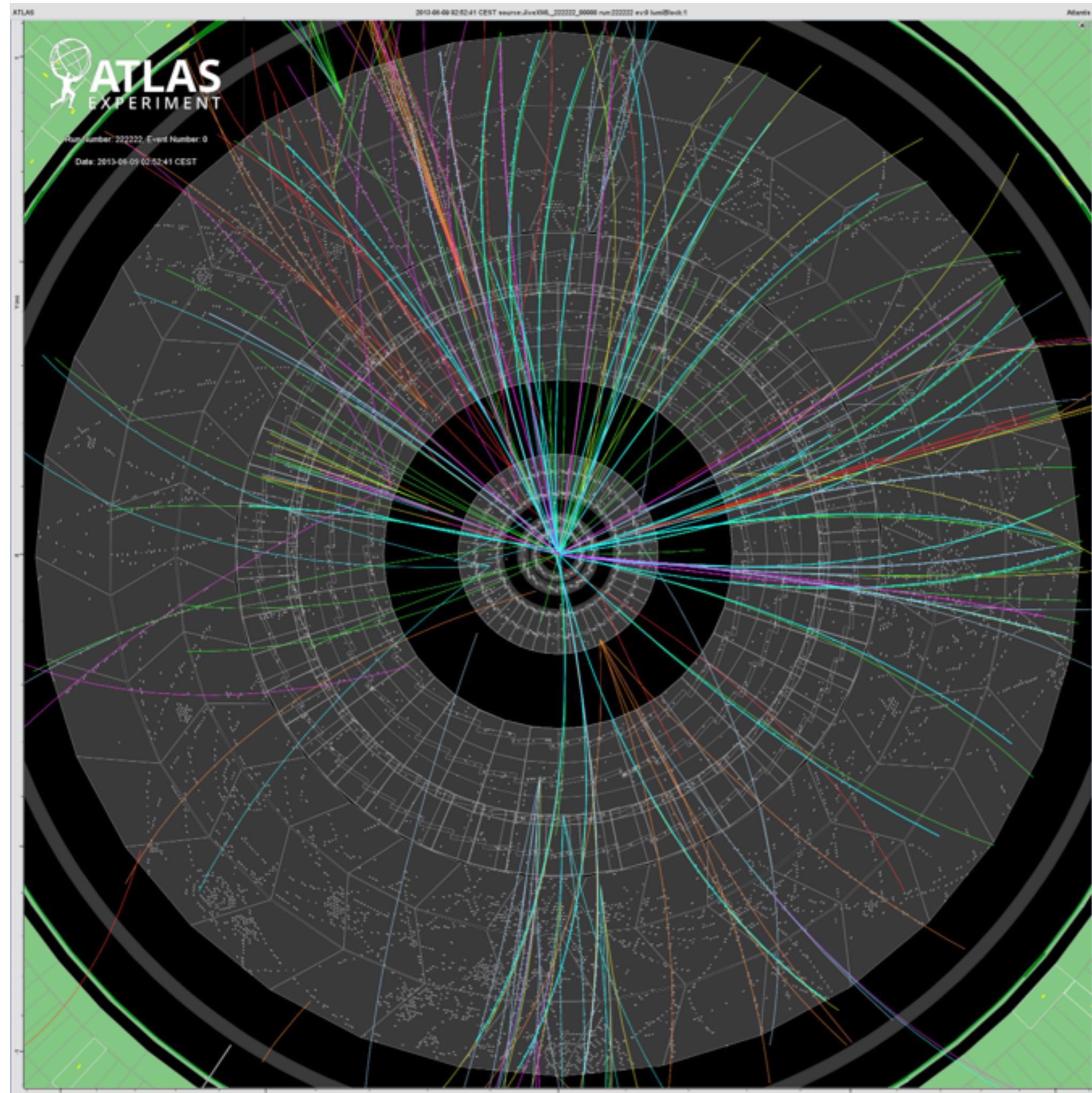
$$\frac{1}{M_X^2} \bar{Q} \gamma_\mu Q \bar{d}_R \gamma^\mu d_R$$

Can use (dark) chiral Lagrangian to estimate:

$$\Gamma(\pi_d \rightarrow \bar{d}d) \approx \frac{f_{\pi_d}^2 m_d^2}{32\pi M_{X_d}^4} m_{\pi_d}$$

$$c\tau_0 \approx 10 \text{ cm} \times \left( \frac{2 \text{ GeV}}{f_{\pi_d}} \right)^2 \left( \frac{100 \text{ MeV}}{m_{\text{down}}} \right)^2 \left( \frac{2 \text{ GeV}}{m_{\pi_d}} \right) \left( \frac{M_{X_d}}{1 \text{ TeV}} \right)^4$$





James Beacham,  
ATLAS Dark Sector  
Workshop, Cosenza,  
Italy, Feb 10, 2016.

# DARK SECTOR

Choose two benchmarks:

	Model A	Model B
$\Lambda_d$	10 GeV	4 GeV
$m_V$	20 GeV	8 GeV
$m_{\pi_d}$	5 GeV	2 GeV
$c \tau_{\pi_d}$	150 mm	5 mm

$$N_c = 3 \text{ and } n_f = 7$$

Dark QCD already in PYTHIA!

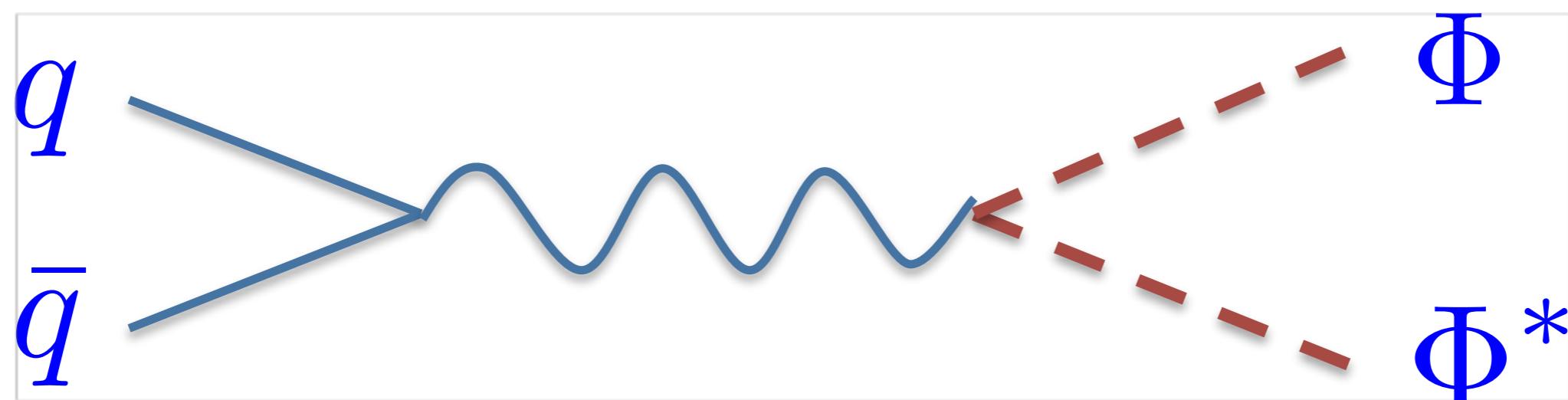
Carloni, Sjorstrand, 2010.

Carloni, Rathsman, Sjorstrand, 2011.

We had to modify Pythia to make hidden sector more realistic, our code is now implemented ( $v > 8.226$ ).

# BENCHMARK MEDiator 1

$$pp \rightarrow \Phi\Phi^\dagger \rightarrow \bar{q} Q_d \overline{Q}_d q$$



# BENCHMARK MEDIATOR 1

$$pp \rightarrow \Phi\Phi^\dagger \rightarrow \bar{q} Q_d \overline{\bar{Q}}_d q$$

Final state is

- 2 QCD jets
- 2 emerging jets

Cross section is stop-like

$$\sigma \approx \text{few} \times \sigma(pp \rightarrow \tilde{t}_1 \tilde{t}_1)$$

$$\sigma(M_\Phi = 1 \text{ TeV}) \approx 10 \text{ fb}$$

@ LHC14

# BENCHMARK MEDIATOR 2

$$pp \rightarrow Z_d \rightarrow Q_d \bar{Q}_d$$

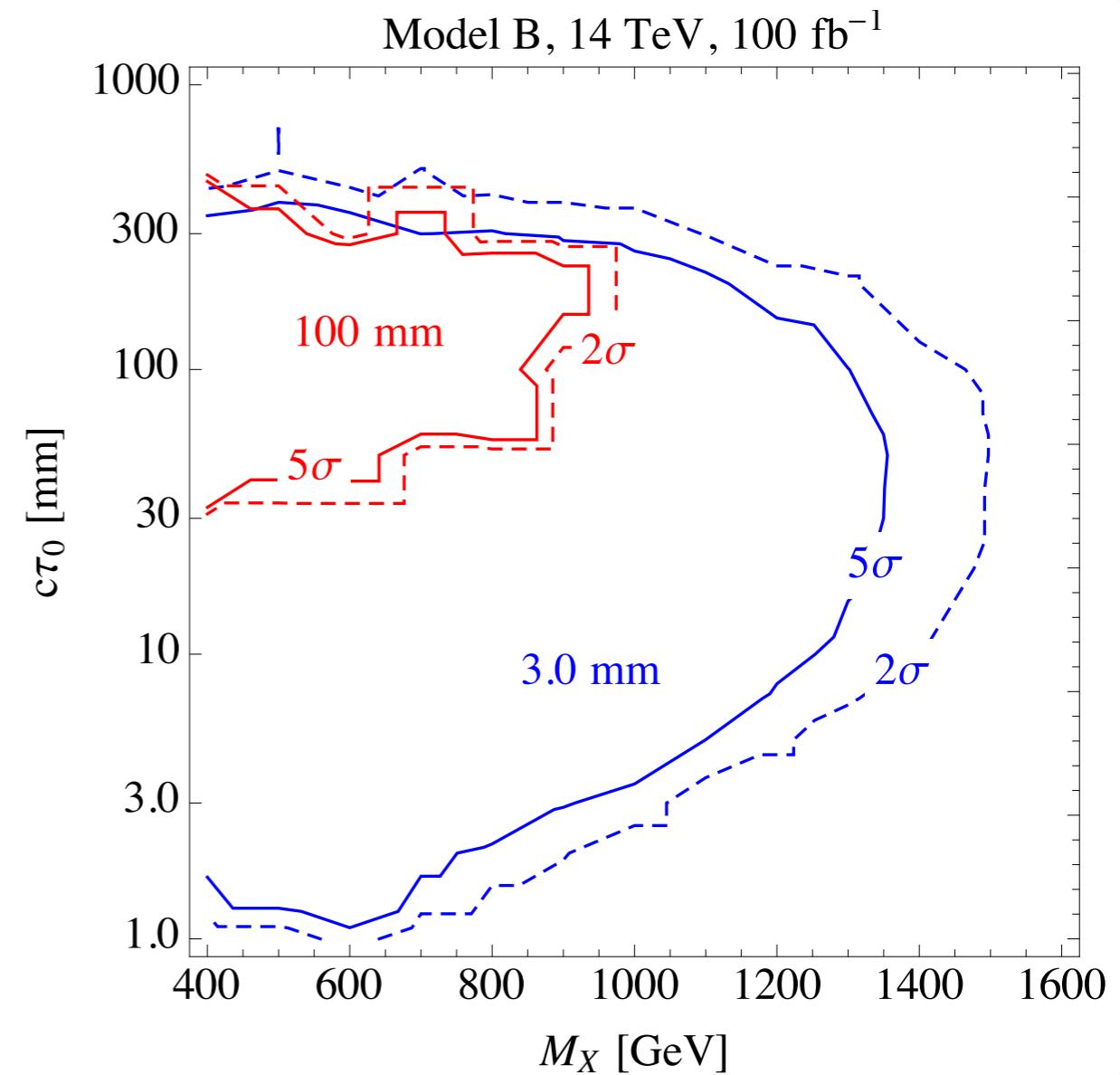
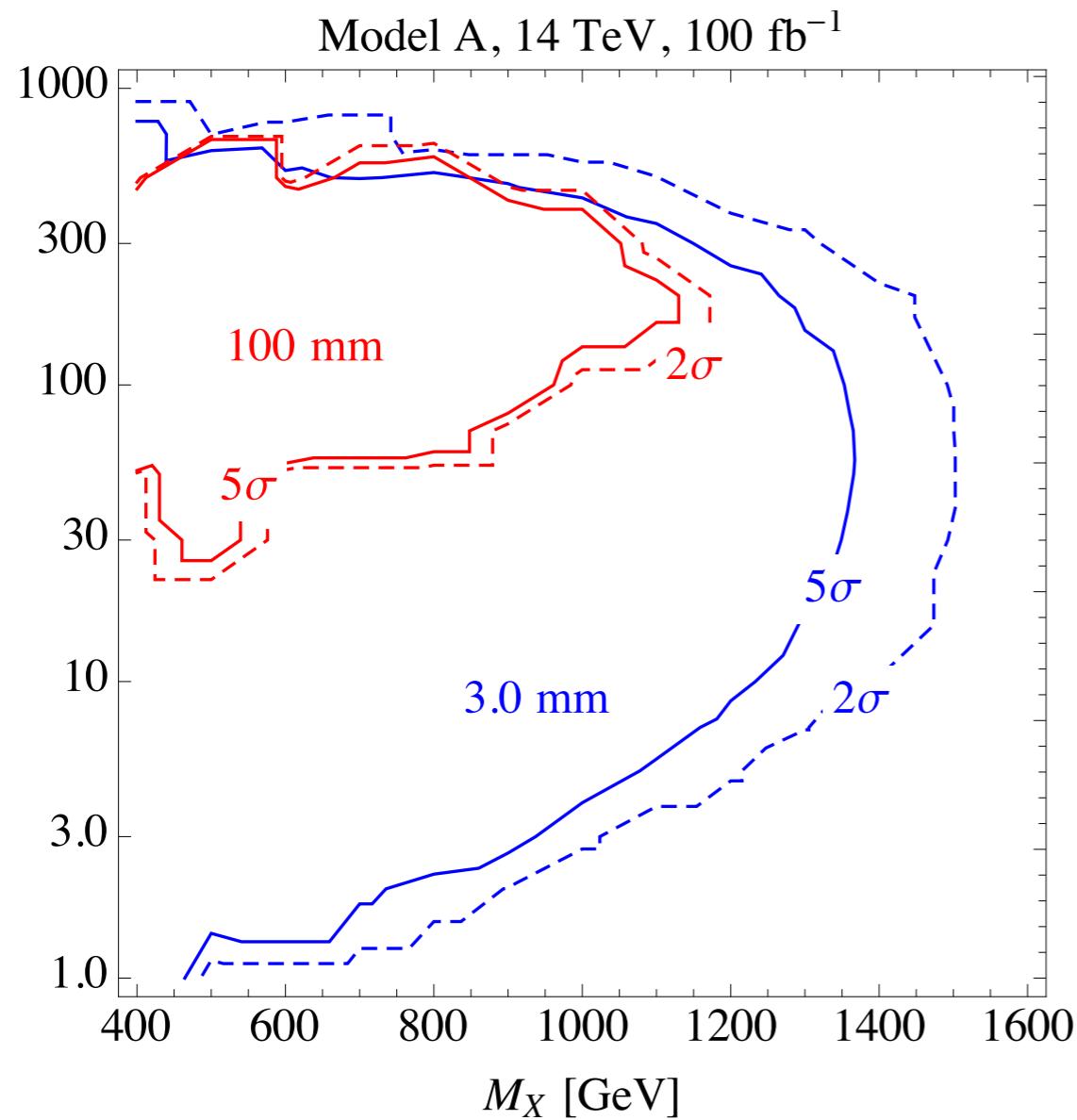
Final state is

- 2 emerging jets

Cross section depends on  
couplings.

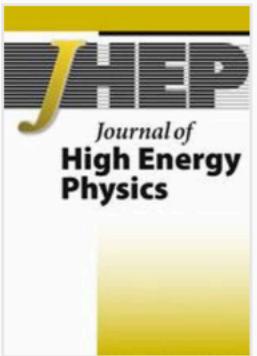
I will come back to this scenario.

# PROJECTED REACH



Schwaller, DS, Weiler, arXiv:1502.0409.

# DEDICATED CMS SEARCH



[Journal of High Energy Physics](#)  
... February 2019, 2019:179 | [Cite as](#)

## Search for new particles decaying to a jet and an emerging jet

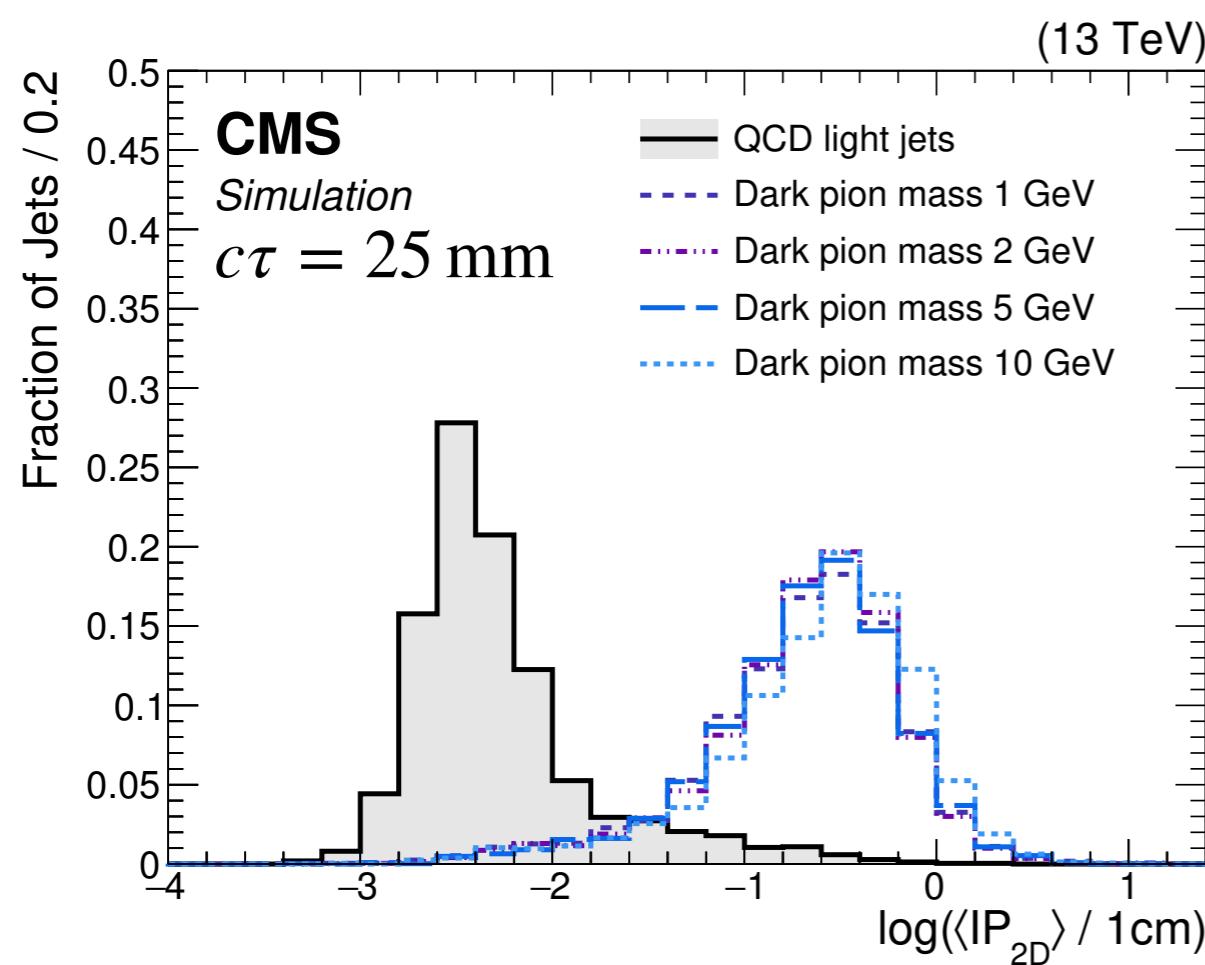
[Authors](#) [Authors and affiliations](#)

The CMS collaboration , A. M. Sirunyan, A. Tumasyan, W. Adam, F. Ambrogi, E. Asilar, T. Bergauer, J. Brandstetter, M. Dragicevic, J. Erö, A. Escalante Del Valle, M. Flechl, R. Frühwirth, V. M. Ghete, J. Hrubec, [show 2286 more](#)

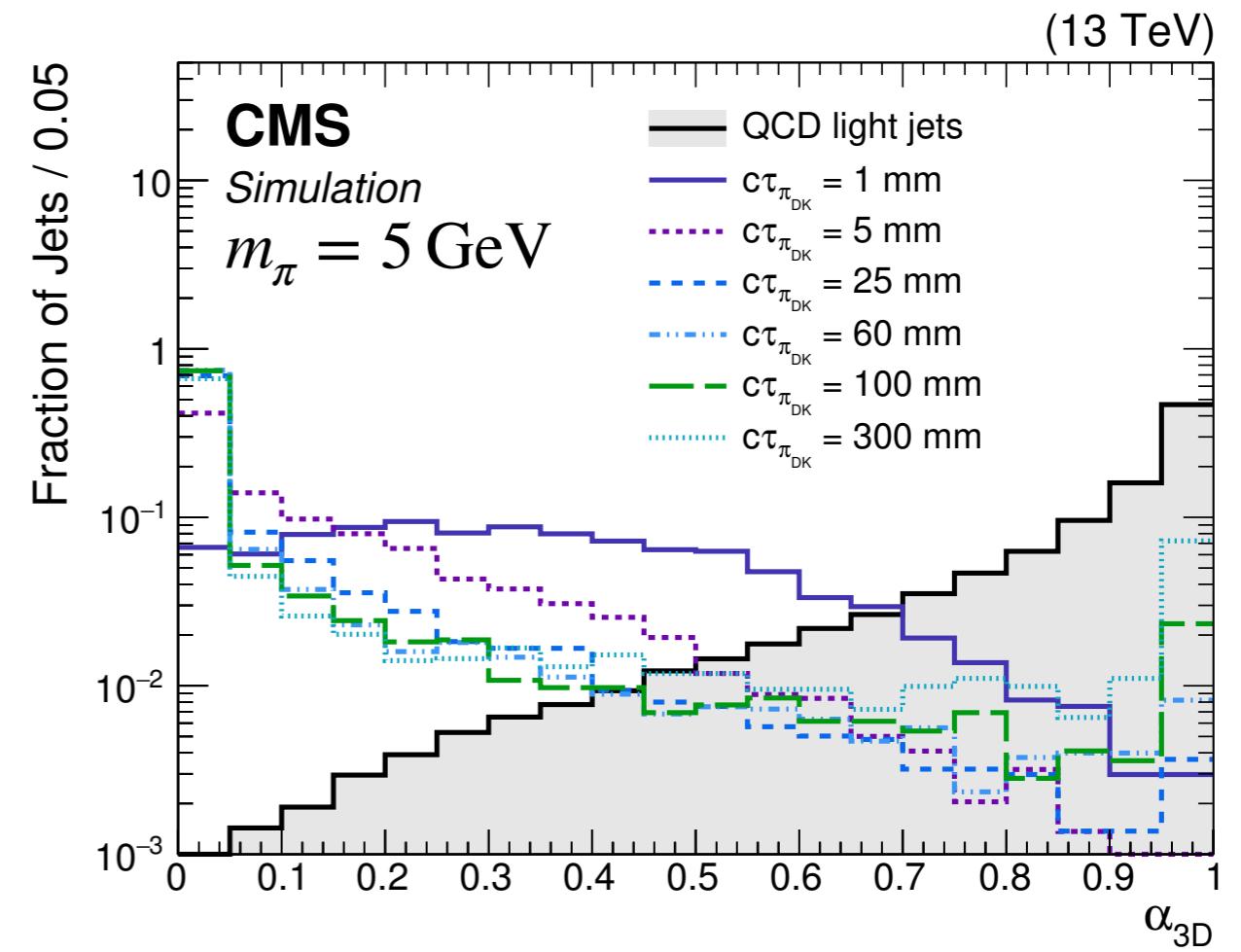
**Open Access** | Regular Article - Experimental Physics  
**First Online:** 26 February 2019

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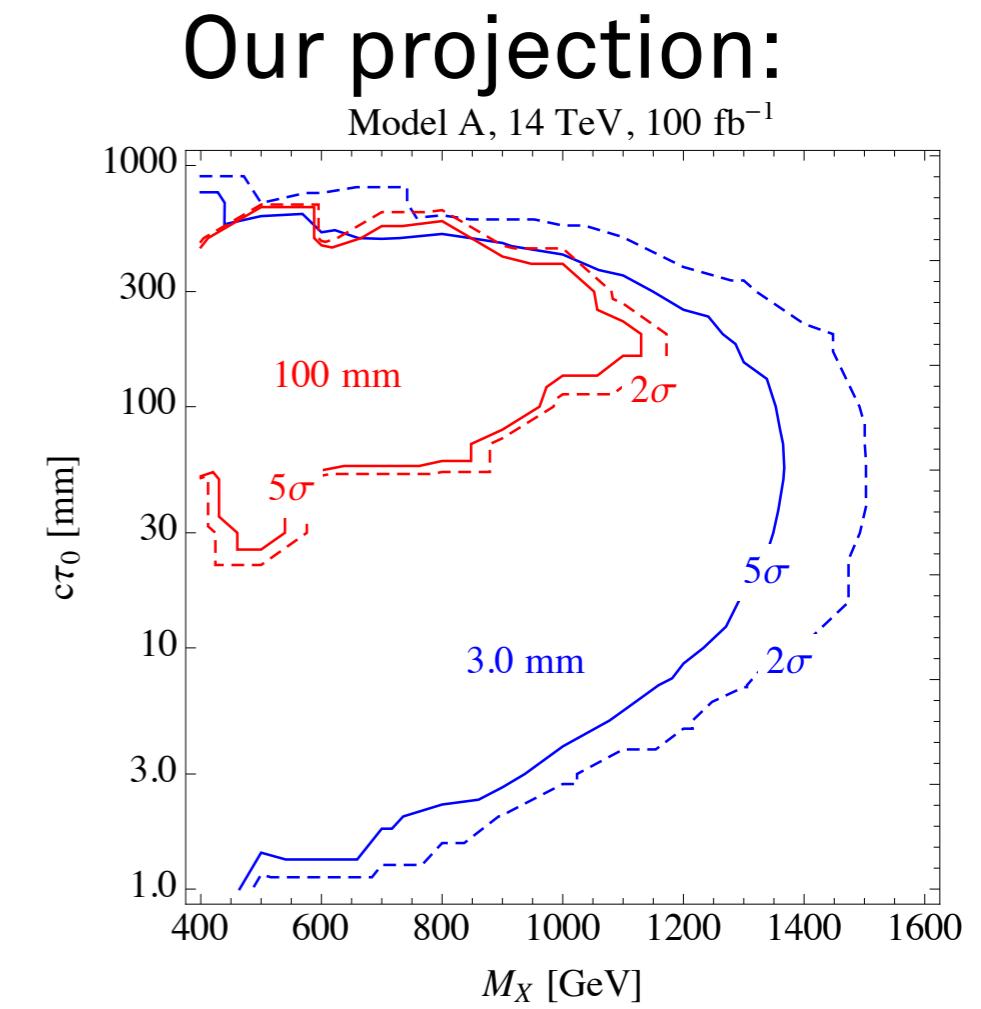
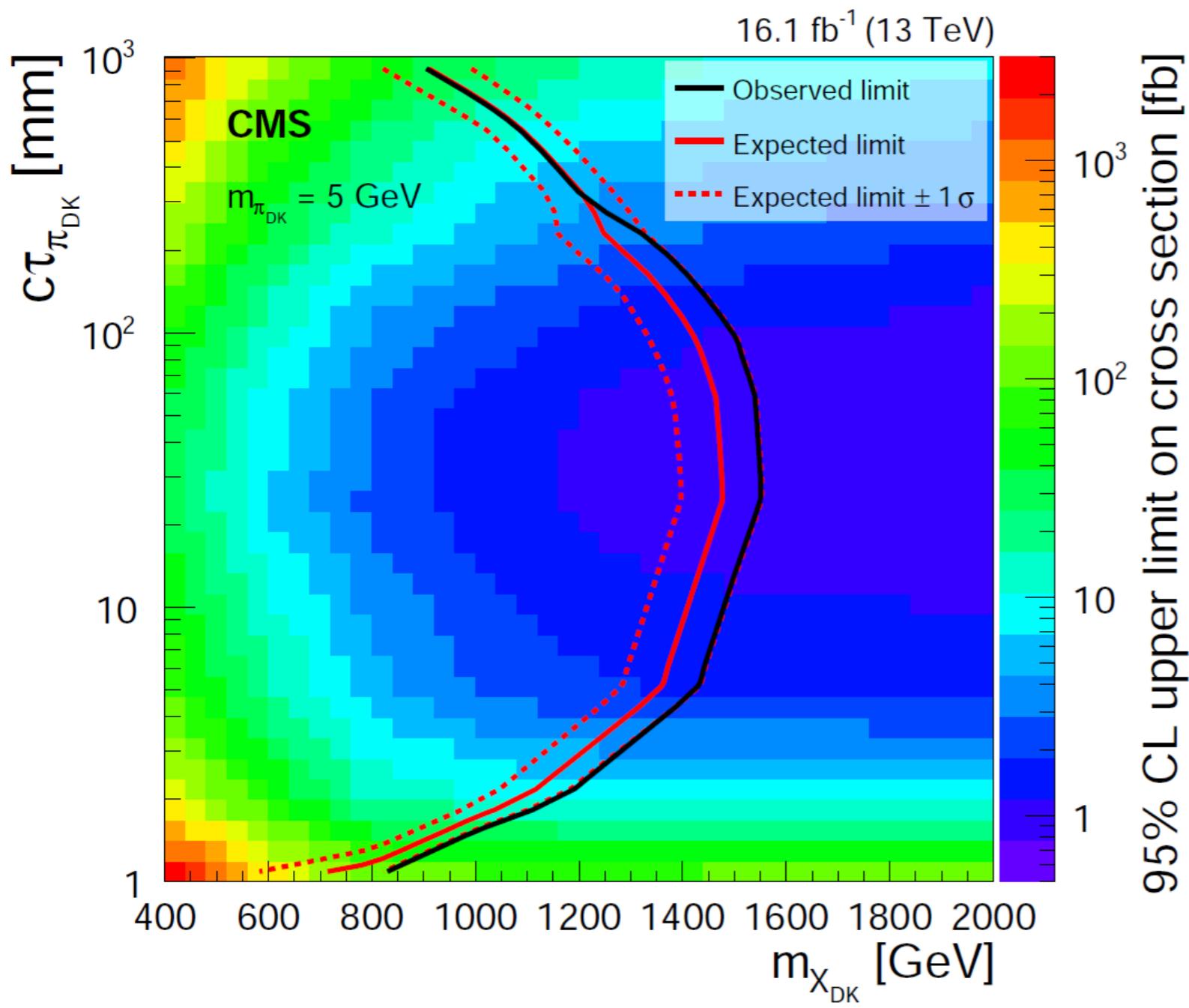
# DISCRIMINATING VARIABLES



Average transverse impact parameter.



Momentum fraction below a given displacement.



# SIGNAL REGIONS

Search used several signal regions.

Set number	Expected	Observed	Signal	Model parameters		
				$m_{X_{DK}}$ [GeV]	$m_{\pi_{DK}}$ [GeV]	$c\tau_{\pi_{DK}}$ [mm]
1	$168 \pm 15 \pm 5$	131	$36.7 \pm 4.0$	600	5	1
2	$31.8 \pm 5.0 \pm 1.4$	47	$(14.6 \pm 2.6) \times 10^2$	400	1	60
3	$19.4 \pm 7.0 \pm 5.5$	20	$15.6 \pm 1.6$	1250	1	150
4	$22.5 \pm 2.5 \pm 1.5$	16	$15.1 \pm 2.0$	1000	1	2
5	$13.9 \pm 1.9 \pm 0.6$	14	$35.3 \pm 4.0$	1000	2	150
6	$9.4 \pm 2.0 \pm 0.3$	11	$20.7 \pm 2.5$	1000	10	300
7	$4.40 \pm 0.84 \pm 0.28$	2	$5.61 \pm 0.64$	1250	5	225

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Me (dumb): Should I be excited about a  $\sim 3\sigma$  excess in set 2?

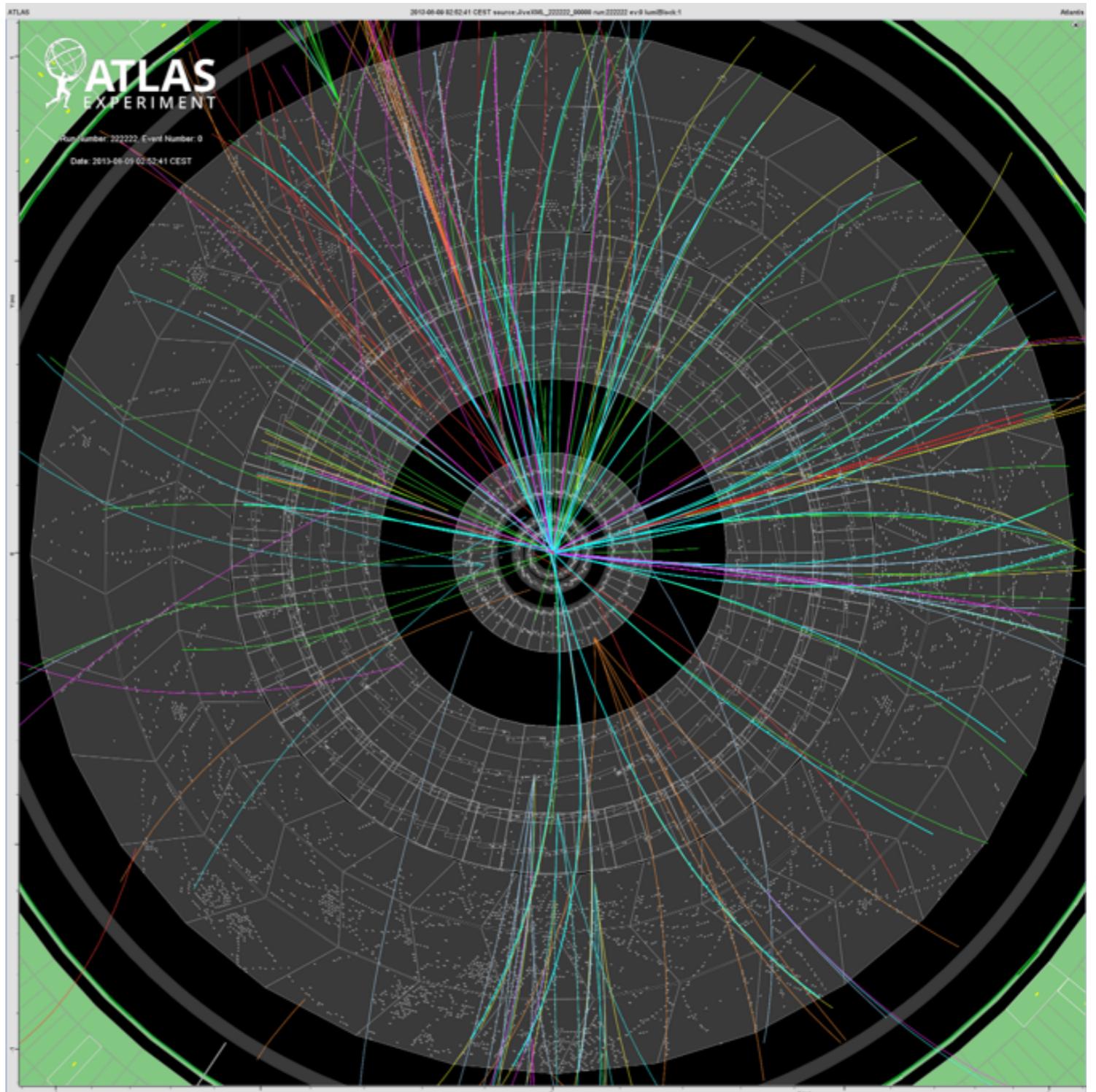
Ted (smart): Good question! I would say, it will be interesting to repeat the analysis next year with close to 10x more data!

# ATLAS DEDICATED SEARCH

ATLAS has a  
dedicated search  
in progress.

Carleton ATLAS  
group is involved.

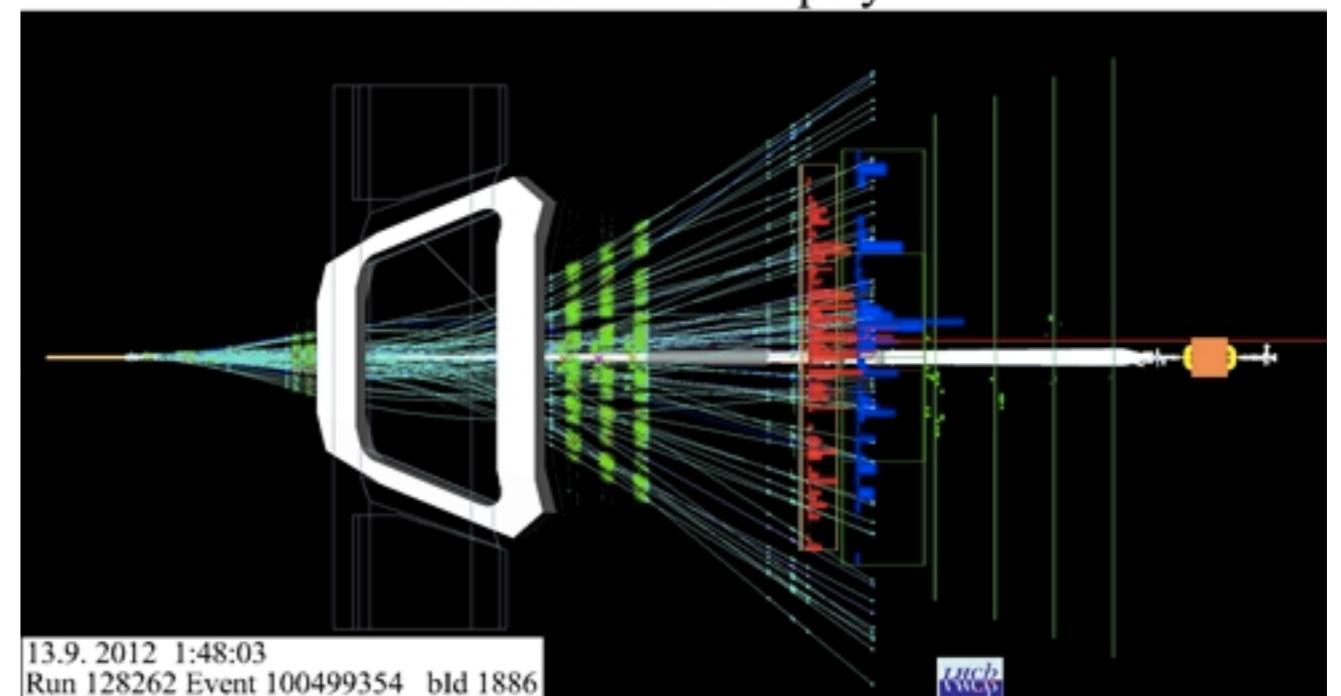
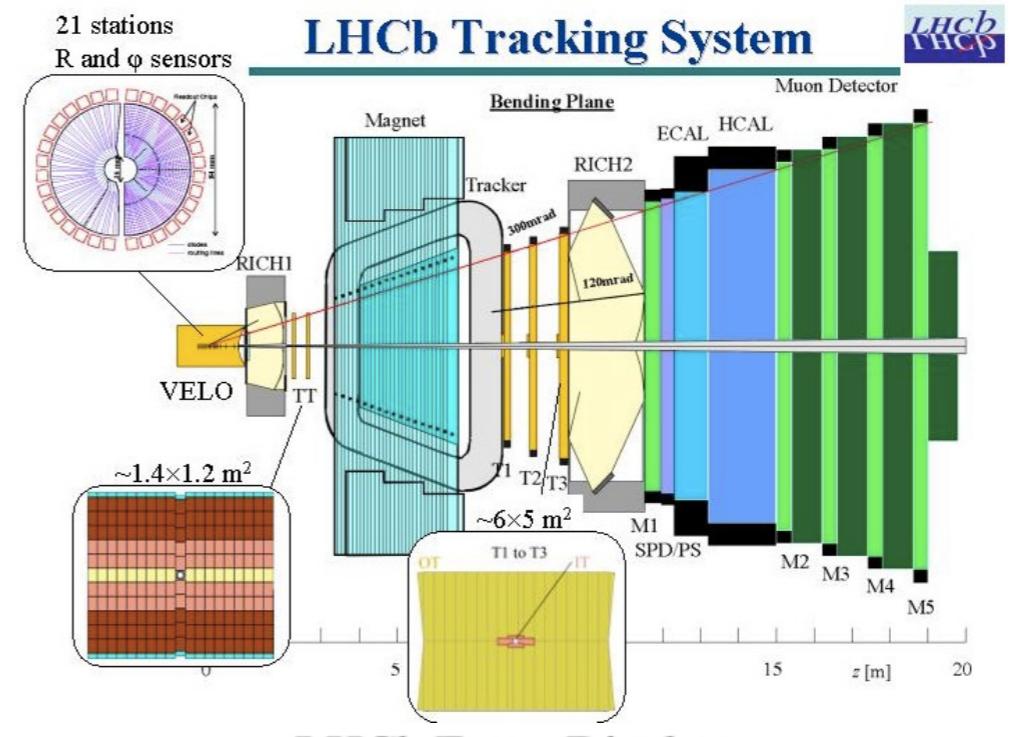
James Beacham,  
ATLAS Dark Sector Workshop,  
Cosenza, Italy, Feb 10, 2016.



# LHCb

LHCb has excellent tracking.

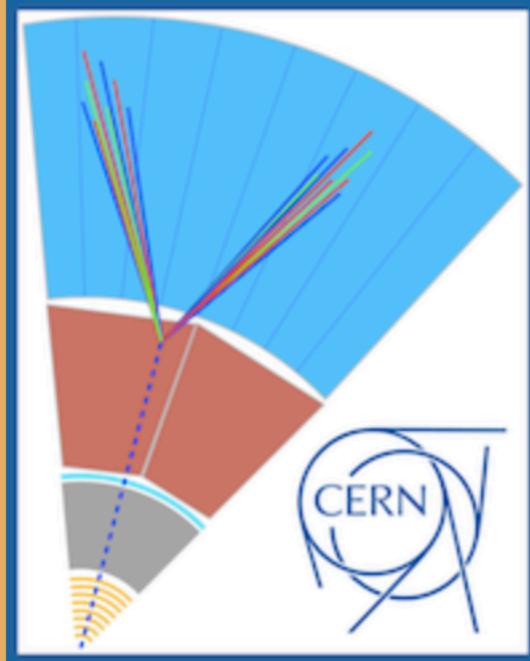
Limited coverage of event.





POTPOURRI

# LLP COMMUNITY



Searching for long-lived particles at the LHC and beyond:  
Thirteenth workshop of the LLP Community

Jun 19 – 23, 2023  
CERN  
Europe/Zurich timezone

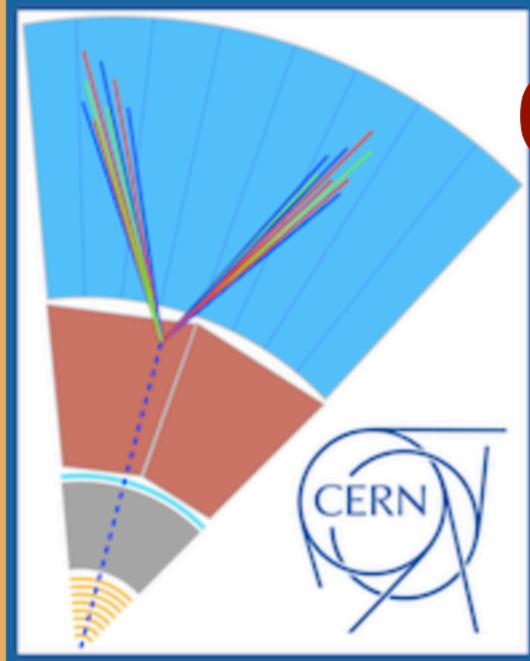
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Joint effort between experimentalists and theorists.

Lots to do if people are interested.

# LLP COMMUNITY



Searching for long-lived particles at the LHC and beyond:  
Thirteenth workshop of the LLP Community

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Joint effort between experimentalists and theorists.

Lots to do if people are interested.

# LJP COMMUNITY

## Searching for long-lived particles beyond the Standard Model at the Large Hadron Collider

March 6, 2019

### Editors:

Juliette Alimena<sup>(1)</sup> (Experimental Coverage, Backgrounds, Upgrades), James Beacham<sup>(2)</sup> (Document Editor, Simplified Models), Martino Borsato<sup>(3)</sup> (Backgrounds, Upgrades), Yangyang Cheng<sup>(4)</sup> (Upgrades), Xabier Cid Vidal<sup>(5)</sup> (Experimental Coverage), Giovanna Cottin<sup>(6)</sup> (Simplified Models, Reinterpretations), Albert De Roeck<sup>(7)</sup> (Experimental Coverage), Nishita Desai<sup>(8)</sup> (Reinterpretations), David Curtin<sup>(9)</sup> (Simplified Models), Jared A. Evans<sup>(10)</sup> (Simplified Models, Experimental Coverage), Simon Knapen<sup>(11)</sup> (Dark Showers), Sabine Kraml<sup>(12)</sup> (Reinterpretations), Andre Lessa<sup>(13)</sup> (Reinterpretations), Zhen Liu<sup>(14)</sup> (Simplified Models, Backgrounds, Reinterpretations), Sascha Mehlhase<sup>(15)</sup> (Backgrounds), Michael J. Ramsey-Musolf<sup>(16,126)</sup> (Simplified Models), Heather Russell<sup>(17)</sup> (Experimental Coverage), Jessie Shelton<sup>(18)</sup> (Simplified Models, Dark Showers), Brian Shuve<sup>(19,20)</sup> (Document Editor, Simplified Models, Simplified Models Library), Monica Verducci<sup>(21)</sup> (Upgrades), Jose Zurita<sup>(22,23)</sup> (Experimental Coverage)

**arXiv:1903.04497 (301 pages)**

# DARK SHOWERS



P. Schwaller

# DARK SHOWERS

Lots more interesting theory and experimental work.

Working towards a thorough experimental program.

Many interested experimentalists.

## *New Frontiers: Dark Showers*

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Chapter editors: Simon Knapen, Jessie Shelton

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# SNOWMASS ACTIVITY

arXiv > hep-ph > arXiv:2203.09503

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## High Energy Physics – Phenomenology

[Submitted on 17 Mar 2022 ([v1](#)), last revised 27 Jun 2022 (this version, v2)]

# Theory, phenomenology, and experimental avenues for dark showers: a Snowmass 2021 report

Guillaume Albouy, Jared Barron, Hugues Beauchesne, Elias Bernreuther, Marcella Bona, Cesare Cazzaniga, Cari Cesarotti, Timothy Cohen, Annapaola de Cosa, David Curtin, Zeynep Demiragli, Caterina Doglioni, Alison Elliot, Karri Folan DiPetrillo, Florian Eble, Carlos Erice, Chad Freer, Aran Garcia–Bellido, Caleb Gemmell, Marie–Hélène Genest, Giovanni Grilli di Cortona, Giuliano Gustavino, Nicoline Hemme, Tova Holmes, Deepak Kar, Simon Knapen, Suchita Kulkarni, Luca Lavezzo, Steven Lowette, Benedikt Maier, Seán Mee, Stephen Mrenna, Harikrishnan Nair, Jeremi Niedziela, Christos Papageorgakis, Nukulsinh Parmar, Christoph Paus, Kevin Pedro, Ana Peixoto, Alexx Perloff, Tilman Plehn, Christiane Scherb, Pedro Schwaller, Jessie Shelton, Akanksha Singh, Sukanya Sinha, Torbjörn Sjöstrand, Aris G.B. Spourdalakis, Daniel Stolarski, Matthew J. Strassler, Andrii Usachov, Carlos Vázquez Sierra, Christopher B. Verhaaren, Long Wang

112 pages.

# FUTURE DETECTORS

Many concepts for additional detectors to detect LLPs produced in LHC collisions.

- milliQan
- SHiP
- MATHUSLA
- MoEDAL-MAPP **taking data**
- FORMOSA
- FASER **under construction**
- CODEX-b
- AL3X
- ANUBIS
- SND

# FUTURE DETECTORS

## A facility to Search for Hidden Particles at the CERN SPS: the SHiP physics case

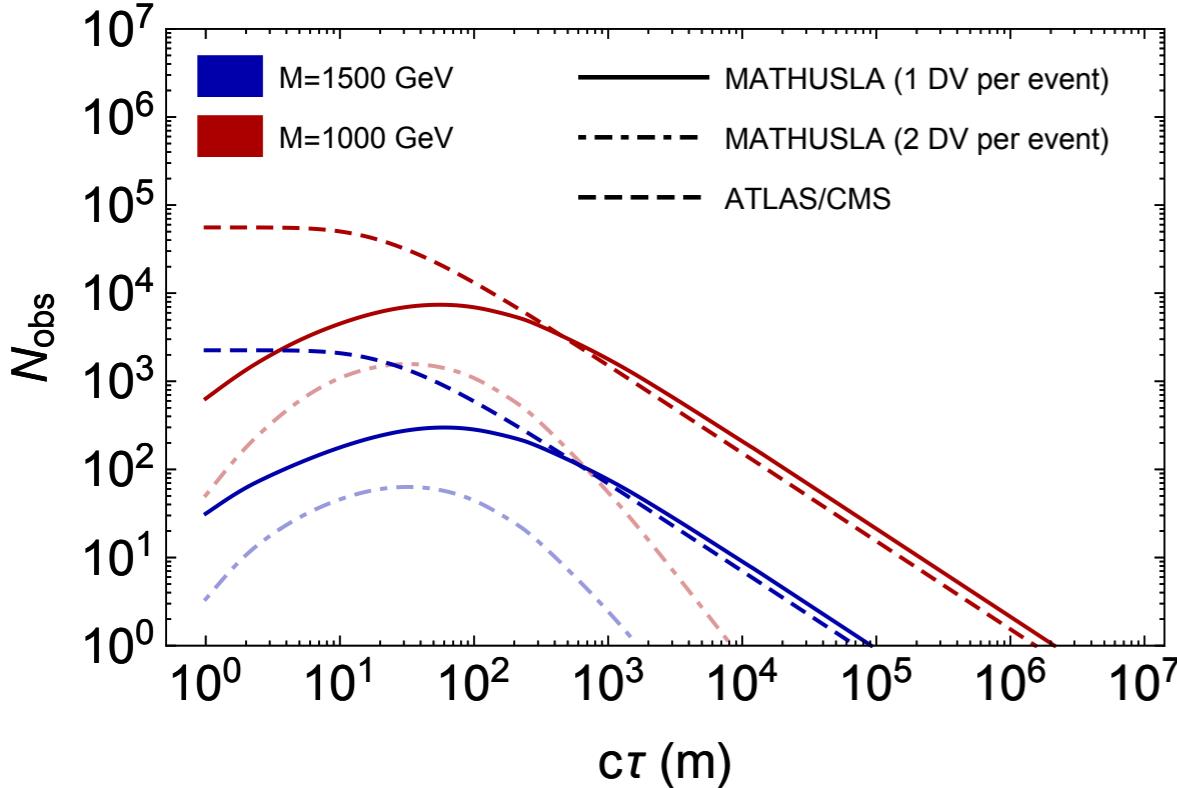
### 3.6 Dark pions

#### 3.6.1 The model and scales

#### 3.6.2 Dark pion lifetime and decay modes

#### 3.6.3 What SHiP could do?

1504.04855

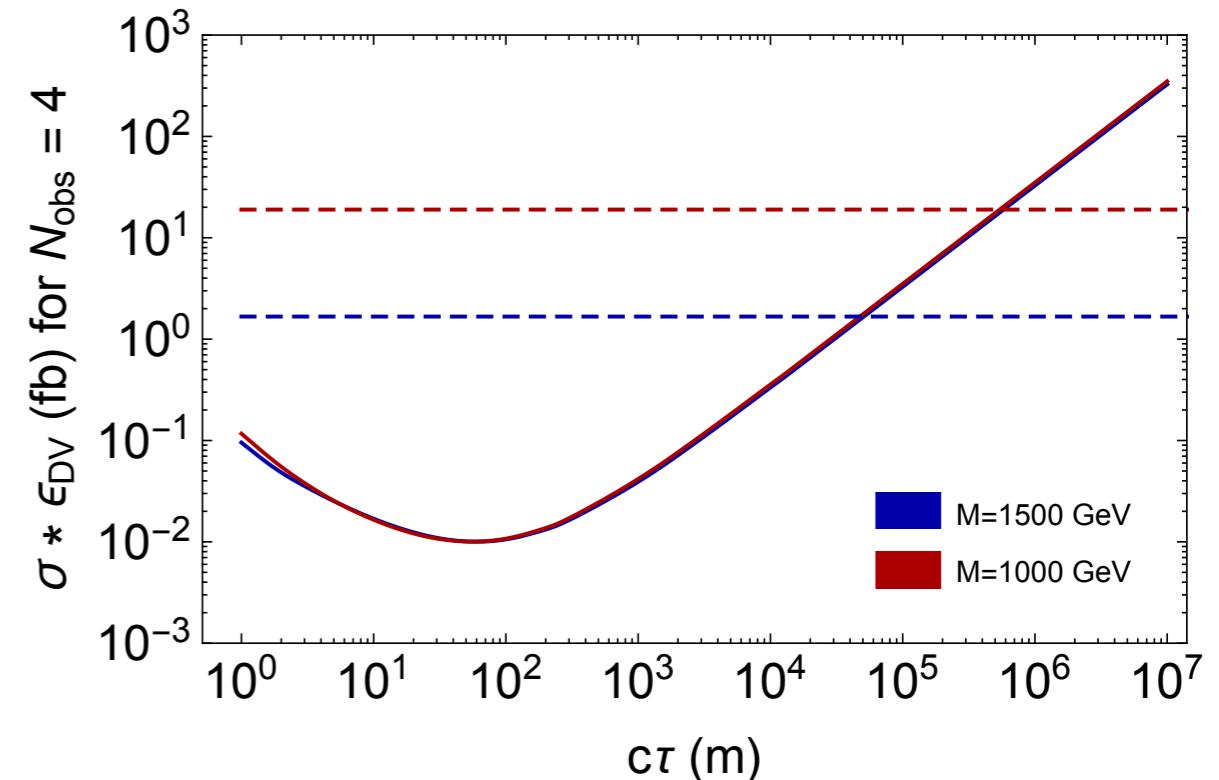


## Long-Lived Particles at the Energy Frontier: The MATHUSLA Physics Case

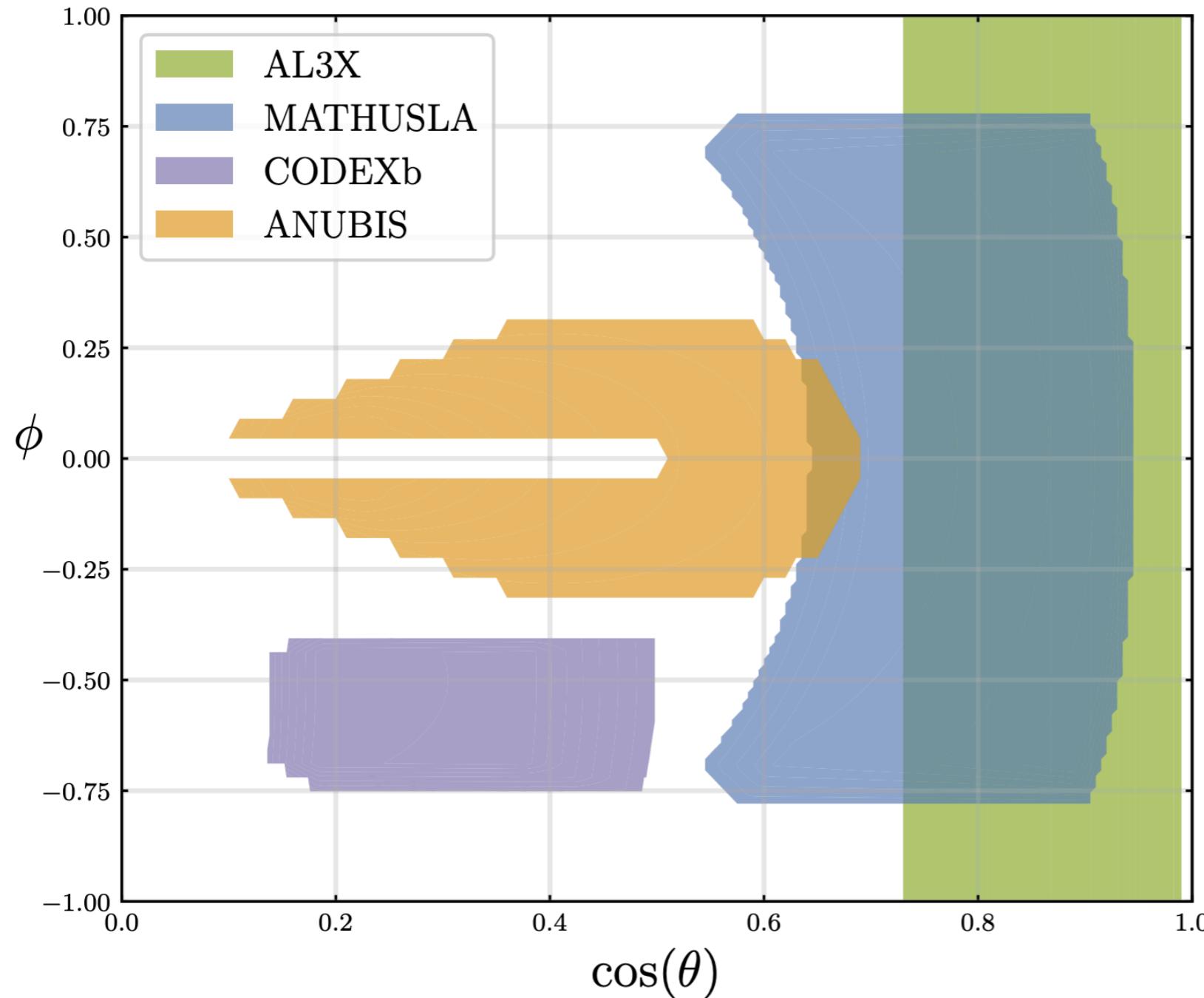
8 Theory Motivation for LLPs: Bottom-Up Considerations

8.1 Hidden Valleys and High Multiplicity Scenarios . . . .

1806.07396



# GEOMETRY

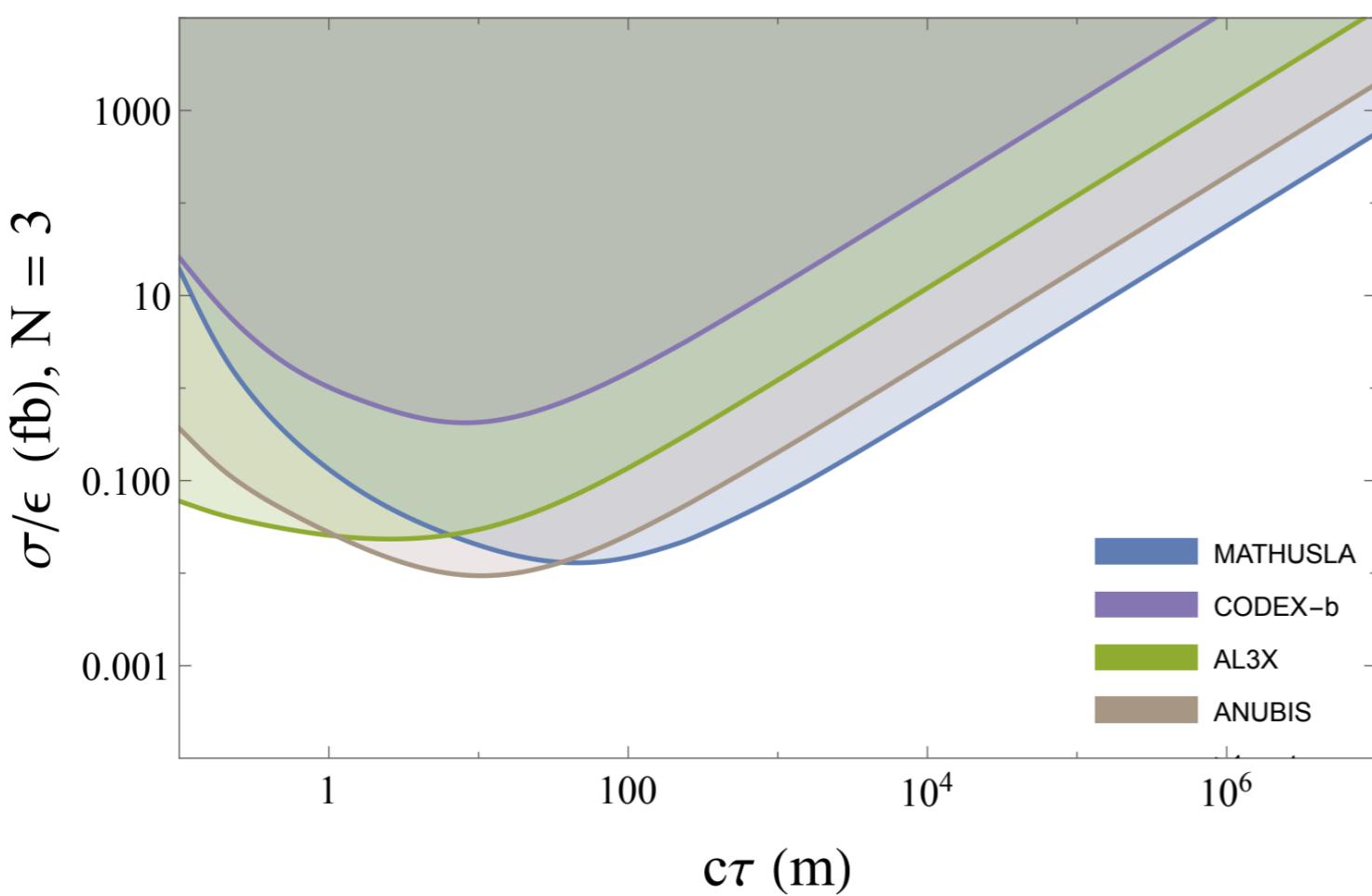


Other detectors are forward, have much smaller efficiency for heavy mediators.

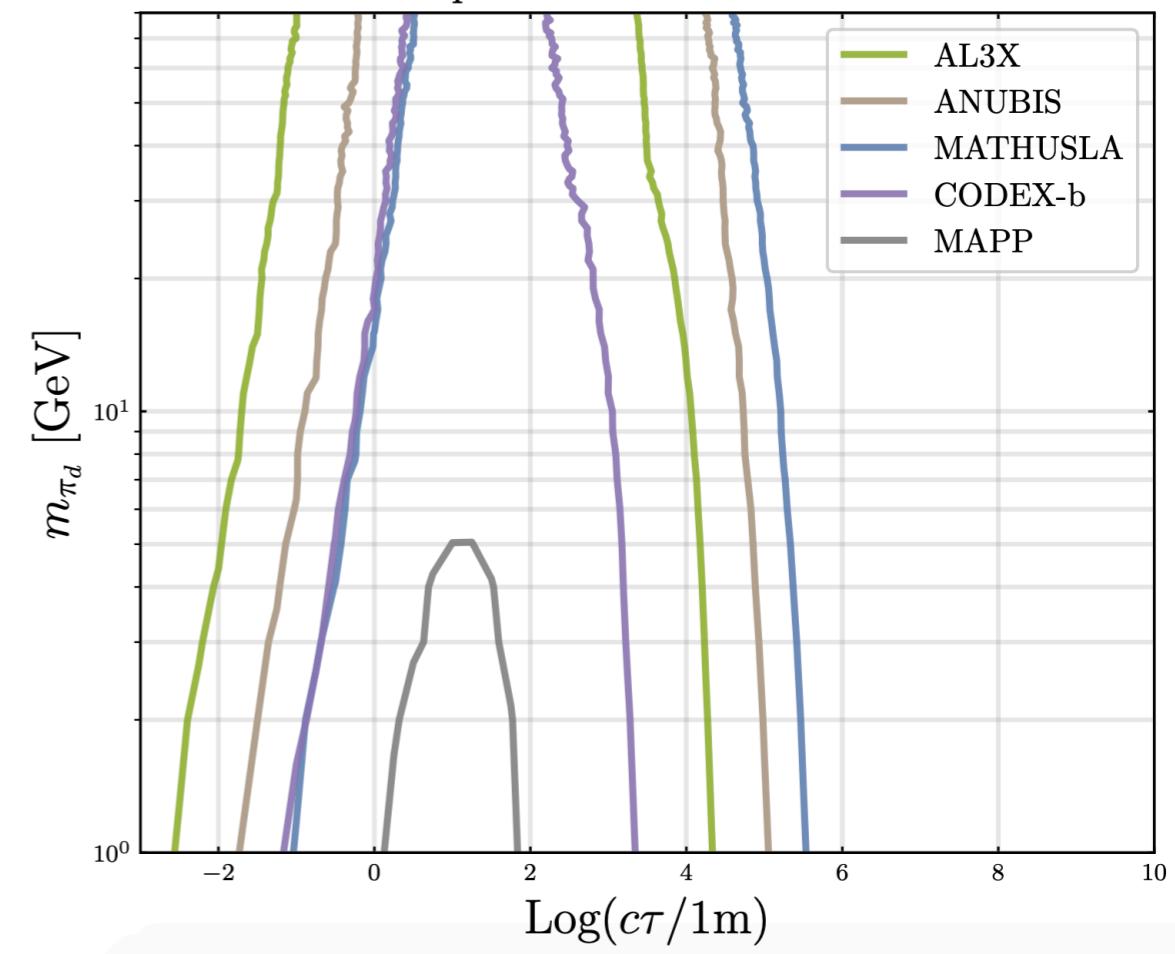
Archer-Smith, Linthorne, DS, arXiv:2112.05690.

# COMPARISON

Limit On Cross Section,  $M_X = 1 \text{ TeV}$ ,  $m_\pi = 1 \text{ GeV}$



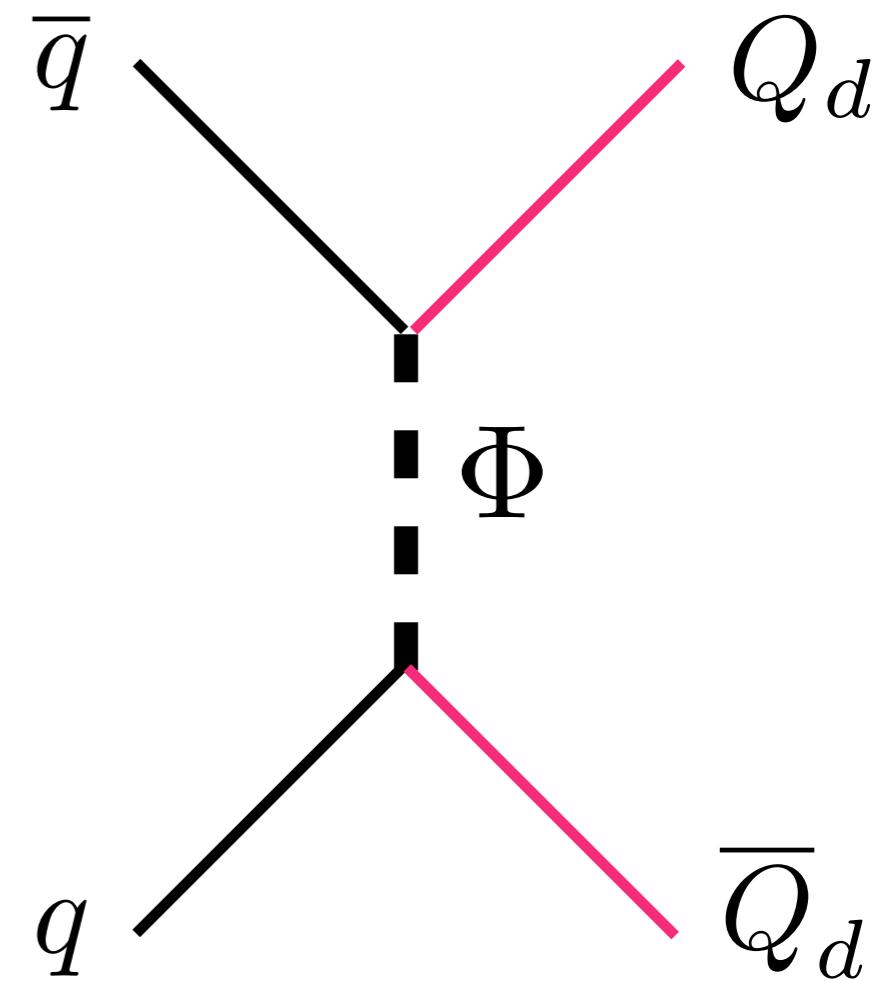
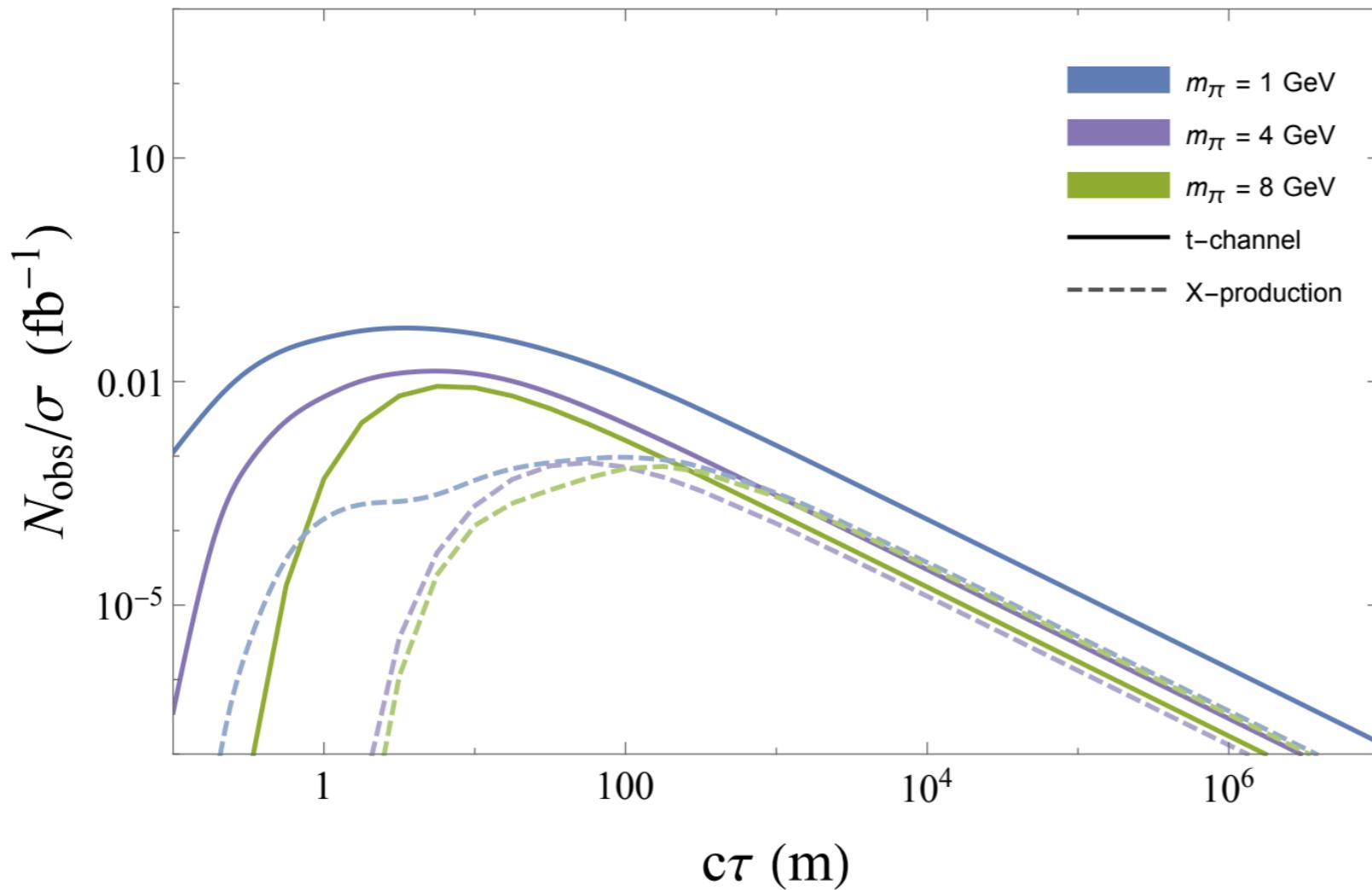
LLP Experiment Contours for  $N = 3$



Archer-Smith, Linthorne, DS, arXiv:2112.05690.

# T-CHANNEL

FASER Detector Acceptance,  $M_X = 1 \text{ TeV}$



Archer-Smith, Linthorne, DS, arXiv:2112.05690.

# TRIGGERING?

LHC experiments need to reduce data rate from  $\sim 10$  MHz to  $\sim 1$  kHz.

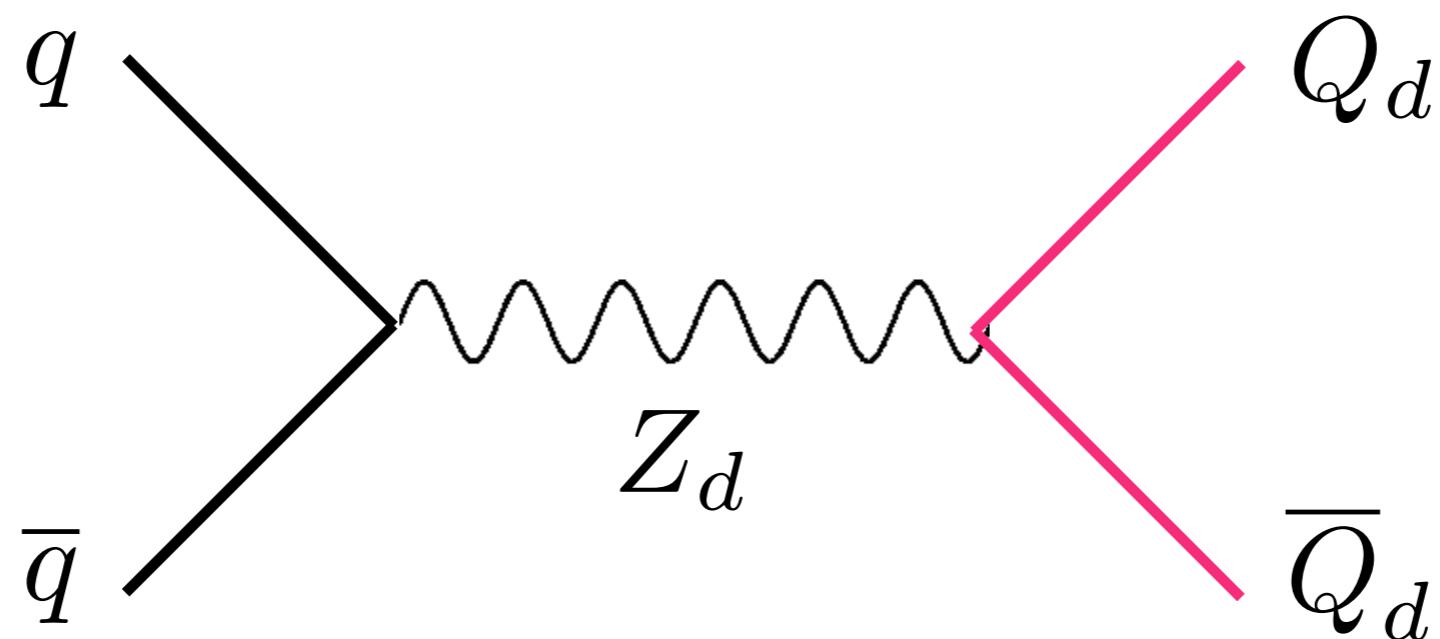
Current triggers look for total energy, jets, leptons, etc.

Not designed for exotic scenarios like emerging jets.

# TRIGGER STRATEGIES

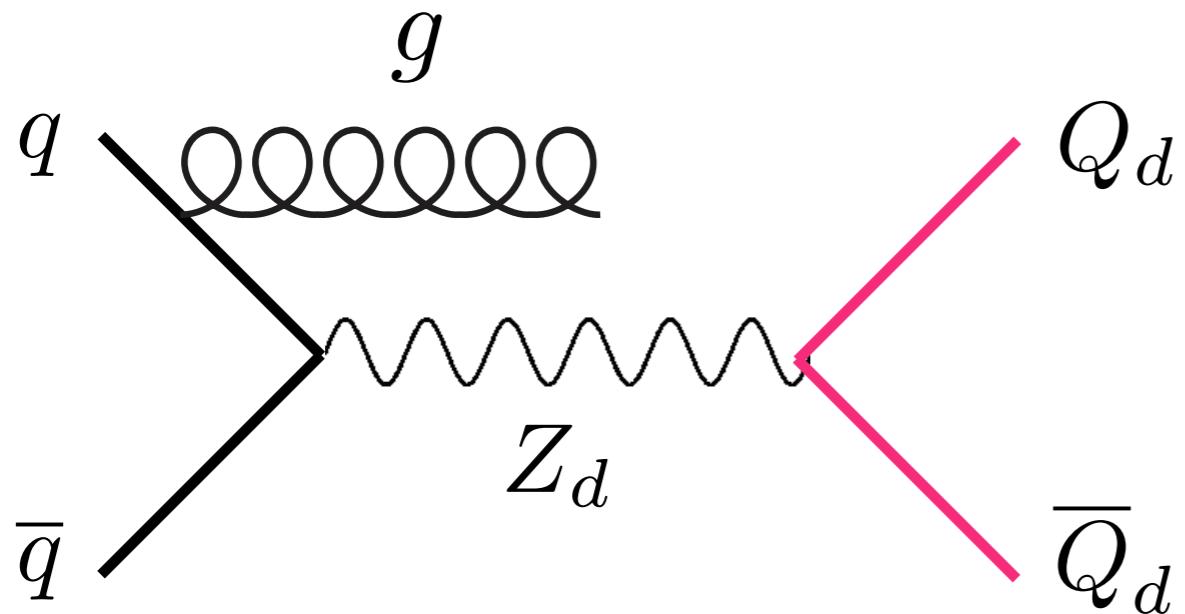
Original study used high mass mediators that produce extra jets: this makes triggering easy.

Can consider more difficult case with vector mediator (with potentially lower mass):



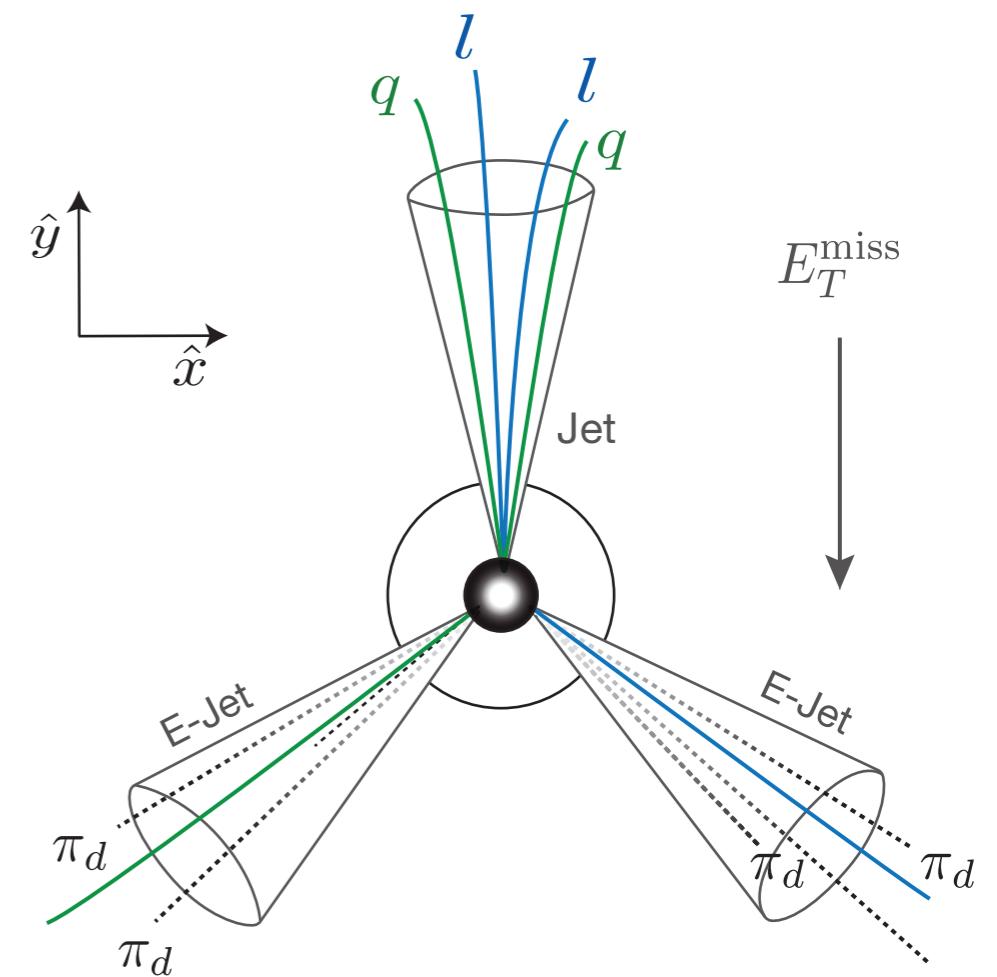
# ISR TRIGGERS

Strategy: Radiate a gluon (or a photon, W/Z, etc).



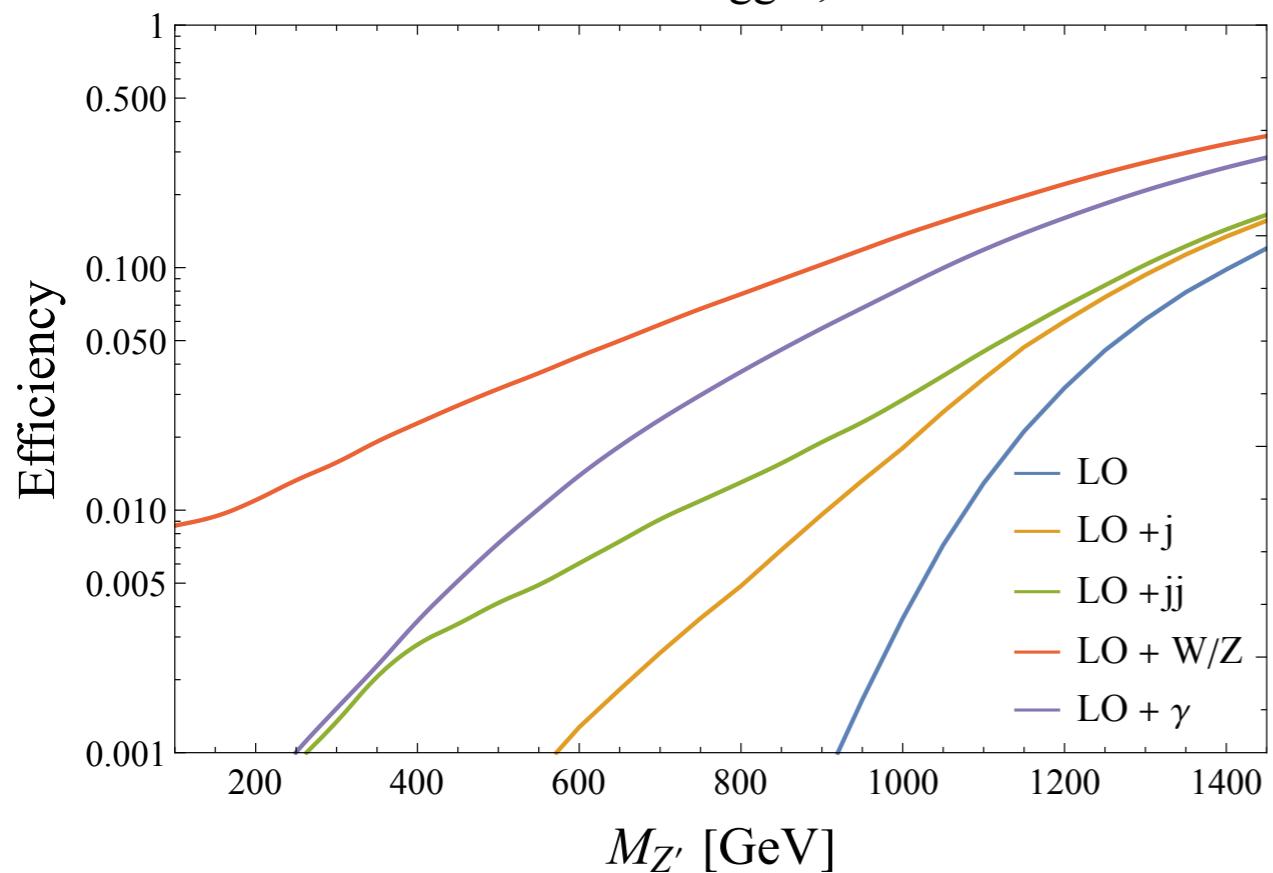
Goodman et.al. 1008.1783.

Fox et. al., 1109.4398.

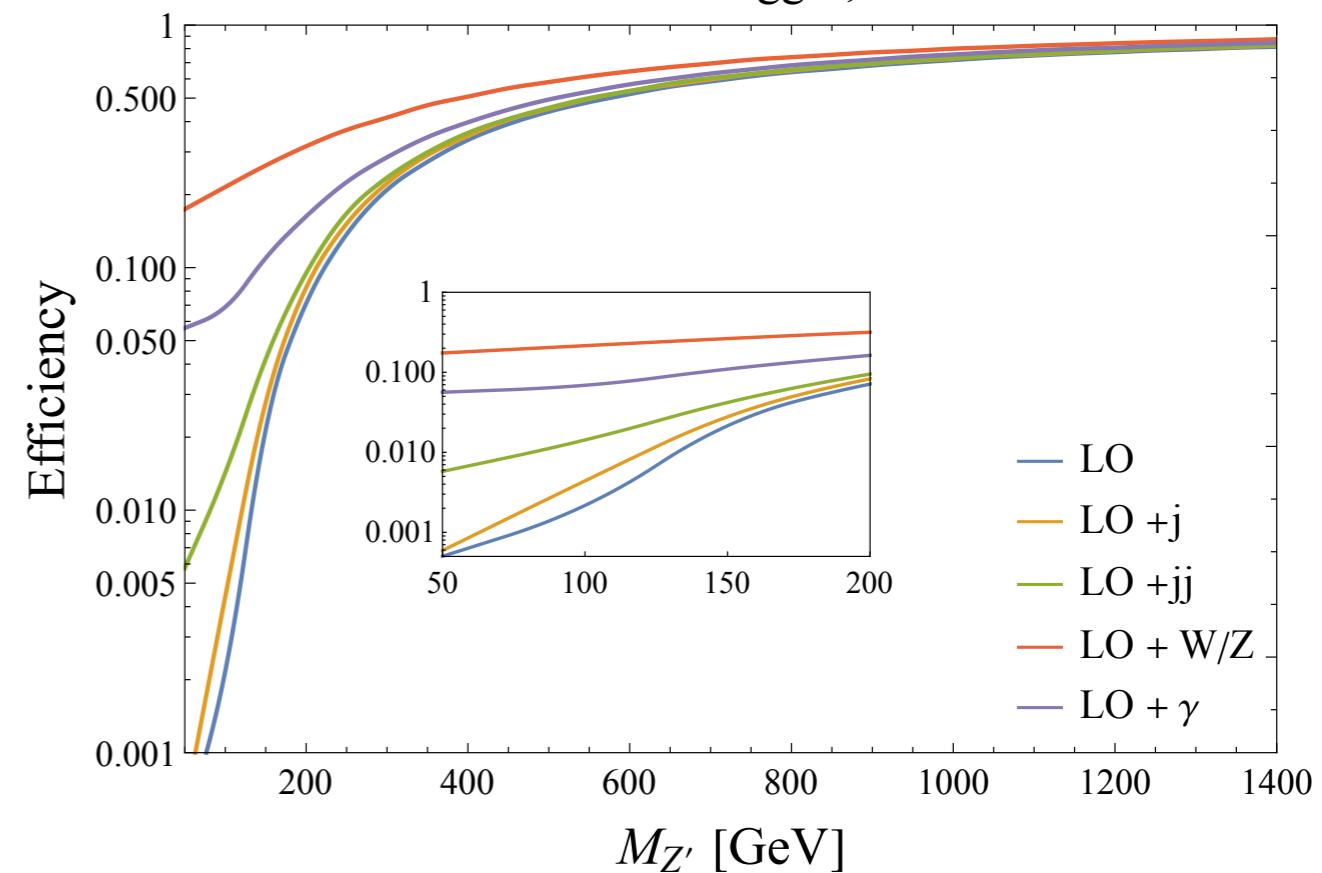


# EFFICIENCY IMPROVEMENT

ATLAS HT Trigger, Model A



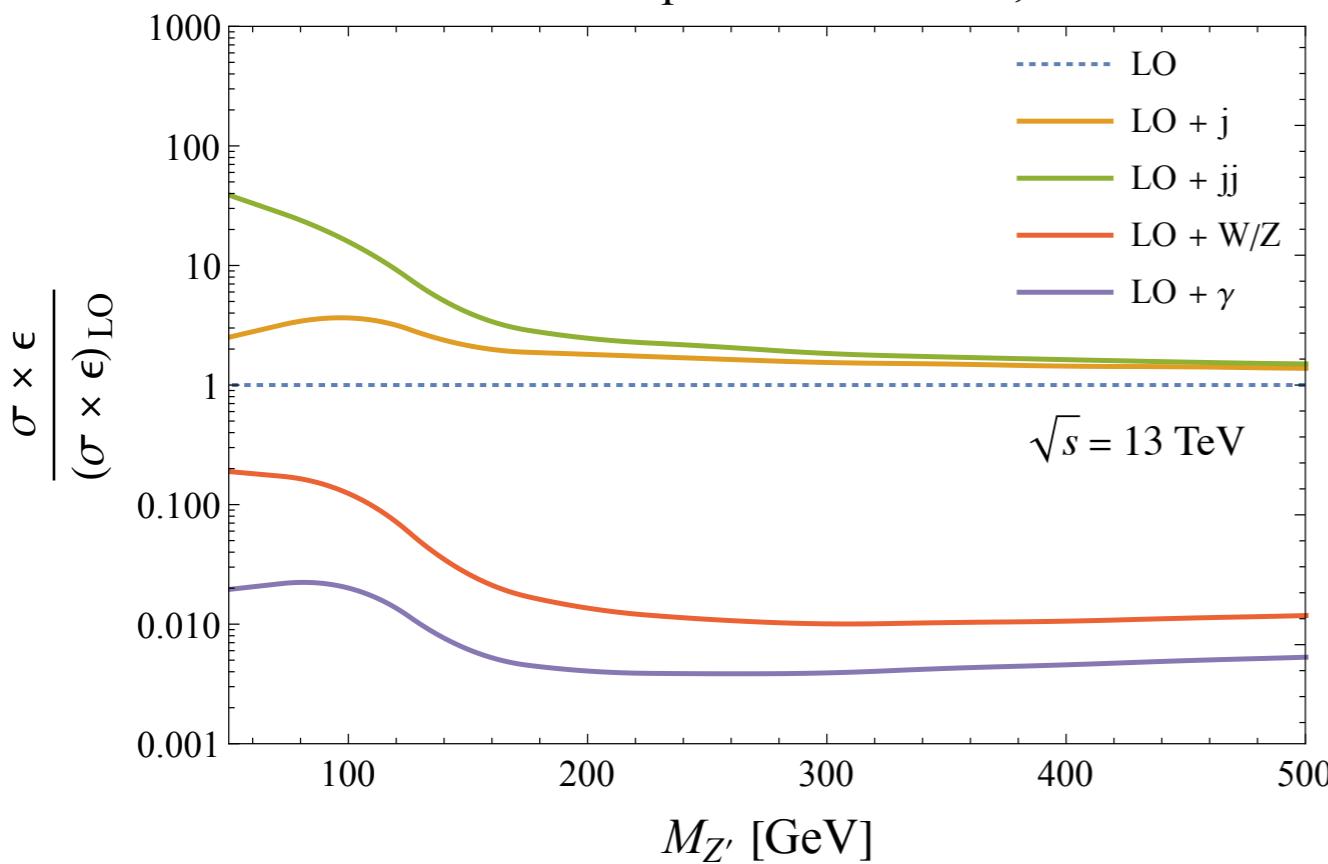
ATLAS MET Trigger, Model A



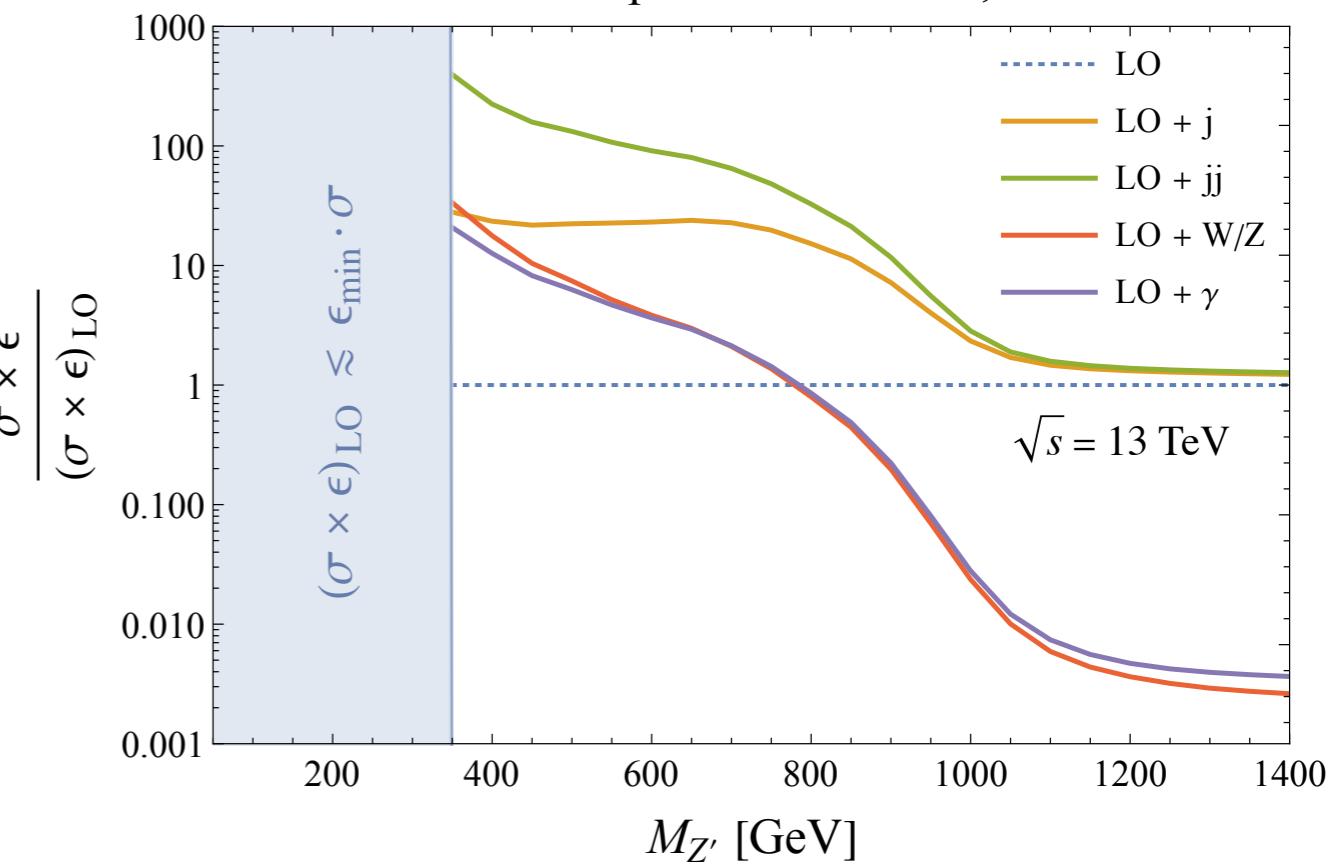
Linthorne, DS, arXiv:2103.08620.

# RATE IMPROVEMENT

ATLAS MET Improvement Factor, Model A



ATLAS HT Improvement Factor, Model B



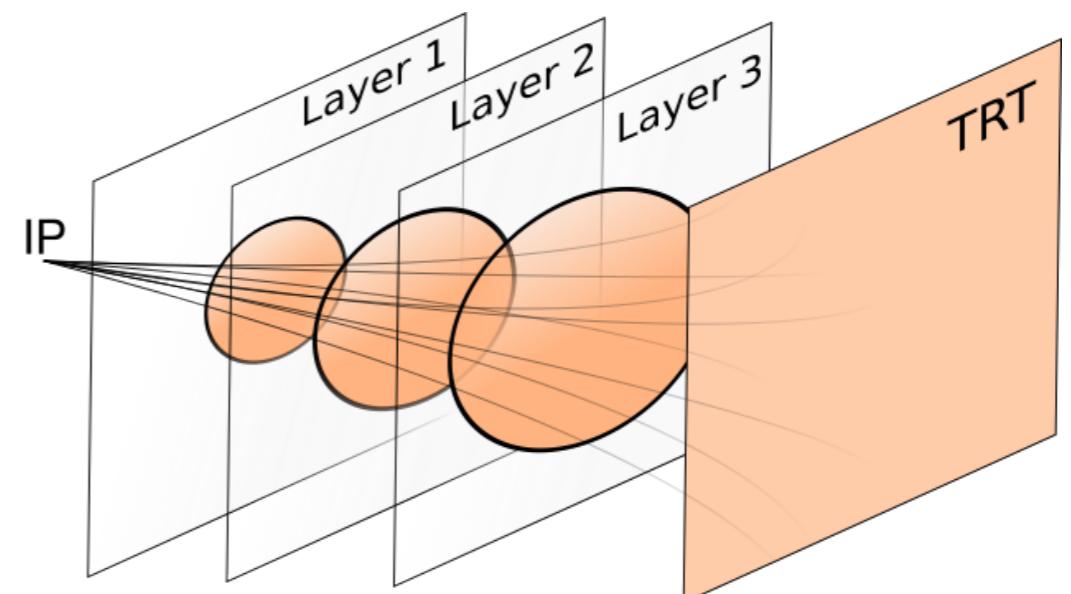
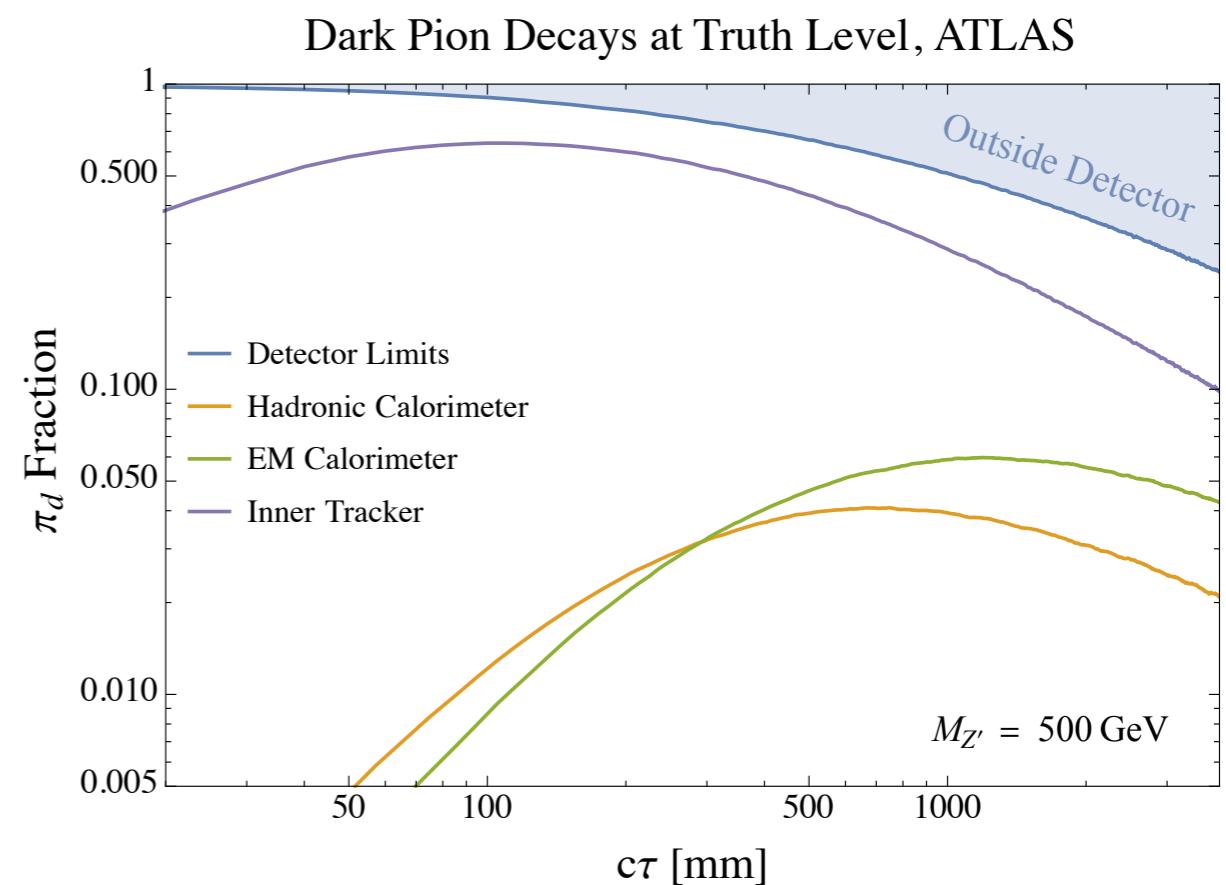
Linthorne, DS, arXiv:2103.08620.

# NEW TRIGGERS?

The tracker has the largest number of decays.

Can we just count hits in different layers of the tracker?

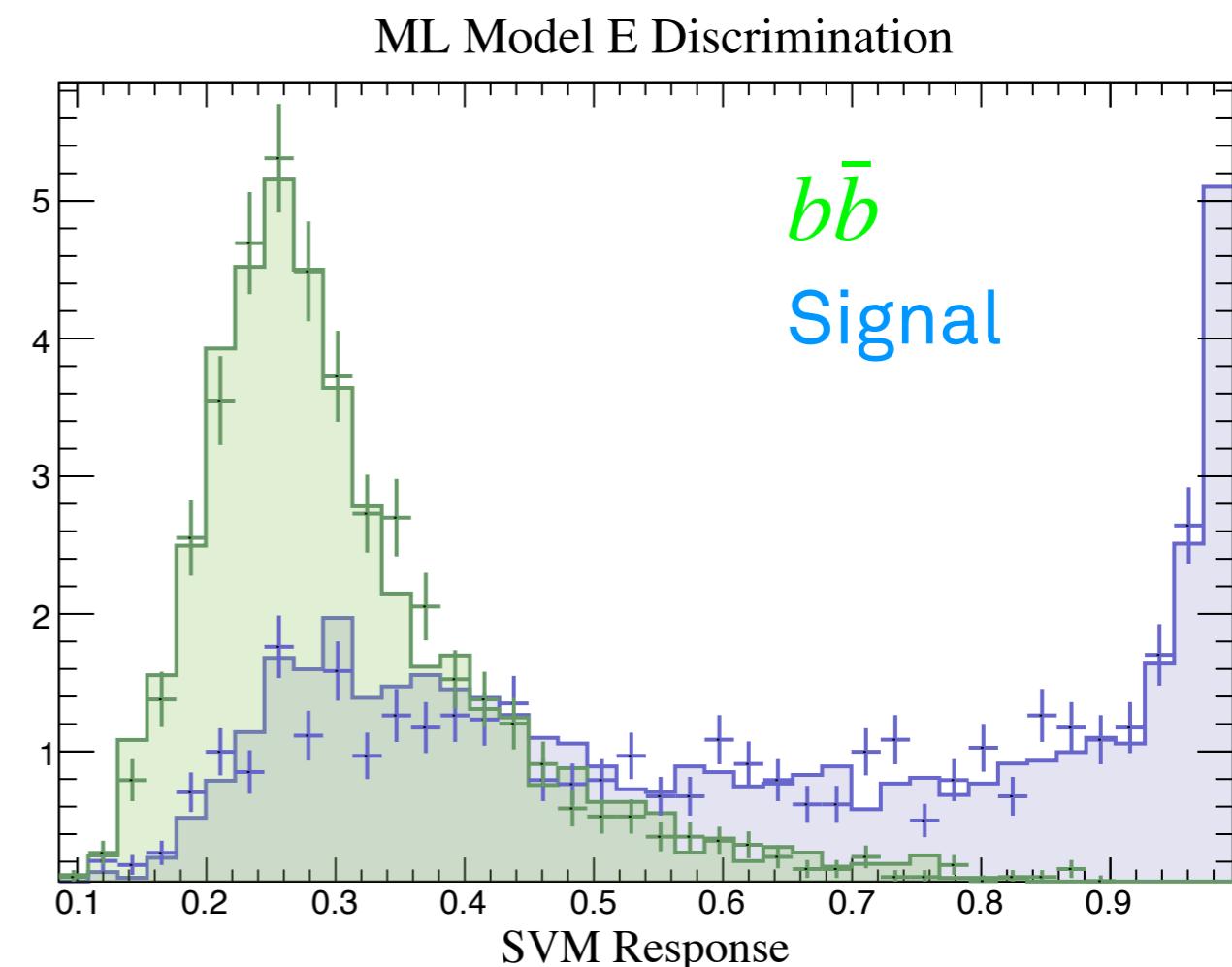
Fast: do not have to reconstruct tracks.



# THE MACHINE LEARNS

Use machine learning to enhance discriminating power.

Downside of ML:  
no physical  
understanding. Not a  
problem for triggers.

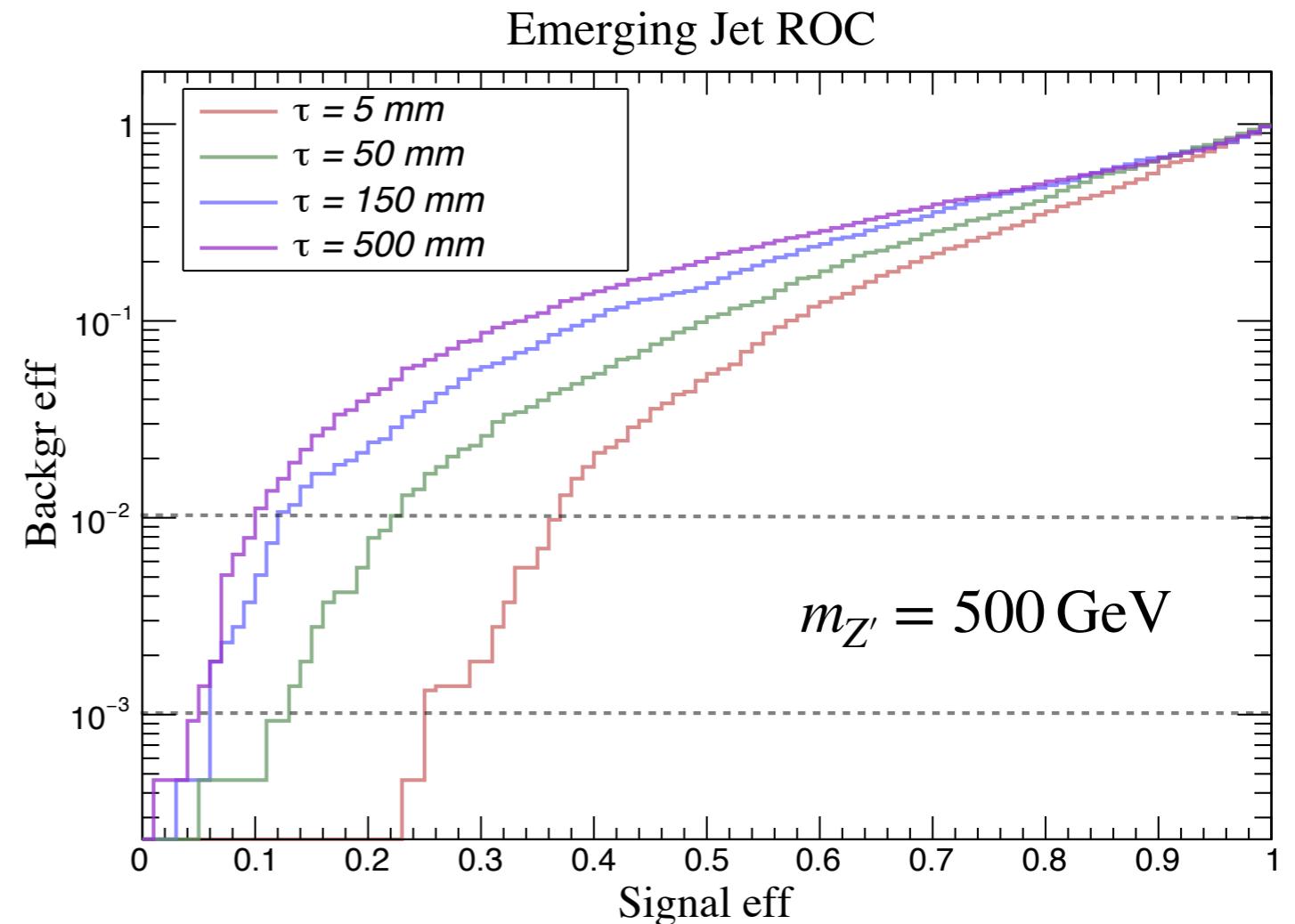


Linthorne, DS, arXiv:2103.08620.

# EFFICIENCY

Need to reject hard QCD at  $\sim 1/1000$ .

Can get efficiencies of  $\sim 10\%$ .

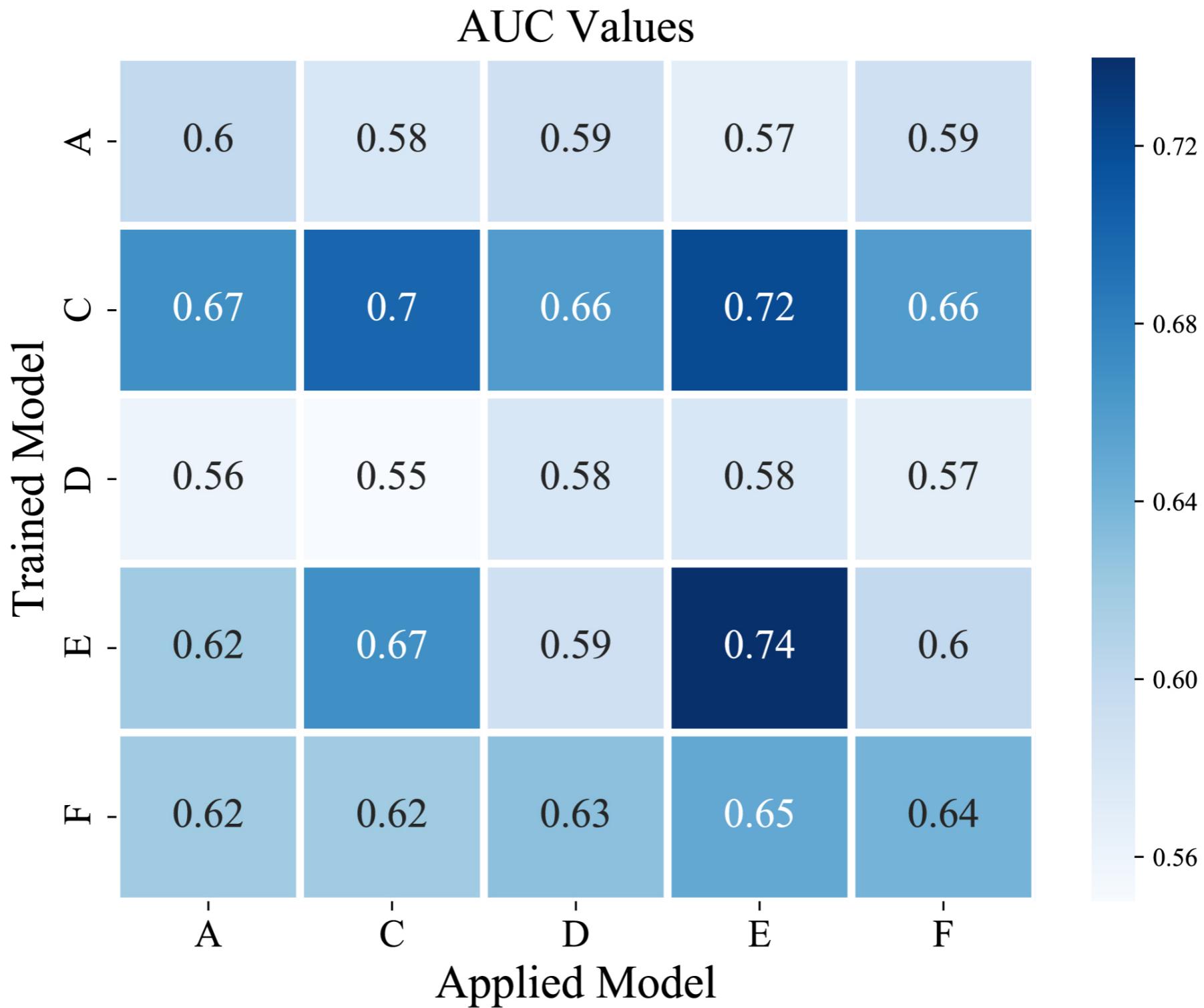


See also:

[Alimena, Iiyama, Kieseler, 2004.10744](#)

and Bernreuther et. al. 2011.06604.

# UNIVERSALITY

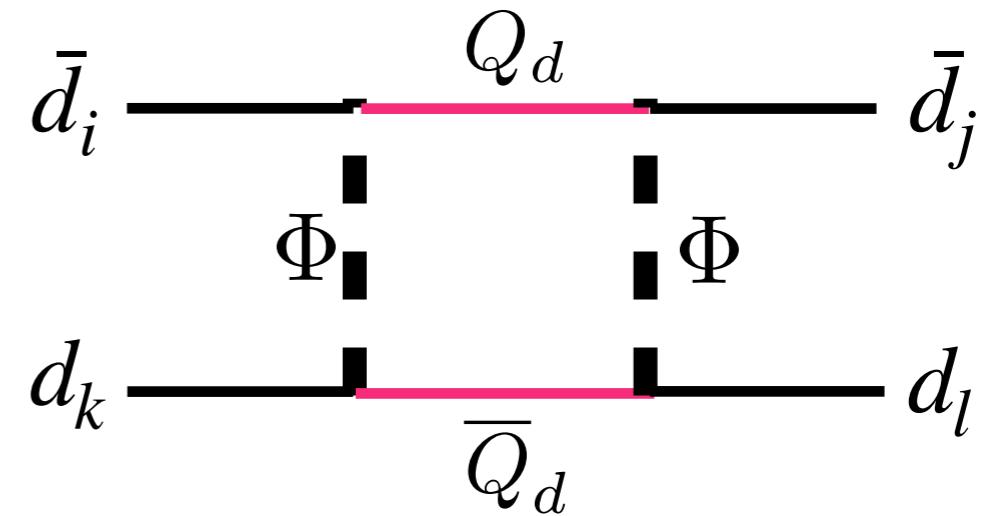
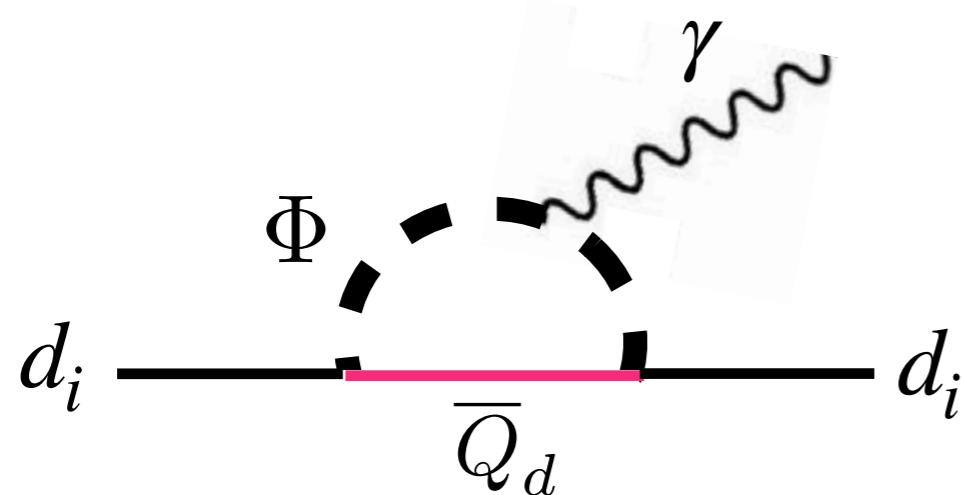


# FLAVOUR CONSTRAINTS

Yukawa couplings to mediator has flavour and dark flavour indices.

$$\lambda_{ij} \bar{Q}_j \Phi d_i \rightarrow \begin{array}{l} \text{SM quark (d, s, b)} \\ \text{bifundamental scalar} \\ \text{dark quark} \end{array}$$

One loop processes can lead to flavour and CP violating transitions.



# ELECTRIC DIPOLE MOMENTS

TeV scale CP violating new physics typically strongly constrained by (C)EDM bounds.

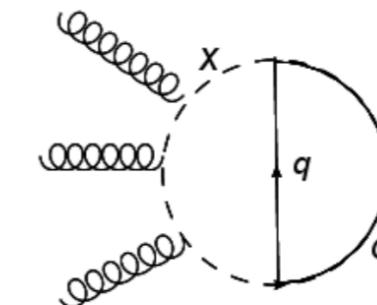
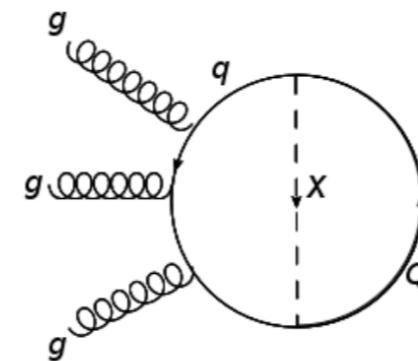
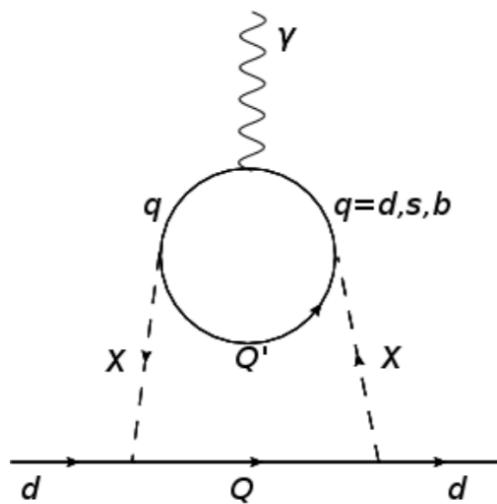
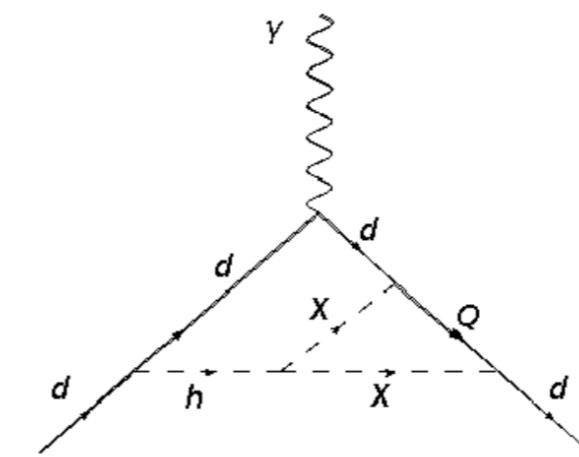
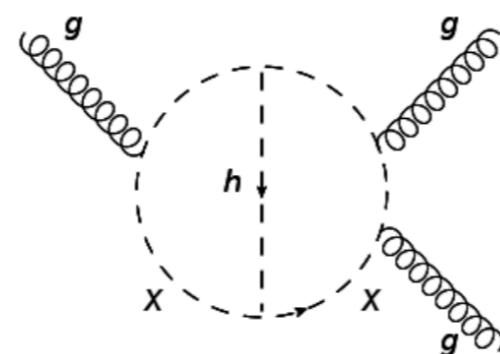
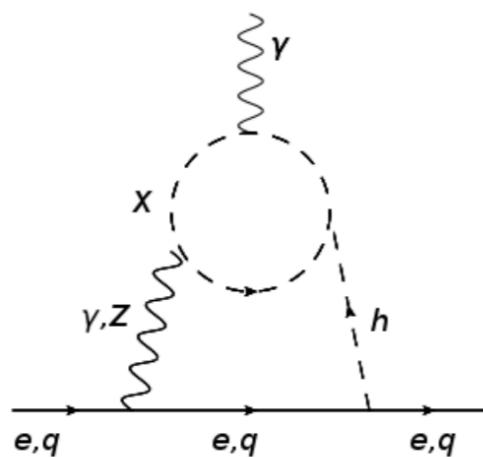
$$\text{Im} [ d_i \rightarrow \bar{Q}_d \Phi \rightarrow \gamma ]$$

$$\propto \text{Im} [\lambda_{dQ} \lambda_{dQ}^*] = 0$$

Bensalem, DS, arXiv:2111.05515.

# ELECTRIC DIPOLE MOMENTS

Two loop diagrams are all real by similar arguments.



Bensalem, DS, arXiv:2111.05515.

# ELECTRIC DIPOLE MOMENTS

Leading contribution to nEDM is at 3 loop!

$$\sim e \lambda_{ij}^4 \left( \frac{1}{16\pi^2} \right)^3 \left( \frac{m_d}{M_X^2} \right) \sim 10^{-29} e \cdot \text{cm} \left( \frac{\text{TeV}}{M_x} \right)^2 < 10^{-26} e \cdot \text{cm}$$

EDMs do not constrain this model (yet).

Bensalem, DS, arXiv:2111.05515.

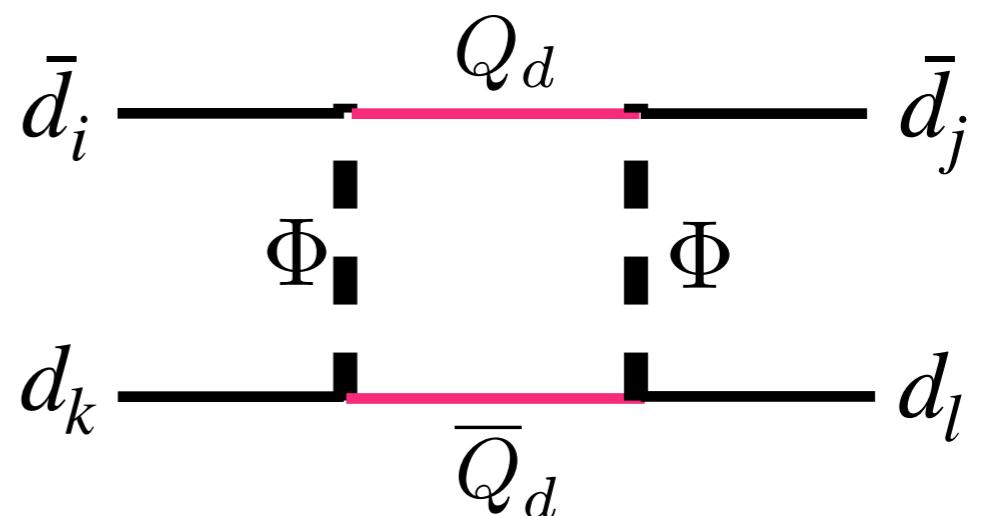
# FLAVOUR CONSTRAINTS

Low energy flavour experiments place constraints on the real and imaginary parts of expressions like:

$$\frac{1}{M_\Phi^2} \sum_{Q Q'} \lambda_{qQ} \lambda_{q'Q'}^* \lambda_{q'Q'} \lambda_{qQ'}^*$$

$q=s, q'=d$ : Kaons

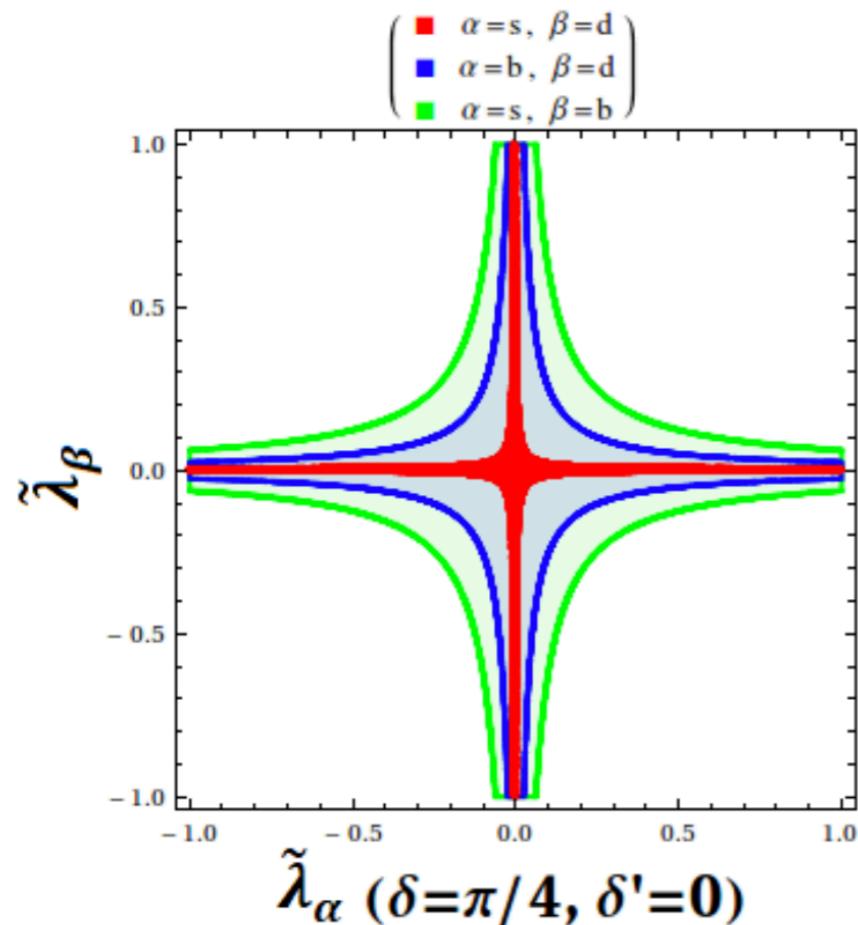
$q=b, q'=d,s$ : B mesons.



See also: Renner, Schwaller, 1803.08080.

# FLAVOUR CONSTRAINTS

Can parameterize  $\lambda$  matrix various different ways to explore allowed space.



Observables	Constraints
$\Delta M_K$	$-6.93 \times 10^{-4} \leq N_c \text{Re} [(\xi_K^*)^2] \left( \frac{1 \text{ TeV}}{M_X} \right)^2 \leq 3.77 \times 10^{-4}$
$\Delta M_d$	$N_c  \xi_{B_d}^2  \left( \frac{1 \text{ TeV}}{M_X} \right)^2 \leq 6.55 \times 10^{-4}$
$\Delta M_s$	$N_c  \xi_{B_s}^2  \left( \frac{1 \text{ TeV}}{M_X} \right)^2 \leq 13.15 \times 10^{-3}$
$ \epsilon_K $	$N_c  \text{Im} [(\xi_K^*)^2]  \left( \frac{1 \text{ TeV}}{M_X} \right)^2 \leq 1.64 \times 10^{-6}$
$S_{\psi K_S}$	$-0.86 \times 10^{-4} \leq N_c \text{Im} [(\xi_{B_d}^*)^2] \left( \frac{1 \text{ TeV}}{M_X} \right)^2 \leq 3.12 \times 10^{-4}$
$S_{\psi \phi}$	$-3.54 \times 10^{-3} \leq N_c \text{Im} [(\xi_{B_s}^*)^2] \left( \frac{1 \text{ TeV}}{M_X} \right)^2 \leq 3.88 \times 10^{-3}$

Bensalem, DS, arXiv:2111.05515.

# MESON DECAY

SM mode	BR(SM)	BR(Exp)	BR(Projected Exp)	NP modes
$B_d \rightarrow \nu\bar{\nu}$ $B_d \rightarrow \nu\bar{\nu}\nu\bar{\nu}$	$1.24 \times 10^{-25}$ [76] $(1.51 \pm 0.28) \times 10^{-16}$ [77]	$< 2.4 \times 10^{-5}$ BaBar [74]	$< 1.5 \times 10^{-6}$ Belle II [75]	$B_d \rightarrow Q\bar{Q}$ $B_d \rightarrow \pi_D\pi_D$
$B_s \rightarrow \nu\bar{\nu}$ $B_s \rightarrow \nu\bar{\nu}\nu\bar{\nu}$	$3.07 \times 10^{-24}$ [76] $(5.48 \pm 0.89) \times 10^{-15}$ [77]	—	$< 1.1 \times 10^{-5}$ Belle II [75]	$B_s \rightarrow Q\bar{Q}$ $B_s \rightarrow \pi_D\pi_D$
$B_d \rightarrow \pi^0\nu\bar{\nu}$	$(1.2 \pm 0.15) \times 10^{-7}$ [86]	$< 0.9 \times 10^{-5}$ Belle [85]	$\mathcal{O}(10^{-7})$ Belle II	$B_d \rightarrow \pi^0 Q\bar{Q}$ $B_d \rightarrow \pi^0\pi_D$
$B_d \rightarrow K_S\nu\bar{\nu}$	$(2.00 \pm 0.25) \times 10^{-6}$ [87]	$< 1.3 \times 10^{-5}$ Belle [85]	$\mathcal{O}(30\%)$ Belle II	$B_d \rightarrow K_S Q\bar{Q}$ $B_d \rightarrow K_S\pi_D$
$B^+ \rightarrow \pi^+\nu\bar{\nu}$	$(2.39 \pm 0.30) \times 10^{-7}$ [86]	$< 1.4 \times 10^{-5}$ Belle [85]	$\mathcal{O}(10^{-7})$ Belle II	$B^+ \rightarrow \pi^+ Q\bar{Q}$ $B^+ \rightarrow \pi^+\pi_D$
$B^+ \rightarrow K^+\nu\bar{\nu}$	$(4.00 \pm 0.50) \times 10^{-6}$ [87]	$< 1.9 \times 10^{-5}$ Belle [85]	$\mathcal{O}(20\%)$ Belle II	$B^+ \rightarrow K^+ Q\bar{Q}$ $B^+ \rightarrow K^+\pi_D$
$K_L \rightarrow \pi^0\nu\bar{\nu}$	$(3.00 \pm 0.31) \times 10^{-11}$ [91]	$< 4.9 \times 10^{-9}$ KOTO [92]	$\mathcal{O}(20\%)$ KOTO Step2 [96]	$K_L \rightarrow \pi^0\pi_D$
$K^+ \rightarrow \pi^+\nu\bar{\nu}$	$(9.11 \pm 0.72) \times 10^{-11}$ [91]	$< 1.78 \times 10^{-10}$ NA62 [93]	—	$K^+ \rightarrow \pi^+\pi_D$

Bensalem, DS, arXiv:2111.05515.

# CP ASYMMETRIES

NP mode	Asym Type	SM	# DF	Asym(NP)
$B_d \rightarrow Q\bar{Q}$	Mixing	$\simeq 0$	1	$-\sin [2(\beta + \delta_{dbQ})] \sin(\Delta M_d t)$
$B_s \rightarrow Q\bar{Q}$	Mixing	$\simeq 0$	1	$-\sin [2(\beta_s + \delta_{sbQ})] \sin(\Delta M_s t)$
$B_d \rightarrow \pi^0 \pi_D$	Mixing	$+ \sin(4\beta) \sin(\Delta M_d t)$	$\geq 2$	$+ \sin [2(\beta + \delta_{dbQ})] \sin(\Delta M_d t)$
$B_d \rightarrow K_S \pi_D$	Mixing	$- \sin [2(\beta + \beta_s)] \sin(\Delta M_d t)$	$\geq 2$	$- \sin [2(\beta + \delta_{sbQ'})] \sin(\Delta M_d t)$
$B^+ \rightarrow \pi^+ \pi_D$	Direct	$\simeq 0$	$\geq 2$	$\propto \sin(\delta_{dbQ_1} - \delta_{dbQ_2})$
$B^+ \rightarrow K^+ \pi_D$	Direct	$\simeq 0$	$\geq 2$	$\propto \sin(\delta_{sbQ_1} - \delta_{sbQ_2})$
$K_L \rightarrow \pi^0 \pi_D$	Grossman-Nir	$(4.3) \sin^2 (\beta_s - \beta)$	$\geq 2$	$(4.3) \sin^2 (\delta_{dsQQ'})$

# CONCLUSIONS

- Confining dark sectors are well motivated extensions of the SM.
- Emerging jets are novel and interesting, LHC experiments are starting a dedicated search program. Future detectors may also have sensitivity to discovery.
- Strategies presented here can reach very low cross sections, sensitive to broad class of displaced models.
- Trigger efficiencies at low mass can be improved by adding radiation effects and possibly with new triggers.
- Lots more to do from both bottom up perspective (other kinds of dark sectors) and top down (flavour structure, unification).

**THANK  
YOU**